Diphoton Analysis: RS graviton search

Pre-blessing

Tracey Pratt
Liverpool University

Ray Culbertson Fermilab

July 10th 2003
Exotics Group meeting

Outline of RS model diphoton search
(shown at Exotics meeting June 26th 03)

- To bring $\gamma\gamma$ up-to-date with ee and $\mu\mu$ channels
  - Data used is from dev_240
    code version 4.10.4 + patches equivalent to 4.11.1
  - Acceptance is calculated using MC generated 4.9.1, as used for blessed ee and $\mu\mu$ searches
  - Efficiencies calculated with MC and compared/corrected for with data/MC, conversions and underlying event effects
  - Signal Region taken as a $\pm 3\sigma$ window around the signal mass, in analogy to the ee search, which has a similar signal invariant mass resolution
  - Background is assumed to be zero, all events taken as signal
  - Systematic uncertainties are included in the limit
  - Limits are obtained using Bayes program, as used for blessed ee and $\mu\mu$ searches

Additions from last presentation

- Acceptance corrected for angular distribution
- Study of conversions on selection criteria efficiencies
- Trigger Efficiency for PHOTON 70 trigger included

Dataset and Trigger

cph108 and cph109 data sets used
From run 138425 (Feb 4th 2002) (108) to 163527 (May 27th 03) (109)
Luminosity: 82 pb^{-1} + 52pb^{-1} = 134 pb^{-1}

The trigger required was PHOTON 70 ($Ultra$) which has no ISO or had/em and Trigger PHOTON 50 ($Super$) which has no ISO cut
PHOTON_50

- Level 1: L1 : One EM tower, with Energy > 8 GeV
- had/em<0.125 unless Et>14
- Level 2 : L2 :
  - Requires an EM cluster, with EM E_T > 40 GeV
  - had/em<0.125
  - No isolation cuts.
- Level 3 : L3 :
  - Require photon to have E_T > 50 GeV
  - HAD/EM<0.125, off at 200 GeV
  - No isolation cuts.

PHOTON_70

- Level 1: L1 : One EM tower, with EM+HAD Et > 10 GeV
- No had/em cut.
- Level 2 : L2 :
  - Requires an EM cluster, with EM E_T > 70 GeV
  - No had/em or isolation cuts.
- Level 3 : L3 :
  - Require photon to have E_T > 70 GeV
  - HAD/EM<0.2 +0.001*E
  - No isolation cuts.

Signal Sample Selection Criteria

Good runs requirements: bit called 'good run core components' set by shift crew, CAL offline and COT offline

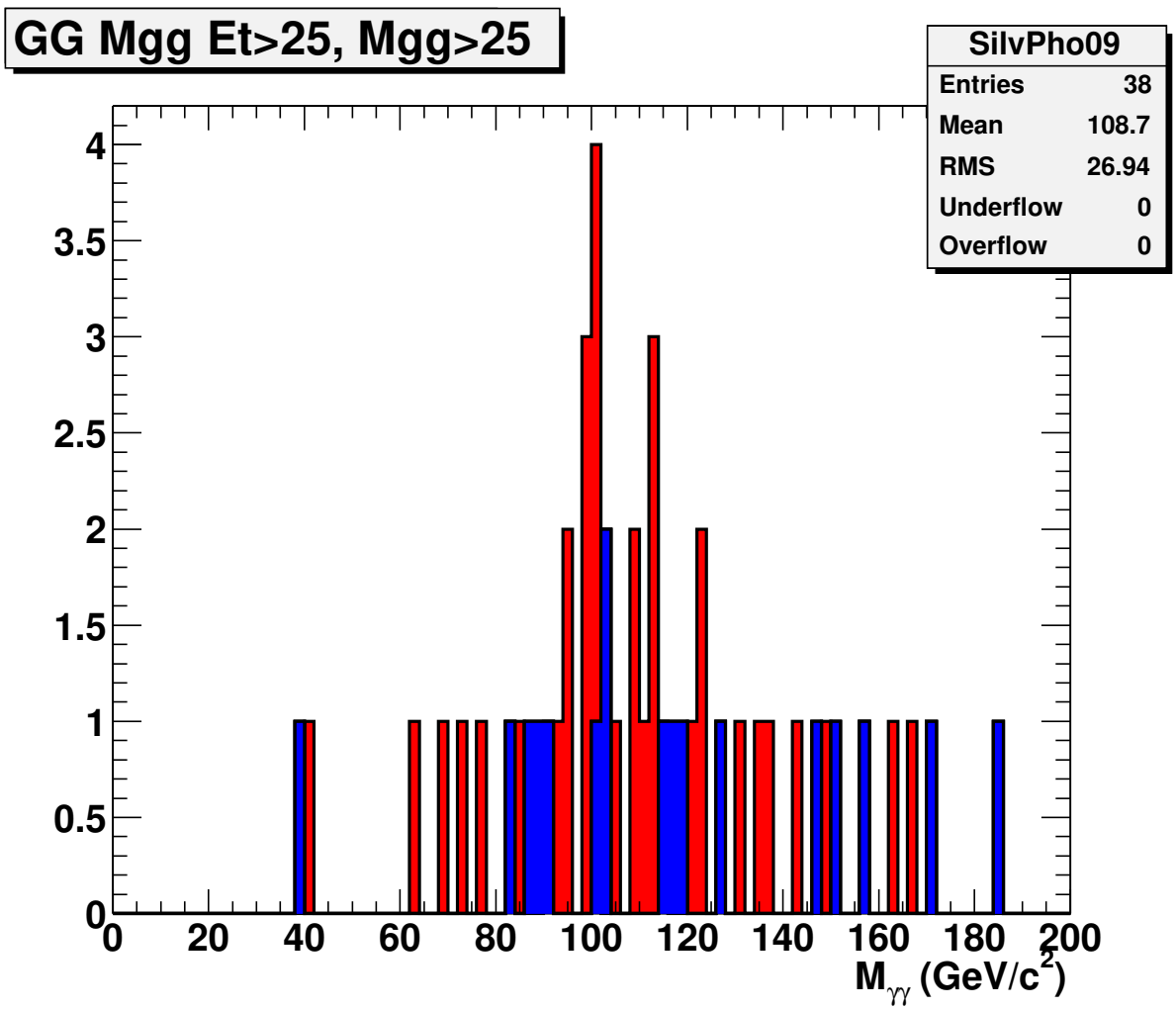
The events were reconstructed using production off-line version 4.10.4
+ patches equivalent to 4.11.1
and ntuplised using the stnuple version dev_240.

Baseline Analysis Cuts for High Pt Photons V2.0	
Variable	Cut
Had/Em	$< 0.055 + 0.00045$ * ECorr OR < 0.125
Chi2 (Strips+Wires)/2.0	< 20
N track (N3D)	≤ 1
Track Pt	$<$ 1+0.005*EtCorr GeV
Cone 0.4 IsoEtCorr	EtCorr<20: $<$ 0.1*EtCorr EtCorr>20: $<$ 2.0+0.02*(EtCorr-20)
Cone 0.4 Track Iso	$<$ 2.0 +0.005*EtCorr
2nd CES cluster	EtCorr<18: $<$ 0.14*EtCorr
(strip and wire)	EtCorr>18: $<$ 2.4+0.01*EtCorr
Fiducial	Ces $\lvert X\rvert$<21 cm, 9< Ces $\lvert Z\rvert$<230 cm

Additional Cuts: require $E_T > 25$ GeV for both central photons
and $M_{\gamma\gamma} > 25$ GeV

Data Sample

After application of selection cuts:
events remaining with 2 good photons with Et>25 and Mgg>25
38 events: in cph108 sample
17 events: in cph109 sample



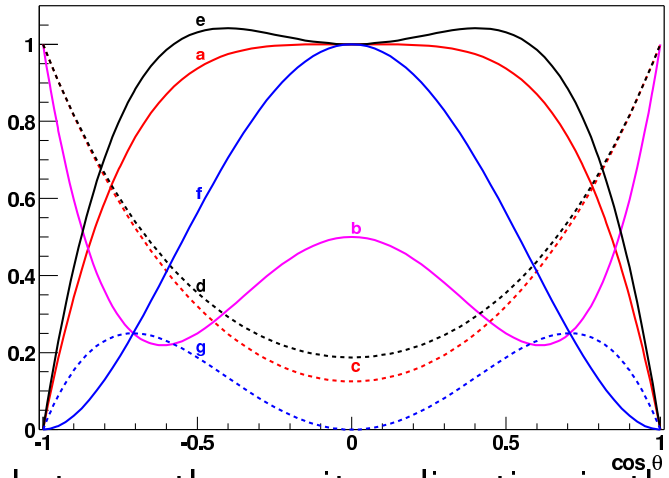
Highest invariant mass event at 185 GeV/c^2

Monte Carlo

The RS MC used was Pythia 6.203
2000 events were generated at each M_G with pdf CTEQ5L
using cdfSim and reconstructed with ProductionExe code version 4.9.1

Pythia assumes that the diphotons decay isotropically in the graviton rest frame

Angular distribution of graviton decay products



cosθ^* is the angle between the graviton direction in the lab frame and one of the photons in the graviton rest frame (Reference: hep-ph 0211205)

Process	Distribution	Plot
Assumed decays	1	*("flat")*
$gg \rightarrow G \rightarrow \gamma\gamma$	$1 + 6\cos^2\theta^* + \cos^4\theta^*$	c *("u shaped")*
$q\,\bar{q} \rightarrow G \rightarrow \gamma\gamma$	$1 - \cos^4\theta^*$	a *("n shaped")*

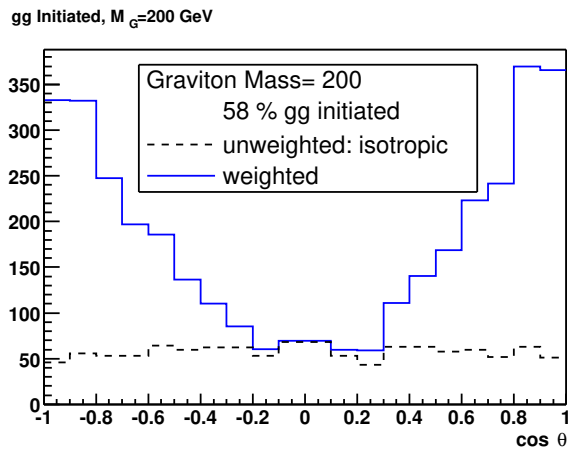
Therefore the events were weighted to give the correct angular distributions

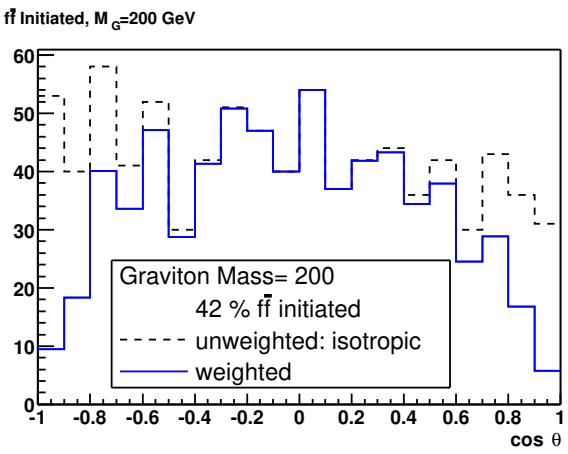
Event weighting

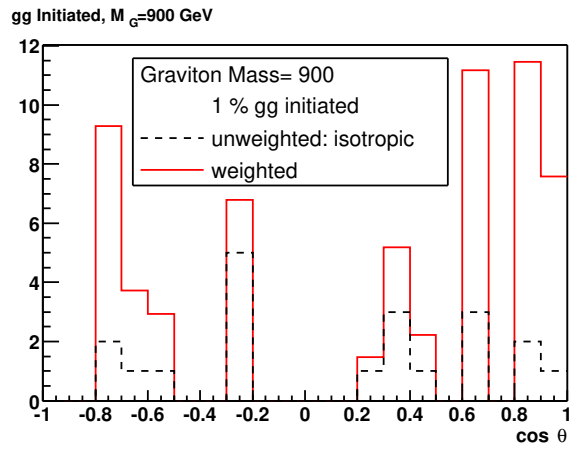
To correct for the isotropic assumption:
Events are weighted according to their correct angular distributions

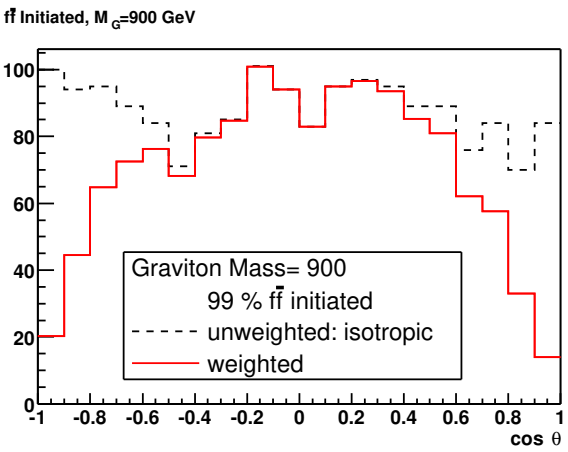
gg $\rightarrow$ G
flat $\rightarrow$ *"u-shaped"*

$q\ \bar{q}\rightarrow$ G
flat $\rightarrow$ *"n-shaped"*





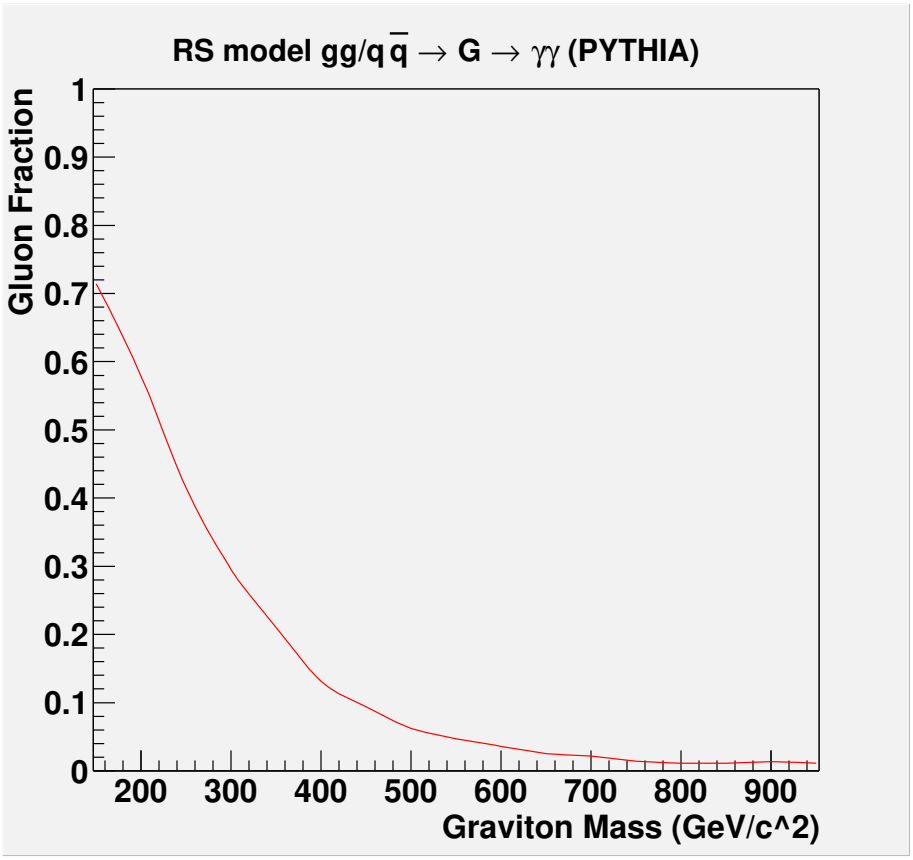




What effect will correct angular distribution have on the $\gamma\gamma$ acceptance ?
Assuming that the graviton decays at rest:
1) gg $\rightarrow$ G $\rightarrow \gamma\gamma$ will have a smaller acceptance (fewer central decays)
2) $q\ \bar{q} \rightarrow$ G $\rightarrow \gamma\gamma$ will have a larger acceptance (more central decays)

Acceptance Correction for angular distribution

Fraction of events graviton events which are gg $\rightarrow$ G $\rightarrow \gamma\gamma$



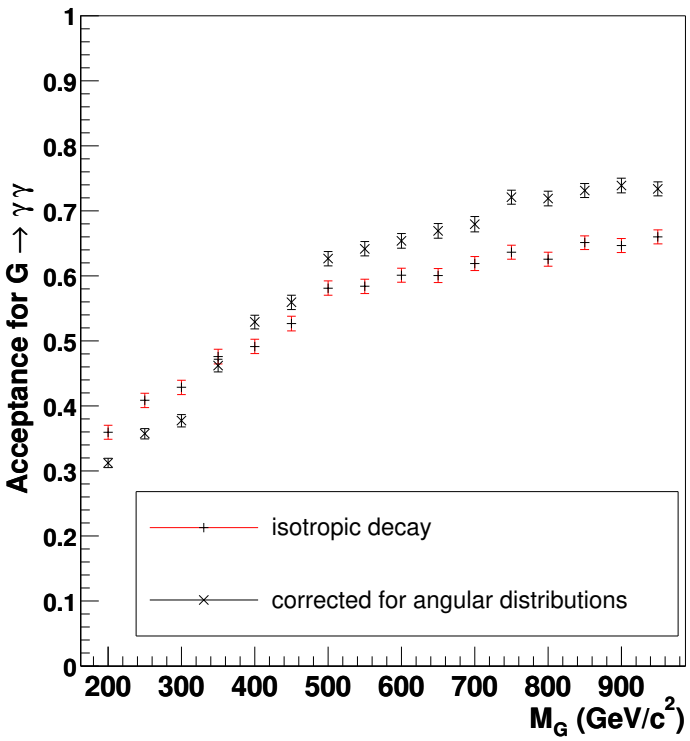
Smaller gg fraction as graviton mass increases.

So assuming that the graviton decays at rest
expect that the acceptance is increased slightly for M_G>400 GeV/c^2

So I had assumed that the acceptance with isotropic decays is a conservative estimate. Now confirmed ...

Acceptance per event

$$\text{acceptance} = \frac{\sum weigthed_events_with_2_central_\gamma s_with_E_T{>}25GeV_and M_{\gamma\gamma}{>}25GeV}{\sum_weighted^{generated}_{events}}$$



Predictions confirmed:

- Larger for large M_G ($>$ 400 GeV), where gg intiated fraction is small
- Here the $f\bar{f}$ fraction is dominant and this decays mostly centrally
- Smaller for lower M_G ($<$ 350 GeV), where gg intiated fraction is large
- Here the gg fraction is dominant and the photons (in the graviton rest frame) are predominantly emitted in the direction the graviton (anti-) direction is travelling

Systematic uncertainty on the acceptance

1) E_T scale dependence: the acceptance was calculated, changing the 25 GeV E_T threshold by $\pm$ 1 % (from the difference of Z mass from MC and data).

M_G (GeV)	Acceptance (%) E_t>25	E_t>24.75	E_t> 27.50	Systematic uncertainty (%)
200	31.24	31.24	31.19	0.17
250	35.72	35.86	35.66	0.16
300	37.73	37.84	37.73	0.00
350	46.20	46.27	46.12	0.17
400	52.91	53.00	52.82	0.17
450	55.92	56.14	55.87	0.08
500	62.65	62.70	62.59	0.08
550	64.17	64.23	64.17	0.00
600	65.38	65.44	65.28	0.16
650	66.93	66.93	66.93	0.00
700	67.94	67.94	67.88	0.09
750	72.10	72.10	72.10	0.00
800	71.88	71.88	71.88	0.00
850	73.13	73.13	73.13	0.00
900	73.89	73.89	73.89	0.00
950	73.39	73.39	73.39	0.00

The systematic uncertainty conservatively taken as 0.2 %
The systematic uncertainty is assumed to be independent of M_G

2) For M_G generate 2000 events with different pdf's

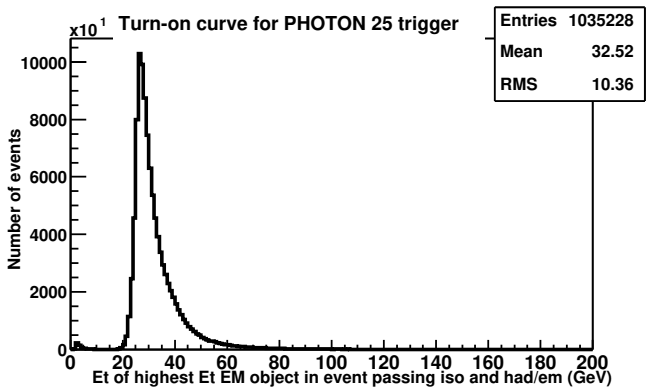
PDF	Acceptance M_G=600 GeV (%)	Relative systematic uncertainty (%)
CTEQ5L (LO)	65.4 $\pm$ 1.1	-
GRV 94L (LO)	65.1 $\pm$ 1.1	-0.43
CTEQ5M (NLO in the MS-bar scheme)	65.1 $\pm$ 1.1	-0.508
GRV 94D (NLO in DIS scheme)	64.2 $\pm$ 1.1	-1.82

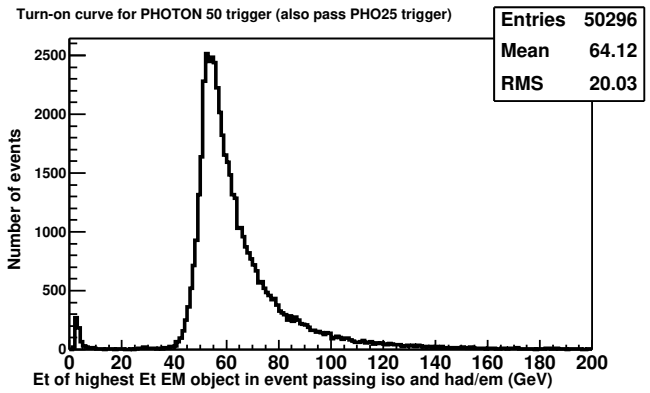
Over estimate the systematic uncertainty on the acceptance as 3 %
(The same as the ee channel)

The turn-on and trigger efficiency for PHOTON 50

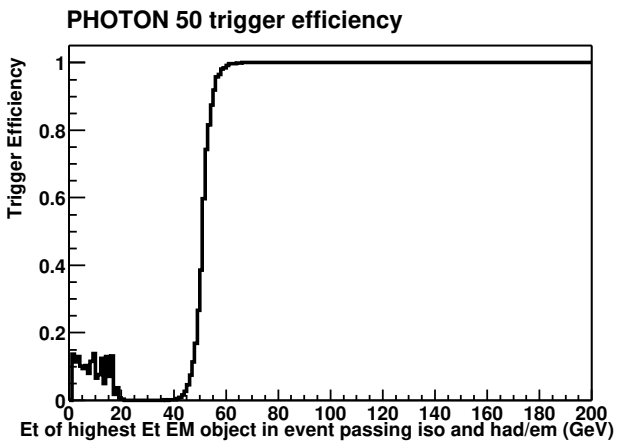
Determined using the PHOTON 25 trigger.

Looped over the cp108 and cph109 samples
Plotted the Et of the highest Et emobject (which passed the iso and had/em cuts) in events which passed the PHOTON 25 trigger (left) and both PHOTON 25 and 50 trigger (right).
The iso and had/em cuts are required in order to reduce the fake photon background





The small peak near zero is could be from events which pass the trigger due to another emobject in the event, which has larger Et, but which does not pass the iso and the had/em cut
The trigger efficiency is obtained by dividing the above two histograms.



Photon 50 Trigger Efficiency for diphoton events

For cph108 and cpho109 samples

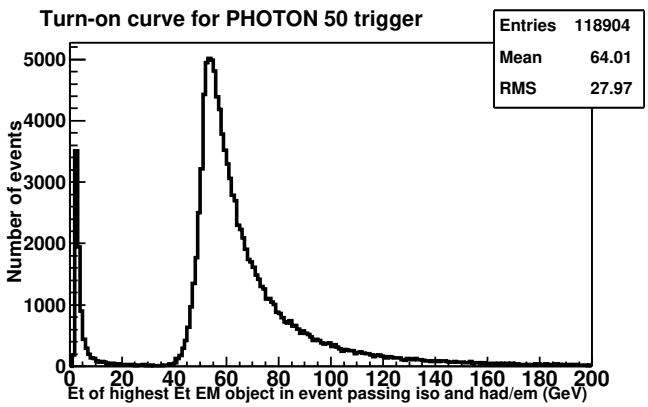
PHOTON 50 Trigger Efficiency above bin minimum				
minimum Bin			Efficiency (%)	
Et (GeV)	$N_{pass25trig}$	$N_{pass25and50trig}$	$(\epsilon^{\gamma}_{trig})$	$(\epsilon^{\gamma\gamma}_{trig} = \epsilon^{\gamma}_{trig} \times (2 - \epsilon^{\gamma}_{trig}))$
50	65819	79783	82.2 ± 0.2	$96.8\ ^{+0.1}_{-0.1}$
55	51584	52815	97.8 ± 0.08	$99.9\ ^{+0.003}_{-0.004}$
60	36006	36140	99.8 ± 0.03	$100\ ^{+0}_{-0.0001}$
65	25531	25586	100 ± 0.01	$100\ ^{+0}_{-4e-06}$
70	18707	18748	100 ± 0.008	$100\ ^{+0}_{-2e-06}$
75	13909	13939	100 ± 0.01	$100\ ^{+0}_{-3e-06}$
80	10484	10506	100 ± 0.01	$100\ ^{+0}_{-5e-06}$
85	8185	8200	100 ± 0.02	$100\ ^{+0}_{-9e-06}$
90	6360	6372	100 ± 0.02	$100\ ^{+0}_{-1e-05}$
95	5020	5029	100 ± 0.03	$100\ ^{+0}_{-2e-05}$
100	4016	4023	100 ± 0.03	$100\ ^{+0}_{-3e-05}$
105	3243	3250	100 ± 0.04	$100\ ^{+0}_{-5e-05}$
110	2632	2637	99.9 ± 0.05	$100\ ^{+0}_{-8e-05}$
115	2197	2201	99.9 ± 0.06	$100\ ^{+0}_{-0.0001}$
120	1794	1796	99.9 ± 0.07	$100\ ^{+0}_{-0.0002}$
125	1465	1467	99.9 ± 0.09	$100\ ^{+0}_{-0.0002}$
130	1236	1238	99.9 ± 0.1	$100\ ^{+0}_{-0.0003}$
135	1026	1028	99.9 ± 0.1	$100\ ^{+0}_{-0.0005}$
140	865	866	99.9 ± 0.1	$100\ ^{+0}_{-0.0006}$
145	733	734	99.8 ± 0.2	$100\ ^{+0}_{-0.0009}$
150	635	636	99.8 ± 0.2	$100\ ^{+0}_{-0.001}$
155	530	531	99.8 ± 0.2	$100\ ^{+0}_{-0.002}$
160	461	462	99.7 ± 0.3	$100\ ^{+0}_{-0.002}$
165	412	413	99.7 ± 0.3	$100\ ^{+0}_{-0.002}$
170	358	359	99.7 ± 0.3	$100\ ^{+0}_{-0.003}$
175	306	307	99.6 ± 0.4	$100\ ^{+0}_{-0.004}$
180	275	276	99.6 ± 0.4	$100\ ^{+0}_{-0.005}$
185	239	240	99.5 ± 0.5	$100\ ^{+0}_{-0.007}$
190	209	210	99.4 ± 0.6	$100\ ^{+0}_{-0.009}$
195	189	190	99.4 ± 0.6	$100\ ^{+0}_{-0.01}$
200	169	170	99.3 ± 0.7	$100\ ^{+0}_{-0.01}$

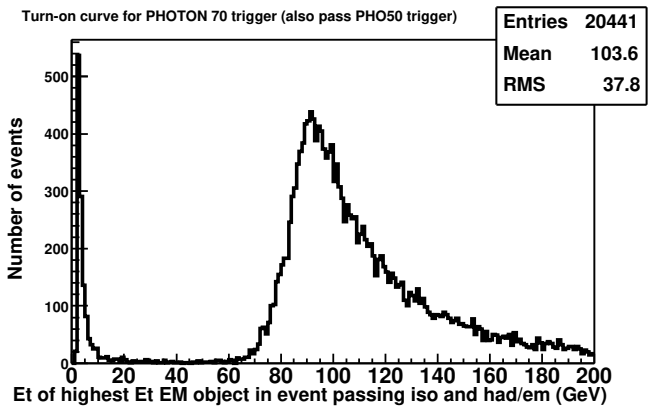
PHOTON 50 trigger efficiency is $100^{+0}_{-0.01}$ % for 2γ above 55 GeV.
(largest error taken, from above 200 GeV)

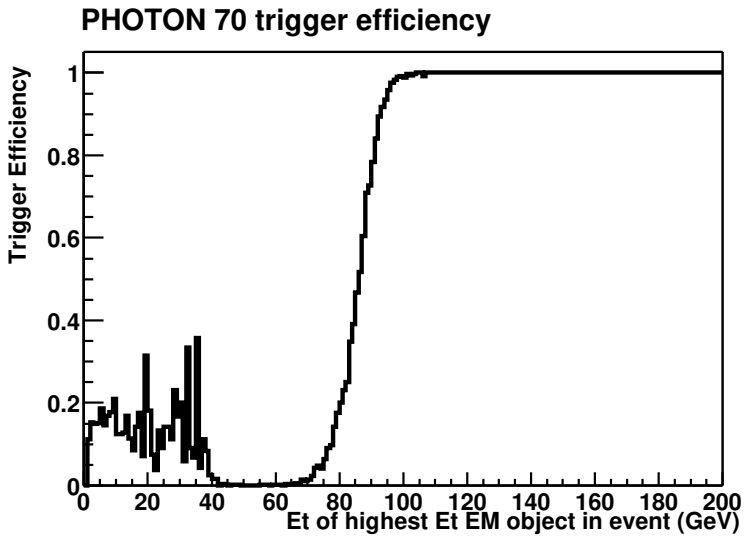
Turn-on and trigger efficiency for PHOTON 70

Determined using the PHOTON 50 trigger
Note that although the PHOTON70 trigger does not have a had/em cut at level 1 or 2,
the PHOTON50 trigger does at Level 1 and 2: had/em$<$0.125
The efficiency of the had/em Level 2 trigger is determined from jet data, looking at the underlying events

Plotted here is the Et of the highest Et emobject (which passed the iso and had/em cuts) in events which passed the
PHOTON 50 trigger (left) and PHOTON 50 and 70 triggers (right).







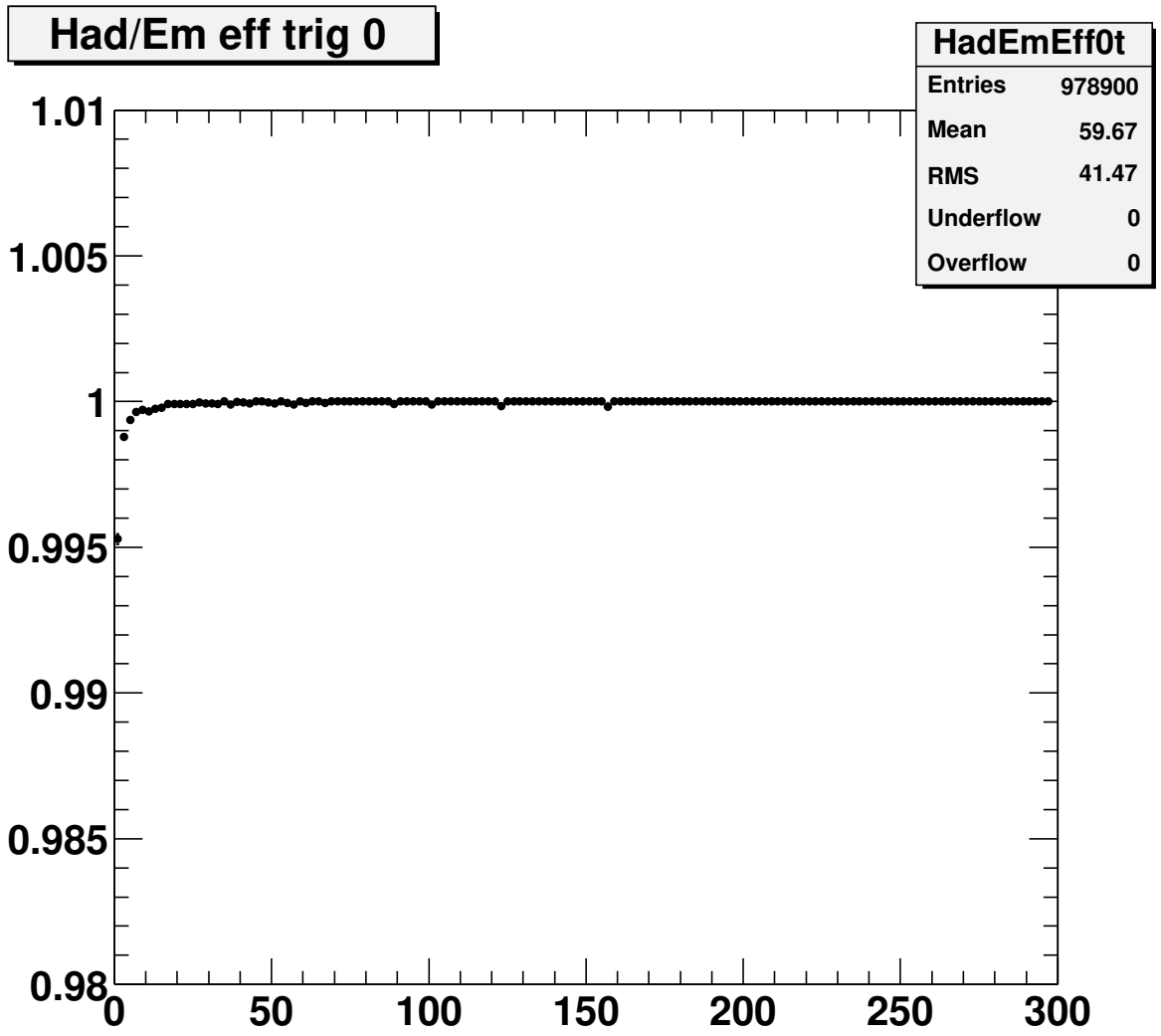
Photon 70 Trigger Efficiency (from cph108)

Photon 70 trigger efficiency for above bin minimum				
minimum Bin	Number of events		Efficiency (%)	
Et (GeV)	$N^{IsoHadEm}_{pass50}$	$N^{IsoHadEm}_{pass50a70}$	$(\epsilon^{\gamma}_{trig})$	$(\epsilon^{\gamma\gamma}_{trig} = \epsilon^{\gamma}_{trig} \times (2 - \epsilon^{\gamma}_{trig}))$
70	60120	97055	52.1 ± 0.3	$77.1\ ^{+0.3}_{-0.3}$
75	55324	72524	65 ± 0.3	$87.7\ ^{+0.2}_{-0.2}$
80	48230	55405	78.3 ± 0.3	$95.3\ ^{+0.1}_{-0.1}$
85	40825	43382	89.5 ± 0.2	$98.9\ ^{+0.05}_{-0.05}$
90	33951	34560	96.8 ± 0.1	$99.9\ ^{+0.0087}_{-0.009}$
95	27711	27812	99.4 ± 0.06	$100\ ^{+0.0007}_{-0.0008}$
100	22818	22836	99.9 ± 0.03	$100\ ^{+5.9e-05}_{-7.8e-05}$
105	18974	18977	100 ± 0.01	$100\ ^{+3.6e-06}_{-7.6e-06}$
110	16022	16022	100 ± 0	$100\ ^{+0}_{0}$
115	13560	13560	100 ± 0	$100\ ^{+0}_{0}$
120	11634	11634	100 ± 0	$100\ ^{+0}_{0}$
125	10018	10018	100 ± 0	$100\ ^{+0}_{0}$
130	8689	8689	100 ± 0	$100\ ^{+0}_{0}$
135	7596	7596	100 ± 0	$100\ ^{+0}_{0}$
140	6689	6689	100 ± 0	$100\ ^{+0}_{0}$
145	5930	5930	100 ± 0	$100\ ^{+0}_{0}$
150	5305	5305	100 ± 0	$100\ ^{+0}_{0}$
155	4748	4748	100 ± 0	$100\ ^{+0}_{0}$
160	4280	4280	100 ± 0	$100\ ^{+0}_{0}$
165	3921	3921	100 ± 0	$100\ ^{+0}_{0}$
170	3595	3595	100 ± 0	$100\ ^{+0}_{0}$
175	3297	3297	100 ± 0	$100\ ^{+0}_{0}$
180	3054	3054	100 ± 0	$100\ ^{+0}_{0}$
185	2812	2812	100 ± 0	$100\ ^{+0}_{0}$
190	2586	2586	100 ± 0	$100\ ^{+0}_{0}$
195	2394	2394	100 ± 0	$100\ ^{+0}_{0}$
200	2238	2238	100 ± 0	$100\ ^{+0}_{0}$

PHOTON 70 trigger efficiency is 100 % for 2γ above 95 GeV.

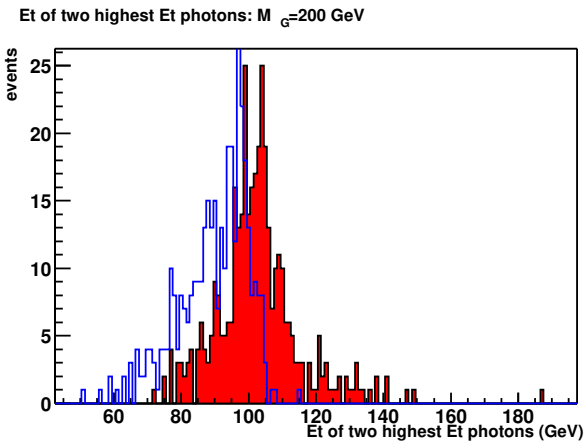
Trigger Efficiency underlying event effects

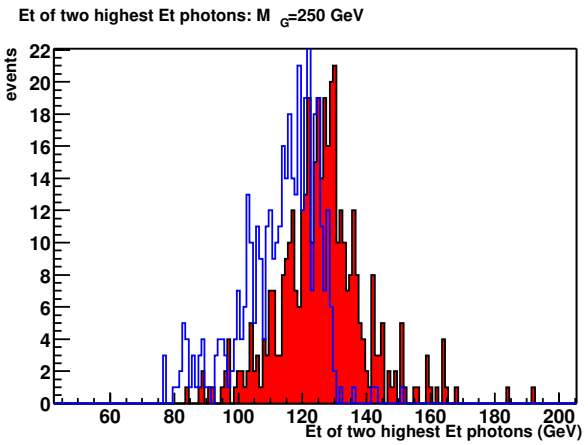
This is the eff for the L2 50gev trigger had/EM¡0.125
the eff is measured for events which pass all offline cuts,
that why it is so efficient

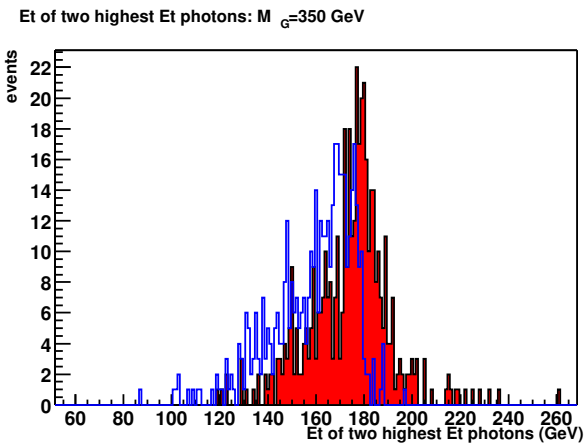


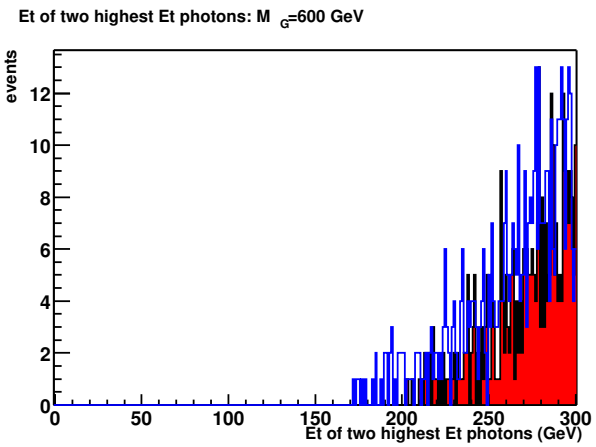
Trigger Efficiency Summary

PHOTON 50 trigger is 100 % efficient for two photons above 55 GeV.
PHOTON 70 trigger is 100 % efficient for two photons above 95 GeV.
The signal photons lie above 55 GeV, shown below:
highest Et signal γ 2nd highest Et signal γ
The trigger efficiency is taken as 100 %









Efficiency for individual selction criteria

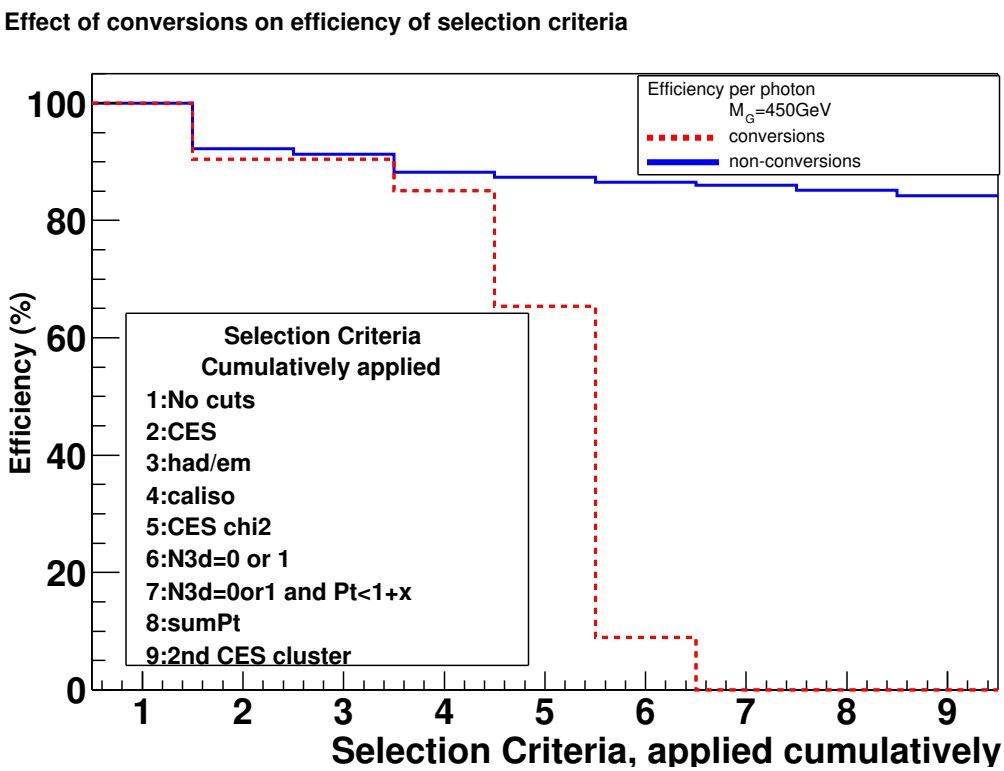
Determined from the RS graviton MC
Corrected using results and systematics obtained from 3 studies:

- Conversions
- Underlying events
- Comparisons of MC to data

Sequential efficiencies of the individual selction criteria for converted and non-converted photons

Studied using RS model Monte Carlo generated as described.

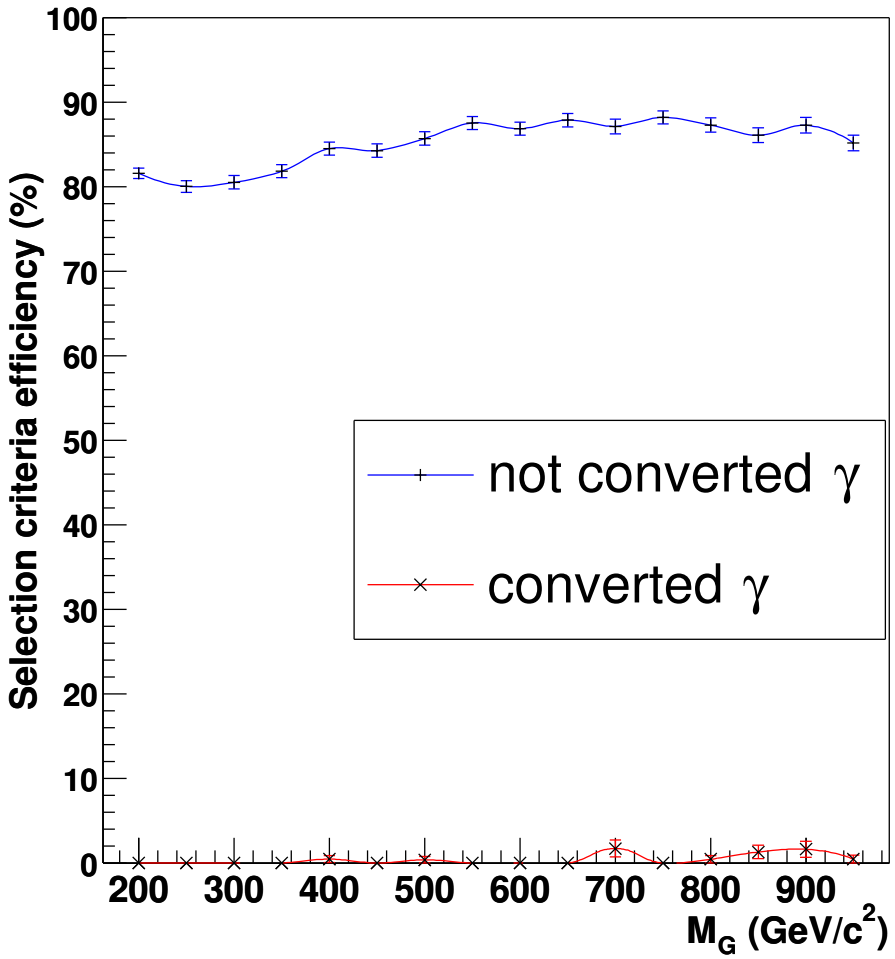
Sequential efficiency $= \frac{N_\gamma_passing_cut_and_all_previously_applied_cuts}{N_\gamma_passing_acceptance_cuts}$



The track cuts remove almost all of the conversion photons, due to the presence of the electron tracks ($\gamma \rightarrow e^+e^-$) (as expected)

Efficiency for all cuts for one photon for a (from the M_G=450 GeV sample): conversion is 0 %, non-conversion is 84.3 $\pm$ 0.8 %

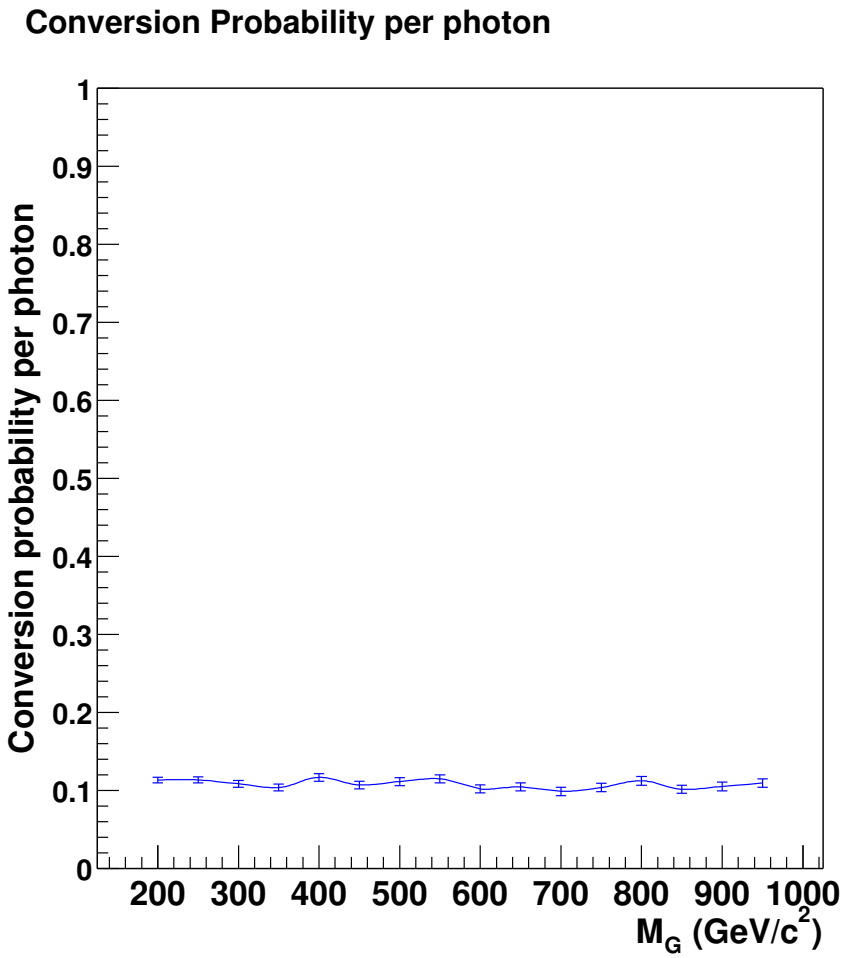
Effect of Conversions on efficiency as a function of M_G



The track cuts are not 100 % efficient at rejecting all conversion events
In some cases one of the converted electrons then emits a hard bremsstrahlung photon, and in these cases the converted events can pass the cuts

Conversion probability per photon

The conversion probability per photon is found to be independent is M_G i.e. the conversion probability is independent of the Et of the photon, as expected



The conversion probability per photon is 10.9 $\pm$ 0.1 %

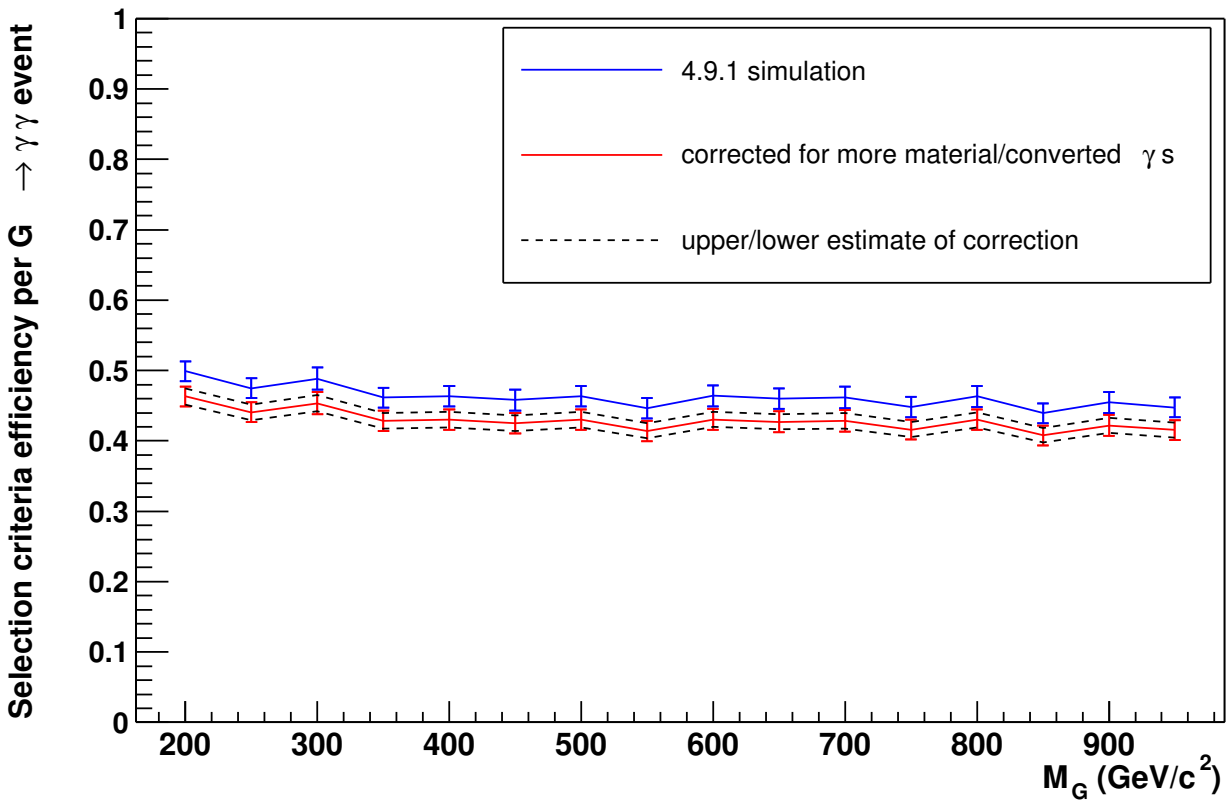
Correction to the efficiency due to conversions

The efficiency per event depends on the conversion probability
This depends on the amount of material in the detector
which is underestimated in the 4.9.1 simulation by 30 $\pm$ 10 $\%$
(Greg Veramendi's estimated the amount of material needed
in the tracking volume from the E/P spectrum:
4.9.1 simulation needs about 4.5 $\pm$ 1.5 $\%$ X_o of Copper on average.)
This uncertainty is probably an overestimate, and should come down in the coming months.

The efficiency per event is corrected for this:

- Assume that the number of conversions is directly proportional to the amount of material
- Increase the fraction of conversions (f_C^γ) by the increase in amount of material (30 $\%$)
- Assume that the conversion of the two gammas are independent
- The efficiency per event is determined from Monte Carlo
  $\epsilon^{\gamma\gamma} = \frac{N_{\gamma\gamma_passing_all_cuts}}{N_{\gamma\gamma_passing_acceptance_cuts}}$
- The efficiency per event ($\varepsilon^{\gamma\gamma}$) depends on f_C^γ
  neither converts, 1st or 2nd photon converts, or both convert
  $\varepsilon^{\gamma\gamma}$=(1-$f_C^\gamma$)*(1-$f_C^\gamma$)*$\varepsilon_{NC}$*$\varepsilon_{NC}$+2*(1-$f_C^\gamma$)*$f_C^\gamma$*$\varepsilon_{NC}$*$\varepsilon_C$+$f_C^\gamma$*$f_C^\gamma$*$\varepsilon_C$*$\varepsilon_C$
  Conservatively assume that converted photons fail the cuts (ε_C=0)
  Therefore: $\varepsilon^{\gamma\gamma}$ = (1-f_C^γ)*(1-f_C^γ)*ε_{NC}*ε_{NC}
  More material changes f_C : so material corrected efficiency
  $\varepsilon_{cor}^{\gamma\gamma} = \frac{(1-f_{Ccor}^\gamma)*(1-f_{Ccor}^\gamma)}{(1-f_C^\gamma)*(1-f_C^\gamma)}$*$\varepsilon^{\gamma\gamma}$: correction of -7.2 $\%$

Efficiency per event corrected for material increase in simulation



Systematic error shown is from the 10 % uncertainty on the amount of material

Underlying Event

Ray has done a study to understand how the underlying event affects the efficiency of the cuts
This could contribute to a systematic uncertainty to the efficiency of the photon cuts

This affects certain photon variables

- Had/em
- Track cuts
- Calorimeter isolation
- 2nd CES

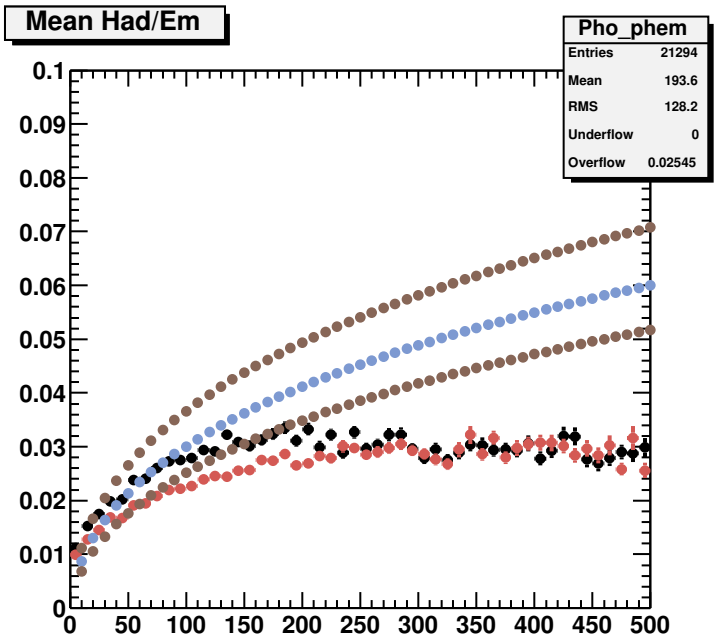
Study performed comparing dijet data to Monte Carlo
Data: sampling 1% of data of from different Et cut samples: 20,50,70,100
MC from dijet sample made for QCD winter group

Method:

Look for a jet
Define jet axis, and a 90^0 region perpdendicular to this axis
Look in cones, one from each wedge this 90^0 region
Ask in random cone:
how often does the quantity pass or fail the photon cuts?

Underlying Event: had/em

What do you expect a had/em for a photon (e) to look like?



black: mean had/em for fake event photons vs Et

red: mean had/em for fake event electrons vs Et

blue: predicted had/em from PDG EM shower model

brown: same as blue, but parameters adjusted to give a nominal variation for the purposes of a systematic error

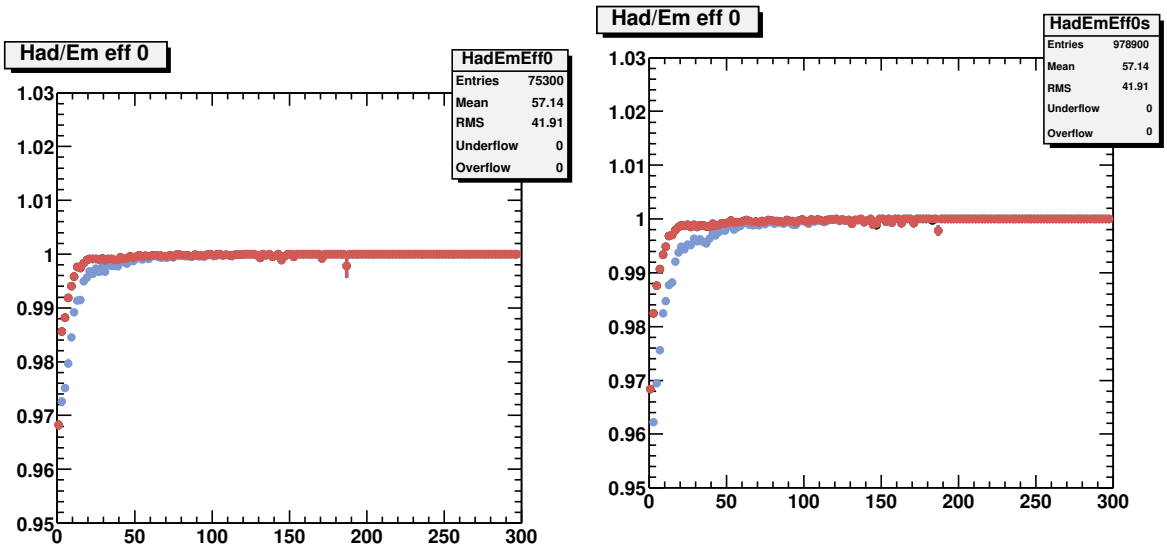
Note: MC underestimates the had/em

Underlying Event: had/em

Take had/em in 2 or 3 towers in the random cone
Leading jet in event defines em Et energy
Plot had/em vs Et of leading jet

ask how often this passes the cut?

Left: had/em efficiency based on simply the underlying event energy



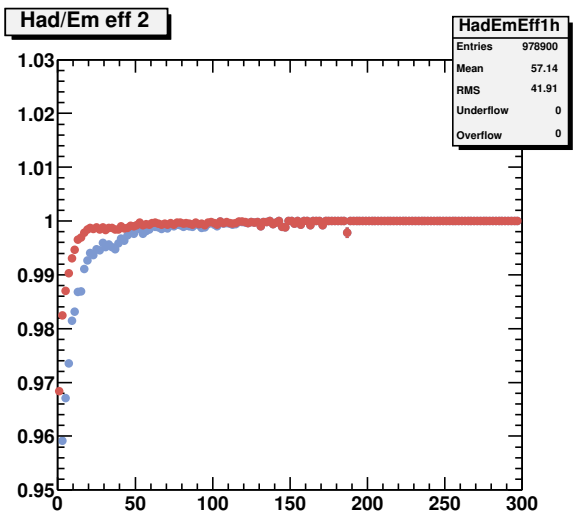
efficiency of had/em<sliding cut
had/em<0.125 or had/em<sliding cut

To find efficiency of had/em
Right: same, but now the predicted photon energy has been added to the underlying event energy.
This is the current best guess of a data-based efficiency.
The eff is slightly lower than upper right.

Underlying event systematic uncertainty: had/em

To estimate a systematic uncertainty
plot had/em vs leading jet E_T again
but now the hadronic energy predicted for the photon has been increased to the upper brown line in the first plot

i.e. it has been increased by an arbitrary amount to help set a systematic uncertainty.

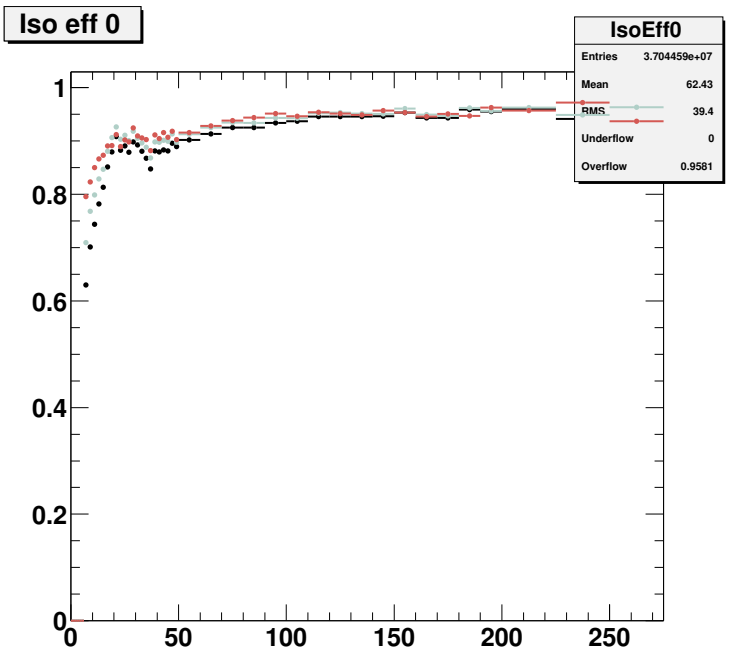


Conclude: at large Et ($>$ 40 GeV):

- there is no inefficiency of the had/em cut due to the underlying event
- the eff of the had/em cut is 100%
- and the systematic uncertainty is negligible

Underlying Event: calorimeter isolation

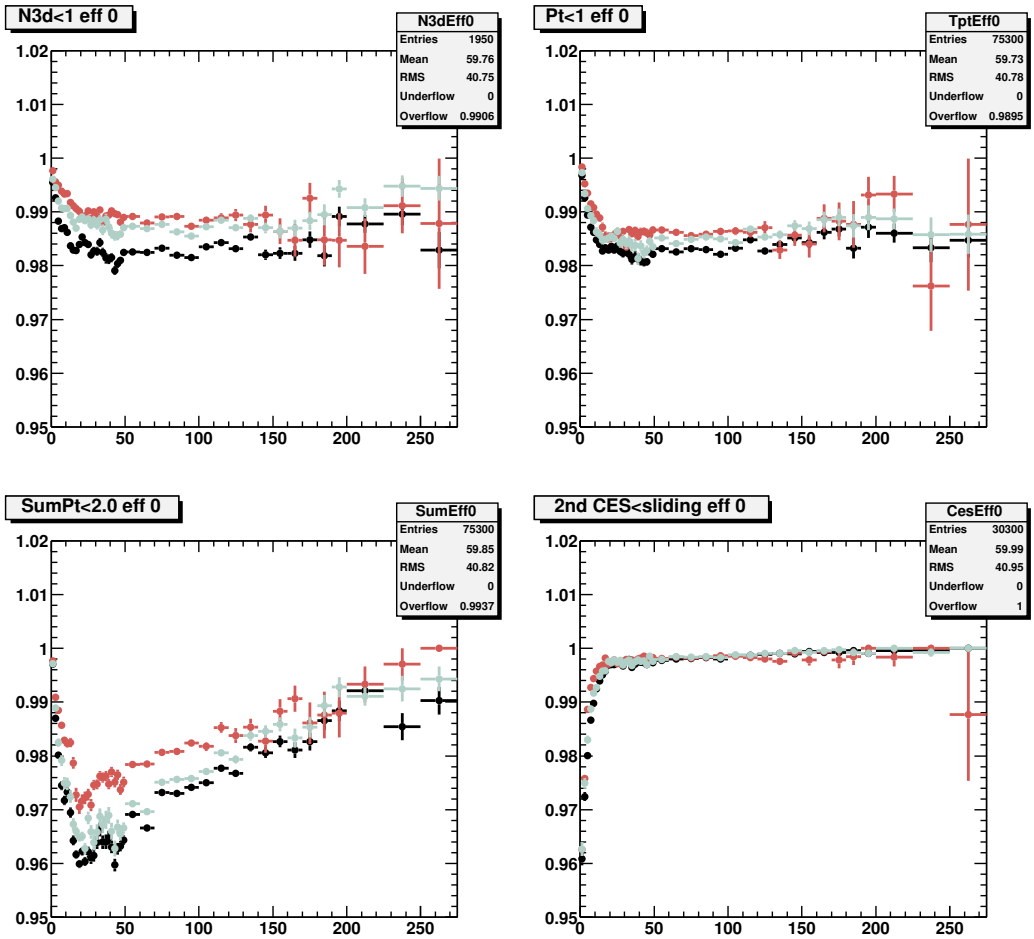
Predicted efficiencies of the isolation cut, in the cut order.
They are a combination of the expected underlying event from the jet data study,
and the leakage from the EM cluster itself, taken from a fake-event photon sample.



The UE dominates.
black = based on jet data
green = based on jet data, exactly one vertex required
red = based on MC jet data

Underlying Event: Track cuts

To study how underlying events affect the track cuts:
n3d, pt of leading track, sum pt is sum pt of tracks of cone of radius 0.4



black: data eff from underlying event in jet data, off axis
green: the same as black, but for events with exactly one vertex
red: same for jet MC

compare green and black to see how the MC is doing
compare read and black to see the overall correction

Relative systematic uncertainties from comparison of Zee MC to data

Method

In electron trigger, find the cluster that passes all offline cuts
If there are two clusters which pass cuts, select one at random,
If there is one, then look for a second cluster which passes Et>20 and 0.9<E/p<1.1 (tight E/p is to reject events which brem.)
Use the second cluster to find efficiencies.
Fit Zee peaks to gaussian and linear background

Finds the effs including the probability that the EM cluster has CES clusters in both views, and the prob that the CES clusters pass fiducial.
This number is really the ratio of the prob to pass CES cluster present (or fiducial) to the prob for making an EM cluster.

Relative systematic uncertainties from comparison of Zee MC to data

"Ces cl" means has strip and wire clusters, "cesfid" means CES clusters pass fiducial

	MC		data	
	cuml. eff	single cut eff	cuml. eff	single cut eff
1 Ces cl	97.0 ± 0.1	97.0 ± 0.1	98.8 ± 0.4	98.8 ± 0.4
2 cesfid	90.2 ± 0.3	93.0 ± 0.2	94.4 ± 0.8	95.5 ± 0.7
here denominator is cluster with Et>20, 0.9<E/p<1.1, "Ces cl", and "cesfid"				
3 HadEm	98.1 ± 0.1	98.1 ± 0.1	100.0 ± 0.3	100.0 ± 0.3
4 Cal Iso	92.9 ± 0.2	94.7 ± 0.2	97.2 ± 0.6	95.5 ± 0.7
5 Ces Chi	92.7 ± 0.2	99.7 ± 0.0	93.4 ± 0.9	96.2 ± 0.7
6 N3D	91.8 ± 0.3	99.0 ± 0.1	92.3 ± 1.0	98.8 ± 0.4
7 Pt	89.3 ± 0.3	97.3 ± 0.1	88.6 ± 1.2	96.0 ± 0.7
8 SumPT	86.4 ± 0.3	96.7 ± 0.2	84.0 ± 1.5	94.7 ± 0.8
9 Ces 2	84.0 ± 0.4	97.2 ± 0.2	81.5 ± 1.6	97.1 ± 0.6
below is the same, but for one vertex				
1 Ces cl	97.0 ± 0.1	97.0 ± 0.1	98.8 ± 0.4	98.8 ± 0.4
2 cesfid	90.2 ± 0.3	93.0 ± 0.2	95.3 ± 0.9	96.4 ± 0.8
3 HadEm	88.5 ± 0.3	98.1 ± 0.1	96.1 ± 0.8	100.0 ± 0.5
4 Cal Iso	83.8 ± 0.4	94.7 ± 0.2	90.6 ± 1.3	94.3 ± 1.0
5 Ces Chi	83.6 ± 0.4	99.7 ± 0.0	87.6 ± 1.5	96.7 ± 0.8
6 N3D	82.8 ± 0.4	99.0 ± 0.1	87.0 ± 1.5	99.3 ± 0.4
7 Pt	80.6 ± 0.4	97.3 ± 0.1	84.2 ± 1.7	96.9 ± 0.8
8 SumPT	77.9 ± 0.4	96.7 ± 0.2	80.1 ± 1.9	95.1 ± 1.0
9 Ces 2	75.7 ± 0.5	97.2 ± 0.2	78.3 ± 2.0	97.7 ± 0.7
here denominator is cluster with Et>20, 0.9<E/p<1.1, "Ces cl", and "cesfid"				
3 HadEm	98.1 ± 0.1	98.1 ± 0.1	100.0 ± 0.5	100.0 ± 0.5
4 Cal Iso	92.9 ± 0.2	94.7 ± 0.2	95.1 ± 0.9	94.3 ± 1.0
5 Ces Chi	92.7 ± 0.2	99.7 ± 0.0	91.9 ± 1.2	96.7 ± 0.8
6 N3D	91.8 ± 0.3	99.0 ± 0.1	91.3 ± 1.3	99.3 ± 0.4
7 Pt	89.3 ± 0.3	97.3 ± 0.1	88.4 ± 1.5	96.9 ± 0.8
8 SumPT	86.4 ± 0.3	96.7 ± 0.2	84.1 ± 1.7	95.1 ± 1.0
9 Ces 2	84.0 ± 0.4	97.2 ± 0.2	82.2 ± 1.9	97.7 ± 0.7

Relative systematic uncertainty on efficiencies summary

Variable	Correction to MC (%)	Rel.Syst.Unc. (%)	Study
Had/Em	+2	1.0	
UE study and Zee's shows had/em very efficent, but MC shows ineff.			
Cal iso	-2	5.0	data/MC in UE & Z
UE study disagrees with Z's by 5%, this sets the syst.			
This is an conservative estimate, since this is comparing different types of event: Z and UE			
The more important number is how much jet MC and jet data disagree			
Chi2 (Strips+Wires)/2.0	-3	2.0	Z study
N track (N3D)	-1	0.5	UE study
Zee study agrees well, but it has artificially lowered eff because the e can brem,which can then convert			
Track Pt	-1	1.0	UE study give 0.5% correction
Zee imply>1.0 but with worse statistics: average to 1			
Cone 0.4 Track Iso	-2	1.0	UE and Zee agree
2nd CES cluster (strip and wire)		-	-
Fiducial	-	3.0	
Conversions	-7.2	2.5	E/p material
Trigger Efficiency	-	-	
Total	-14.2	7.0	

Corrections added linearly: 14.2 %

Add systematics in quadrature: 7.2%.

Systematic uncertainty on efficiencies are assumed to be independent of energy

Systematic Uncertainties

Systematic uncertainties considered

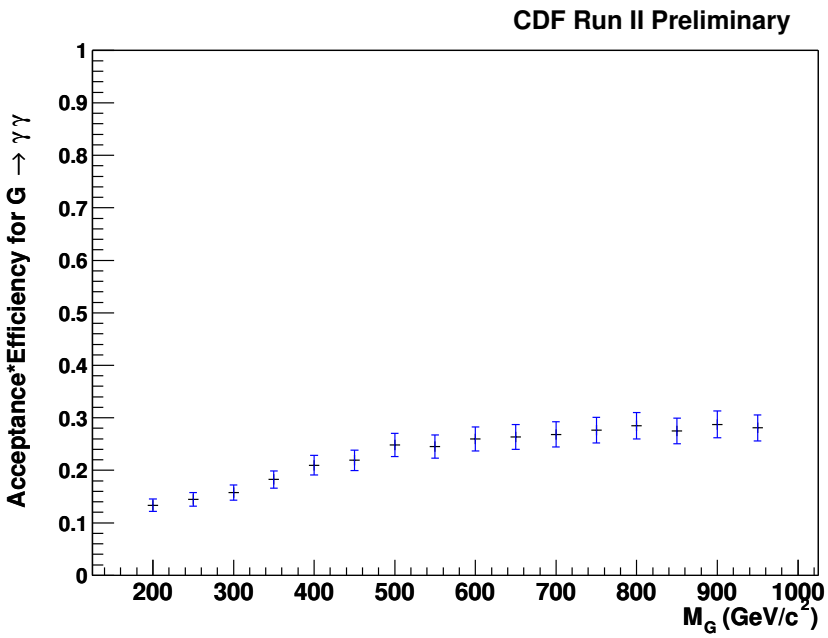
- Acceptance
  - E_T scale
  - PDFs
- Efficiency
  - underlying events
  - comparison of MC to data
  - conversions
- Luminosity

CDF Run II Preliminary (134 pb^{-1})

	Relative systematic uncertainties (%)			Statistical(%)	Combined
M_G	Efficiency	Acceptance	Luminosity	Eff*Acc	(%)
200	7.2	3.0	6.0	3.6	10
250	7.2	3.0	6.0	3.7	11
300	7.2	3.0	6.0	4.1	11
350	7.2	3.0	6.0	3.8	11
400	7.2	3.0	6.0	3.7	11
450	7.2	3.0	6.0	3.8	11
500	7.2	3.0	6.0	3.6	10
550	7.2	3.0	6.0	3.7	11
600	7.2	3.0	6.0	3.6	10
650	7.2	3.0	6.0	3.6	10
700	7.2	3.0	6.0	3.8	11
750	7.2	3.0	6.0	3.5	10
800	7.2	3.0	6.0	3.5	10
850	7.2	3.0	6.0	3.6	10
900	7.2	3.0	6.0	3.7	11
950	7.2	3.0	6.0	3.5	10

Acceptance*Efficiency

weighted for MC angular distribution
corrected for conversions, underlying events and MC/data discrepancies



Errors shown are statistical and systematic

Background

Using zero as our background estimate produces a conservative limit
Method not conservative for discovery

Unlike the ee and $\mu\mu$ channel there is no Drell-Yan background

The main background at high invariant mass is from jets faking photons
The probability of a jet faking a photon is 10^{-4} above E_t>100 GeV/c^2

Also taking signal region $\pm 3\sigma$ windows (c.f. ee),
rather than one large mass window above X GeV/c^2 (c.f. $\mu\mu$),
so expected background in signal window is negligible

Therefore 0 is a resonable background estimate for the high invariant mass region

Present

A conservative limit is obtain if the expected background is zero

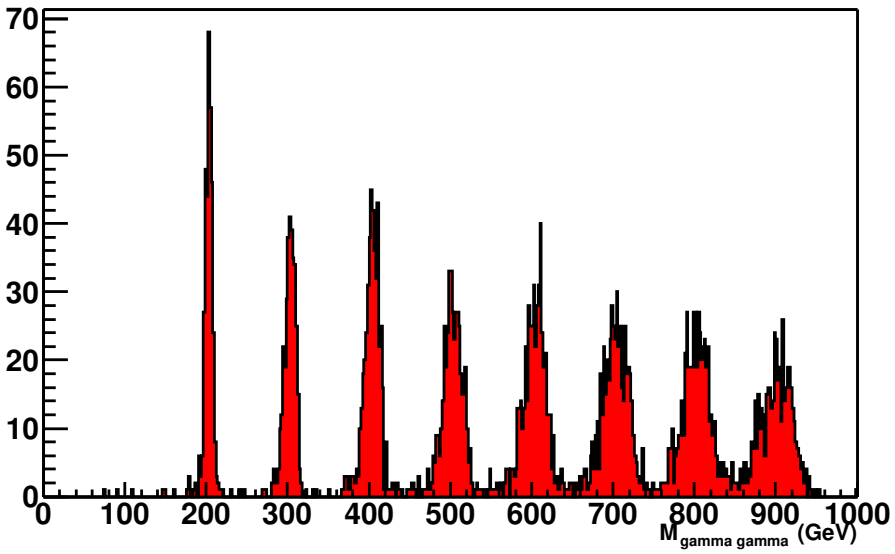
Future

Include background fits, to enable potential discovery

Signal Regions

Taking 3 σ around the reconstructed invariant mass for each M_G

Widths increase with M_G due to the EM energy resolution decreasing at large mass



M_G	Mean	Sigma (σ_M)	$3\times\sigma_M$
200	203	4.2	12.5
250	253	5.3	15.8
300	303	6.3	19.0
350	353	7.7	23.0
400	403	7.3	21.7
450	452	10.5	31.4
500	503	12.0	36.1
550	552	12.9	38.8
600	603	13.3	40.0
650	652	15.1	45.2
700	703	15.4	46.2
750	751	16.0	48.0
800	800	16.9	50.7
850	846	20.0	60.0
900	896	22.6	67.7
950	943	22.3	66.8

Summary Limits Table

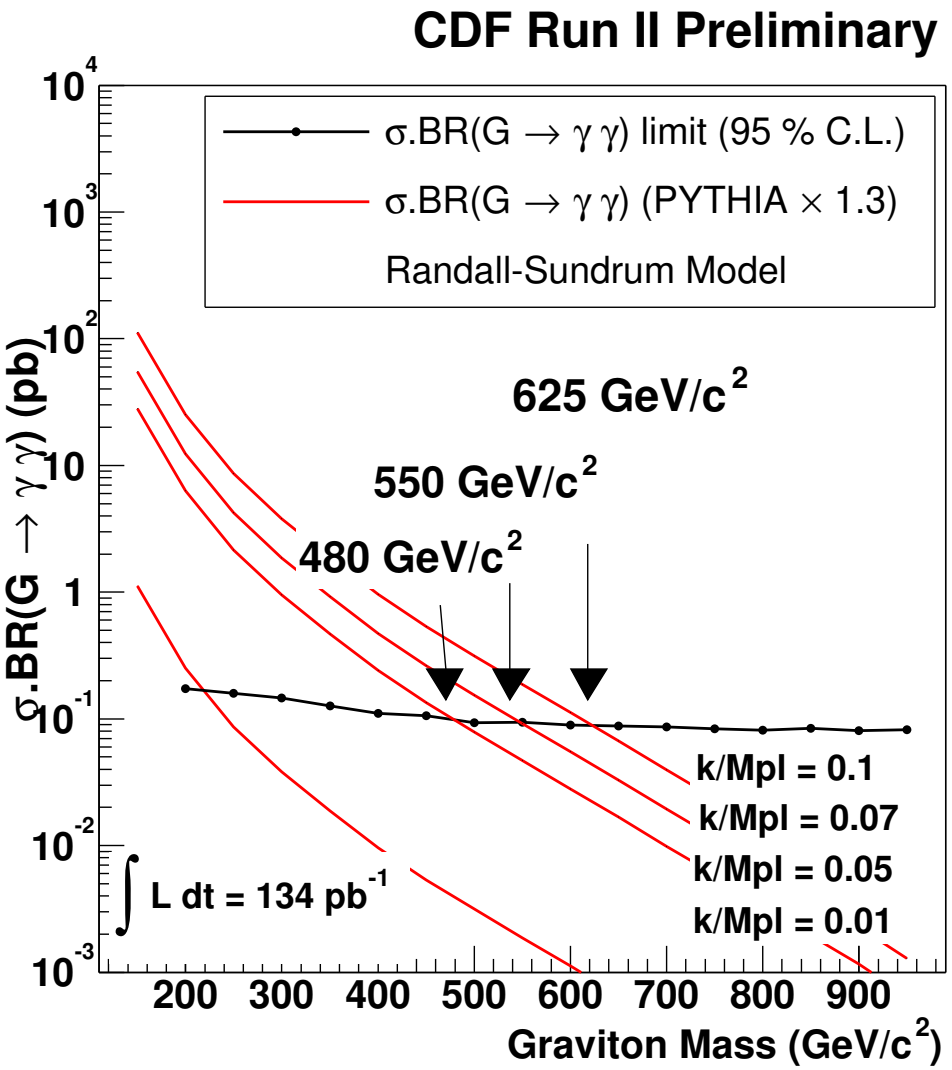
Use Bayes limit program to obtain upper limit $N_{95\%}$

$$\sigma_{95\%} \cdot B(G \to \gamma\gamma) = \frac{N_{95\%}}{\mathcal{L} \times \epsilon \times A(M)} \quad (1)$$

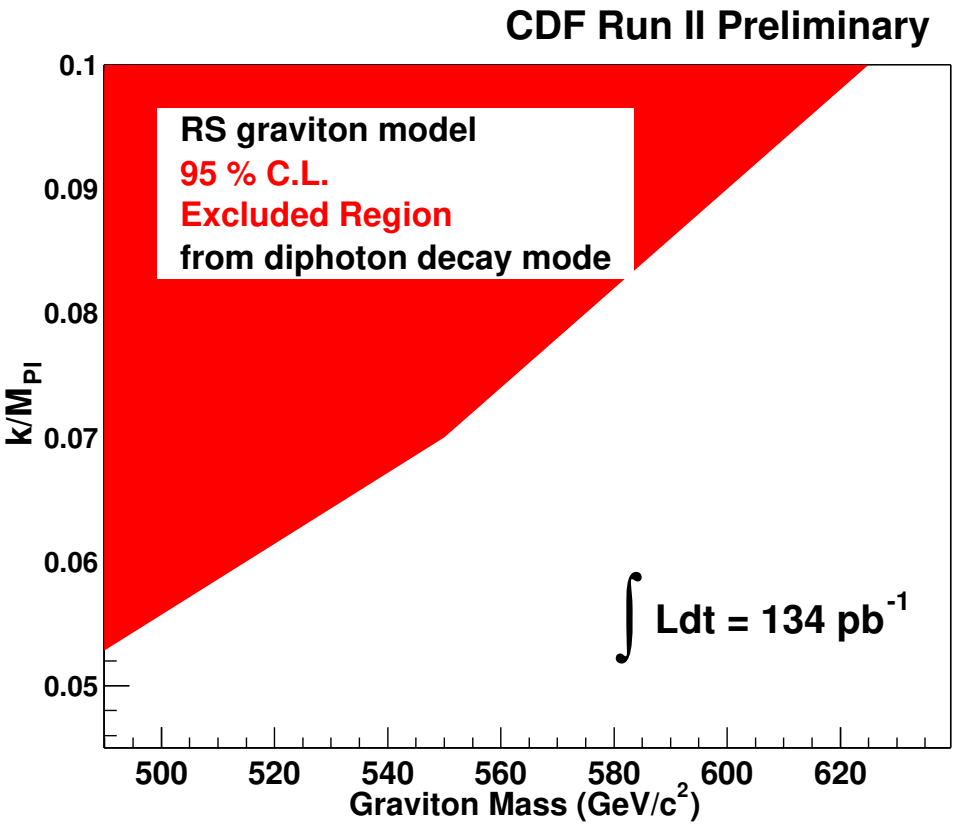
CDF Run II Preliminary (134 pb^{-1})

M_G	$3\times\sigma_{M_G}$	N_{signal}	$N_{background}$	N_{data}	$N_{95\%}$	$\sigma_{95\%}$
(GeV/c^2)	(GeV/c^2)					(pb)
200	13	4.5 ± 0.5	0	0	3.1	0.17
250	16.	1.7 ± 0.2	0	0	3.1	0.16
300	19.	0.80 ± 0.08	0	0	3.1	0.15
350	23.	0.46 ± 0.05	0	0	3.1	0.13
400	22.	0.27 ± 0.03	0	0	3.1	0.11
450	31.	0.16 ± 0.02	0	0	3.1	0.11
500	36.	0.10 ± 0.01	0	0	3.1	0.093
550	39.	0.062 ± 0.006	0	0	3.1	0.094
600	40.	0.039 ± 0.004	0	0	3.1	0.089
650	45.	0.024 ± 0.002	0	0	3.1	0.088
700	46.	0.014 ± 0.001	0	0	3.1	0.086
750	48.	0.0087 ± 0.0009	0	0	3.1	0.084
800	51.	0.0054 ± 0.0005	0	0	3.1	0.081
850	60.	0.0029 ± 0.0003	0	0	3.1	0.084
900	68.	0.0018 ± 0.0002	0	0	3.1	0.080
950	67.	0.00099 ± 1.e-04	0	0	3.1	0.082

95 % cross-section limit



Exclusion Plane



Future Plans

Aim to bless at next Exotics meeting!
Combine results with other RS model search channels

Extra plots:

Efficiency for a single photon from MC

The cuts were then placed sequentially, in the order shown below.

Sequential efficiency = $\frac{N_\gamma_passing_cut_and_all_previously_applied_cuts}{N_\gamma_passing_acceptance_cuts}$

Sequential efficiency for applying the selection criteria per photon (includes conversions and non-conversions)								
M_G	CES X,Z	Had/em	CalIso	CES χ^2	N3D	Pt<1	Sum Pt	2nd CES
200	88.8±0.5	87.5±0.5	83.1±0.5	78.5±0.6	73.7±0.6	72.1±0.7	71.1±0.7	70.3±0.7
250	88.7±0.5	85.8±0.5	82.0±0.6	77.2±0.7	72.0±0.7	70.7±0.7	68.8±0.7	68.1±0.7
300	86.7±0.6	83.9±0.6	80.0±0.7	76.4±0.7	72.0±0.8	70.4±0.8	69.2±0.8	68.6±0.8
350	85.6±0.6	82.2±0.7	78.9±0.7	74.9±0.7	69.9±0.8	68.9±0.8	67.4±0.8	66.5±0.8
400	85.1±0.6	83.0±0.7	79.6±0.7	77.1±0.7	70.3±0.8	68.8±0.8	68.5±0.8	68.0±0.8
450	83.6±0.7	81.3±0.7	77.5±0.8	75.3±0.8	69.3±0.8	67.7±0.9	66.9±0.9	66.2±0.9
500	83.7±0.7	81.2±0.7	78.6±0.8	75.9±0.8	68.9±0.9	67.2±0.9	66.1±0.9	65.6±0.9
550	81.8±0.7	79.5±0.8	77.0±0.8	74.7±0.8	68.0±0.9	66.4±0.9	65.7±0.9	65.1±0.9
600	84.2±0.7	81.9±0.7	79.0±0.8	76.0±0.8	68.9±0.9	67.6±0.9	66.9±0.9	66.5±0.9
650	83.4±0.7	81.1±0.7	78.5±0.8	75.8±0.8	68.3±0.9	66.9±0.9	66.3±0.9	65.7±0.9
700	81.2±0.8	79.6±0.8	76.0±0.9	73.4±0.9	66.9±0.9	65.2±1.0	65.0±1.0	64.3±1.0
750	81.6±0.7	79.8±0.8	76.9±0.8	74.0±0.8	67.7±0.9	65.5±0.9	65.0±0.9	64.5±0.9
800	81.6±0.7	80.2±0.8	77.3±0.8	74.3±0.8	67.5±0.9	65.1±0.9	64.6±0.9	64.2±0.9
850	81.5±0.7	80.1±0.8	77.4±0.8	73.2±0.8	67.1±0.9	65.0±0.9	64.5±0.9	63.7±0.9
900	82.2±0.8	80.4±0.8	77.1±0.8	73.1±0.9	66.5±0.9	64.3±1.0	63.8±1.0	63.3±1.0
950	80.0±0.7	78.4±0.8	76.1±0.8	71.5±0.8	65.5±0.9	62.8±0.9	62.3±0.9	61.9±0.9

It is interesting to consider the photons into those that do/ do not convert.

Efficiency for individual selction criteria

The efficiency of applying each selection criterion to a single photon was determined at each graviton mass using the RS model Monte Carlo.
For each Monte Carlo sample the events were looped over to find photons which passed the acceptance cuts (central EM objects with E_T>25 GeV).
The cuts were then placed sequentially, in the order shown below.
Efficiency of cut given photon passes previous cuts =
$$\frac{N_\gamma_passing_cut_and_all_previously_applied_cuts}{N_\gamma_passing_all_previous_cuts}$$

Sequential efficiency for applying the selection criteria								
M_G	CES X,Z	Had/em	CalIso	CES χ^2	N3D	Pt<1	Sum Pt	2nd CES
200	88.8±0.5	98.5±0.2	95.0±0.3	94.4±0.4	94.0±0.4	97.8±0.2	98.5±0.2	98.9±0.2
250	88.7±0.5	96.7±0.3	95.6±0.3	94.1±0.4	93.3±0.4	98.2±0.2	97.3±0.3	99.0±0.2
300	86.7±0.6	96.7±0.3	95.4±0.4	95.5±0.4	94.2±0.5	97.8±0.3	98.3±0.3	99.1±0.2
350	85.6±0.6	96.1±0.4	96.0±0.4	94.9±0.4	93.3±0.5	98.6±0.2	97.9±0.3	98.7±0.2
400	85.1±0.6	97.5±0.3	96.0±0.4	96.8±0.3	91.1±0.6	98.0±0.3	99.5±0.1	99.2±0.2
450	83.6±0.7	97.3±0.3	95.4±0.4	97.1±0.3	92.0±0.6	97.6±0.3	98.9±0.2	98.9±0.2
500	83.7±0.7	97.0±0.3	96.9±0.4	96.5±0.4	90.8±0.6	97.6±0.3	98.3±0.3	99.2±0.2
550	81.8±0.7	97.2±0.3	96.9±0.4	97.1±0.4	90.9±0.6	97.7±0.3	98.9±0.2	99.2±0.2
600	84.2±0.7	97.4±0.3	96.4±0.4	96.3±0.4	90.6±0.6	98.1±0.3	99.0±0.2	99.4±0.2
650	83.4±0.7	97.2±0.3	96.9±0.4	96.6±0.4	90.0±0.7	98.0±0.3	99.1±0.2	99.2±0.2
700	81.2±0.8	98.0±0.3	95.5±0.5	96.5±0.4	91.2±0.7	97.4±0.4	99.6±0.2	98.9±0.3
750	81.6±0.7	97.8±0.3	96.4±0.4	96.2±0.4	91.6±0.6	96.7±0.4	99.2±0.2	99.1±0.2
800	81.6±0.7	98.2±0.3	96.4±0.4	96.1±0.4	90.8±0.6	96.4±0.4	99.2±0.2	99.4±0.2
850	81.5±0.7	98.2±0.3	96.6±0.4	94.6±0.5	91.7±0.6	96.9±0.4	99.2±0.2	98.8±0.3
900	82.2±0.8	97.8±0.3	95.9±0.4	94.8±0.5	91.0±0.7	96.6±0.4	99.3±0.2	99.1±0.2
950	80.0±0.7	98.0±0.3	97.1±0.4	94.0±0.5	91.5±0.6	96.0±0.5	99.1±0.2	99.3±0.2

Interesting to split efficiency per photon into those that convert and those that do not.

Efficiency for individual selction criteria- for conversions

The efficiency of applying each selection criterion to a single photon which converted was determined at each graviton mass using the RS model Monte Carlo.

For each Monte Carlo sample the events were looped over to find photons which passed the acceptance cuts (central EM objects with E_T>25 GeV). The cuts were then placed sequentially, in the order shown below.

Efficiency of cut given photon passes previous cuts =
$$\frac{N_\gamma_passing_cut_and_all_previously_applied_cuts}{N_\gamma_passing_all_previous_cuts}$$

Sequential efficiency for applying the selection criteria to converted photons

M_G	CES X,Z	Had/em	CalIso	CES χ^2	N3D	Pt<1	Sum Pt	2nd CES
200	92.8 ± 1.	100. ± 0.	91.4 ± 1.	51.8 ± 3.	21.3 ± 3.	0.00 ± 0.	-	-
250	96.5 ± 0.9	100. ± 0.	93.2 ± 1.	53.9 ± 3.	14.2 ± 3.	0.00 ± 0.	-	-
300	91.5 ± 2.	99.7 ± 0.3	93.0 ± 2.	55.4 ± 4.	17.2 ± 4.	0.00 ± 0.	-	-
350	95.9 ± 1.	100. ± 0.	94.2 ± 1.	54.5 ± 3.	10.7 ± 3.	0.00 ± 0.	-	-
400	92.9 ± 2.	100. ± 0.	99.0 ± 0.6	76.5 ± 3.	11.1 ± 2.	5.98 ± 5.	100. ± 0.	100. ± 0.
450	90.4 ± 2.	100. ± 0.	94.1 ± 2.	76.9 ± 3.	13.6 ± 3.	0.00 ± 0.	-	-
500	96.0 ± 1.	100. ± 0.	98.0 ± 0.9	76.9 ± 3.	11.0 ± 2.	4.68 ± 5.	100. ± 0.	100. ± 0.
550	91.3 ± 2.	100. ± 0.	98.3 ± 0.9	84.7 ± 3.	13.6 ± 3.	0.00 ± 0.	-	-
600	94.5 ± 2.	100. ± 0.	97.1 ± 1.	81.1 ± 3.	10.9 ± 2.	0.00 ± 0.	-	-
650	97.1 ± 1.	100. ± 0.	95.7 ± 1.	88.3 ± 2.	9.60 ± 2.	0.00 ± 0.	-	-
700	93.1 ± 2.	100. ± 0.	97.1 ± 1.	90.3 ± 2.	12.6 ± 3.	16.7 ± 9.	100. ± 0.	100. ± 0.
750	93.0 ± 2.	100. ± 0.	96.1 ± 1.	89.7 ± 2.	16.2 ± 3.	0.00 ± 0.	-	-
800	94.2 ± 2.	100. ± 0.	99.1 ± 0.6	92.1 ± 2.	16.7 ± 3.	3.14 ± 3.	100. ± 0.	100. ± 0.
850	97.1 ± 1.	100. ± 0.	98.1 ± 1.0	90.7 ± 2.	19.0 ± 3.	7.91 ± 5.	100. ± 0.	100. ± 0.
900	95.6 ± 2.	100. ± 0.	99.3 ± 0.6	93.1 ± 2.	19.7 ± 3.	9.36 ± 5.	100. ± 0.	100. ± 0.
950	96.5 ± 1.	100. ± 0.	98.9 ± 0.7	94.0 ± 2.	22.6 ± 3.	2.28 ± 2.	100. ± 0.	100. ± 0.

Efficiency for individual selction criteria- for non-conversions

The efficiency of applying each selection criterion to a single photon which did not convert was determined at each graviton mass using the RS model Monte Carlo.

For each Monte Carlo sample the events were looped over to find photons which passed the acceptance cuts (central EM objects with E_T>25 GeV). The cuts were then placed sequentially, in the order shown below.

Efficiency of cut given photon passes previous cuts =
$$\frac{N_\gamma_passing_cut_and_all_previously_applied_cuts}{N_\gamma_passing_all_previous_cuts}$$

Sequential efficiency for applying the selection criteria to non-converted photons

M_G	CES X,Z	Had/em	CalIso	CES χ^2	N3D	Pt<1	Sum Pt	2nd CES
200	89.7±0.5	98.6±0.2	96.4±0.3	99.5±0.1	99.5±0.1	99.3±0.2	98.4±0.2	98.9±0.2
250	90.7±0.5	97.4±0.3	96.0±0.4	99.2±0.2	99.1±0.2	99.6±0.1	97.4±0.3	99.0±0.2
300	89.9±0.6	97.1±0.4	96.9±0.4	0.0477±0.05	0.00±0.	divide 0	98.1±0.3	99.2±0.2
350	91.6±0.6	96.9±0.4	97.1±0.4	99.5±0.2	99.1±0.2	99.8±0.09	97.8±0.3	98.7±0.2
400	92.0±0.6	98.7±0.2	96.5±0.4	99.1±0.2	99.0±0.2	99.6±0.2	99.5±0.2	99.2±0.2
450	92.2±0.6	99.0±0.2	96.6±0.4	99.1±0.2	99.0±0.2	99.4±0.2	99.0±0.2	99.0±0.2
500	92.4±0.6	99.6±0.1	97.7±0.4	99.0±0.2	99.4±0.2	99.5±0.2	98.0±0.3	99.3±0.2
550	93.4±0.6	99.7±0.1	98.3±0.3	98.7±0.3	99.3±0.2	99.8±0.1	98.7±0.3	99.1±0.2
600	94.0±0.6	99.9±0.07	97.7±0.4	98.0±0.3	98.7±0.3	99.8±0.1	98.9±0.3	99.3±0.2
650	94.0±0.6	99.9±0.08	98.4±0.3	97.4±0.4	99.8±0.1	99.8±0.1	99.0±0.3	99.2±0.2
700	93.8±0.6	99.9±0.07	98.0±0.4	97.0±0.5	99.3±0.2	99.9±0.07	99.5±0.2	99.0±0.3
750	94.6±0.5	100.±0.	98.3±0.3	97.1±0.4	99.3±0.2	99.9±0.09	99.2±0.2	99.3±0.2
800	93.9±0.6	100.±0.	98.4±0.3	96.6±0.5	99.3±0.2	99.9±0.09	99.1±0.3	99.5±0.2
850	94.2±0.6	100.±0.	98.5±0.3	95.0±0.6	99.6±0.2	100.±0.	98.9±0.3	99.0±0.3
900	96.1±0.5	100.±0.	98.5±0.3	94.8±0.6	98.8±0.3	99.9±0.1	99.2±0.3	99.2±0.3
950	94.6±0.6	99.4±0.2	99.0±0.3	93.8±0.7	99.5±0.2	99.8±0.1	98.9±0.3	99.2±0.2