

Nuclear Structure I: Basic Facts and Models

J. Engel

University of North Carolina

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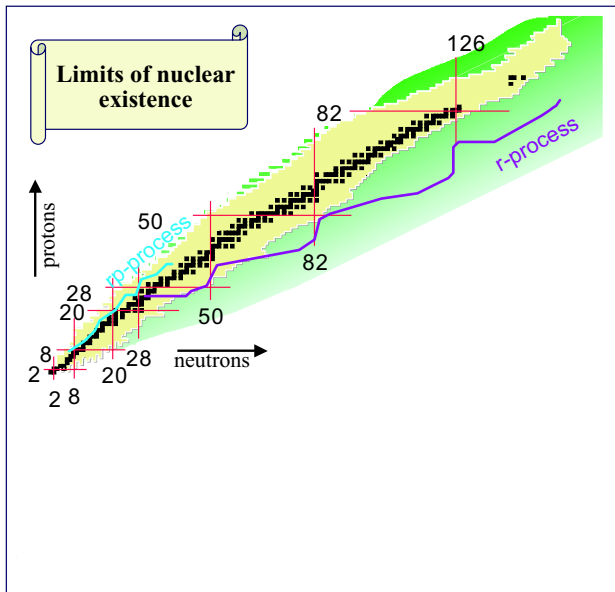
Outline

- 1 Intro
- 2 Fermi-Gas
- 3 Shell Model
- 4 Drop Model
- 5 Early Microscopy

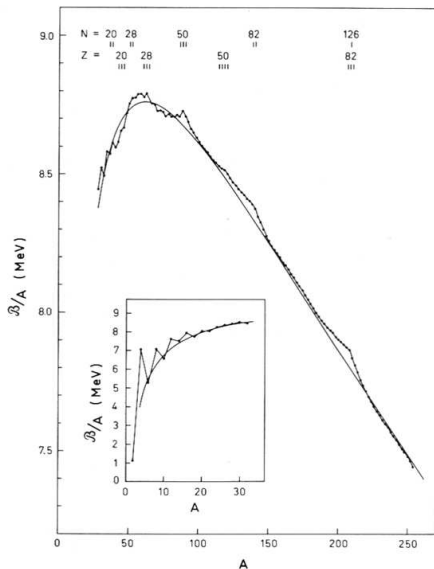
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The Nuclear Landscape



Binding Energies



Bethe-Weizsacker Formula

$$\begin{aligned}
 B(Z, A) = & a_V A - a_S A^{2/3} \\
 & - a_C \frac{Z^2}{A^{1/3}} - a_{\text{sym}} \frac{(Z-N)^2}{A}
 \end{aligned}$$

with

$$a_V = 15.65 \text{ MeV}$$

$$a_S = 17.23 \text{ MeV}$$

$$a_C = .698 \text{ MeV}$$

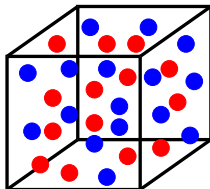
$$a_{\text{sym}} = 28.1 \text{ MeV}$$

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Simplest Possible Model: Nucleons in a "Box"

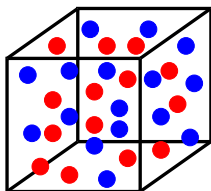
Periodic boundary conditions:



$$k_x = \frac{2\pi n_x}{L}, \quad \Delta k_x = \frac{2\pi}{L}, \quad \Delta^3 k = \frac{(2\pi)^3}{L^3} = \frac{(2\pi)^3}{V}$$

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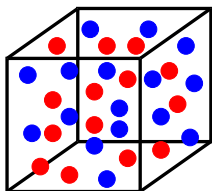


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$$\rho_0 = (1/V) \sum_{\vec{k}, s} \theta(k_F - k) = (1/V) \sum_{\vec{k}, s} \frac{V}{(2\pi)^3} \Delta^3 k \theta(k_F - k)$$

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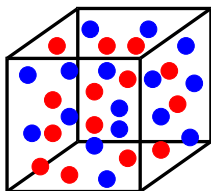


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From $\rho \approx 1.72 \text{ fm}^{-3}$ get $k_F = 1.36 \text{ fm}^{-1}$, $\epsilon_F = 38 \text{ MeV}$.

Energy density:

$$\mathcal{E}_0 = \frac{2}{(2\pi)^3} 4\pi \int_0^{k_F} \frac{k^4}{2m} dk = \frac{k_F^5}{10m\pi^2} = \frac{3}{5} \frac{k_F^2}{2m} \rho_0 = 3/5 \epsilon_F \rho_0$$

and $\bar{\epsilon} \equiv \mathcal{E}_0/\rho_0 = (3/5)\epsilon_F \approx 23$ MeV.

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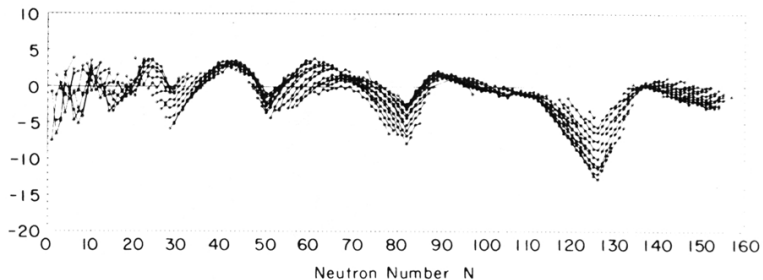
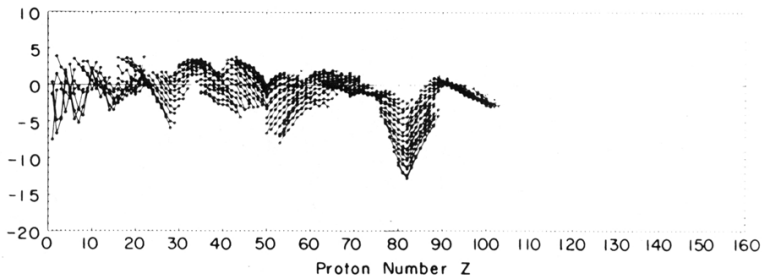
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Homework problem: Derive the value of a_{sym} above.

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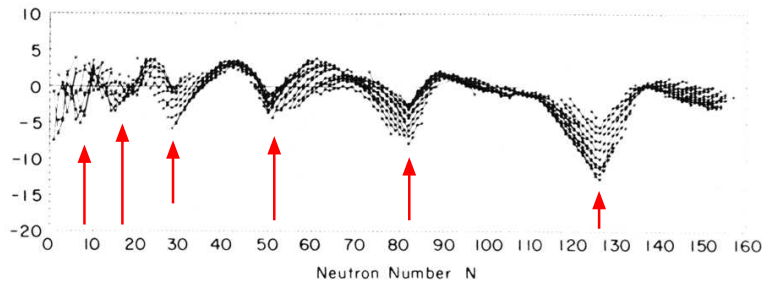
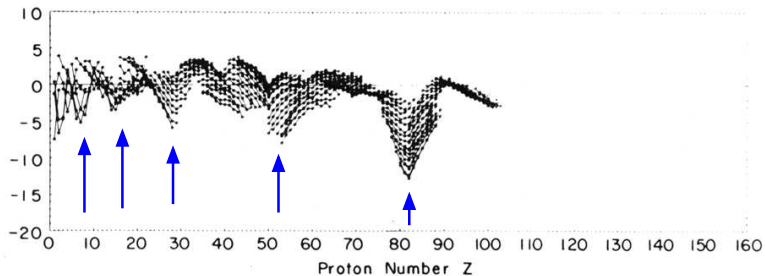
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Binding-Energy Deviations



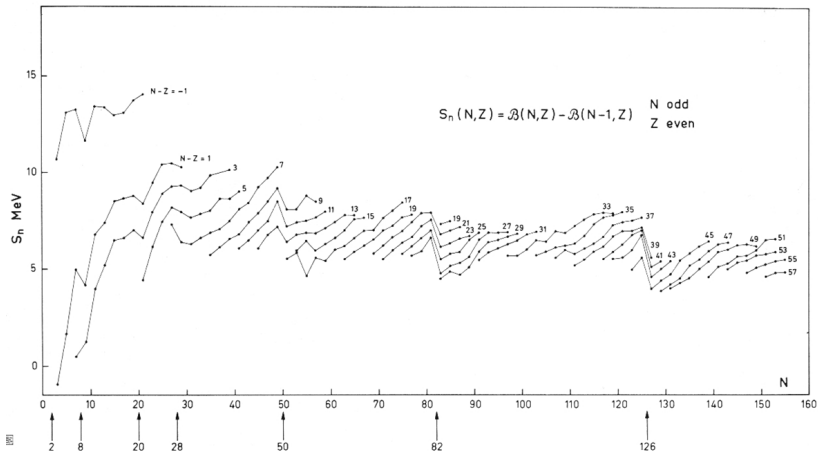
From W.D. Meyers and W.J. Swiatecki, Nucl. Phys. **81**, 1 (1966).

Binding-Energy Deviations



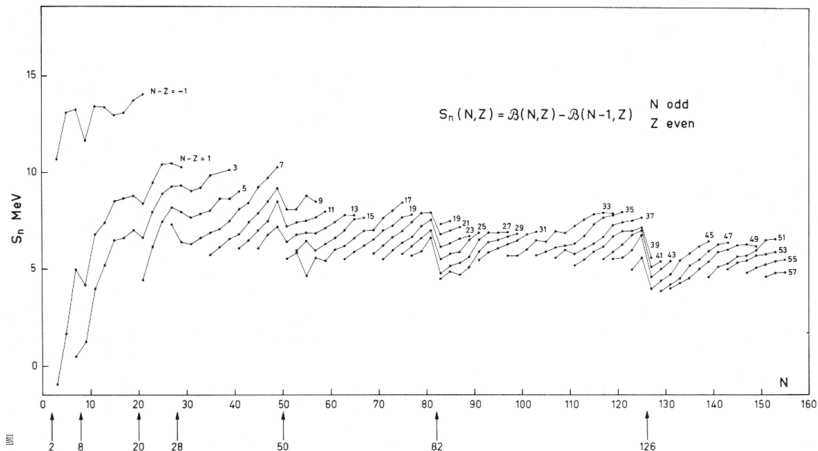
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Neutron Separation Energies



From Bohr and Mottelson, *Nuclear Structure*, v.1, p. 193.

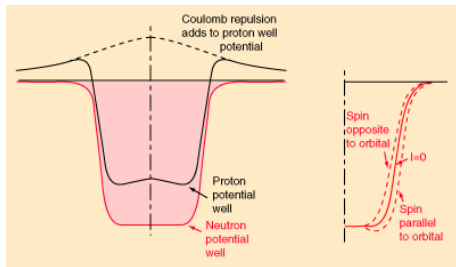
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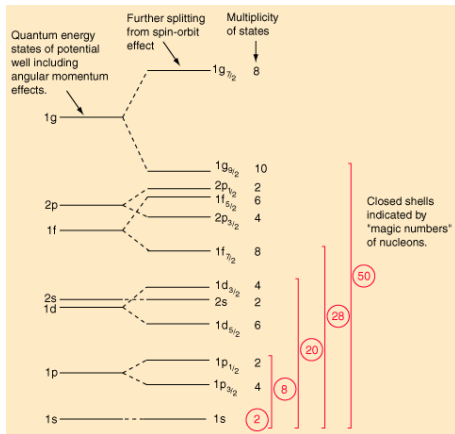
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Levels not evenly spaced: **shell gaps** at “magic numbers”.

Shell Model of Nucleus

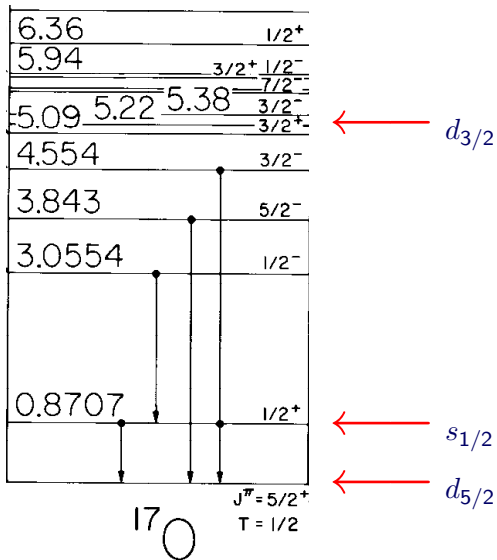


<http://hyperphysics.phy-astr.gsu.edu/hbase/nuclear/shell.html>

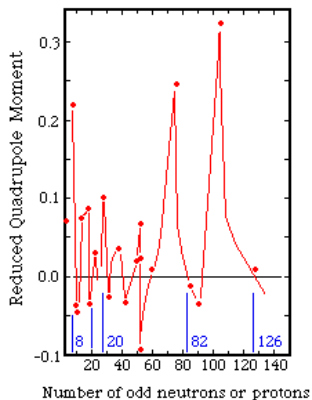


Replace spherical box with more sophisticated potential. Get magic numbers: 2, 8, 20, 28, 50, 126

An Example

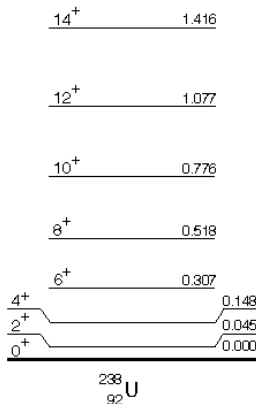
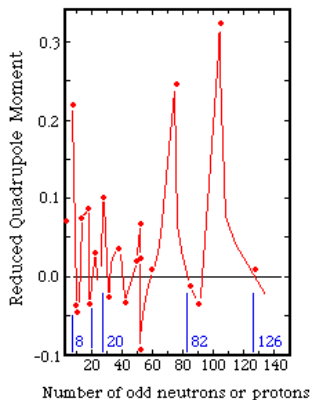


Shell Model Doesn't Explain Collective Rotation...



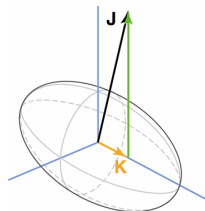
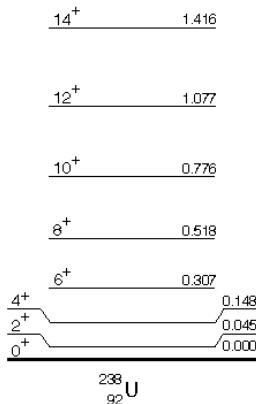
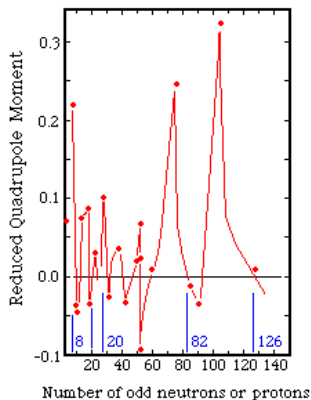
From Booth and Combey, <http://www.shef.ac.uk/physics/teaching/phy303/phy303-3.html> and
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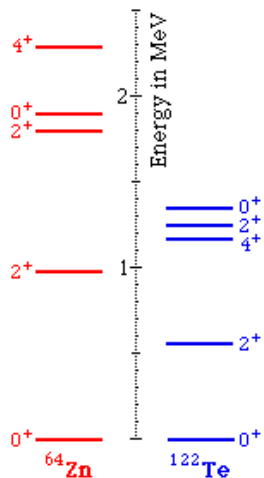
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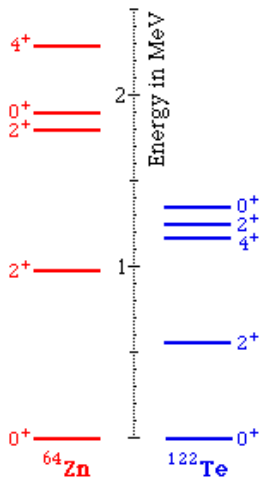
Some sort of collective rotation going on between magic numbers!

And Doesn't Explain Collective Vibrations

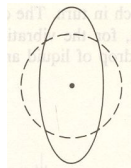


From Booth and Combey, <http://www.shef.ac.uk/physics/teaching/phy303/phy303-3.html> and <http://www.fen.bilkent.edu.tr/~aydinli/Collective%20Model.ppt>

And Doesn't Explain Collective Vibrations



Two vibrational "phonons" with angular momentum 3 give states with angular momentum 0, 2, or 4.



From Booth and Combey, <http://www.shef.ac.uk/physics/teaching/phy303/phy303-3.html> and <http://www.fen.bilkent.edu.tr/~aydinli/Collective%20Model.ppt>

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Alternative: “Liquid Drop” Model

Low-Lying Quadrupole Collective States

Protons and neutrons move together; volume is conserved, surface changes shape.

Ansatz for surface:

$$R(\theta, \phi) = R_0 (1 + \sum_m \alpha_m Y_{2,m}(\theta, \phi))$$

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The 5 α 's are **collective variables**. For vibrations, Hamiltonian obtained e.g. from classical fluid model and Bethe-Weizsacker:

$$H \approx 1/2B \sum_m |\dot{\alpha}_m|^2 + 1/2C \sum_m |\alpha_m|^2$$

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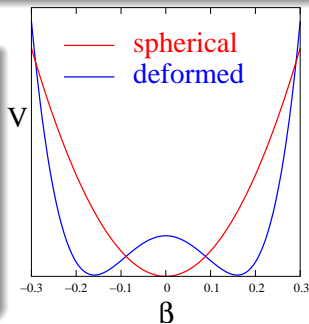
This reproduces the average trend of $E = \omega$ with A, but real life is more complicated, with frequencies depending significantly on how nearly magic the nucleus is.

Deformation in Liquid Drop Model

Restoring parameter C becomes negative if Coulomb effects overcome surface tension. Nucleus can then deform. Go to “intrinsic frame,” replacing 5 intrinsic-frame α 's with

$$\beta \equiv \sqrt{\alpha_0^2 + 2\alpha_2^2} \quad \text{and} \quad \gamma \equiv \tan^{-1}[\sqrt{2}\alpha_2/\alpha_0]$$

Wave function: $|\Psi\rangle \approx D_{MK}^{J*}(\theta, \phi, \psi)\Phi_{\text{int.}}(\beta, \gamma)$.

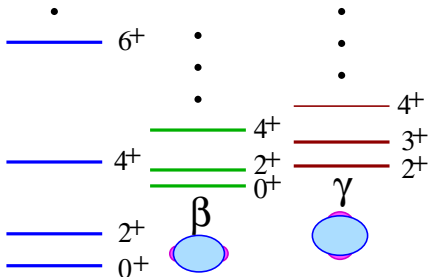
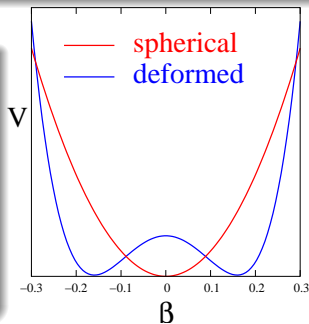


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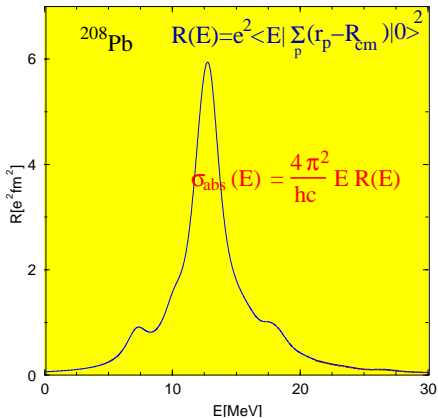


Low-lying states in deformed nuclei

- 1 Rotations of deformed nucleus with moment of inertia between that of rigid body and “irrotational flow”
- 2 Surface vibrations along or against symmetry axis

Density Oscillations

Photoabsorption cross section proportional to “isovector” dipole strength, which lies mainly at higher energy than surface modes.

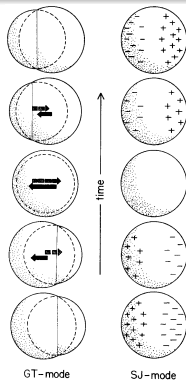
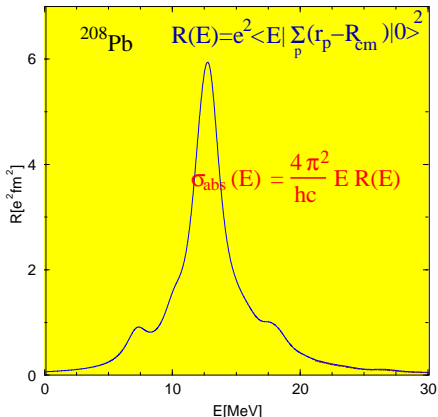


From N. Paar, <http://www1.physik.tu-muenchen.de/~npaar/pygmy/>

Giant dipole resonance

Density Oscillations

Photoabsorption cross section proportional to “isovector” dipole strength, which lies mainly at higher energy than surface modes.



From W D. Myers et al., Phys. Rev. C 15, 2032 (1977).

Hydro for SJ mode gives

$$E \approx 2\hbar \sqrt{\frac{8a_{\text{sym}}NZ}{mA^3R_0^2}} \approx 12 \text{ MeV in } ^{208}\text{Pb}.$$

From N. Paar, <http://www1.physik.tu-muenchen.de/~npaar/pygmy/>

Giant dipole resonance

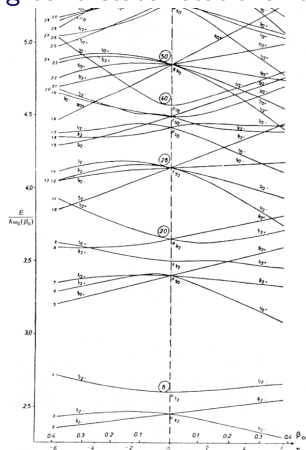
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Early Attempts at “Unification”: Nilsson Model

$$|\Psi\rangle \approx D_{m\Omega}^{j*}(\theta, \phi, \psi) |N, l, m_l, \Omega\rangle_{\text{int.}}$$

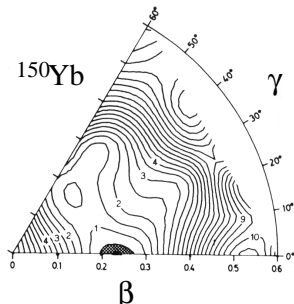
Ω is like K . Valence intrinsic level is ground-state rotational bandhead.



From Eisenberg and Greiner, *Nuclear Models* p. 542

Shell model in deformed potential

- Part of deformation energy from “deformed Fermi gas”.
- Can map out “energy surface” to determine ground-state deformation and energies of vibrations.



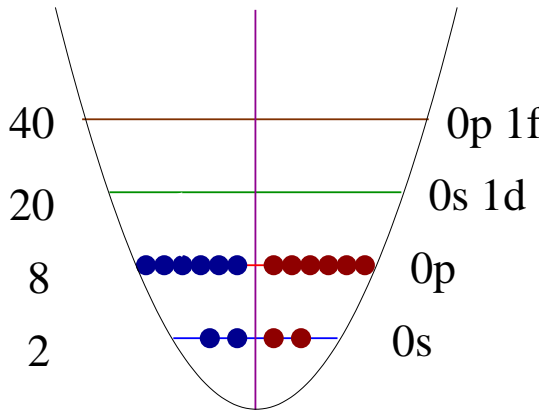
From C.G. Andersson et al, *Nucl. Phys.* **A268**, 205 (1976)

Microscopic Goldhaber-Teller Picture

Assume potential is harmonic oscillator

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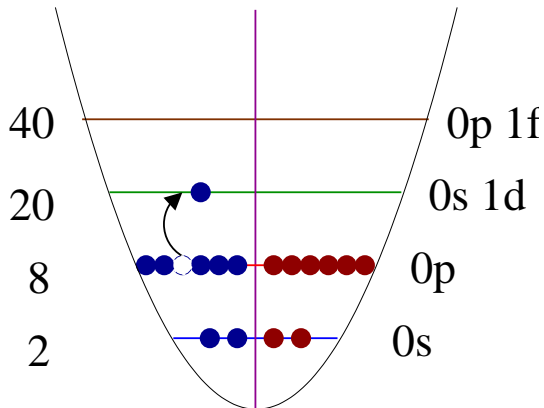


The operator $\sum_p (\vec{r}_p - \vec{R}_{cm})$
acts on the ^{16}O ground state

...

Microscopic Goldhaber-Teller Picture

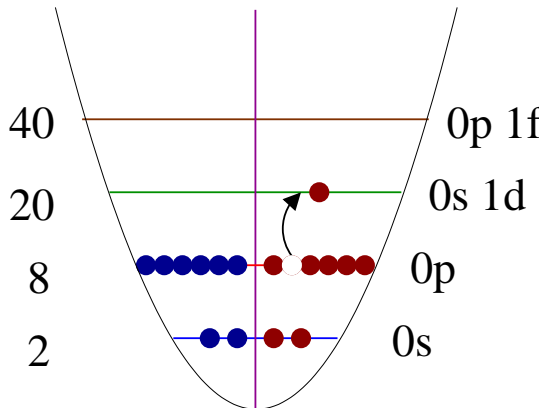
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promoting a proton up one oscillator shell ...

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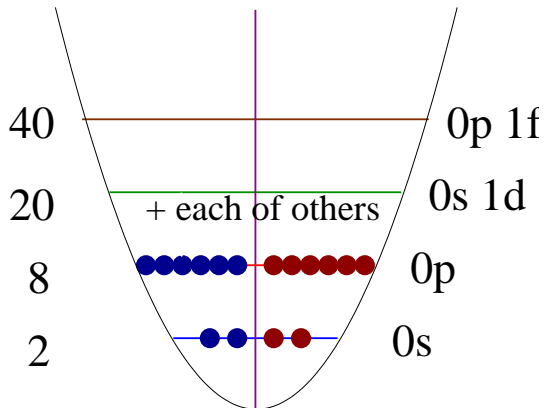
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and a neutron ...

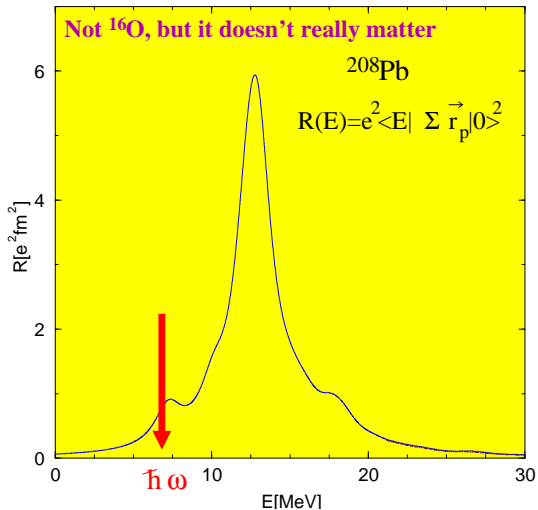
Microscopic Goldhaber-Teller Picture

Assume potential is harmonic oscillator



and doing this “coherently”
 (with a common amplitude for
 protons and another for
 neutrons) to all the p -shell
 particles...

Microscopic Goldhaber-Teller Picture



creating the resonant state
with energy about $1 \hbar\omega$.

⋮

Well, though I haven't shown
it to you, at least the size of
the peak turns out about right.
Something's still missing from
the energy, though.

Next Time ...

These early attempts to merge the shell-model and liquid-drop pictures — to understand the “microscopy” of collective nuclear motion — provide insight but are crude. Next time we’ll start to examine new methods that let us do the job better, starting from the nucleon-nucleon interaction. Then it’s on to the the drip lines, and after that to the role nuclear structure in the search for new fundamental physics.