

Introduction to MiniBooNE and  $\nu_\mu$   
Charged Current Quasi-Elastic  
(CCQE) Results

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For the MiniBooNE collaboration

# The MiniBooNE Collaboration

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**Embry Riddle University**  
**Fermi National Accelerator Laboratory**  
**Indiana University**

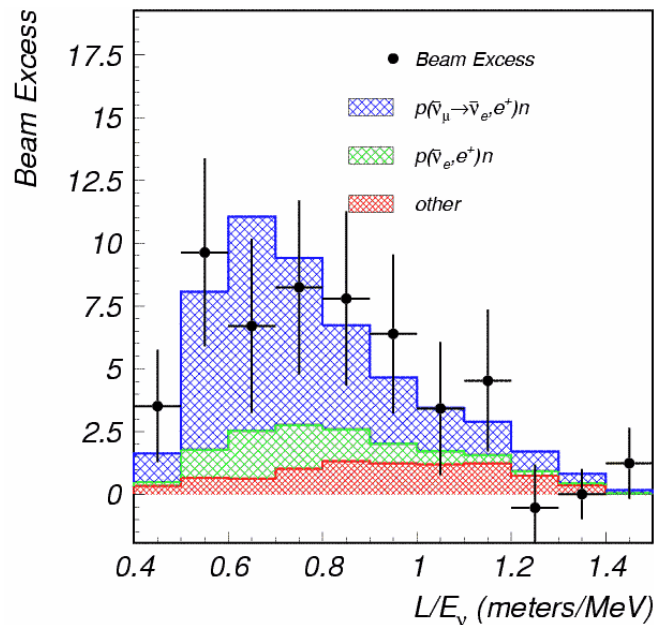
**Los Alamos National Laboratory**  
**Louisiana State University**  
**University of Michigan**  
**Princeton University**  
**Saint Mary's University of Minnesota**  
**Virginia Polytechnic Institute**  
**Western Illinois University**  
**Yale University**

74 people, 16 Institutions

MiniBooNE was approved in 1998,  
with the goal of addressing the LSND anomaly:

an excess of  $\bar{\nu}_e$  events in a  $\bar{\nu}_\mu$  beam,  
 **$87.9 \pm 22.4 \pm 6.0$  ( $3.8\sigma$ )**

which can be interpreted as  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations:



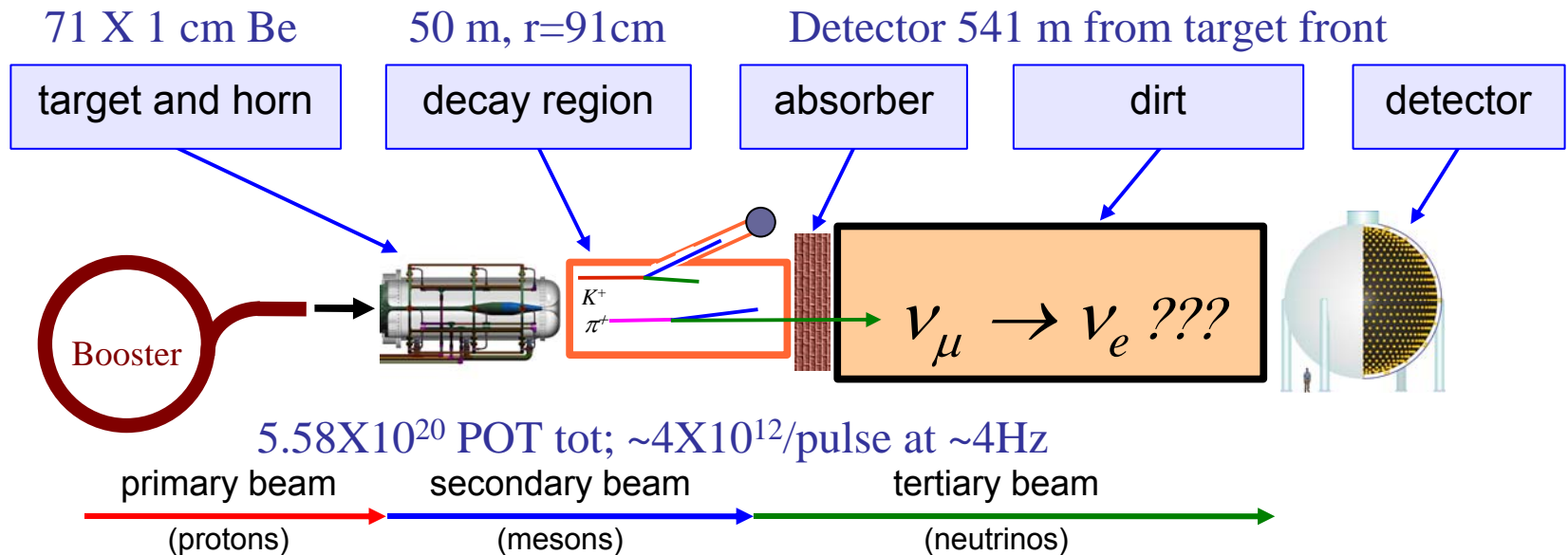
Points -- LSND data  
Signal (blue)  
Backgrounds (red, green)

*LSND Collab, PRD 64, 112007*

# MiniBooNE's Design Strategy...

Keep L/E same  
while changing systematics, energy & event signature

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L/E)$$

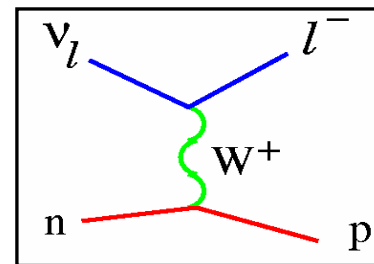
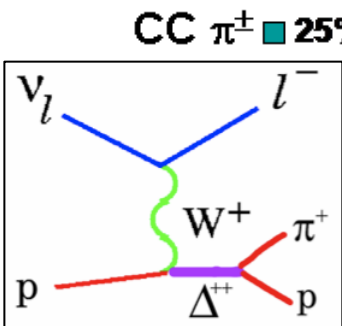
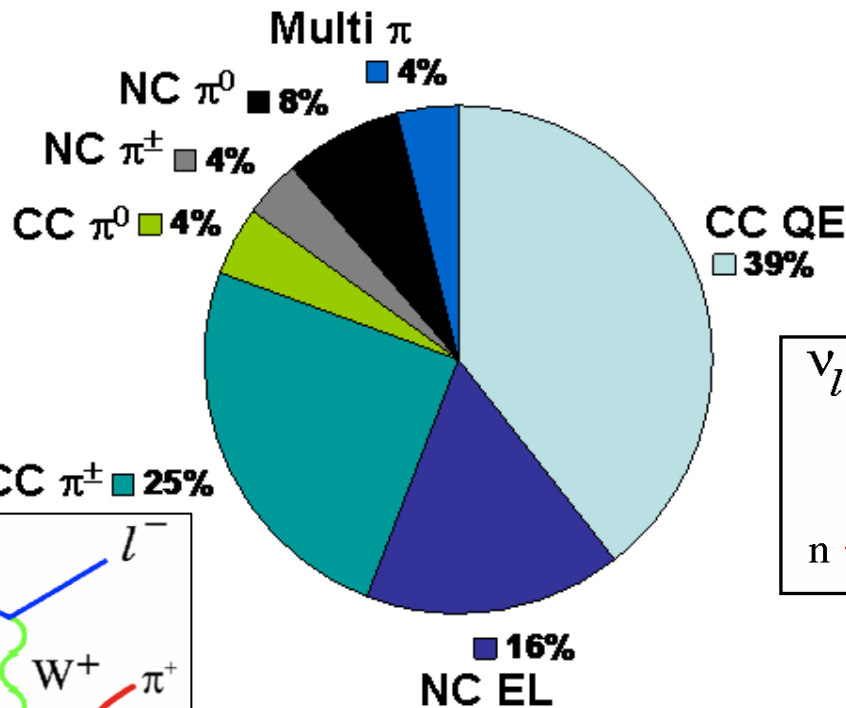
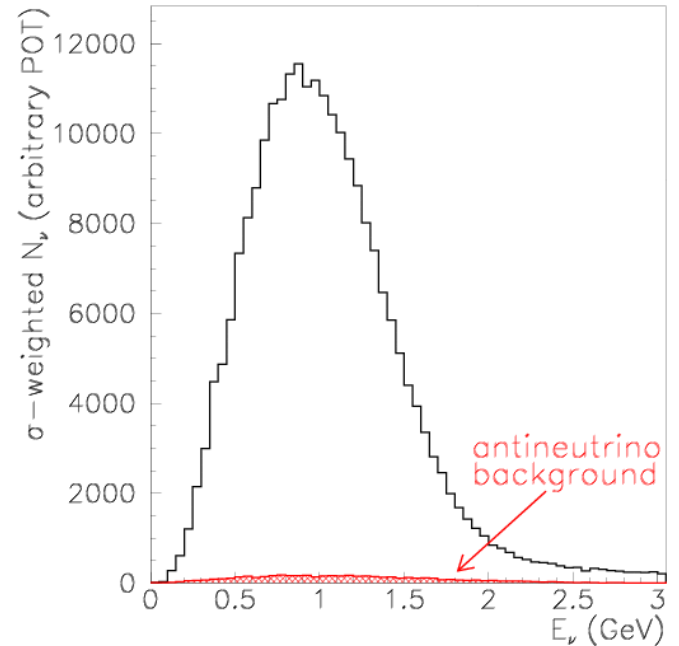
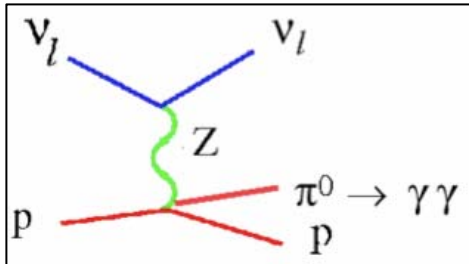


Order of magnitude  
higher energy (~500 MeV)  
than LSND (~30 MeV)

Order of magnitude  
longer baseline (~500 m)  
than LSND (~30 m)

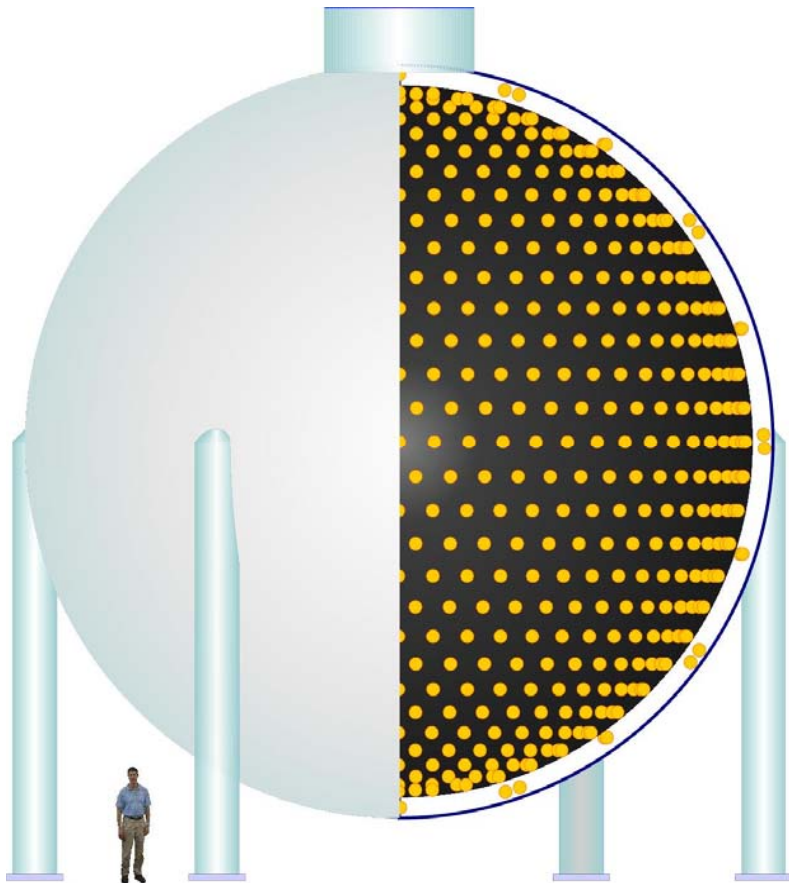
# Predicted event rates before cuts (NUANCE Monte Carlo)

D. Casper, NPS, 112 (2002) 161



$\nu_e/\nu_\mu = 0.5\%$ ; anti- $\nu = 6\%$   
Most  $\nu_e$  from  $\mu$ , K decays

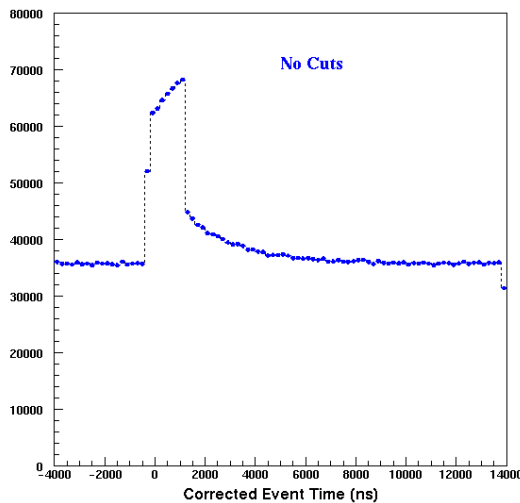
# The MiniBooNE Detector



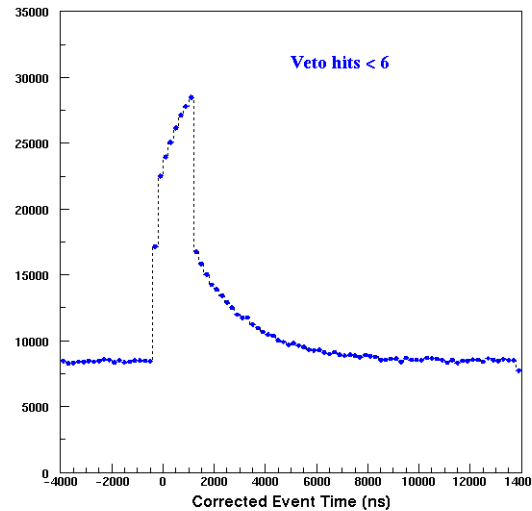
- 541 meters downstream of target
- 3 meter overburden of dirt
- 12 meter diameter sphere  
(10 meter “fiducial” volume)
- Filled with 800 t of pure mineral oil ( $\text{CH}_2$ --  
density 0.86,  $n=1.47$ )
  - (Fiducial volume: 450 t)
  - 1280 inner 8” phototubes-10% coverage,  
240 veto phototubes  
(Less than 2% channels failed during run)

Progressively introducing cuts (19.2  $\mu\text{s}$  time window starting 4  $\mu\text{s}$  before beam)

Phototubes have 1.7 ns ( $\sim 75\%$ ) and 1.2 ns time resolutions



Raw data

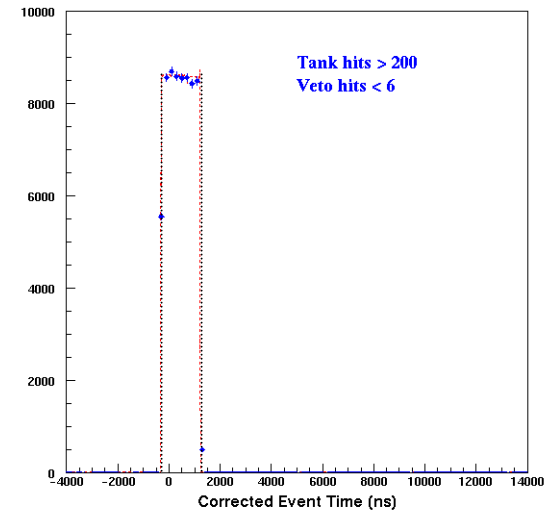


Veto < 6 removes through-going cosmics ( $\sim 2$  CR in entire oscillation set)

This leaves

“Michel electrons”

( $\mu \rightarrow \nu_\mu \nu_e e$ ) from cosmics



Tank Hits > 200 (equivalent to energy) removes Michel electrons, which have 52 MeV endpoint

# Subevents; Kinds of Light

- 100 ns bins for subevents (separate mu-decays)
- Cherenkov/scintillation light about 8/1.  
Cherenkov comes at fixed angle to track direction and is prompt. Scintillation light and light scattered by fluorescence is delayed.
- Fluorescence and attenuation important and functions of frequency; prompt/delayed light at phototubes is about 10/1 on the average.



## The types of particles these events produce:

Muons:

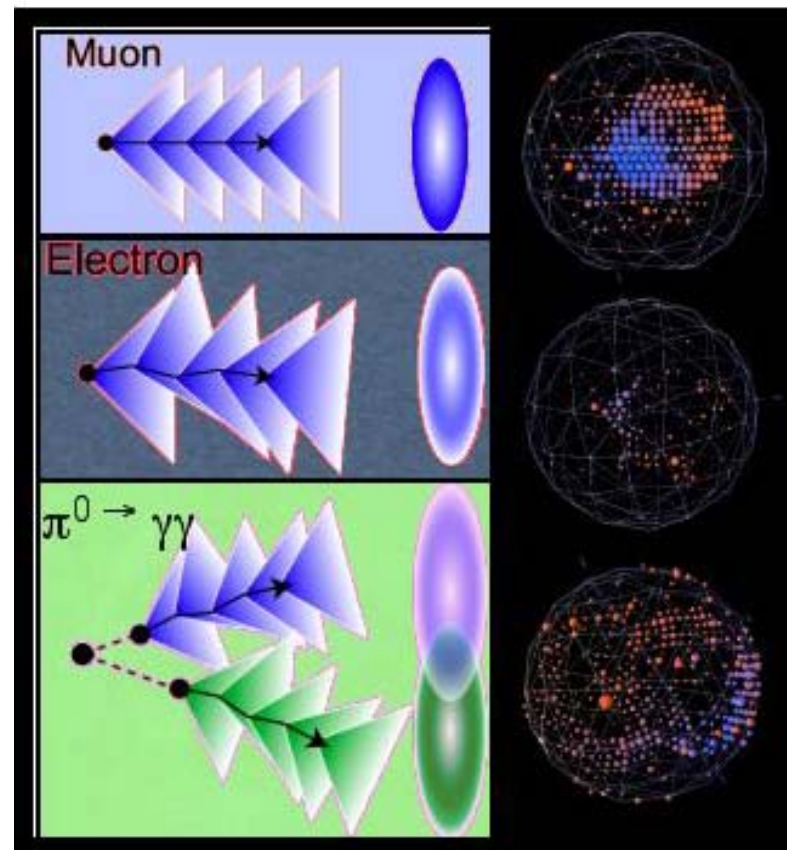
Produced in most CC events.  
Usually 2 subevents (only 8%  $\mu^-$  capture) or exiting.

Electrons:

Tag for  $\nu_\mu \rightarrow \nu_e$  CCQE signal.  
1 subevent

$\pi^0$ s:

Can form a background if one photon is weak or exits tank.  
In NC case, 1 subevent.



# Reconstruction

- Initial guess. Position mainly from timing of hits; angle from a grid of possibilities using prompt (Cherenkov) light
- Final fit. Minuit fits to hypotheses
  - a. One outgoing muon track
  - b. One outgoing electron track
  - c. Two tracks (aimed at  $\pi^0$  events)

# Two Analysis Chains

For most of analysis had two equal reconstructions, sfitter, rfitter

- Toward end of analysis, a new more powerful reconstruction based on sfitter—the pfitter became available. Better especially on 2 track fits (22 cm position error,  $2.8^\circ$  1 track angle error,  $\sim 20$  MeV  $\pi^0$  mass resolution)—BUT takes about 10 times more computer time.
- rfitter dropped, sfitter and pfitter retained.

# Simulations

- Use measured proton cross sections (Harp, BNL910, earlier experiments)
- Geant4 for following produced particles through magnetic horn, decay region...
- V3 Nuance for neutrino cross sections (mod. by MiniBooNE measurements and other improvements.)
- Detailed optical model for detector using GEANT3. (39 model parameters--obtained from measurements)

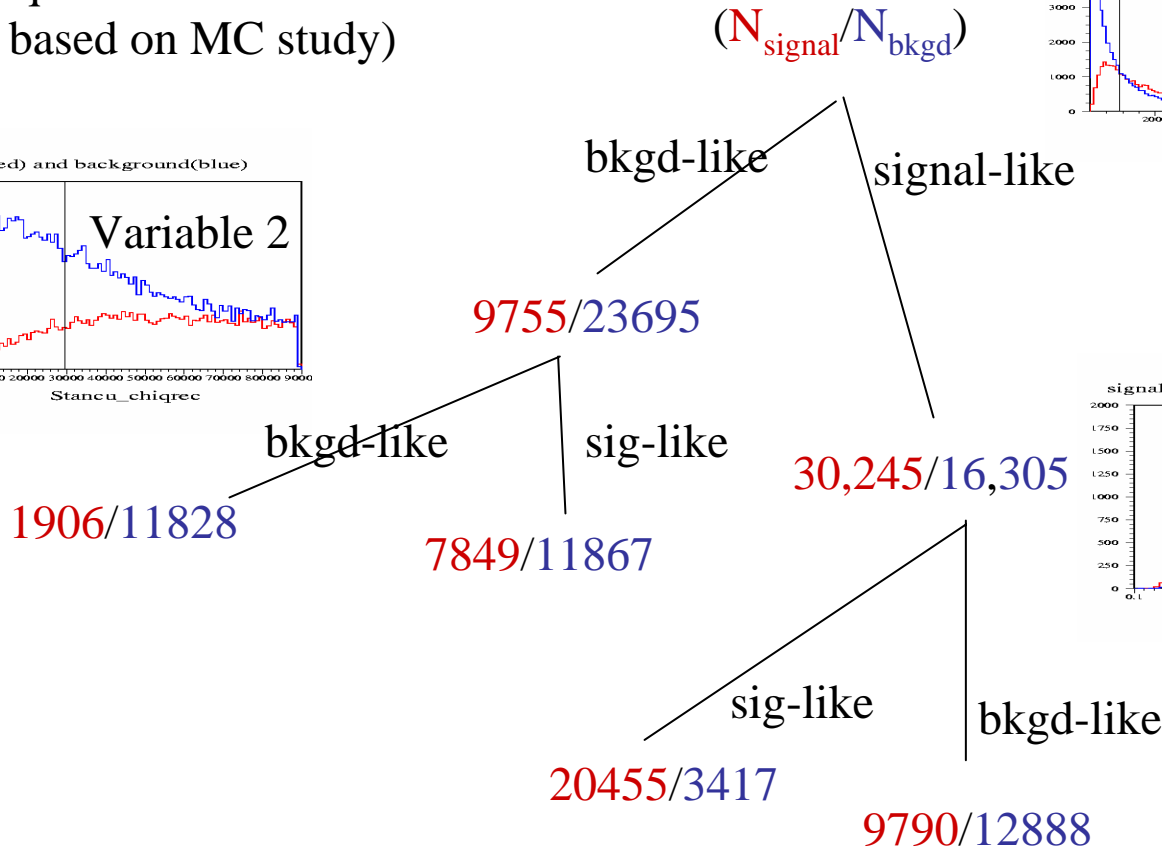
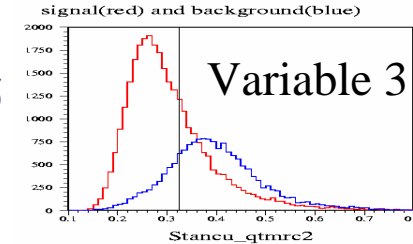
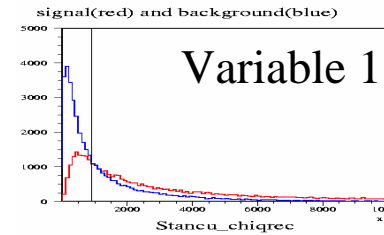
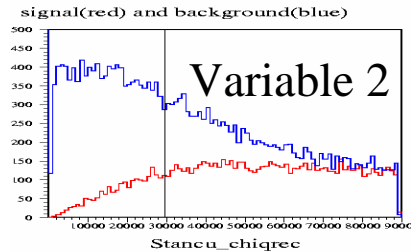
# Plan

- First discuss  $\nu_e$  CCQE selection for the oscillation analysis
- Then present  $\nu_\mu$  CCQE cross section results.

# Event Classification Schemes for Oscillation Measurement

- Signal events were defined as  $\nu_e$  CCQE events
- Pfitter used simple cuts (TB--“Track based analysis”) to separate these events based on:
  - a. Likelihood of 1 track e-fit vs 1 track  $\mu$ -fit
  - b. Likelihood of 1 track e-fit vs 2 track fit
  - c. Mass of  $\pi^0$  in 2 track fit
- Sfitter used a method new to physics— boosted decision trees (BDT) with many variables (172)

# A Decision Tree (sequential series of cuts based on MC study)

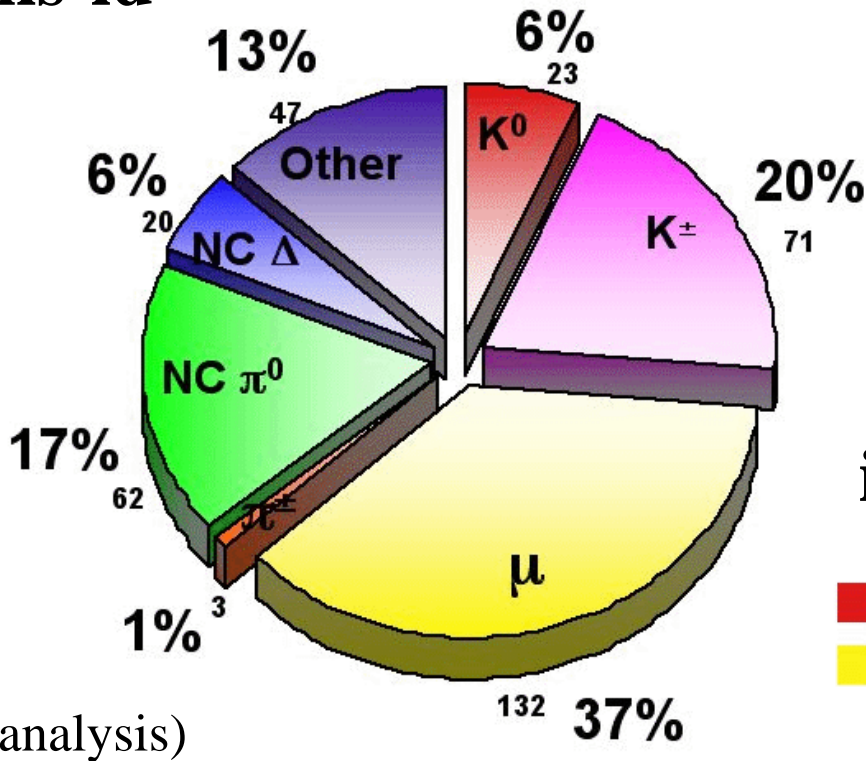


etc.

*Weight events misclassified higher and make new “boosted tree”. Continue 100’s of times; sum results of each tree: 1 if signal leaf, -1 if background leaf*

We have two categories of backgrounds:

$\nu_\mu$  mis-id



(TB analysis)

intrinsic  $\nu_e$



Predictions of the backgrounds are among the nine sources of significant error in the analysis



Source of Uncertainty On $\nu_e$ background	Track Based /Boosted Decision Tree error in %	Checked or Constrained by MB data	Further reduced by tying $\nu_e$ to $\nu_\mu$
Flux from $\pi^+/\mu^+$ decay	6.2 / 4.3*	✓	✓
Flux from $K^+$ decay	3.3 / 1.0	✓	✓
Flux from $K^0$ decay	1.5 / 0.4	✓	✓
Target and beam models	2.8 / 1.3	✓	
$\nu$ -cross section	12.3 / 10.5*	✓	✓
NC $\pi^0$ yield	1.8 / 1.5	✓	
External interactions (“Dirt”)	0.8 / 3.4	✓	
Optical model	6.1 / 10.5	✓	✓
DAQ electronics model	7.5 / 10.8*	✓	

\* Errors quoted are before constraints from measured  $\nu_\mu$  flux which strongly reduces them

# Charged Current $\nu_\mu$ Quasi Elastic Events

- Close to 2 o.m. more events than any previous experiment
- 39% of all neutrino interactions before cuts
- 193,709 events asking for 2 subevents and that the second subevent be consistent with  $\mu$  decay in position and have  $<200$  hits. 60% eff.
- KE resolution 7% at 0.3 GeV, angular res.  $\sim 5^\circ$
- 74% pure—mostly  $\pi$  backgrounds
- Mainly  $0 < Q^2 < 1 \text{ GeV}^2$

# Standard Parameters Don't Work

- Relativistic Fermi Gas nuclear model
- $P_F=220$  MeV/c;  $E_B=34$  MeV;  $F_V$  from electron experiments.
- Axial Vector FF =  $g_A/(1+ Q^2/M_A^2)^2$  with  $g_A =1.2671$  and  $M_A= 1.03$  GeV from previous low statistics  $\nu$  expts mostly on lighter targets.

Discrepancy tends to follow lines of constant  $Q^2$  rather than lines of constant energy

# Correction to Pauli Blocking Term

## Smith & Moniz model

- Carbon is described by the collection of incoherent Fermi gas particles.
- all complications come from hadronic tensor;

$$(W_{\mu\nu})_{\text{lab}} = \int_{E_{\text{lo}}}^{E_{\text{hi}}} f(\vec{k}, \vec{q}, \omega) T_{\mu\nu} : \text{hadronic tensor}$$

$f(\vec{k}, \vec{q}, \omega)$  : density function (energy conservation, state distribution)

$T_{\mu\nu} = T_{\mu\nu}(F_1, F_2, F_A, F_P)$  : nucleon tensor

$E_{\text{hi}}$  : the highest energy state of nucleon =  $\sqrt{(PF^2 + M^2)}$

$E_{\text{lo}}$  : the lowest energy state of nucleon (for QE interaction =  $\sqrt{(PF^2 + M^2)} - \omega_{\text{eff}}$ )

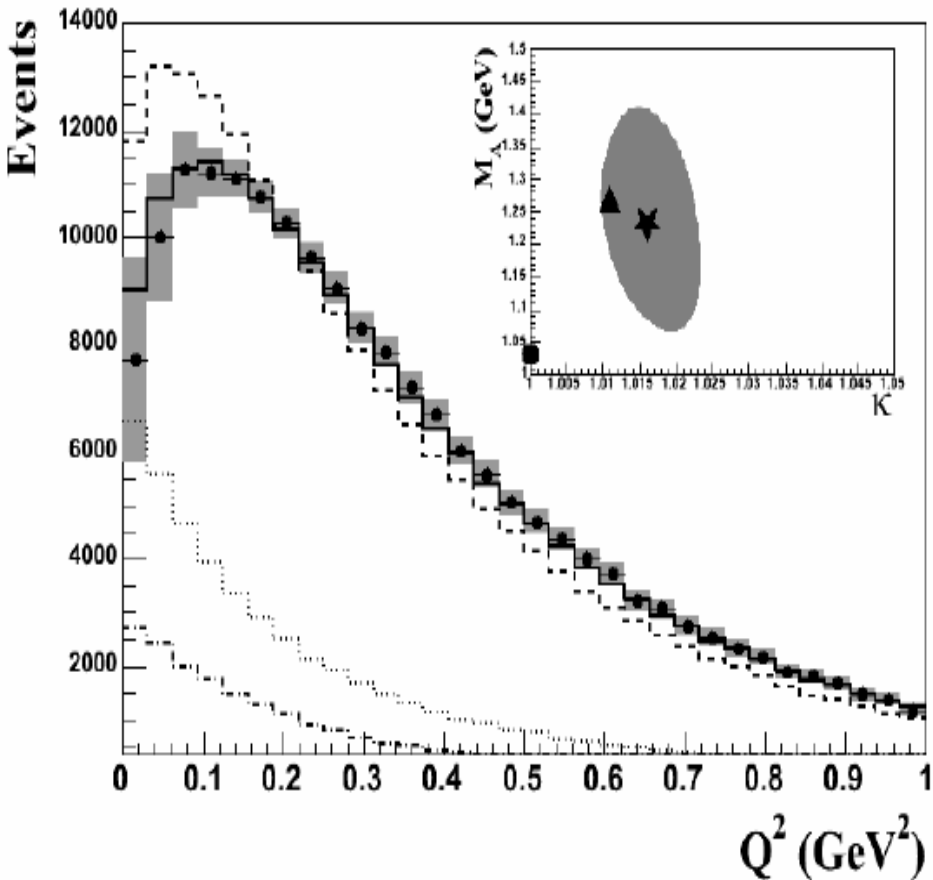
$\omega$  = energy transfer

New term: Scale  $E_{\text{lo}}$ —multiply by  $\kappa$ . (Default 1) Effectively changing energy level distribution.

Best fit is  $M_A = 1.23 \pm 0.20$ ;  $\kappa = 1.019 \pm 0.011$

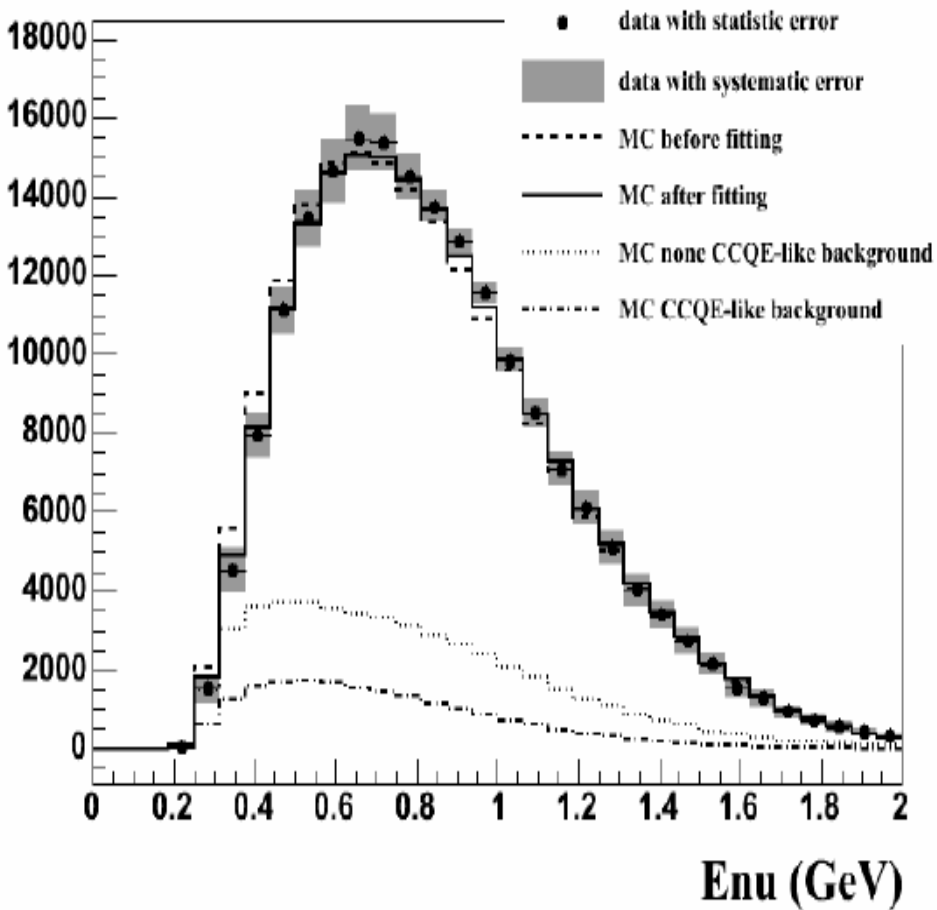
arXiv:0706.0926 (hep-ex), submitted to PRL.

# Results



- Dashed—before fit
  - Solid—after fit
  - Dotted—background
  - Dash dotted CCQE-like background (only  $\mu$  in final state)
  - Dots—data with error
  - Star—best fit point
  - Circle—Original values
  - Triangle—Best varying CCPIP background
- $\chi^2/\text{dof}$  58.1 before  
32.8 after fit for 30 d.f.

# CCQE Energy Distribution



- The new variable,  $\kappa$ , is empirical. It corresponds to a change in the nuclear energy levels.
- This data should provide a guide leading to a better nuclear model.
- The fitted distribution was critical for normalization for the oscillation analysis: 5.6% increase in pred.  $\nu_{\mu}$  CCQE events

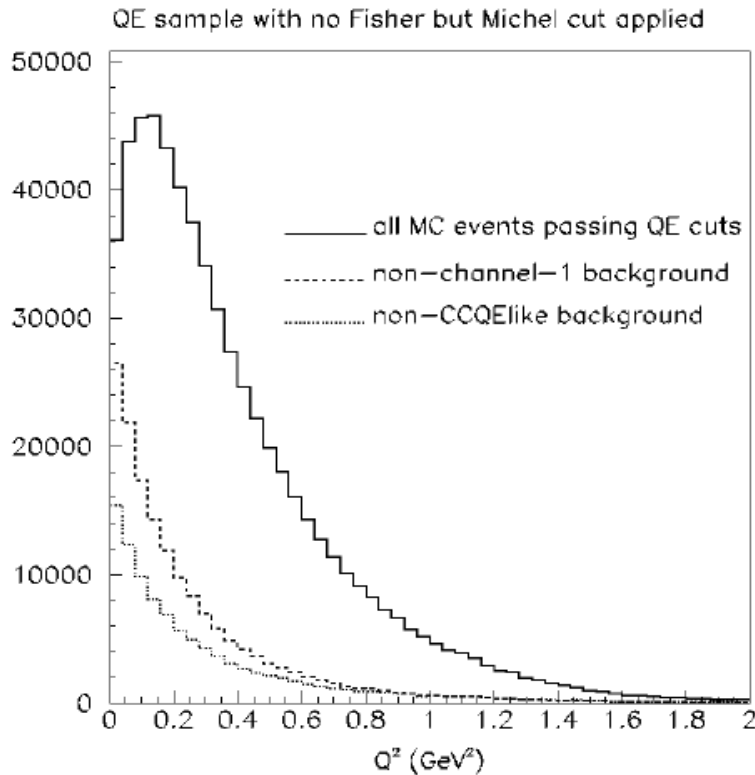
# BACKUP

# Modifications to V3 NUANCE

- MiniBooNE measured CCQE results
- MiniBooNE measured p dependence of  $\pi^0$  production
- MiniBooNE measured coherent pion production
- Tuned final state interaction model
- Explicit nuclear de-excitation photon emission model
- Angular correlation for Delta (1232) to agree with Rein-Sehgal model



# Charged Current Quasi-Elastic Events



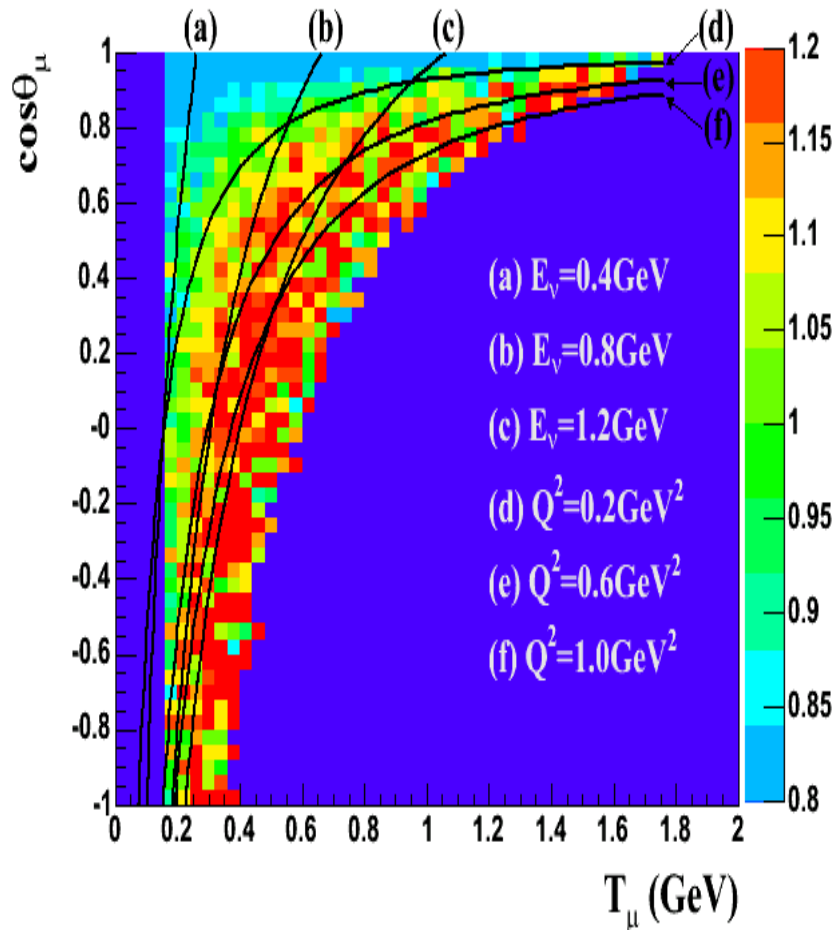
Close to 2 o.m. larger sample than any previously  
193,709 CCQE events  
asking 2 subevents and 2<sup>nd</sup> vertex consistent with decay  
& <200 hits (60% eff.)

KE res 7% at 0.3 GeV;  
angular res.  $\sim 5^\circ$

74% pure—mostly  $\pi$  backgrounds

- $0 < Q^2 < 1 \text{ GeV}^2$

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- $p_F=220$ ,  $E_B=34$  MeV,  
 $F_V$  (from electron expts)
- AV FF  $M_A=1.03\text{GeV}$ ;  
 $g_A=1.2671$  (from  
previous  $\nu$  expts)  
 $F_A=g_A/(1+Q^2/M_A^2)^2$
- Discrepancy follows  
lines of constant  $Q$  more  
than constant  $E$