## X-Ray Data Booklet

### 4.3 GRATINGS AND MONOCHROMATORS

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## A. DIFFRACTION PROPERTIES

## A. 1 Notation and sign convention

We adopt the notation of Fig. 4.6 in which $\alpha$ and $\beta$ have opposite signs if they are on opposite sides of the normal.

## A. 2 Grating equation

The grating equation may be written

$$
\begin{equation*}
m \lambda=d_{0}(\sin \alpha+\sin \beta) \tag{1}
\end{equation*}
$$

The angles $\alpha$ and $\beta$ are both arbitrary, so it is possible to impose various conditions relating them. If this is done, then for each $\lambda$, there will be a unique $\alpha$ and $\beta$. The following conditions are used:
(I) ON-BLAZE CONDITION:

$$
\begin{equation*}
\alpha+\beta=2 \theta_{B} \tag{2}
\end{equation*}
$$

where $\theta_{B}$ is the blaze angle (the angle of the sawtooth). The grating equation is then

$$
\begin{equation*}
m \lambda=2 d_{0} \sin \theta_{B} \cos \left(\beta+\theta_{B}\right) . \tag{3}
\end{equation*}
$$



Fig. 4-6. Grating equation notation.
(II) FIXED IN AND OUT DIRECTIONS:

$$
\begin{equation*}
\alpha-\beta=2 \theta, \tag{4}
\end{equation*}
$$

where $2 \theta$ is the (constant) included angle. The grating equation is then

$$
\begin{equation*}
m \lambda=2 d_{0} \cos \theta \sin (\theta+\beta) . \tag{5}
\end{equation*}
$$

In this case, the wavelength scan ends when $\alpha$ or $\beta$ reaches $90^{\circ}$, which occurs at the horizon wavelength $\lambda_{\mathrm{H}}=2 d_{0} \cos ^{2} \theta$

## (III) CONSTANT INCIDENCE ANGLE: EQUATION (1) GIVES $\beta$ DIRECTLY.

(IV) CONSTANT FOCAL DISTANCE (OF A PLANE GRATING):

$$
\begin{equation*}
\frac{\cos \beta}{\cos \alpha}=\text { a constant } c_{f f} \tag{6}
\end{equation*}
$$

leading to a grating equation

$$
\begin{equation*}
1-\left(\frac{m \lambda}{d_{0}}-\sin \beta\right)^{2}=\frac{\cos ^{2} \beta}{c_{f f}^{2}} \tag{7}
\end{equation*}
$$

Equations (3), (5), and (7) give $\beta$ (and thence $\alpha$ ) for any $\lambda$. Examples of the above $\alpha-\beta$ relationships are as follows:
(i) Kunz et al. plane-grating monochromator (PGM) [1], Hunter et al. double PGM[2], collimated-light SX700 [3]
(ii) Toroidal-grating monochromators (TGMs) [4, 5], spherical-grating monochromators (SGMs, "Dragon" system) [6], Seya-Namioka [7, 8] most aberration-reduced holographic SGMs [9], variable-angle SGM[10], PGMs [11, 12, 13]
(iii) Spectrographs, "Grasshopper" monochromator [14]
(iv) SX700 PGM [15] and variants [10, 16, 3]

## B. FOCUSING PROPERTIES [17]

The study of diffraction gratings[18, 19] goes back more than a century and has included plane, spherical [20, 21, 22], toroidal [23] and ellipsoidal[24] surfaces and groove patterns made by classical ("Rowland") ruling [25], holography [26, 27, 28] and variably-spaced ruling [29]. In recent years the optical design possibilities of holographic groove patterns [30, 31, 32] and variably-spaced rulings [13] have been extensively developed. Following normal practice, we provide an analysis of the imaging properties of gratings by means of the path function $F$ [32]. For this purpose we use the notation of Fig. 4.7, in which the zeroth groove (of width $d_{0}$ ) passes through the grating pole O , while the $n$th groove passes through the variable point $\mathrm{P}(w, l)$. The holographic groove pattern is supposed to be made using two coherent point sources C and D with cylindrical polar coordinates $\left(r_{\mathrm{C}}, \gamma, z_{\mathrm{C}}\right),\left(r_{\mathrm{D}}, \delta, z_{\mathrm{D}}\right)$ relative to O . The lower (upper) sign in eq. (9) refers


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Fig. 4-7*. Focusing properties notation.
to C and D both real or both virtual (one real and one virtual) for which case the equiphase surfaces are confocal hyperboloids (ellipses) of revolution about CD. The grating with varied line spacing $d(w)$ is assumed to be ruled according to $d(w)=d_{0}\left(1+v_{1} w+v_{2} w^{2}+\ldots\right)$. We consider all the gratings to be ruled on the general surface $x=\mathcal{L}_{i j} a_{i j} w l$ and the $a_{i j}$ coefficients[33] are given for the important cases in Tables 1 and 2.

## B. 1 Calculation of the path function $F$

$F$ is expressed as

$$
\begin{equation*}
F=\sum_{i j k} F_{i j k} w^{i} l^{j} \tag{8}
\end{equation*}
$$

where $F_{i j k}=z^{k} C_{i j k}(\alpha, r)+z^{\prime k} C_{i j k}\left(\beta, r^{\prime}\right)+\frac{m \lambda}{d_{0}} f_{i j k}$.
and the $f_{i j k}$ term, originating from the groove pattern, is given by one of the following expressions.

$$
f_{i j k}= \begin{cases}\delta_{(i-1) j k} & \text { Rowland }  \tag{9}\\ \frac{d_{0}}{\lambda_{0}}\left\{z_{\mathrm{C}}^{k} C_{i j k}\left(\gamma, r_{\mathrm{C}}\right) \pm z_{\mathrm{D}}^{k} C_{i j k}\left(\delta, r_{\mathrm{D}}\right)\right\} & \text { holographic } \\ n_{i j k} & \text { varied line spacing }\end{cases}
$$

The coefficient $F_{i j k}$ is related to the strength of the $i, j$ aberration of the wavefront diffracted by the grating. The coefficients $C_{i j k}$ and $n_{i j k}$ are given up to sixth order in Tables 3 and 4 in which the following notation is used:
$T=T(r, \alpha)=\frac{\cos ^{2} \alpha}{r}-2 a_{20} \cos \alpha \quad S=S(r, \alpha)=\frac{1}{r}-2 a_{02} \cos \alpha$

Table 1: Ellipse coefficients $Q_{i j}$ from which the $a_{i j}$ 's are obtained ${ }^{a}[33]$

| $\boldsymbol{j}$ | $\mathbf{0}$ | $\mathbf{1}$ | $\mathbf{2}$ | $\mathbf{3}$ | $\mathbf{4}$ | $\mathbf{5}$ | $\mathbf{6}$ |
| ---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\boldsymbol{i}$ |  |  | 0 | 1 | 0 | $C / 4$ | 0 |
| $\mathbf{0}$ | 0 | 0 | $A$ | 0 | $3 A C / 4$ | 0 | $*$ |
| $\mathbf{1}$ | 0 | 0 | $\left(2 A^{2}+C\right) / 2$ | 0 | $3 C\left(4 A^{2}+C\right) / 8$ | 0 | $*$ |
| $\mathbf{2}$ | 1 | 0 | $A\left(2 A^{2}+3 C\right) / 2$ | 0 | $*$ | 0 | $*$ |
| $\mathbf{3}$ | $\left(4 A^{2}+C\right) / 4$ | 0 | $\left(8 A^{4}+24 A^{2} C+3 C^{2}\right) / 8$ | 0 | $*$ | 0 | $*$ |
| $\mathbf{4}$ | $A\left(4 A^{2}+3 C\right) / 4$ | 0 | $*$ | 0 | $*$ | 0 | $*$ |
| $\mathbf{5}$ | $\left(8 A^{4}+12 A^{2} C+C^{2}\right) / 8$ | 0 | $*$ | 0 | $*$ | 0 | $*$ |

${ }^{\text {a }}$ If $r, r^{\prime}$ and $\theta$ are the object distance, image distance, and incidence angle to the normal, respectively, then
$a_{i j}=a_{20} \frac{Q_{i j}}{\cos ^{j} \theta} \quad$ where $\quad a_{20}=\frac{\cos \theta}{4}\left(\frac{1}{r}+\frac{1}{r^{\prime}}\right), \quad A=\frac{\sin \theta}{2}\left(\frac{1}{r}-\frac{1}{r^{\prime}}\right), \quad C=A^{2}+\frac{1}{r r^{\prime}}$
The $a_{i j}$ 's for spheres, circular cylinders, paraboloids and hyperboloids can also be obtained from Tables 1 and 2 by suitable choices of the input parameters $r, r^{\prime}$ and $\theta$.

Table 2: Toroid ${ }^{a} a_{i j}$ 's [33]

| $\boldsymbol{j}$ | $\mathbf{0}$ | $\mathbf{1}$ | $\mathbf{2}$ | $\mathbf{3}$ | $\mathbf{4}$ | $\mathbf{5}$ | $\mathbf{6}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\boldsymbol{i}$ |  |  |  |  |  |  |  |
| $\mathbf{0}$ | 0 | 0 | $1 /(2 \rho)$ | 0 | $1 /\left(8 R^{3}\right)$ | 0 | $1 /\left(16 \rho^{5}\right)$ |
| $\mathbf{1}$ | 0 | 0 | 0 | 0 | 0 | 0 | $*$ |
| $\mathbf{2}$ | $1 /(2 R)$ | 0 | $1 /\left(4 \rho R^{2}\right)$ | 0 | $(2 \rho+R) /\left(16 \rho^{3} R^{3}\right)$ | 0 | $*$ |
| $\mathbf{3}$ | 0 | 0 | 0 | 0 | $*$ | 0 | $*$ |
| $\mathbf{4}$ | $1 /\left(8 R^{3}\right)$ | 0 | $3 /\left(16 \rho R^{4}\right)$ | 0 | $*$ | 0 | $*$ |
| $\mathbf{5}$ | 0 | 0 | $*$ | 0 | $*$ | 0 | $*$ |
| $\mathbf{6}$ | $1 /\left(16 R^{5}\right)$ | 0 | $*$ | 0 | $*$ | 0 | $*$ |

${ }^{\text {a }} R$ and $\rho$ are the major and minor radii of the bicycle-tire toroid we are considering.

## Table 3: Coefficients $C_{i j k}$ of the expansion of $F^{a}$

$$
\begin{aligned}
& C_{011}=-\frac{1}{r} \\
& C_{020}=\frac{S}{2} \\
& C_{022}=-\frac{S}{4 r^{2}}-\frac{1}{2 r^{3}} \\
& C_{031}=\frac{S}{2 r^{2}} \\
& C_{040}=\frac{4 a_{02}^{2}-S^{2}}{8 r}-a_{04} \cos \alpha \\
& C_{042}=\frac{a_{04} \cos \alpha}{2 r^{2}}+\frac{3 S^{2}-4 a_{02}^{2}}{16 r^{3}}+\frac{3 S}{4 r^{4}} \\
& C_{100}=-\sin \alpha \\
& C_{102}=\frac{\sin \alpha}{2 r^{2}} \\
& C_{111}=-\frac{\sin \alpha}{r^{2}} \\
& C_{120}=\frac{S \sin \alpha}{2 r}-a_{12} \cos \alpha \\
& C_{131}=-\frac{a_{12} \cos \alpha}{r^{2}}+\frac{3 S \sin \alpha}{2 r^{3}} \\
& C_{122}=\frac{a_{12} \cos \alpha}{2 r^{2}}-\frac{3 S \sin \alpha}{4 r^{3}}-\frac{3 \sin \alpha}{2 r^{4}} \\
& C_{200}=\frac{T}{2} \quad C_{140}=-a_{14} \cos \alpha+\frac{1}{2 r}\left(2 a_{02} a_{12}+a_{12} S \cos \alpha-a_{04} \sin 2 \alpha\right)+\frac{\sin \alpha}{8 r^{2}}\left(4 a_{02}^{2}-3 S^{2}\right) \\
& C_{202}=-\frac{T}{4 r^{2}}+\frac{\sin ^{2} \alpha}{2 r^{3}} \\
& C_{211}=\frac{T}{2 r^{2}}-\frac{\sin ^{2} \alpha}{r^{3}} \\
& C_{013}=\frac{1}{2 r^{3}} \\
& C_{300}=-a_{30} \cos \alpha+\frac{T \sin \alpha}{2 r} \\
& C_{220}=-a_{22} \cos \alpha+\frac{1}{4 r}\left(4 a_{20} a_{02}-T S-2 a_{12} \sin 2 \alpha\right)+\frac{S \sin ^{2} \alpha}{2 r^{2}} \\
& C_{222}=\frac{1}{2 r^{2}} a_{22} \cos \alpha+\frac{1}{8 r^{3}}\left(3 S T-4 a_{02} a_{20}+6 a_{12} \sin 2 \alpha\right)+\frac{3}{4 r^{4}}\left(T-2 S \sin ^{2} \alpha\right)-\frac{3 \sin ^{2} \alpha}{r^{5}} \\
& C_{231}=-\frac{1}{r^{2}} a_{22} \cos \alpha+\frac{1}{4 r^{3}}\left(-3 S T+4 a_{02} a_{20}-6 a_{12} \sin 2 \alpha\right)+\frac{3 S \sin ^{2} \alpha}{r^{4}} \\
& C_{240}=-a_{24} \cos \alpha+\frac{1}{2 r}\left(a_{12}^{2} \sin ^{2} \alpha+2 a_{04} a_{20}+a_{22} S \cos \alpha+a_{04} T \cos \alpha-a_{14} \sin 2 \alpha+2 a_{02} a_{22}\right) \\
& +\frac{1}{16 r^{2}}\left(-4 a_{02}^{2} T-8 a_{02} a_{20} S+12 a_{12} S \sin 2 \alpha+3 T S^{2}+16 a_{02} a_{12} \sin \alpha-8 a_{04} \sin 2 \alpha\right)+\frac{\sin ^{2} \alpha}{4 r^{3}}\left(2 a_{02}^{2}-3 S^{2}\right) \\
& C_{302}=\frac{a_{30} \cos \alpha}{2 r^{2}}-\frac{3 T \sin \alpha}{4 r^{3}}+\frac{\sin ^{3} \alpha}{2 r^{4}} \\
& C_{311}=-\frac{a_{30} \cos \alpha}{r^{2}}+\frac{3 T \sin \alpha}{2 r^{3}}-\frac{\sin ^{3} \alpha}{r^{4}}
\end{aligned}
$$

*Table 3: Coefficients $C_{i j k}$ of the expansion of $F^{a}$ (continued)

$$
\begin{aligned}
C_{320}= & -a_{32} \cos \alpha+\frac{1}{2 r}\left(2 a_{20} a_{12}+2 a_{30} a_{02}+a_{30} S \cos \alpha+a_{12} T \cos \alpha-a_{22} \sin 2 \alpha\right) \\
& +\frac{1}{4 r^{2}}\left(4 a_{20} a_{02} \sin \alpha-3 S T \sin \alpha-4 a_{12} \cos \alpha \sin ^{2} \alpha\right)+\frac{S \sin ^{3} \alpha}{2 r^{3}} \\
C_{400}= & -a_{40} \cos \alpha+\frac{1}{8 r}\left(4 a_{20}^{2}-T^{2}-4 a_{30} \sin 2 \alpha\right)+\frac{T \sin ^{2} \alpha}{2 r^{2}} \\
C_{402}= & -\frac{1}{16 r^{3}}\left(4 a_{20}^{2}+3 T^{2}+12 a_{30} \sin 2 \alpha\right)+\frac{a_{40} \cos \alpha}{2 r^{2}}-\frac{3 T \sin ^{2} \alpha}{2 r^{4}}+\frac{\sin ^{4} \alpha}{2 r^{5}} \\
C_{411}= & -\frac{a_{40} \cos \alpha}{r^{2}}+\frac{1}{8 r^{3}}\left(4 a_{20}^{2}-3 T^{2}-12 a_{30} \sin 2 \alpha\right)+\frac{3 T \sin ^{2} \alpha}{r^{4}}-\frac{\sin ^{4} \alpha}{r^{5}} \\
C_{420}= & -a_{42} \cos \alpha+\frac{1}{2 r}\left(2 a_{20} a_{22}+2 a_{12} a_{30} \sin ^{2} \alpha+2 a_{02} a_{40}-a_{32} \sin ^{2} \alpha+a_{40} S \cos \alpha+a_{22} T \cos \alpha\right) \\
& +\frac{1}{16 r^{2}}\left(-4 a_{20}^{2} S-8 a_{02} a_{20} T+3 S T^{2}+12 \sin 2 \alpha\left(a_{30} S+a_{12} T\right)+8 \sin \alpha\left(2 a_{02} a_{30}-2 a_{22} \sin 2 \alpha+2 a_{12} a_{20}\right)\right) \\
& +\frac{1}{2 r^{3}}\left(2 a_{02} a_{20} \sin ^{2} \alpha-3 S T \sin ^{2} \alpha-2 a_{12} \cos \alpha \sin ^{3} \alpha\right)+\frac{S \sin ^{4} \alpha}{2 r^{4}} \\
C_{500}= & -a_{50} \cos \alpha+\frac{1}{2 r}\left(2 a_{20} a_{30}+a_{30} T \cos \alpha-a_{40} \sin ^{2} \alpha\right)+\frac{\sin ^{2} \alpha}{2 r^{2}}\left(a_{20}^{2}-a_{30} \sin 2 \alpha\right)-\frac{3 T^{2} \sin \alpha}{8 r^{2}}+\frac{T \sin ^{3} \alpha}{2 r^{3}} \\
C_{600}= & -a_{60} \cos \alpha+\frac{1}{2 r}\left(a_{30}^{2} \sin ^{2} \alpha+2 a_{20} a_{40}+a_{40} T \cos \alpha-a_{50} \sin 2 \alpha\right) \\
& +\frac{1}{16 r^{2}}\left(-4 a_{20}^{2} T+T^{3}+16 a_{20} a_{30} \sin \alpha+12 a_{30} T \sin ^{2} 2 \alpha-16 a_{40} \cos \alpha \sin 2 \alpha\right) \\
& +\frac{1}{4 r^{3}}\left(2 a_{20}^{2} \sin ^{2} \alpha-3 T^{2} \sin ^{2} \alpha-4 a_{30} \cos \alpha \sin ^{3} \alpha\right)+\frac{T \sin ^{4} \alpha}{2 r^{4}}
\end{aligned}
$$

${ }^{\text {a The coefficients for which } i \leq 6, j \leq 4, k \leq 2, i+j+k \leq 6, j+k=\text { even are included in this table. The }}$ only addition to those is $C_{013}$, which has some interest, because, when the system is specialized to be symmetrical about the $x$ axis, it represents a Seidel aberration, namely distortion.

Table 4: Coefficients $n_{i j k}$ of the expansion of $F$ for a grating with variable line spacing

$$
\begin{array}{ll}
\hline n_{i j k}=0 \text { for } j, k \neq 0 & \\
n_{100}=1 & n_{400}=\left(-v_{1}^{3}+2 v_{1} v_{2}-v_{3}\right) / 4 \\
n_{200}=-v_{1} / 2 & n_{500}=\left(v_{1}^{4}-3 v_{1}^{2} v_{2}+v_{2}^{2}+2 v_{1} v_{3}-v_{4}\right) / 5 \\
n_{300}=\left(v_{1}^{2}-v_{2}\right) / 3 & n_{600}=\left(-v_{1}^{5}+4 v_{1}^{3} v_{2}-3 v_{1} v_{2}^{2}-3 v_{1}^{2} v_{3}+2 v_{2} v_{3}+2 v_{1} v_{4}-v_{5}\right) / 6
\end{array}
$$

## B. 2 Determination of the Gaussian image point

By definition the principal ray $\mathrm{AOB}_{0}$ arrives at the Gaussian image point $\mathrm{B}_{0}\left(r_{0}^{\prime}, \beta_{0}, z_{0}^{\prime}\right)$ (Fig. 4.7) and its direction is given by Fermat's principal which implies
$[\sigma F / \sigma w]_{w=0, l=0}=0, \quad\lfloor\sigma F / \sigma l\rfloor_{w=0, l=0}=0$ whence

$$
\begin{equation*}
\frac{m \lambda}{d_{0}}=\sin \alpha+\sin \beta_{0}, \quad \frac{z}{r}+\frac{z_{0}^{\prime}}{r_{0}^{\prime}}=0 \tag{11}
\end{equation*}
$$

The tangential focal distance $r_{0}^{\prime}$ is obtained by setting the focusing term $F_{200}$ equal to zero and is given by

$$
T(r, \alpha)+T\left(r_{0}^{\prime}, \beta_{0}\right)= \begin{cases}0 & \text { Rowland }  \tag{12}\\ -\frac{m \lambda}{\lambda_{0}}\left\{T\left(r_{\mathrm{C}}, \gamma\right) \pm T\left(r_{\mathrm{D}}, \delta\right)\right\} & \text { holographic } \\ \frac{v_{1} m \lambda}{d_{0}} & \text { varied line spacing }\end{cases}
$$

Equations (11) and (12) determine the Gaussian image point $\mathrm{B}_{0}$, and in combination with the sagittal focusing condition ( $F_{020}=0$ ), describe the focusing properties of grating systems under the paraxial approximation.

For a Rowland spherical grating the focusing condition (Eq. (12)) is

$$
\begin{equation*}
\left(\frac{\cos ^{2} \alpha}{r}-\frac{\cos \alpha}{R}\right)+\left(\frac{\cos ^{2} \beta}{r_{0}^{\prime}}-\frac{\cos \beta}{R}\right)=0 \tag{13}
\end{equation*}
$$

which has important special cases. (i) plane grating, $R=\infty$ implying $r_{0}^{\prime}=-r \cos ^{2} \alpha / \cos ^{2} \beta$, (ii) object and image on the Rowland circle, or $r=R \cos \alpha, r_{0}^{\prime}=R \cos \beta$ and $M=1$ and (iii) $\beta=90^{\circ}$ (Wadsworth condition). The focal distances of TGMs and SGMs are also determined by eq. (13).

## B. 3 Calculation of ray aberrations

In an aberrated system, the outgoing ray will arrive at the Gaussian image plane at a point $B_{R}$ displaced from the Gaussian image point $\mathrm{B}_{0}$ by the ray aberrations $\Delta y^{\prime}$ and $\Delta z^{\prime}$ (Fig. 4.7). The latter are given by $[34,35,36]$

$$
\begin{equation*}
\Delta y^{\prime}=\frac{r_{0}^{\prime}}{\cos \beta_{0}} \frac{\partial F}{\partial w}, \quad \Delta z^{\prime}=r_{0}^{\prime} \frac{\partial F}{\partial l} \tag{14}
\end{equation*}
$$

where $F$ is to be evaluated for $\mathrm{A}=(r, \alpha, z), \mathrm{B}=\left(r_{0}^{\prime}, \beta_{0}, z_{0}^{\prime}\right)$. By means of the expansion of $F$, these equations allow the ray aberrations to be calculated separately for each aberration type.

$$
\begin{equation*}
\Delta y_{i j k}^{\prime}=\frac{r_{0}^{\prime}}{\cos \beta_{0}} F_{i j k} i w^{i-1} l^{j}, \quad \Delta z_{i j k}^{\prime}=r_{0}^{\prime} F_{i j k} w^{i}{ }_{j} l^{j-1} \tag{15}
\end{equation*}
$$

Moreover, provided the aberrations are not too large, they are additive, so that they may either reinforce or cancel.

## C. DISPERSION PROPERTIES

## C. 1 Angular dispersion

$$
\begin{equation*}
\left(\frac{\partial \lambda}{\partial \beta}\right)_{\alpha}=\frac{d \cos \beta}{m} \tag{16}
\end{equation*}
$$

## C. 2 Reciprocal linear dispersion

$$
\begin{equation*}
\left(\frac{\partial \lambda}{\partial\left(\Delta y^{\prime}\right)}\right)_{\alpha}=\frac{d \cos \beta}{m r^{\prime}} \equiv \frac{10^{-3} d[\AA] \cos \beta}{m r^{\prime}[\mathrm{m}]} \AA / \mathrm{mm} \tag{17}
\end{equation*}
$$

## C. 3 Magnification (M)

$$
\begin{equation*}
M(\lambda)=\frac{\cos \alpha}{\cos \beta} \frac{r^{\prime}}{r} \tag{18}
\end{equation*}
$$

## C. 4 Phase-space acceptance ( $\varepsilon$ )

$$
\begin{equation*}
\varepsilon=N \Delta \lambda_{S_{1}}=N \Delta \lambda_{S_{2}}\left(\text { assuming } S_{2}=M S_{1}\right) \tag{19}
\end{equation*}
$$

where $N$ is the number of participating grooves.

## D. RESOLUTION PROPERTIES

The following are the main contributions to the width of the instrumental line spread function. The actual width is the vector sum.
(I) ENTRANCE SLIT (WIDTH $S_{1}$ ):

$$
\begin{equation*}
\Delta \lambda_{S_{1}}=\frac{S_{1} d \cos \alpha}{m r} . \tag{20}
\end{equation*}
$$

(II) EXIT SLIT (WIDTH $S_{2}$ ):

$$
\begin{equation*}
\Delta \lambda_{S_{2}}=\frac{S_{2} d \cos \beta}{m r^{\prime}} . \tag{21}
\end{equation*}
$$

(III) ABERRATIONS (OF PERFECTLY MADE GRATING):

$$
\begin{equation*}
\Delta \lambda_{A}=\frac{\Delta y^{\prime} d \cos \beta}{m r^{\prime}}=\frac{d}{m}\left(\frac{\partial F}{\partial w}\right) . \tag{22}
\end{equation*}
$$

(IV) SLOPE ERROR $\Delta \phi$ (OF IMPERFECTLY MADE GRATING):

$$
\begin{equation*}
\Delta \lambda_{S E}=\frac{d(\cos \alpha+\cos \beta) \Delta \phi}{m}, \tag{23}
\end{equation*}
$$

Note that, provided the grating is large enough, diffraction at the entrance slit always guarantees a coherent illumination of enough grooves to achieve the slit-width limited resolution and a diffraction contribution to the width need not be added to the above.

## 1. EFFICIENCY

The most accurate way to calculate grating efficiencies is by the full electromagnetic theory [37, 38] for which a code is available from Neviere. However, approximate scalar-theory calculations are often useful and, in particular, provide a way to choose the groove depth $(h)$ of a laminar grating. According to Bennett [39], the best value of the groove-width-to-period ratio $(r)$ is the one for which the usefully illuminated groove area is equal to the land area. The scalar theory efficiency of a laminar grating with $r=0.5$ is given by [40]
$E_{0}=\frac{R}{4}\left\{1+2(1-P) \cos \left(\frac{4 \pi h \cos \alpha}{\lambda}\right)+(1-P)^{2}\right\}$
$E_{m}= \begin{cases}\frac{R}{m^{2} \pi^{2}}\left\{1-2 \cos Q^{+} \cos \left(Q^{-}+\delta\right)+\cos ^{2} Q^{+}\right\} & m=\text { odd } \\ \frac{R}{m^{2} \pi^{2}} \cos ^{2} Q^{+} & m=\text { even }\end{cases}$
where
$P=\frac{4 h \tan \alpha}{d_{0}}, Q^{ \pm}=\frac{m \pi h}{d_{0}}(\tan \alpha \pm \tan \beta), \quad \delta=\frac{2 \pi h}{\lambda}(\cos \alpha+\cos \beta)$
and $R$ is the reflectance at angle $\sqrt{\alpha \beta}$.

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