

The WFI for IXO

OUTLOOK

- o the WFI
- o the HXI
- o the HTRS

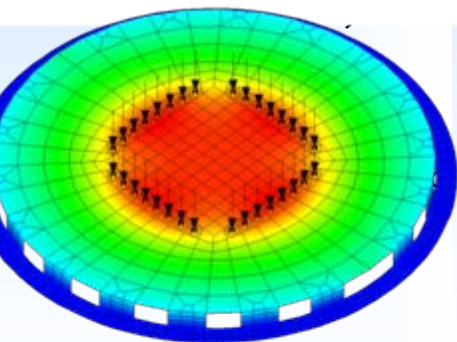
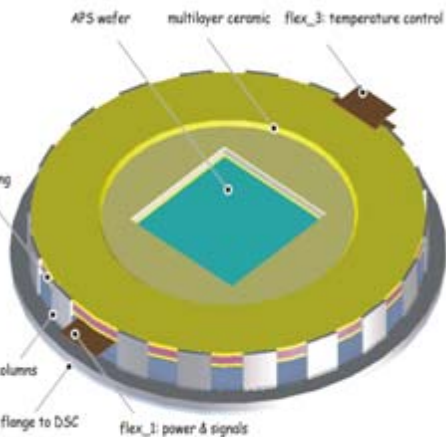
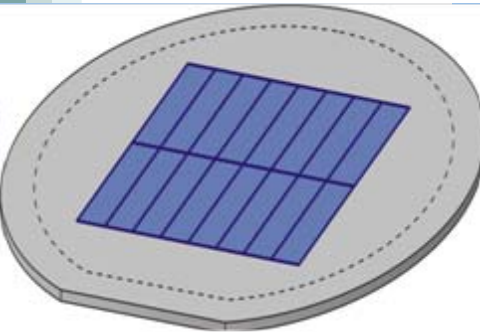
with contributions from

- MPE-HLL
- Politecnico di Milan
- University of Leicester

8.2 cm

8.2 cm

The Wide Field Imager



■ specifications

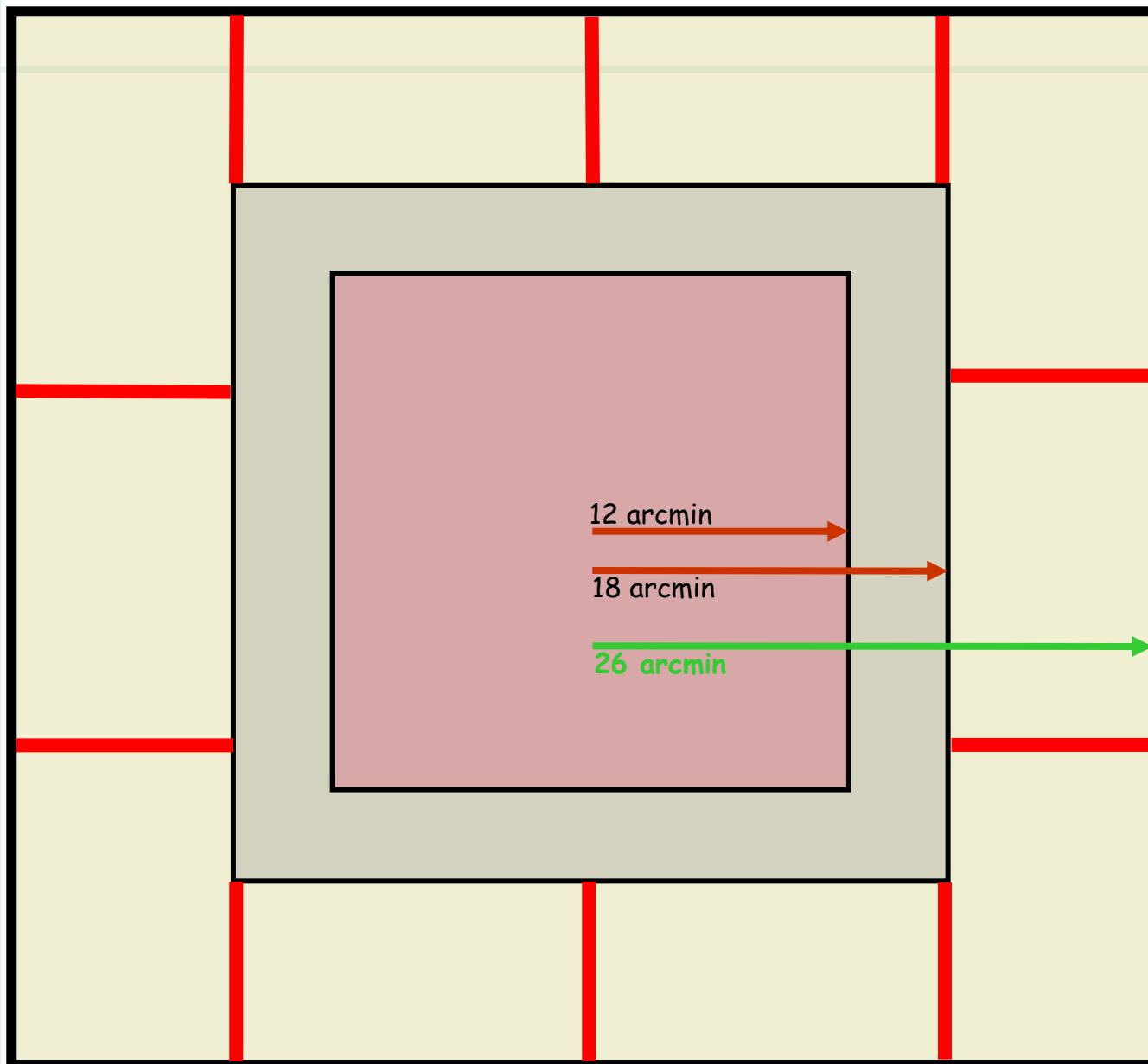
- field of view (min.)
 $> 12 \text{ arcmin } \varnothing$
 $> 7.2 \text{ cm } \varnothing$
- angular resolution
 $\leq 5 \text{ arcsec @ } 20 \text{ m}$
 point spread function
 $500 \mu\text{m}$
- energy range
 $0.1 \dots 15 \text{ keV}$
- energy resolution
 $< 130 \text{ eV @ } 6 \text{ keV}$
- count rate capability
 10 kcps
 $< 1\% \text{ pileup}$
- hard X-ray camera option
- spacecraft constraints

■ WFI parameters

- detector format
 $1024 \times 1024 \text{ pix}$
 $14 \text{ arcmin @ } 25\text{m}$
 $18 \text{ arcmin @ } 20\text{m}$
- pixel size
 $100 \times 100 \mu\text{m}^2$
 $75 \times 75 \mu\text{m}^2$
- thin entrance window
- detector thickness
 $450 \mu\text{m}$
- low electronic noise
 $\ll 4 \text{ el. ENC}$
- fast readout
 $2 \mu\text{sec} / \text{pixel-row}$
- window mode
 $32 \times 1024 \text{ pixel}$
- monolithic device
- low power
 $< 10 \text{ W}$
- high temperature
 $\geq -60 \text{ }^\circ\text{C}$

- min. - max. pixel size
 $24 \times 24 \mu\text{m}^2$
 $500 \times 500 \mu\text{m}^2$
- thin entrance window
 for $E > 50 \text{ eV}$
- integrated optical blocking filter
 for $E > 50 \text{ eV}$
- detector thickness
 $150 - 450 \mu\text{m}$
- low electronic noise
 $\ll 3.5 \text{ el. ENC}$
 $\Rightarrow \Delta E = 130 \text{ eV (FWHM)}$
- fast readout
 $2 \mu\text{sec} / \text{row}$
 $\text{@ } 1000 \text{ frames per second}$
- window mode
 $32 \times 1024 \text{ pixel, i.e.}$
 $32 \mu\text{s per frame}$
- monolithic device
- low power
 $< 8 \text{ W}$
- high temperature
 $\geq -40 \text{ }^\circ\text{C}$

Potential WFI layouts (FL 20 m)



WFI expansion with e.g.
MOS-type CCDs:
pixel size: $100 \times 100 \mu\text{m}^2$,
format: 1024×512 pixel
sensitive thickness: $200 \mu\text{m}$

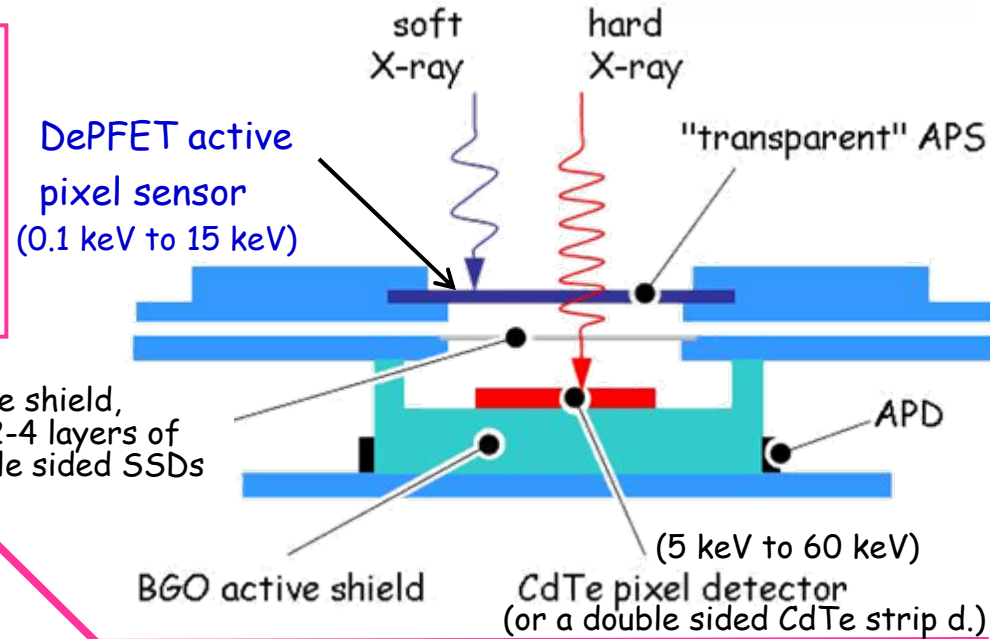
„actual“ WFI monolithic
DePFET APS, BI,
sensitive thickness: $450 \mu\text{m}$,
format: 1024×1024 pixel,
FOV: 18 arcmin @ 20 m FL ,
pixel size: $100 \times 100 \mu\text{m}^2$

„original“ WFI: monolithic
DePFET APS, BI,
sensitive thickness: $450 \mu\text{m}$
format: 712×712 pixel
FOV: 12 arcmin @ 20 m FL ,
pixel size: $100 \times 100 \mu\text{m}^2$

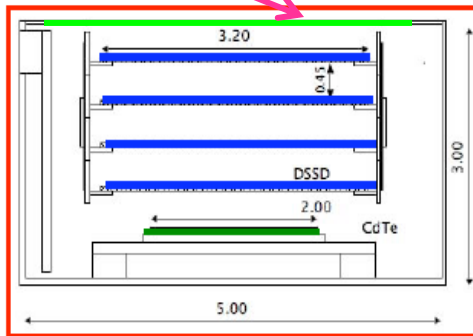
WFI and HXI on XEUS



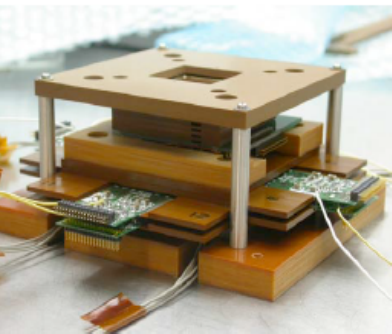
one possibility of accomodation the DePFET WFI and the HXI within a single camera housing:
 For X-rays above typ. 20 keV the DePFET becomes transparent and the X-rays will eventually interact with the double sided SSD or the CdTe detector



wide field imager (DePFET)



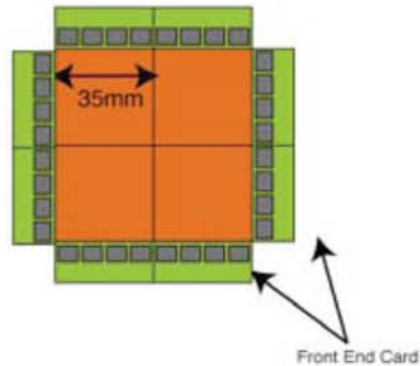
Test set-up for the combined double sided SSD and the CdTe detector. The distance between the WFI and the CdTe detector of the HXI is approximately 2 cm.



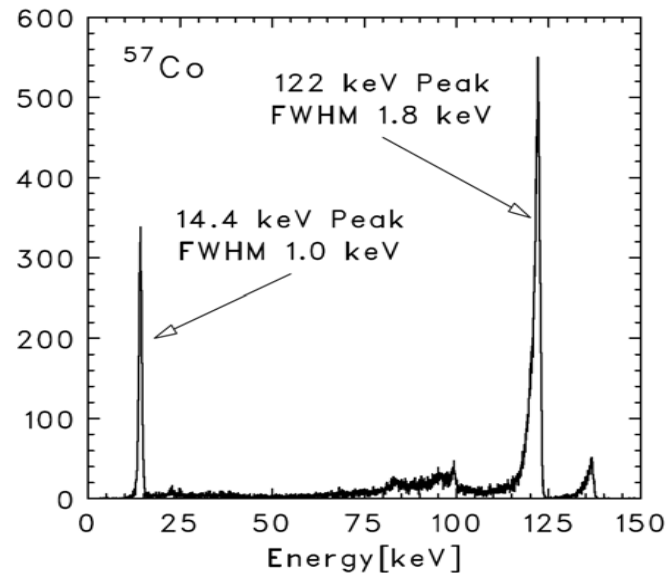
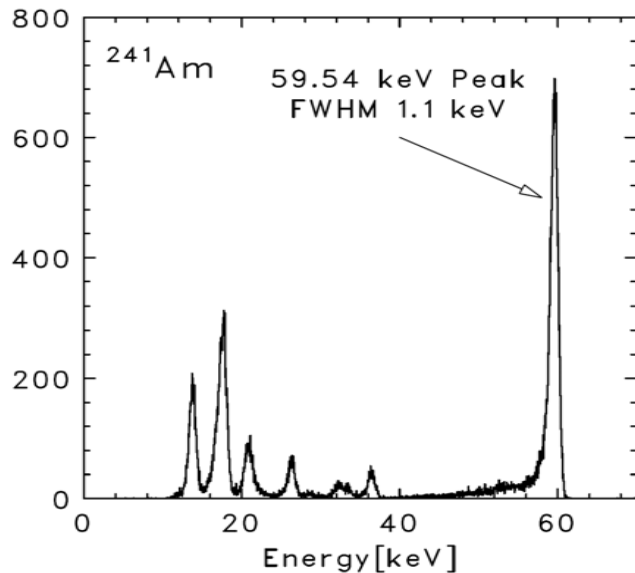
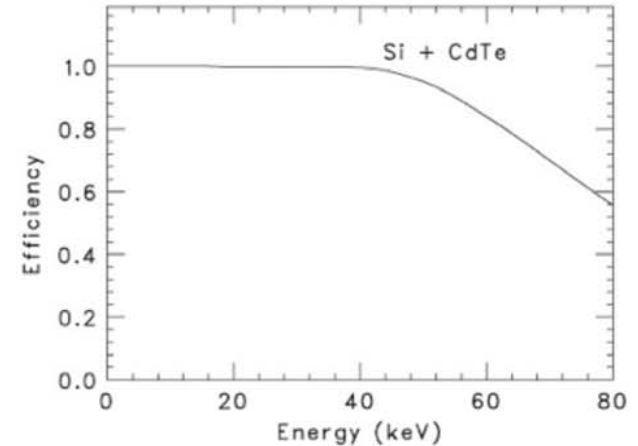
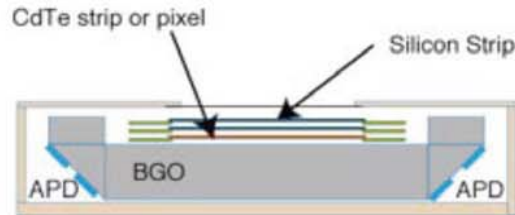
- Double Side Silicon Strip (4 layers)
- Double Side CdTe Strip (1 layer)

HXI layout and performance

Top View



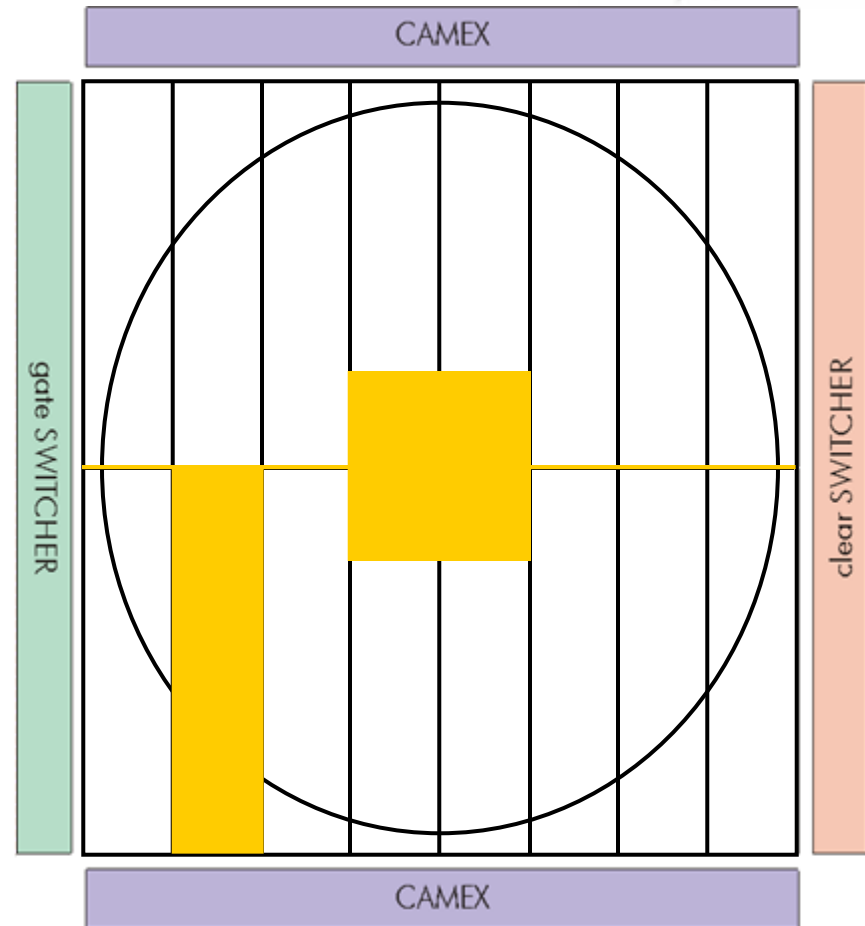
Side View



DEPFETs for the XEUS WFI



1. Flexible operating modes
2. low power dissipation (less than 2 W in 100 cm², DePFETs only)
3. Fano limited energy resolution from 0.5 keV to 30 keV
4. Spatial resolution better than 20 μm @ 100 μm pixel size
5. Homogeneous radiation entrance window
6. Intrinsic radiation hardness, no charge transfer needed
7. ENC was lowered to 0.2 e⁻ rms with RNDR
8. Thin optical "Blocking Filter" can be directly integrated
9. Operation at "warm temperatures", e.g. - 40 °C

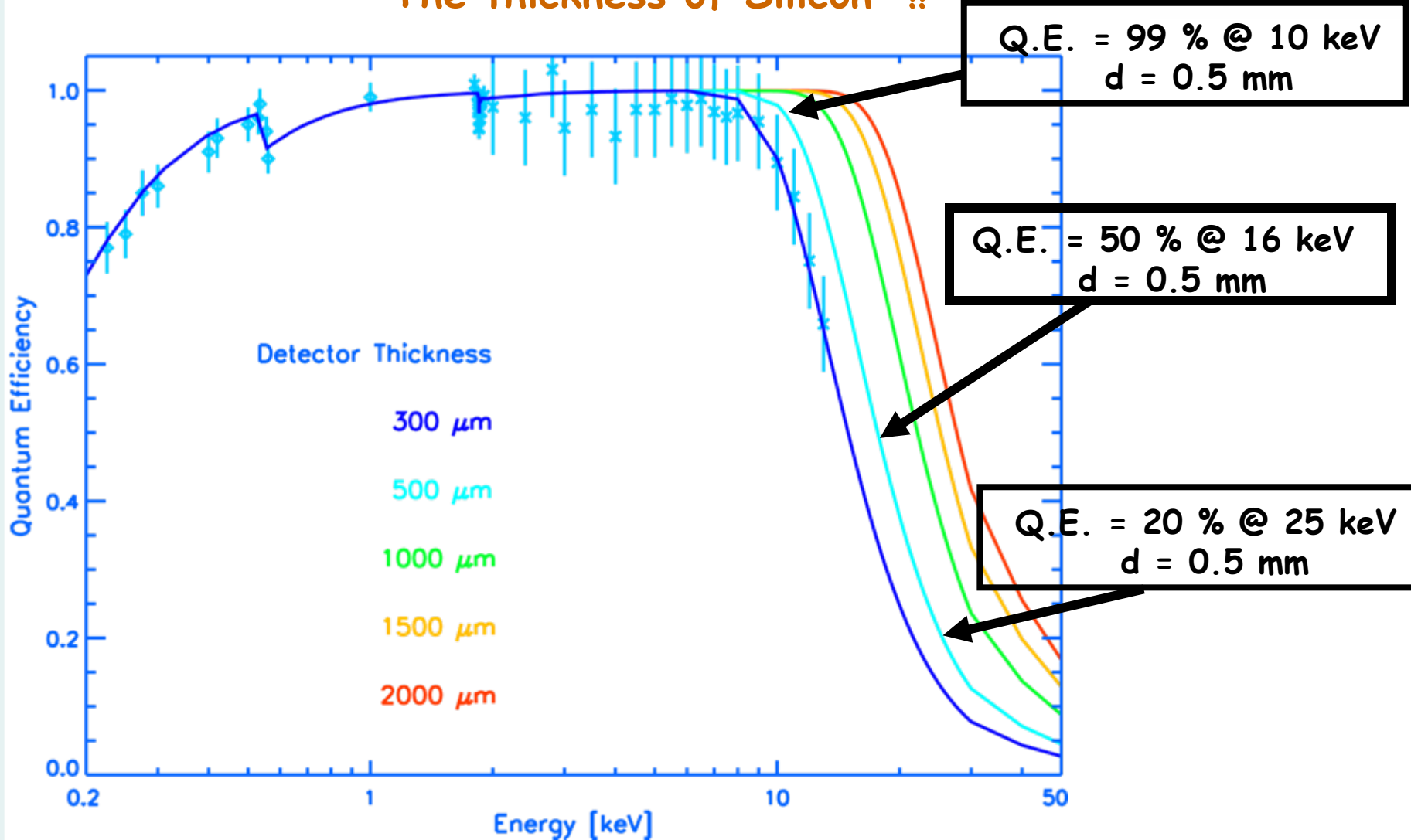


focal length of 35 m

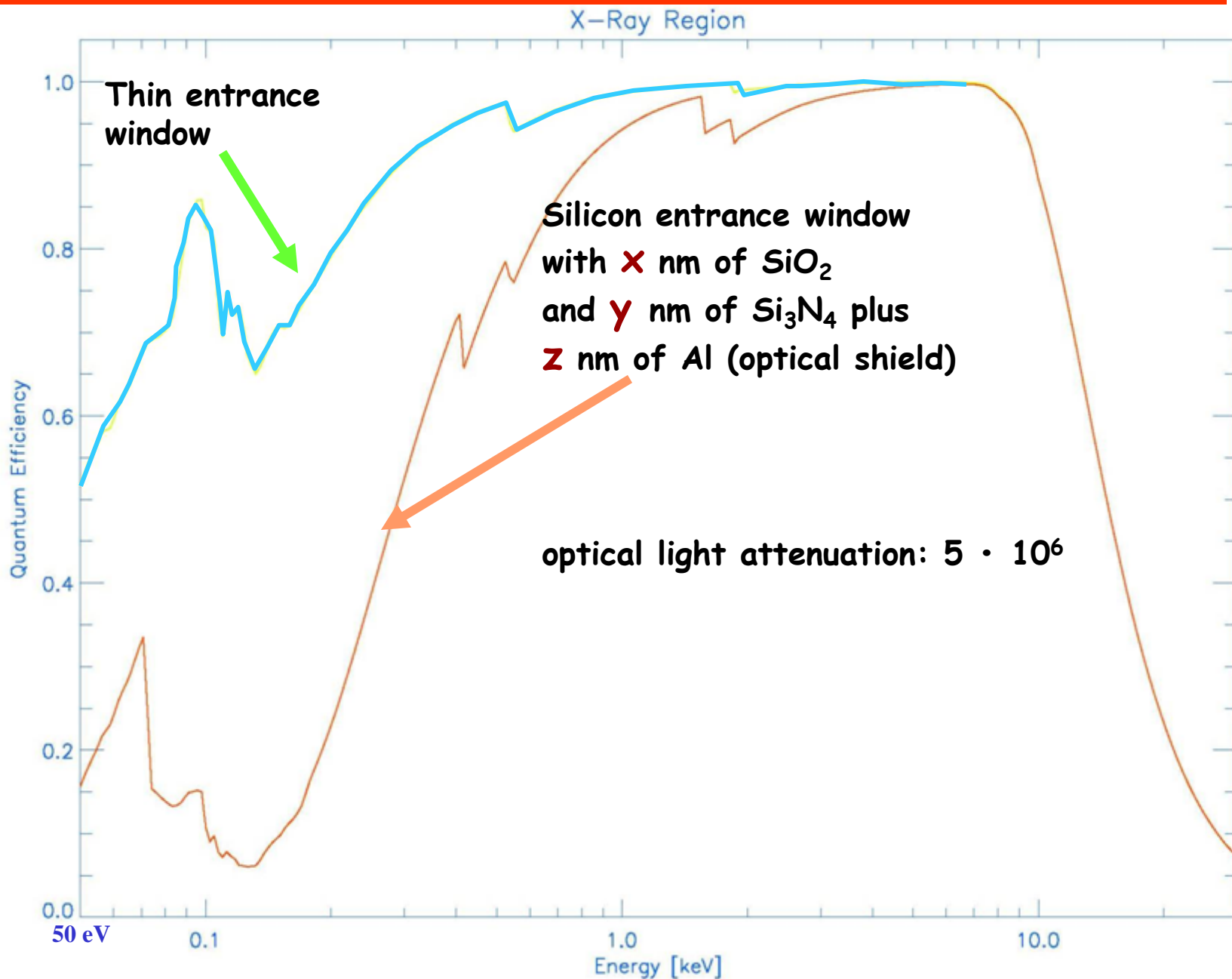
What is limiting the quantum efficiency for photons ?



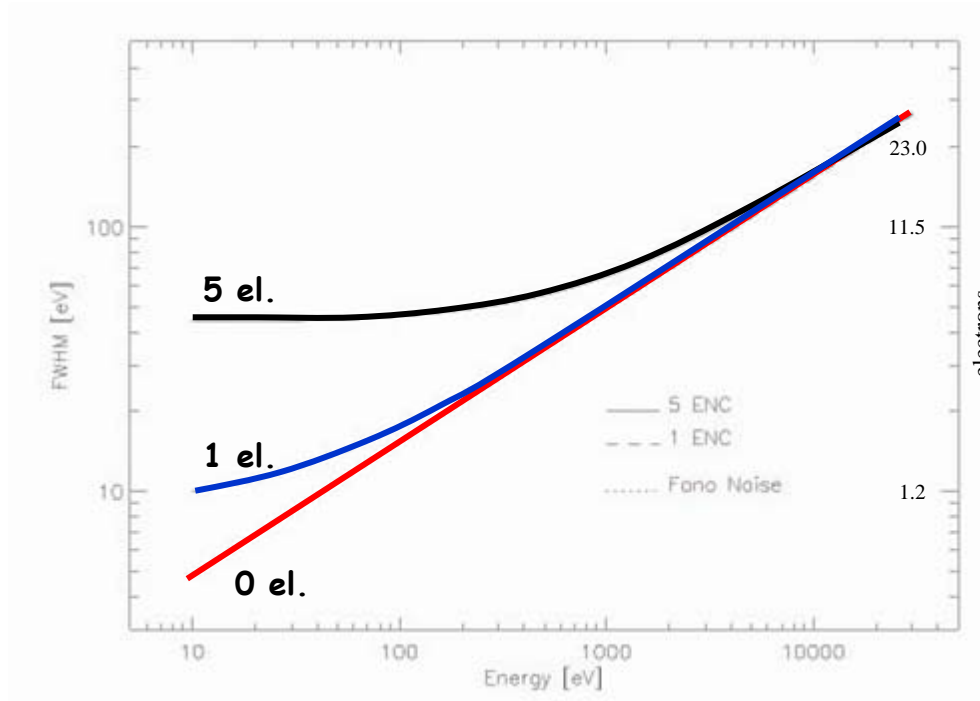
The thickness of Silicon !!



Monolithic Implementation of optical blocking filters



Limits of energy resolution of e.g. X-rays



$$ENC_{\text{fano}}^2 = \frac{F \cdot E}{w}$$

$$ENC^2 = A_1 \left(\alpha \frac{2kT}{g_m} C_{\text{tot}}^2 \right) \frac{1}{\tau} + \left[A_2 \left(2\pi a_f C_{\text{tot}}^2 + \frac{b_f}{2\pi} \right) \right] + A_3 \left(qI_l + \frac{2kT}{R_f} \right) \tau$$

$$ENC_{\text{tot}}^2 = ENC_{\text{el}}^2 + ENC_{\text{fano}}^2 + \dots = (0.2 \text{ e}^- - 5 \text{ e}^-)^2$$

DePFET Active Pixel Sensors

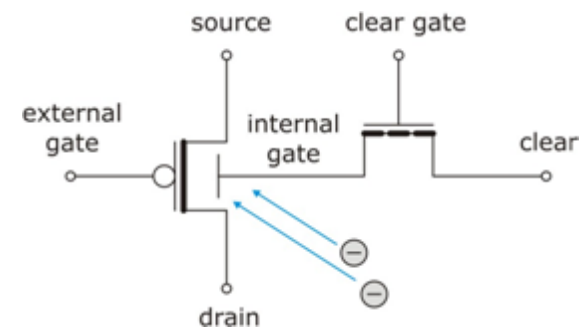
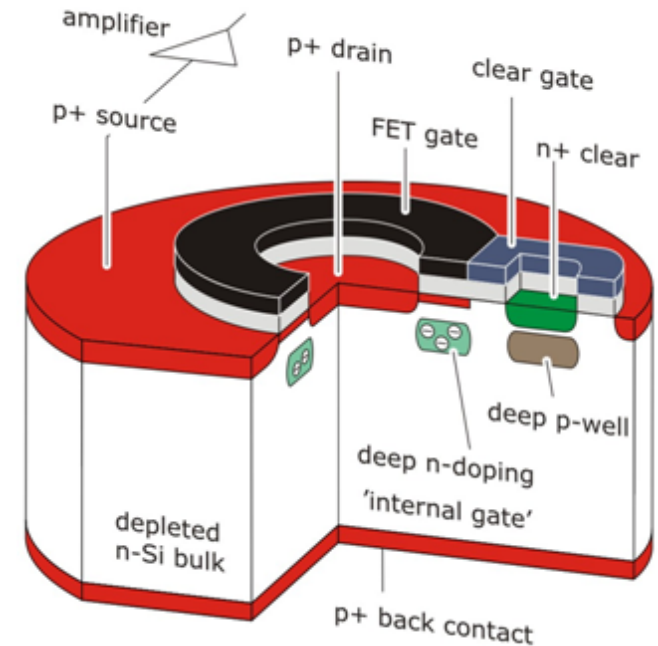


p-FET on depleted n-bulk

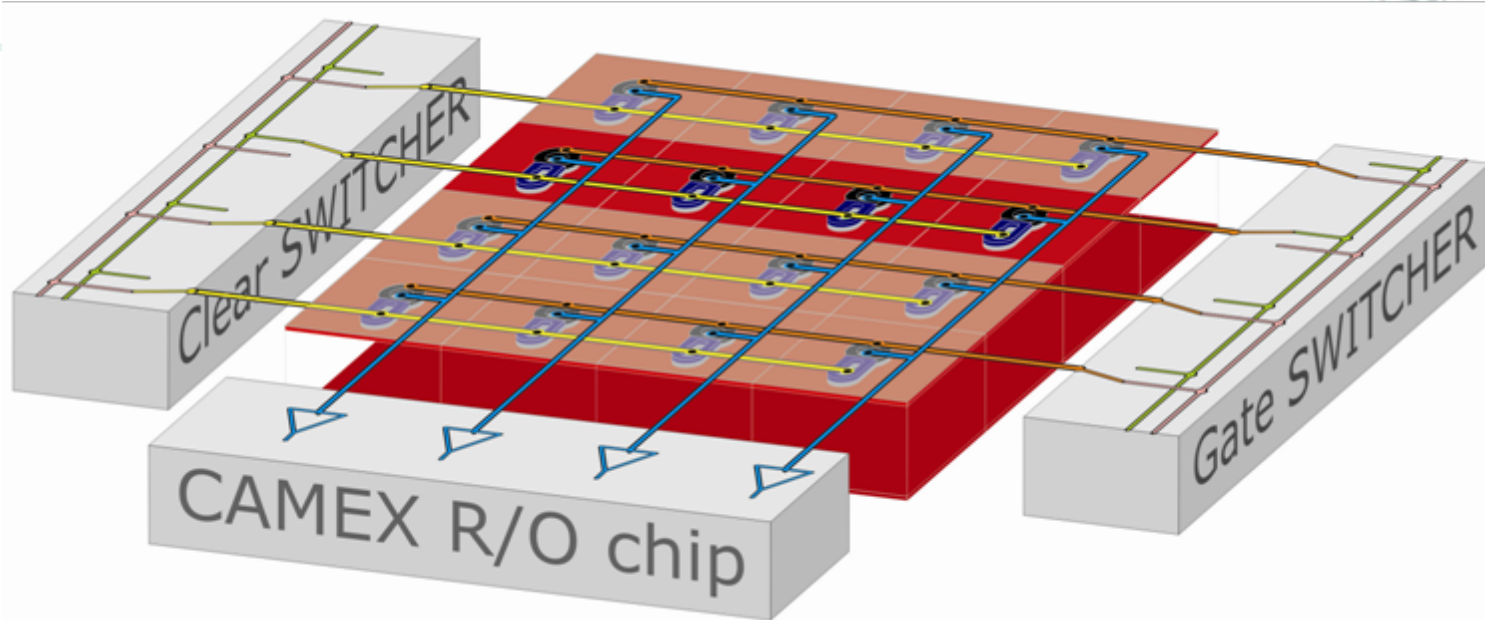
- signal charge collected in potential minimum below FET channel
- steers transistor current (1 el. \sim 300 pA)

combined sensor & amplifier

- low capacitance and noise
 - » **excellent spectroscopic performance**
- complete clearing of signal charge
 - » **no reset noise**
- non-destructive readout
 - » **potential of repetitive readout**
- charge storage capability
 - » **readout on demand**
- full depletion
 - » **backside illumination**
 - » **thin entrance window**



DEPFET Active Pixel Sensor



◆ matrix organisation

- common back contact
 - » thin, homogeneous entrance window
 - » fill factor 100 %
- row-wise connection of gate, clear, clear gate
- column-wise connection of source / drain
 - » individually addressable pixels
 - » windowing option

◆ operation philosophy

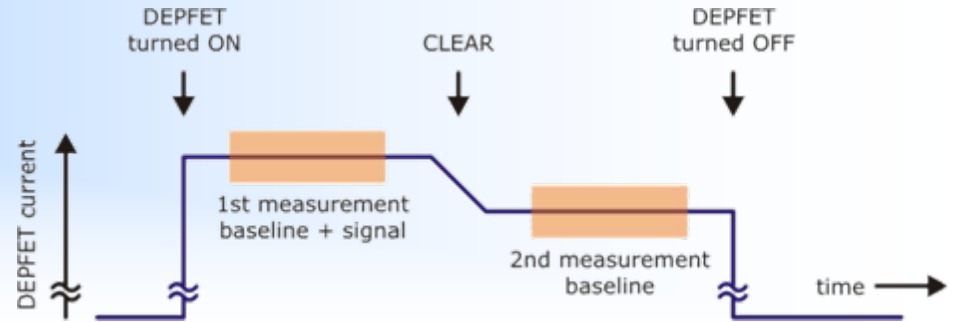
- one active row
- all other pixels turned off
 - » low power consumption
- all operations in a row in parallel
 - » fast processing

DEPFET readout



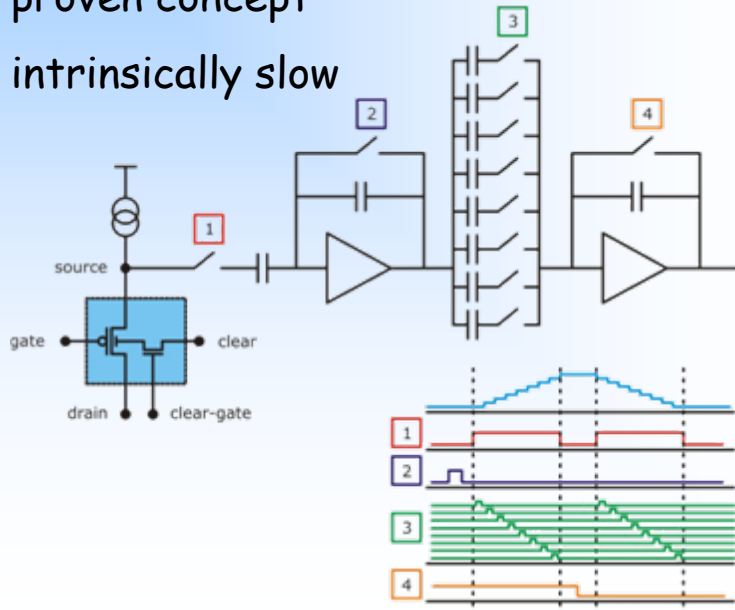
principle

- measure signal + baseline current
- 'clear' = remove signal electrons
- measure baseline current



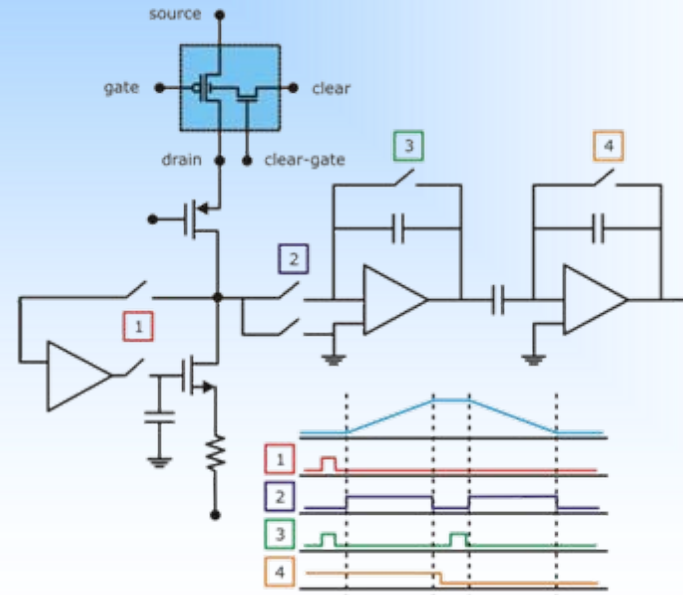
CAMEX

- source follower, 8-fold CDS
- CMX64 in operation, CMX128 in design
- ✓ proven concept
- ✗ intrinsically slow

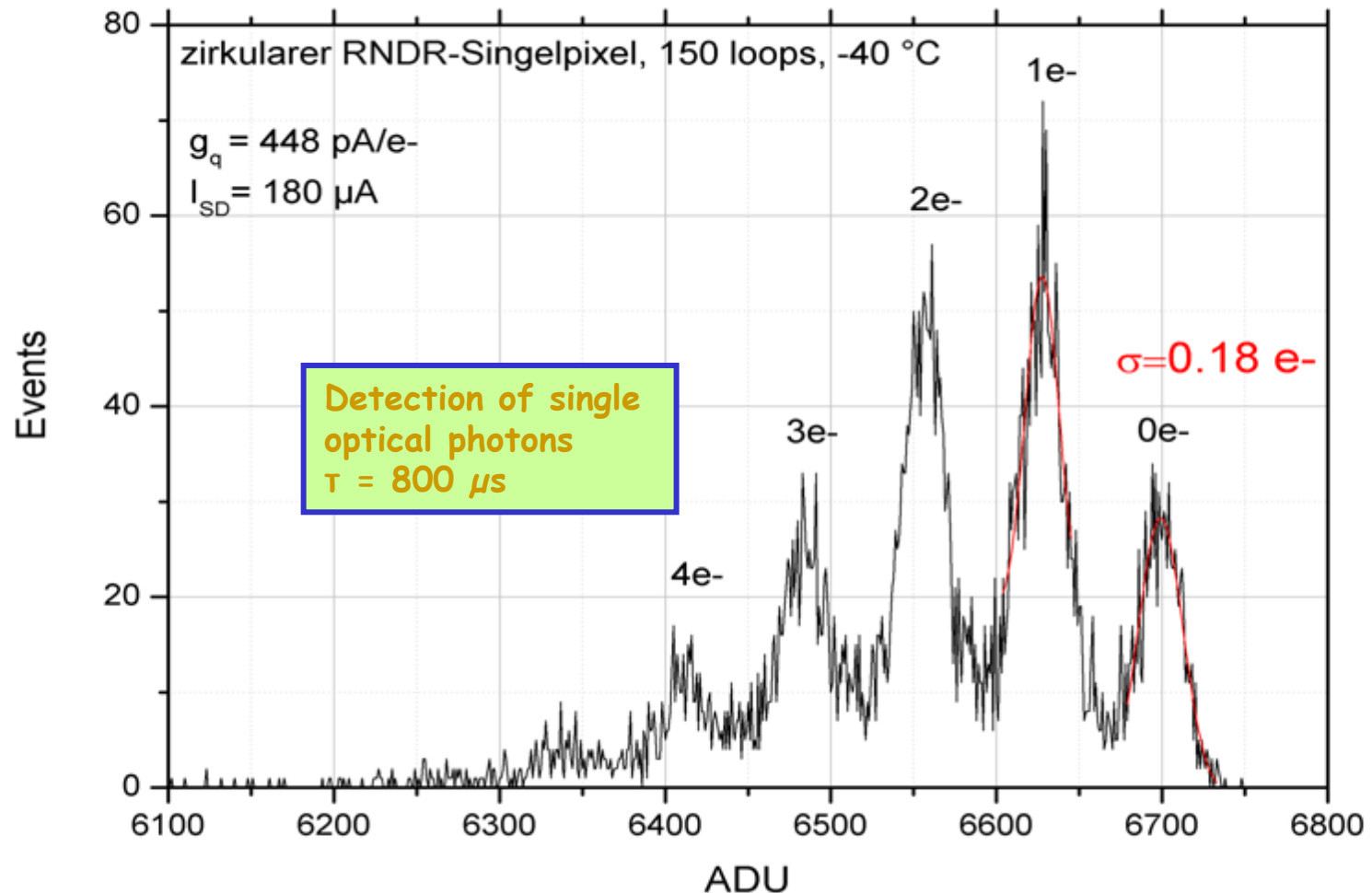


VELA (in collaboration with Politecnico di Milano)

- drain current integration/deintegration
- 64-channel prototype under test
- ✓ high readout speed
- ✗ more sensitive to flatband shifts



First Results with 64-channel Vela ASIC



DEPFET APS - prototypes

◆ DEPFET pixel $75 \mu\text{m}$ \square

• geometry

$$W = 47 \mu\text{m}$$

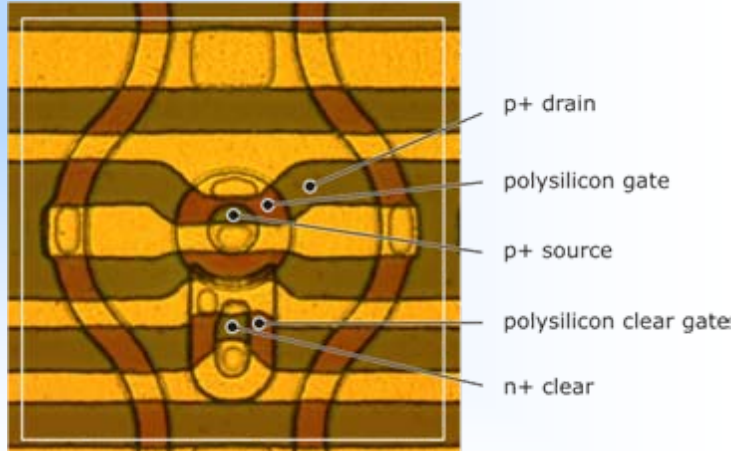
$$L = 5 \mu\text{m}$$

• current

$$I = 30 \mu\text{A}$$

• sensitivity

$$g_Q = 300 \text{ pA/el}$$



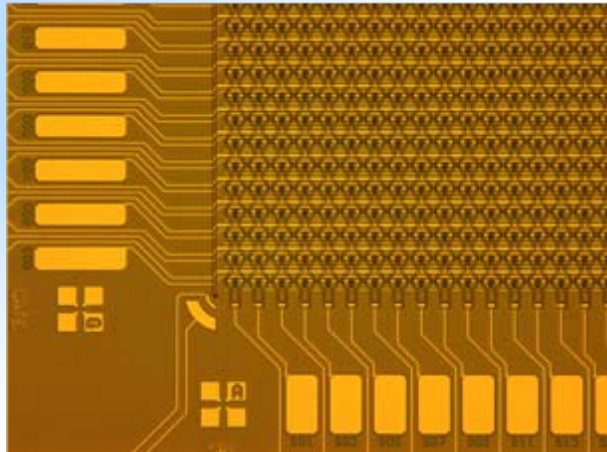
◆ 256 x 256 APS

dedicated technology

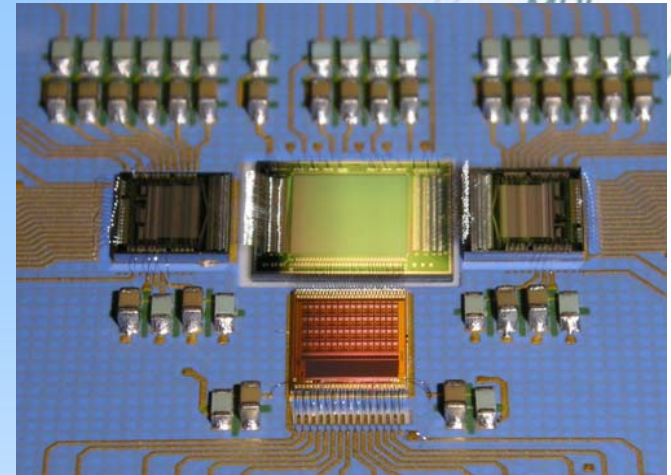
- 2 polysilicon layers
- 2 metal layers

leakage current level

- 100 pA/cm^2
- 16 fA/pixel



◆ APS hybrid



CAMEX64 readout ASIC

- 64 channel low noise voltage amplifier
- 64 channel 8-fold CDS filter
- 64/1 analog multiplexer
- source follower gain $3.7 \mu\text{V/el}$.

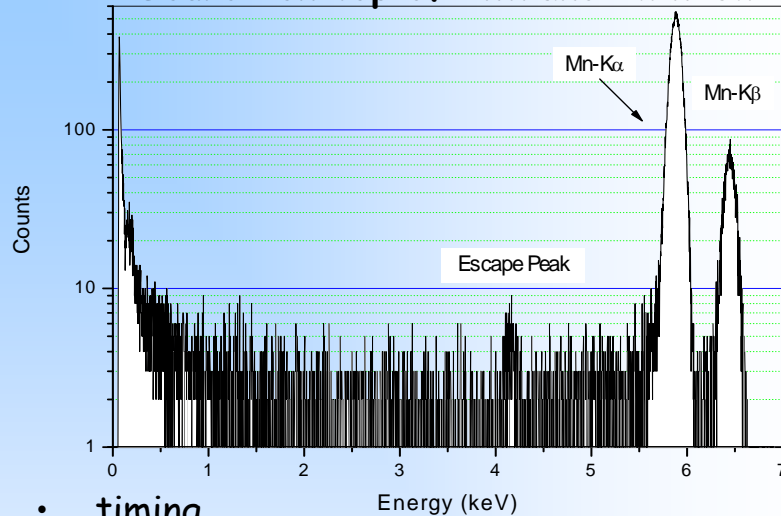
2 x Switcher II control ASIC

- 64 channel control chip
- 2 ports / channel
- integrated sequencer
- high voltage CMOS process
> 20 V p-p
- 50 MHz clock

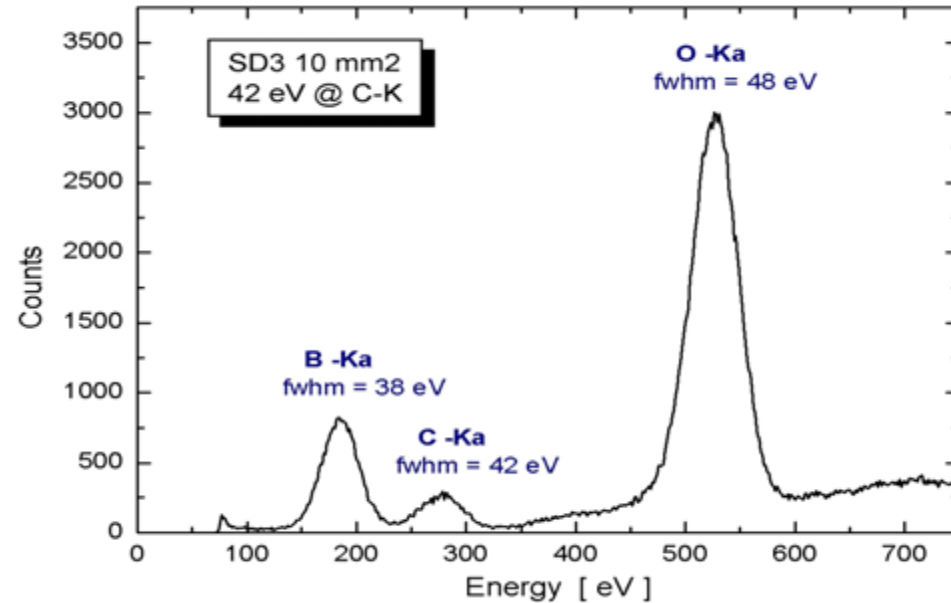
DEPFET APS - performance

low energy response

"Backside" illumination:
Source on top of entrance window



- timing
 - 4 $\mu\text{sec}/\text{row}$ \leftrightarrow 64 $\mu\text{sec}/32 \times 512$ sensor
- room temperature
 - 220 eV FWHM @ 5.9 keV (singles)
- moderate cooling -40 °C
 - 126 eV FWHM @ 5.9 keV (singles)
 - 132 eV FWHM @ 5.9 keV (all events)
- extrinsic speed & resolution limitations



yield & homogeneity

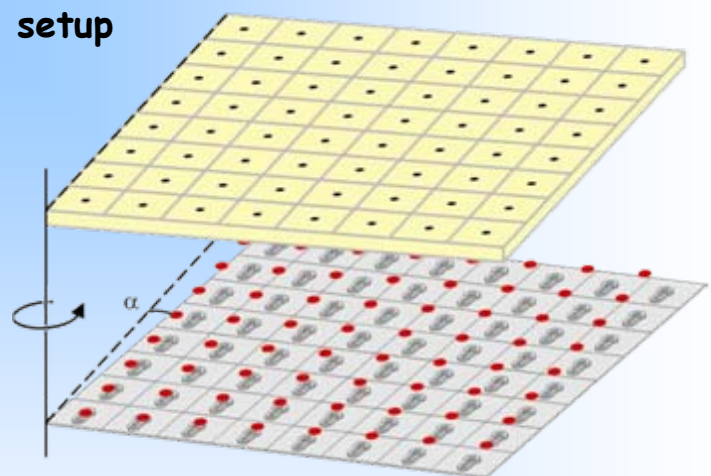
- defect pixels
 - 2 in 45 devices ($> 10^6$ pixels)
 - pixel yield > 0.99999
- dispersions
 - offset $< 2\%$ (of Mn-K α)
 - gain $< 5\%$
 - noise $< 10\%$

DEPFET APS - mesh experiment

method

- irradiation through tilted periodic mesh
- Moire pattern
- X-ray interaction position with subpixel resolution

setup

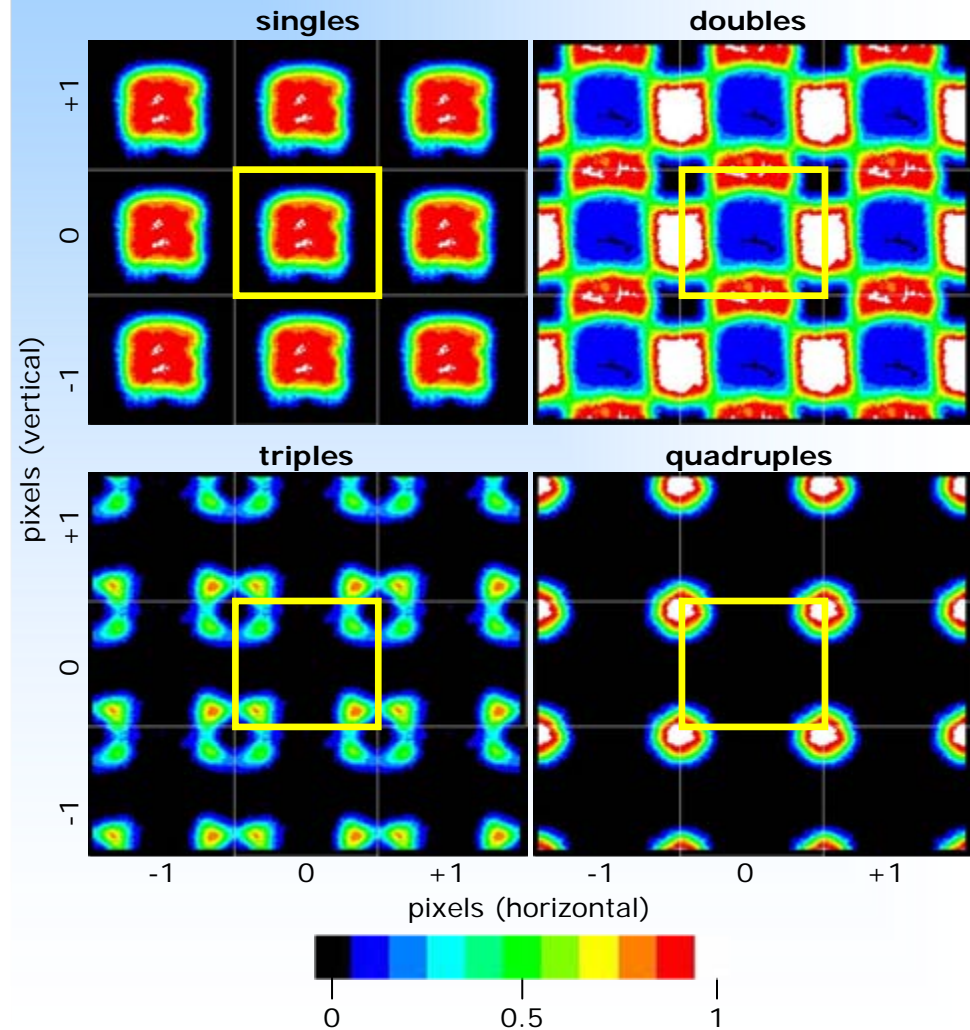


- mesh 10 μm gold
5 μm holes
150 μm pitch
- X-rays Cr-K $_{\alpha}$ (5.4 keV)

example

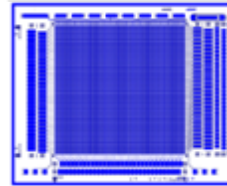
- variation of multiple pixel hit patterns with back contact voltage
- V_{back} =

-400 V

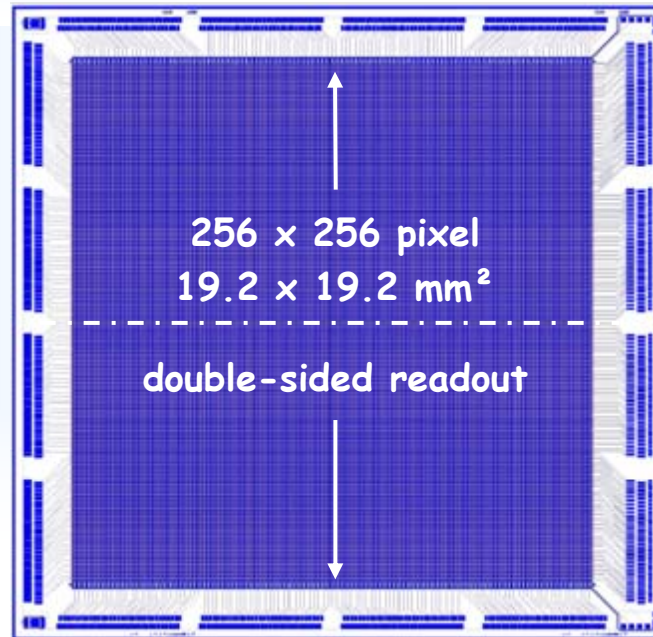


DEPFET APS - next generation

- ◆ **new production**
 - first results scheduled summer 2008
- ◆ **minor pixel modifications**
 - faster charge collection
 - low-voltage clear process
- ◆ **large formats**
 - yield & homogeneity studies
 - readout & control system adaptation
 - effect of long signal lines

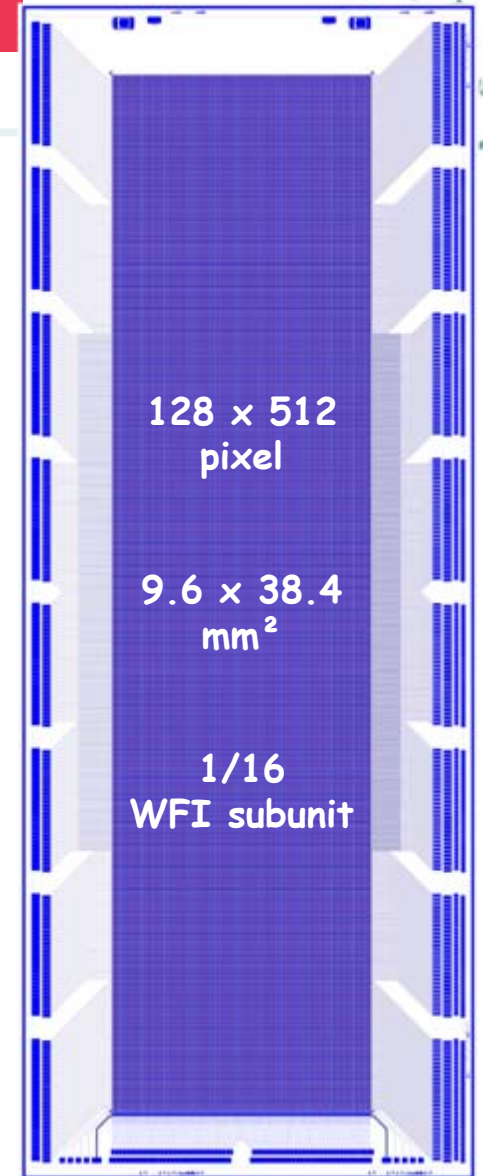


64 x 64 pixel
4.8 x 4.8 mm²



256 x 256 pixel
19.2 x 19.2 mm²

double-sided readout

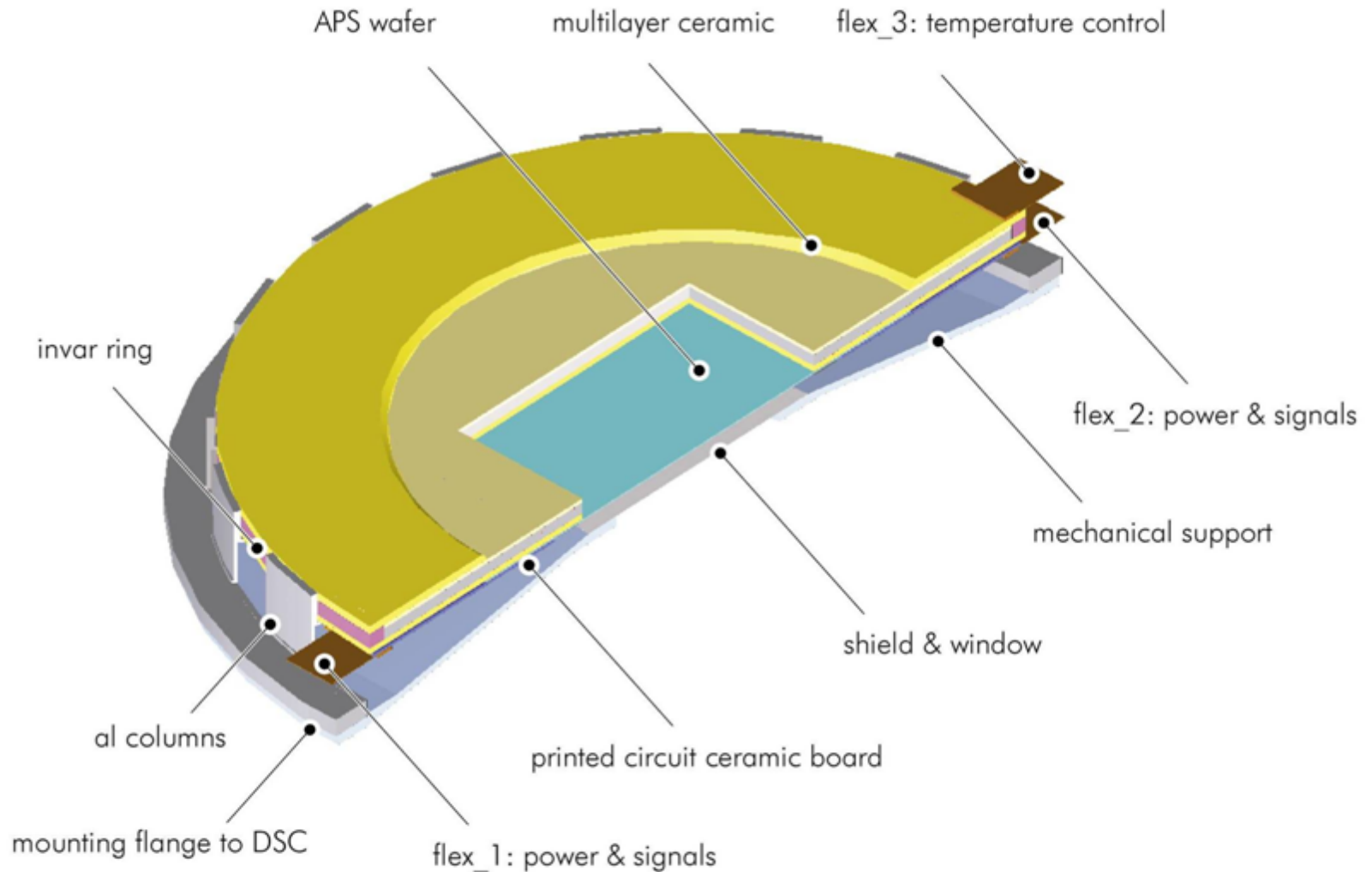


128 x 512
pixel

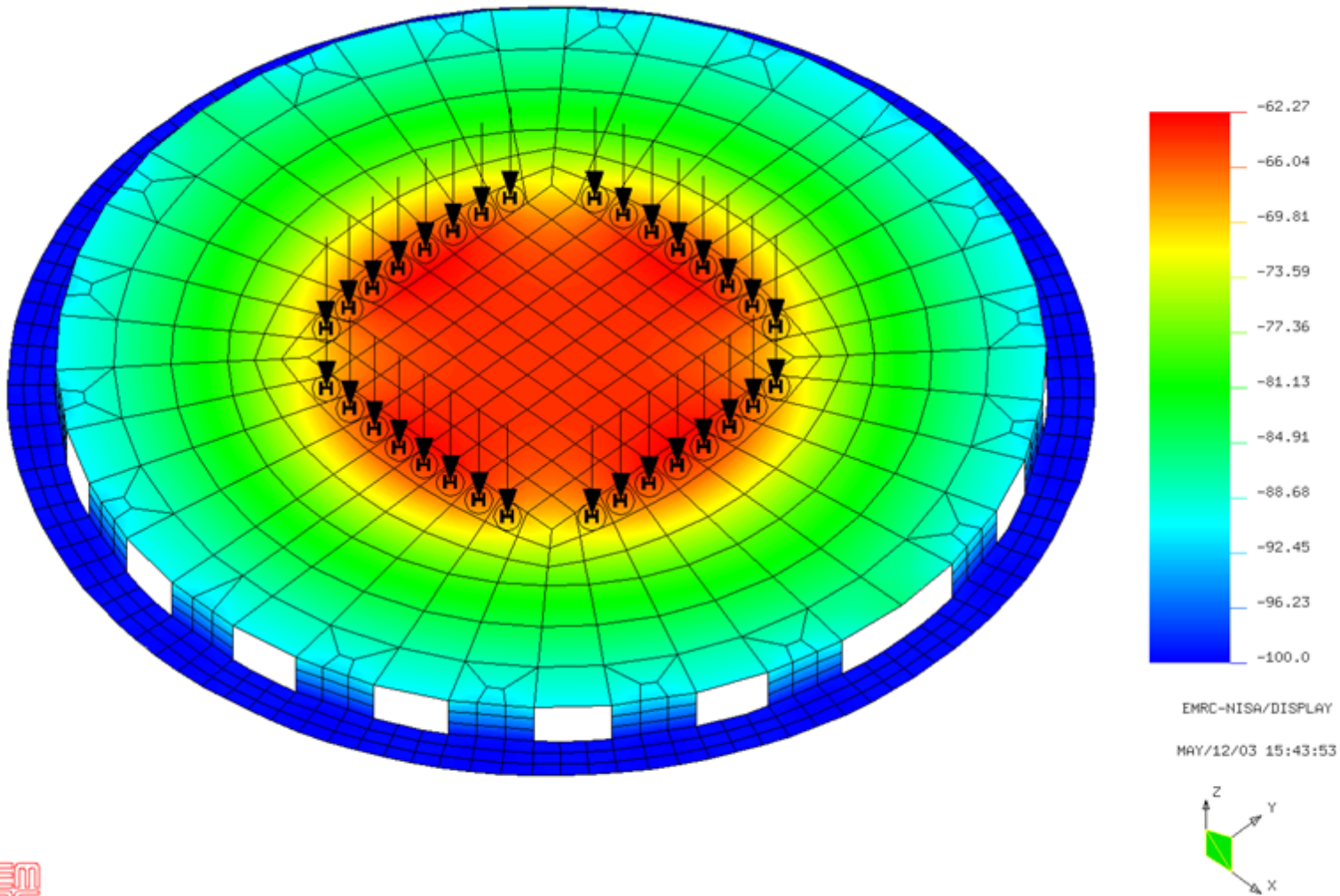
9.6 x 38.4
mm²

1/16
WFI subunit

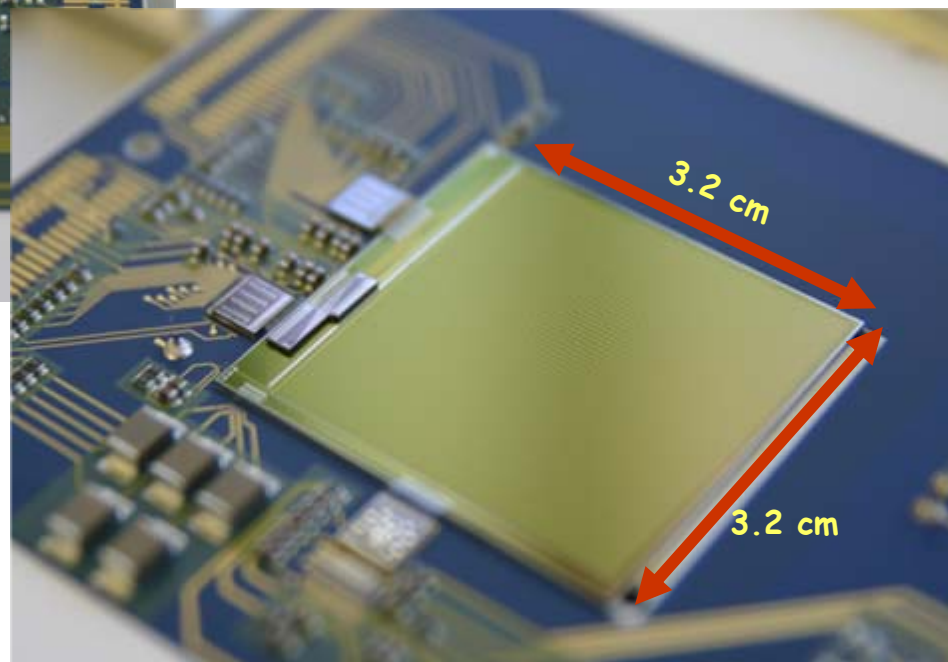
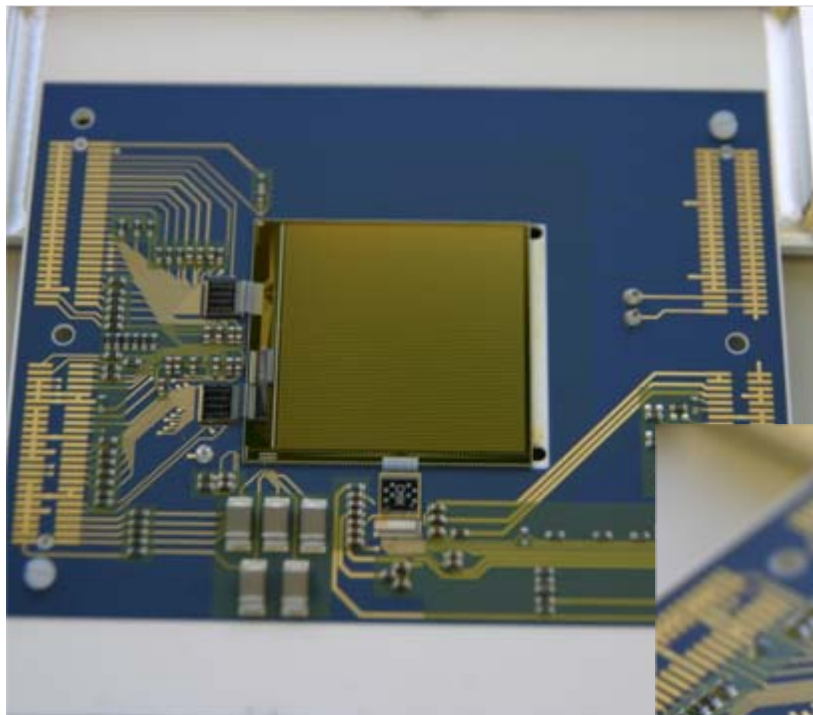
Mechanical structure of APS



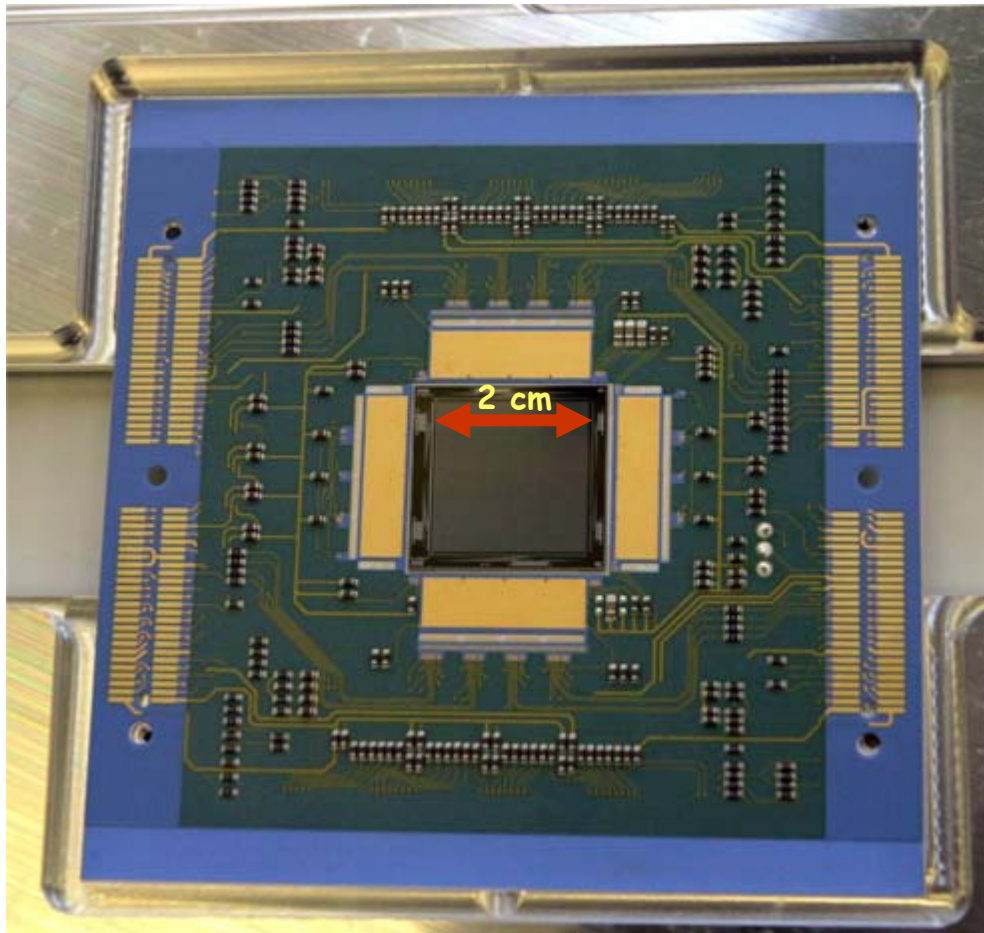
Thermal Analysis of the APS System



SIMBOL-X-Hybrid



BepiColombo Chip and Hybrid



**Breadboard module test assembly
of the BepiColombo system**

Potential WFI Team:

| | |
|--------------------------|---------|
| MPE, Garching, IAAT | Germany |
| LU, Leicester | UK |
| Politecnico di Milano, | Italy |
| CAS, Tsinghua University | China |
| PNSensor GmbH | Germany |
| University of Osaka | Japan |

US teams are welcome

8.2 cm

8.2 cm

HTRS requirements (CESR, Didier Barret)



| PARAMETER | REQUIREMENT (GOAL) | SCIENCE DRIVER |
|--|--|--|
| Effective area (m ²) | 1 (1.5) @ 0.2 keV | WHIM, early BHs, clusters |
| | 5 @ 1 keV | Clusters, WHIM, early BHs |
| | 2 @ 7 keV | EOS, gravity in strong fields |
| | 1 @ 10 keV (0.1) @ 30 keV | EOS, acceleration, early BHs Acceleration, early BHs, EOS |
| Energy range (keV) | 0.1–40 | BHs, acceleration, clusters |
| Angular resolution (arc-sec) | 5 (2) @ < 10 keV 10 @ 40 keV | Clusters, early BHs, WHIM Early BHs |
| Field of view (arc-min) | 7 (10) diameter: WFI, HXI 1.7 diameter: NFI | Clusters, early BHs, acceleration Clusters, enrichment, galaxy evolution |
| Spectral resolution (eV)(FWHM) | 2 (1) @ 0.5 keV: NFI 2 @ <2 keV: NFI 6 (3) (@ 6 keV: NFI 150 @ 6 keV: WFI 1000 @ 40 keV: HXI | WHIM Clusters Clusters, enrichment, galaxy evolution Early BHs Early BHs |
| Point source detection sensitivity, erg cm ⁻² s ⁻¹ | (3 x 10 ⁻¹⁸) @ 0.2–8 keV; 4σ | Early BHs |
| Time Resolution (s) | 10 ⁻⁵ : HTRS | EOS studies |
| Count rate capability (s ⁻¹) | 2 10 ⁶ : HTRS | EOS studies |
| Polarimetry (MDP, 3σ-confidence in 10 ks) | 2% at 10 ⁻² Crab: XPOL | EOS studies |
| Observing constraints | >2 weeks visibility each 6m | EOS studies |
| | ToO response in (<1 day) | EOS studies |
| | 10 ³ (5.10 ⁴) s cont. observ. | EOS studies, strong gravity |
| | ±5° (±15°) range Sun angle | EOS studies |

SDDs for astrophysics the HTRS on XEUS



HTRS (High Time Resolution Spectrometer)

possible choice: 19-cell SDD array

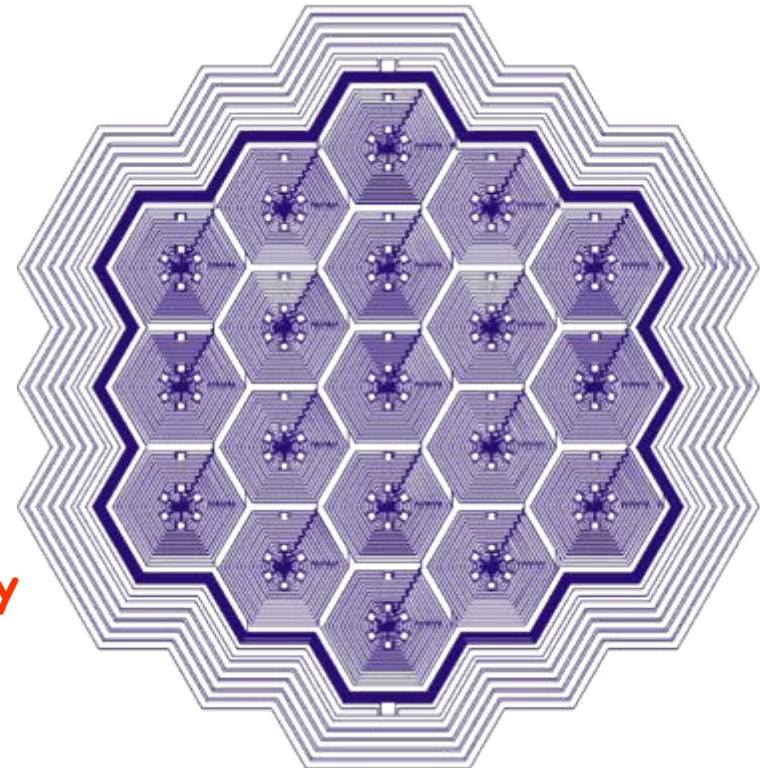
timing and spectroscopy

1- 10 μsec / 150 eV / 20 Mcps

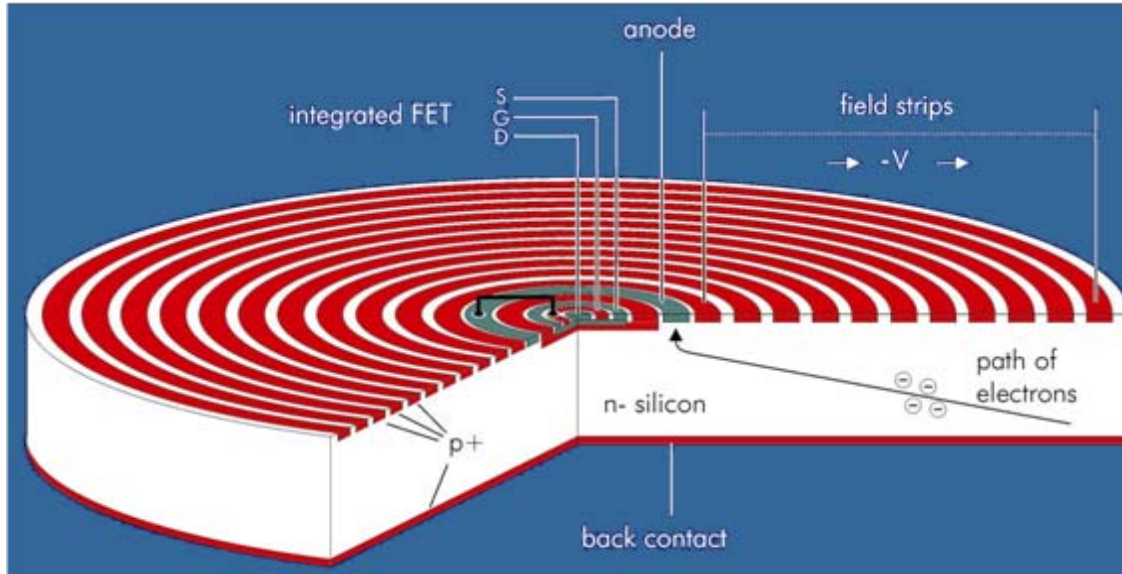
operation out of focus typ. 10 cm to 20 cm

every SDD cell has a count rate capability
of typically 1.000.000 cps

$$19 \times 5 \text{ mm}^2 = 95 \text{ mm}^2$$



SDDs in astrophysics: the HTRS on XEUS



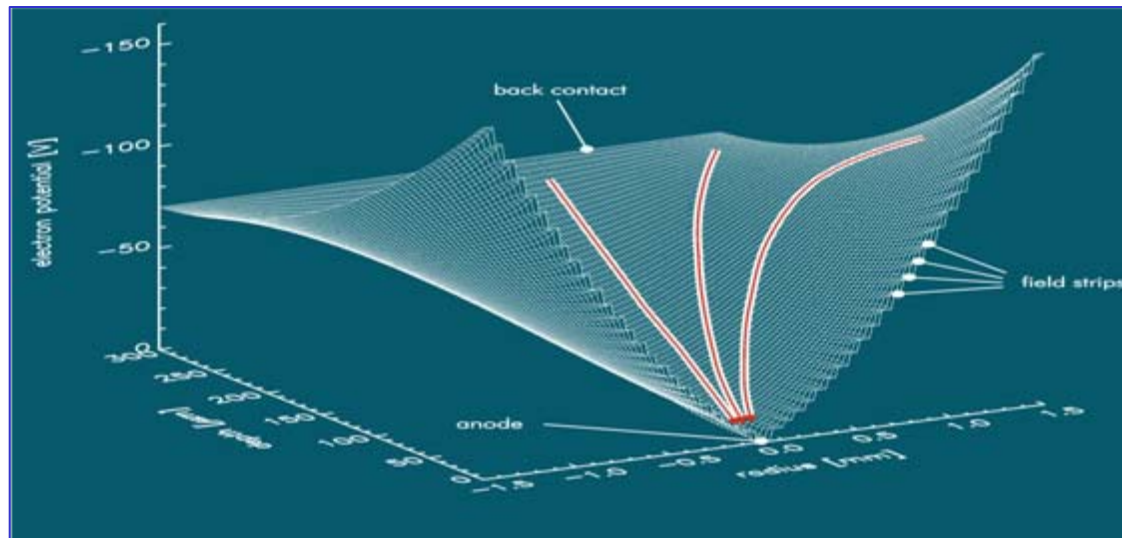
**SDD with
integrated
SSJFET or
DePFET**

time to drift from
device edge to
readout node:

$$\begin{aligned} \Delta t &= \Delta x / v \\ &= \Delta x / \mu E \\ &= 100 \text{ ns} \end{aligned}$$

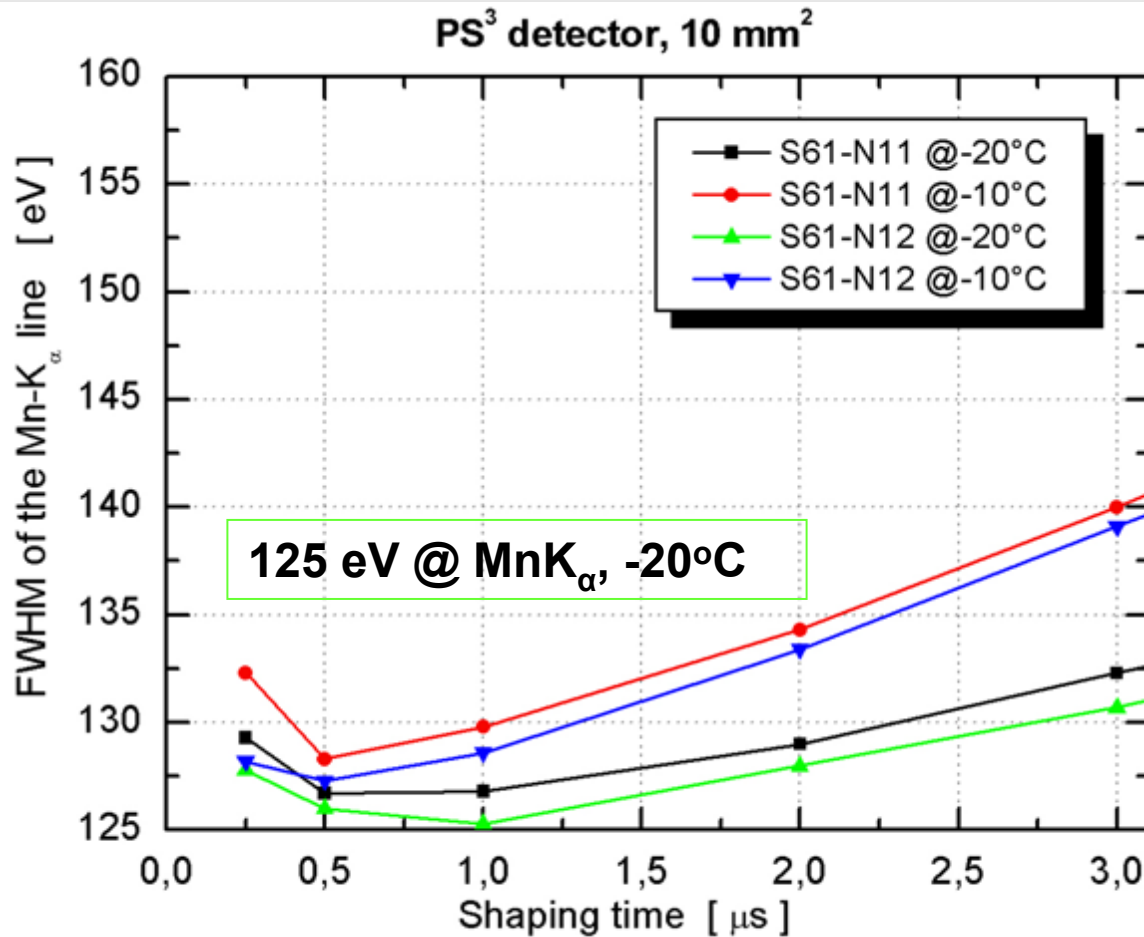
$$\sigma_{\Delta t} \leq 8 \text{ ns}$$

for $\Delta r = 1.3 \text{ mm}$
i.e. $A = 6 \text{ mm}^2$
and $E = 600 \text{ V/cm}$



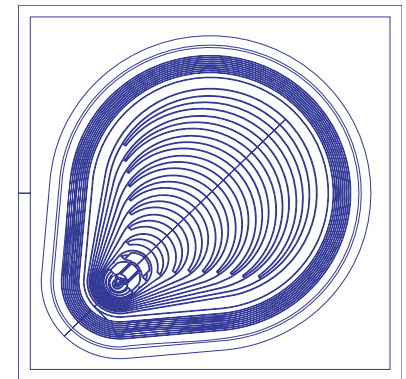
**Electrical
Potential in
a circular SDD**

Energy resolution as a function of τ_s



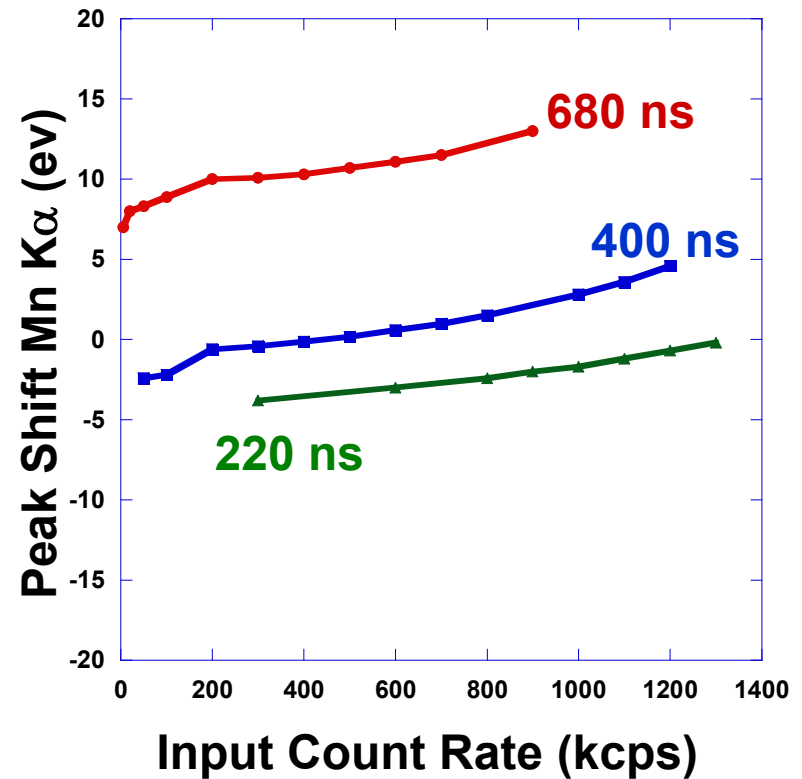
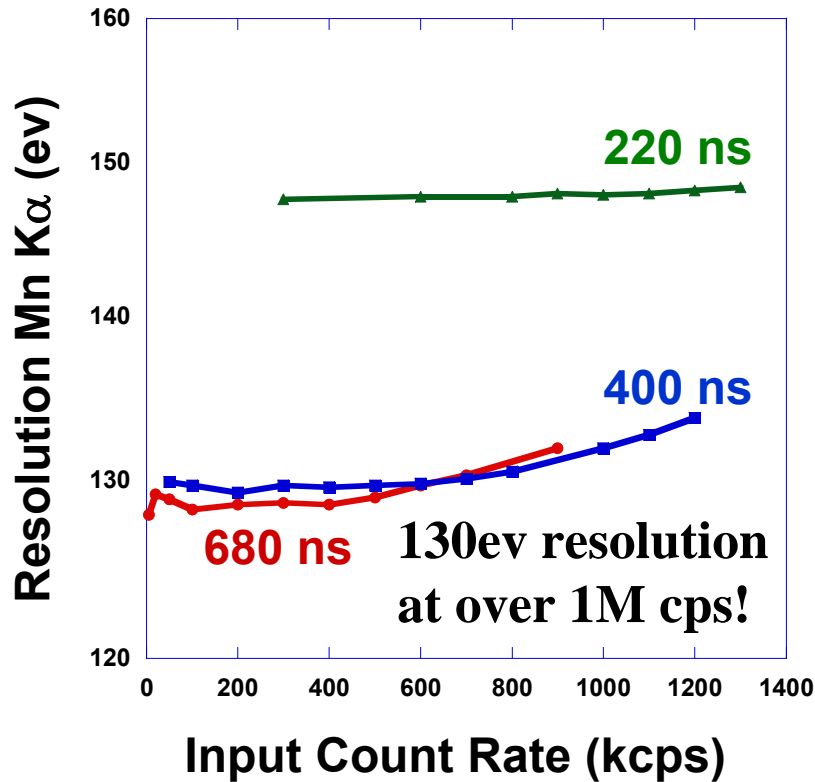
SD³

in polysilicon technology



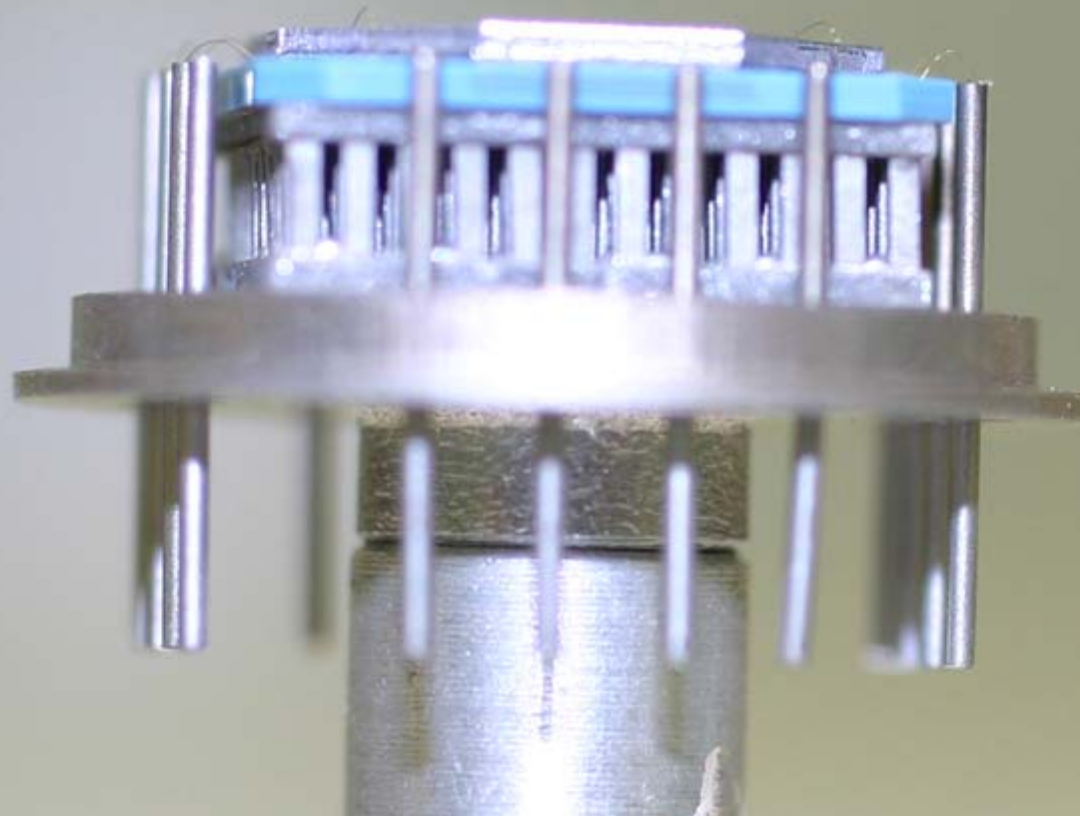
$C_{\text{tot}} = 40 \text{ fF}$

Pulsed-reset circuitry: Spectacular Performance



Nearly flat response of resolution and peak position with count-rate for all pulse-processors up to over 1M counts/sec

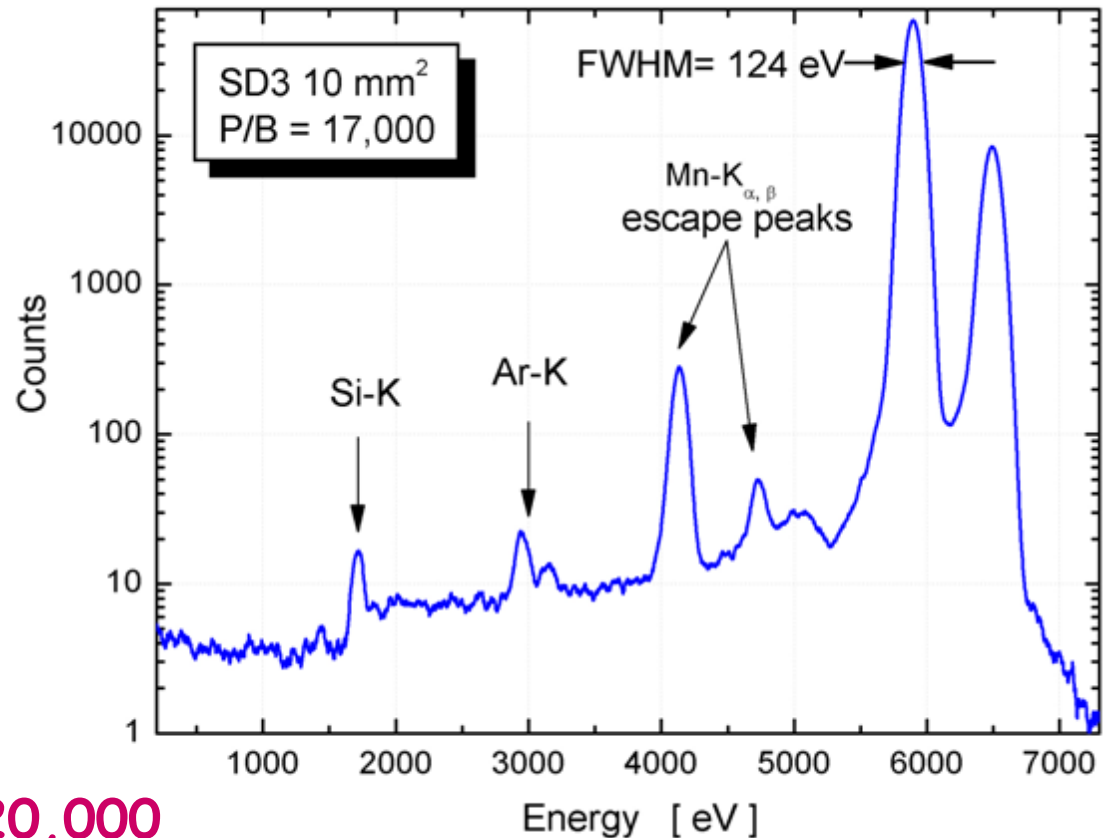
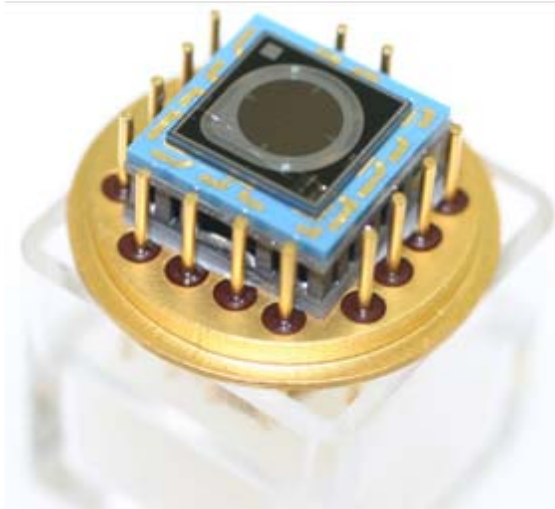
flying on - MER (SPIRIT, OPPORTUNITY)
- ROSETTA
- EXOMARS (soon)



Spectroscopic performance of 5 and 10 mm² SD3 detectors

- ◆ operation temperature **-20°C**
- ◆ 1-stage Peltier cooler
- ◆ optimum shaping time **0.5 μs**
- ◆ pulsed reset operation

Optimized detector entrance window
for light element detection - **pnWindow**



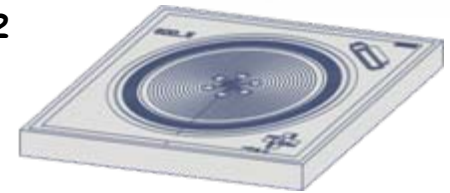
- ◆ Light element detection
FWHM @ B-K = 38 eV
- ◆ **P/B = 13,000 ÷ 20,000**

A single SDD can process up to 1 M counts per sec below 170 eV

Small and Large Area SDDs



Classic Round SDDs with sensitive area of 5, 10 and 20 and 30 mm² up to 1cm²



SDD 5 mm²
chip 5 x 5 x 0.45 mm³



SDD 10 mm²
chip 6 x 6 x 0.45 mm³



SDD 20 mm²
chip 8 x 8 x 0.45 mm³



SDD 30 mm²
chip 9 x 9 x 0.45 mm³

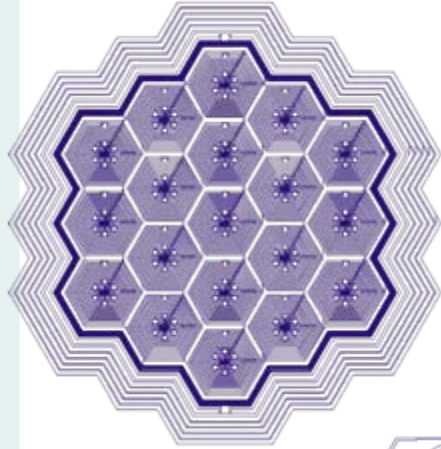


SDD 100 mm²
chip 14 x 14 x 0.45 mm³

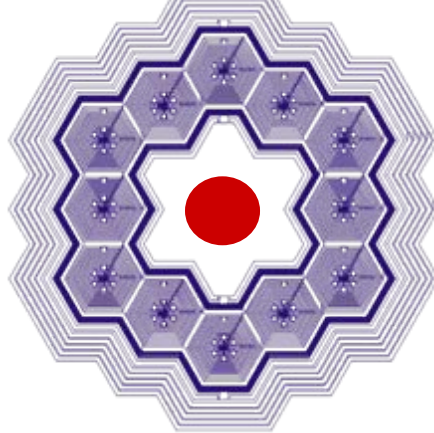
Multichannel SDDs



$19 \times 5 \text{ mm}^2 = 95 \text{ mm}^2$



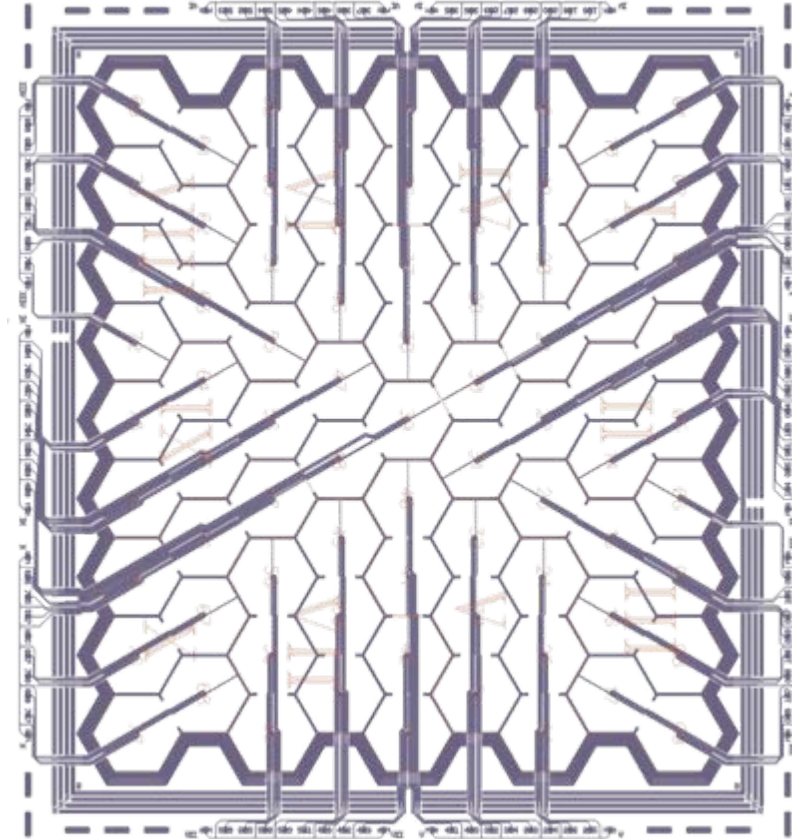
$12 \times 5 \text{ mm}^2 = 60 \text{ mm}^2$



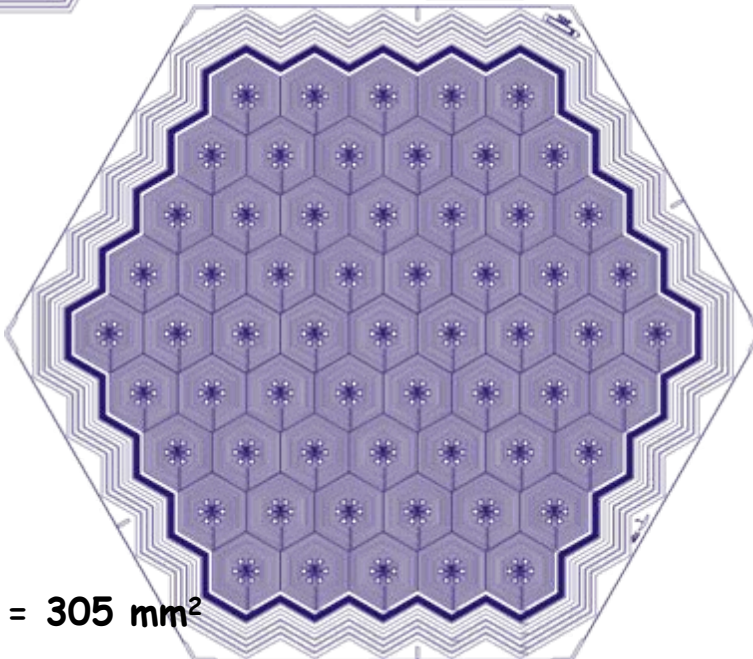
$6 \times 5 \text{ mm}^2 = 30 \text{ mm}^2$



$77 \times 7 \text{ mm}^2 = 539 \text{ mm}^2$



$61 \times 5 \text{ mm}^2 = 305 \text{ mm}^2$



Conclusions

DePFET is ready for

- ❑ fast and slow readout
- ❑ thick and thin depletion layers
- ❑ for large and small pixels
- ❑ for small and large monolithic fields of view

radiation hard and defect free

