

X-RAY SCATTERING FROM OPTICAL SURFACES



- Diffraction from an imperfect surface
- The general scattering equation
- The Debye-Waller-factor treatment
- Conventional versus fractal surfaces
- Specifications and performance estimates
- Effect of errors on the image
 - Strehl factor
 - Image width factor
- Natural breakdown into “figure” and “finish”
- Practical example

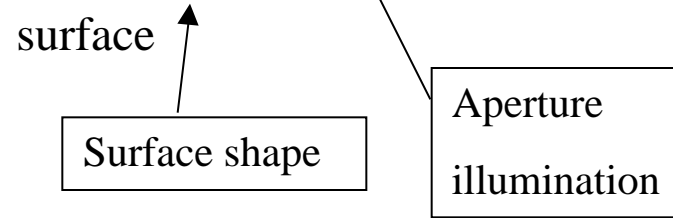
DIFFRACTION FROM AN IMPERFECT SURFACE



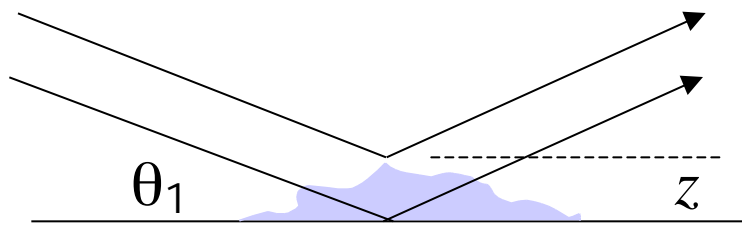
Fraunhofer diffraction integral:

S. Sinha, 1988; E. L. Church, various publications

$$U_1 = \frac{U_0 r(\theta_1) \sin(\theta_1)}{\lambda z} \int e^{i\Delta\phi(\mathbf{x})} A(\mathbf{x}) e^{2\pi i \mathbf{x} \cdot \mathbf{f}} d\mathbf{x}$$



$$\frac{1}{I_0} \frac{dI}{d\Omega} = \frac{R(\theta_1) \sin(\theta_1)}{\lambda^2} \left| \int \Omega(\mathbf{x}) e^{2\pi i \mathbf{x} \cdot \mathbf{f}} d\mathbf{x} \right|^2$$



$$\begin{aligned} \Delta\phi(\mathbf{x}) &= \frac{2\pi}{\lambda} 2z \sin\theta_1 \\ &= 2\pi f_z z(\mathbf{x}) \end{aligned}$$

INTERPRETATION OF THE INTEGRAL



$$\begin{aligned}
 \frac{1}{I_0} \frac{dI}{d\Omega} &= \frac{R(\theta_1) \sin(\theta_1)}{\lambda^2} \left| \int_{-}^{+} \Omega(\mathbf{x}) e^{2\pi i \mathbf{x} \cdot \mathbf{f}} d\mathbf{x} \right|^2 \\
 &= \frac{R(\theta_1) \sin(\theta_1)}{\lambda^2} \int e^{2\pi i \mathbf{f} \cdot (\mathbf{x}_1 - \mathbf{x}_2)} \Omega(\mathbf{x}_1) \Omega(\mathbf{x}_2) d\mathbf{x}_1 d\mathbf{x}_2 \\
 &= \frac{R(\theta_1) \sin(\theta_1)}{\lambda^2} \int e^{2\pi i \mathbf{f} \cdot \boldsymbol{\tau}} [\Omega \quad \Omega]_{\boldsymbol{\tau}} d\boldsymbol{\tau}
 \end{aligned}$$

$\boldsymbol{\tau}$ means the “lag” distance, $[\Omega \quad \Omega]_{\boldsymbol{\tau}}$ means correlation

MEANING OF THE AUTOCORRELATION OF Ω



$$\Omega(\mathbf{x}) = A(\mathbf{x}) e^{2\pi i f_z z(\mathbf{x})}$$

$$[\Omega \quad \Omega]_{\tau} = A(\mathbf{x} + \tau) A(\mathbf{x}) e^{2\pi i f_z \{z(\mathbf{x} + \tau) - z(\mathbf{x})\}} d\mathbf{x}$$

$$\begin{aligned} \langle [\Omega \quad \Omega]_{\tau} \rangle &= A(\mathbf{x} + \tau) A(\mathbf{x}) d\mathbf{x} \left\langle e^{2\pi i f_z \{z(\mathbf{x} + \tau) - z(\mathbf{x})\}} \right\rangle \\ &= AA(\tau) \cdot \chi_2(2\pi f_x, -2\pi f_x; \tau) \end{aligned}$$

- Autocorrelation of A
- System OTF

- Joint characteristic function of $z(\mathbf{x} + \tau)$ and $z(\mathbf{x})$
- Roughness OTF

DIGRESSION ON THE CHARACTERISTIC FUNCTION



For a probability distribution p , the characteristic function is

$$\chi_1(\omega) = \langle e^{i\omega z} \rangle = \int_{-\infty}^{+\infty} p(z) e^{i\omega z} dz$$

$$\chi_2(\omega_1, \omega_2) = \langle e^{i(\omega_1 z_1 + \omega_2 z_2)} \rangle = \int_{-\infty}^{+\infty} p(z_1, z_2) e^{i(\omega_1 z_1 + \omega_2 z_2)} dz_1 dz_2$$

We want $\langle (e^{i\omega z_1})(e^{i\omega z_2}) \rangle = \chi_2(\omega, -\omega)$ for the Gaussian bivariate distribution

$$p(Z_1, Z_2) = \frac{1}{2\pi\sigma^2 \sqrt{1-\rho^2}} \exp \left[-\frac{Z_1^2 - 2\rho Z_1 Z_2 + Z_2^2}{2(1-\rho^2)} \right] \quad \text{where } Z_1 = \frac{z_1}{\sigma} \text{ etc}$$

and $\rho = \langle Z_1 Z_2 \rangle$

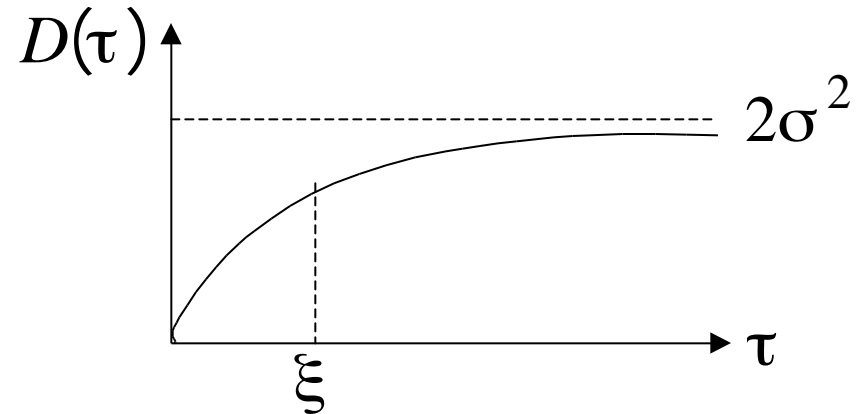
$$[\chi_2(\omega, -\omega)]_{\text{GB}} = e^{-\frac{1}{2}\omega^2 D(\tau)} \leftarrow \text{Structure function}$$

STRUCTURE FUNCTION



The structure function $D(\tau)$ appears naturally in scattering theory

$$\begin{aligned} D(\tau) &= \langle (z_1 - z_2)^2 \rangle \\ &= \langle z_1^2 - 2z_1z_2 + z_2^2 \rangle \end{aligned}$$



Correlation
length

In one or two dimensions:

$$D(\tau) = \text{FT}\{S_1(f_x)\} \quad \text{or} \quad \text{FT}\{S_2(\mathbf{f})\}$$

where S_1 and S_2 are the 1D and 2D
power spectra of the roughness

THE BASIC SURFACE-SCATTERING EQUATION



$$\left\langle \frac{1}{I_0} \frac{dI}{d\Omega} \right\rangle = \frac{\sin(\theta_1)}{\lambda^2} R(\theta_1) AA(\tau) \exp -\frac{1}{2} (2\pi f_z)^2 D(\tau) e^{2\pi i \mathbf{f} \cdot \tau} d\tau$$

- System OTF
- Coherence width W , where $1/W$ is a critical spatial wavelength
- WL's $> W$ diffract into the system image
- WL's $< W$ diffract outside the system image

- Joint characteristic function of $z(\mathbf{x} + \tau)$ and $z(\mathbf{x})$
- Roughness OTF

DEBYE-WALLER FACTOR TREATMENT

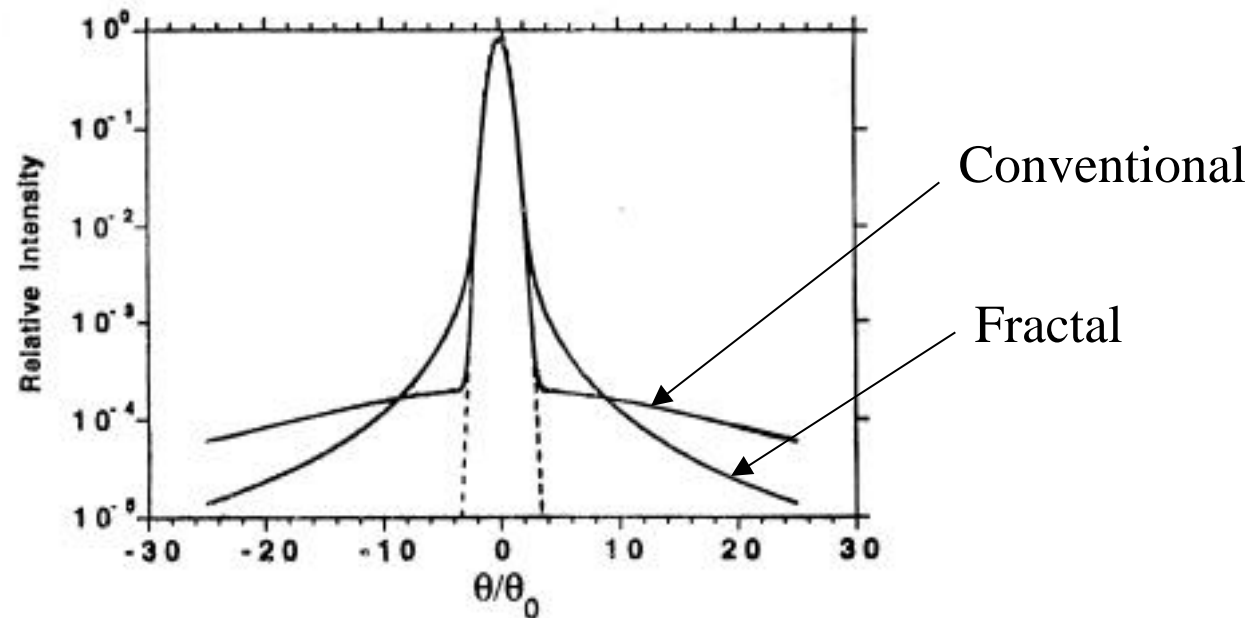


$$\left\langle \frac{1}{I_0} \frac{dI}{d\Omega} \right\rangle = \frac{\sin(\theta_1)}{\lambda^2} R(\theta_1) \int_{-W/2}^{W/2} AA(\tau) \exp \left[-\frac{1}{2} (2\pi f_z)^2 D(\tau) \right] e^{2\pi i f \cdot \tau} d\tau$$

- $AA(\tau)$ limits the range of τ to a width W
- $D(\tau)$ has a correlation length ξ
- If $\xi \ll W$ then most of the included roughness is uncorrelated: “conventional” surface
- In such case $D(\xi) = 2\sigma_z^2$ and the roughness OTF goes outside the integral

$$\left\langle \frac{1}{I_0} \frac{dI}{d\Omega} \right\rangle = \frac{1}{I_0} \frac{dI}{d\Omega} \text{ perfect mirror } e^{-(2\pi f_z \sigma_z)^2}$$

CONVENTIONAL VS REAL OPTICAL SURFACES



- For a good optical surface made on a large lap, the roughness correlations are often long range so x-ray mirrors normally have fractal rather than conventional surfaces
- Fractal surfaces have large ξ and PSD's of the form $S_1 = K_n / f^n$
- In the smooth surface limit the Gaussian can be approximated by its first two terms (valid for non-Gaussian surfaces as well)

SPECIFICATIONS AND PERFORMANCE ESTIMATES



$$\frac{I(\theta)}{I_s(0)} = \frac{1}{\sqrt{2\pi}} \int_0^{f_z \Theta} \cos(st) e^{-\frac{t^2}{8}} \left(1 - \frac{1}{2} f_z^2 D \frac{2t}{f_z \Theta}\right) dt$$

- Simplified 1D form for a mirror length L and magnification M

- $s = \frac{\theta}{\Theta}$

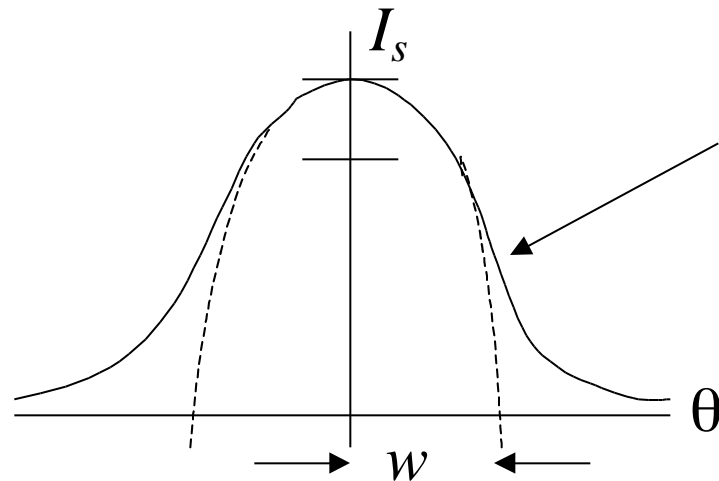
- $t = f_z \Theta \tau / 2$

- $D(\tau) = 4 \int_0^{f_x} S_1(f_x) \sin^2(\pi f_x \tau) df_x$

- $I_s(\theta) = I_s(0) \exp \left[-2 \left(\frac{\theta}{\Theta} \right)^2 \right]$

where $\Theta^2 = \frac{\lambda}{L \sin(\theta_1)} + (M \theta_{\text{source}})^2 + (\theta_{\text{detector}})^2$ and $W = \frac{\sqrt{2} \lambda}{\Theta \sin(\theta_1)}$

EFFECT OF ERRORS ON THE IMAGE



$$I_s(\theta) = I_s(0) \left[1 - a\theta^2 + \dots \right]$$

(without errors)

- For a real mirror (with errors) there will be a similar curve with reduced height and increased width
- Define two measures of error:
- Strehl factor: (Born and Wolf para 9.3) equals $I(0)/I_s(0)$
- Image width factor:

$$\frac{w_i}{w_{si}} = \frac{d^2 I(\theta)}{d\theta^2} \bigg|_0^{-1/2} \bigg/ \frac{d^2 I_s(\theta)}{d\theta^2} \bigg|_0^{-1/2}$$

STREHL FACTOR



(Marechal 1947)

$$\frac{I(0)}{I_s(0)} = 1 - \frac{4\pi \sin\theta_i}{\lambda} \sigma_{eff}^2 \quad \text{where} \quad \sigma_{eff}^2 = \int_0 S_1(f_x) \{1 - \exp(-W^2 f_x^2)\} df_x$$

$$\frac{I(0)}{I_s(0)} = 1 - \frac{8\mu_1^2}{\Theta^2} - \frac{4\pi \sin\theta_i}{\lambda} \sigma_1^2$$

$$\mu_1^2 = (2\pi)^2 \int_{1/L}^{1/W} S_1(f_x) f_x^2 df_x \quad \text{and} \quad \sigma_1^2 = \int_{1/W} S_1(f_x) df_x$$

RMS slope error
(figure)

RMS height error
(finish)

Acceptable value? traditionally 80%, say 10% figure error, 10% finish error

IMAGE-WIDTH FACTOR



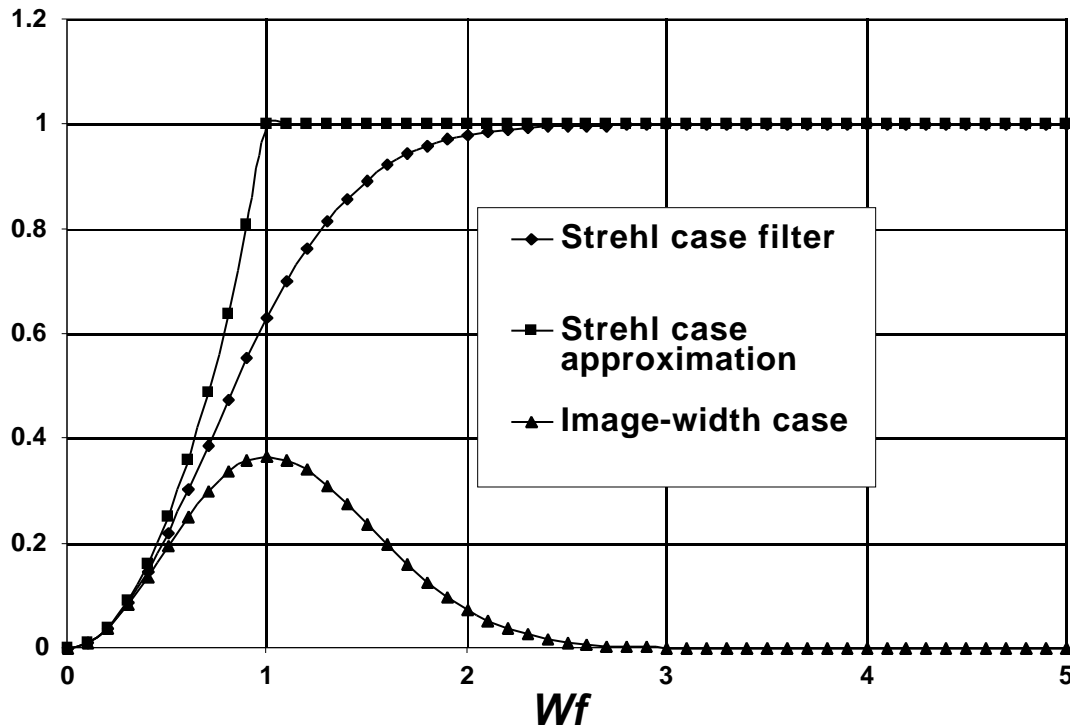
$$\frac{w_i}{w_{si}} = 1 + \frac{4\pi \sin \theta_i}{\lambda}^2 \sigma_{iw}^2$$

where

$$\sigma_{iw}^2 = \int_0 S_1(f_x) \left\{ W^2 f_x^2 \exp(-W^2 f_x^2) \right\} df_x$$

Acceptable value? let's say 120%

SPATIAL-FREQUENCY FILTER FUNCTIONS



$$1 - \exp(-W^2 f_x^2)$$

(Strehl factor)

$$W^2 f_x^2 \exp(-W^2 f_x^2)$$

(Image-width factor)

Approximation:

$$1 - \exp\left[-(W f_x)^2\right] \quad (W f_x)^2 \quad \text{when } 0 \leq f_x \leq 1/W \text{ (figure)}$$

$$1 \quad \text{otherwise (finish)}$$

FRACTAL POWER SPECTRA



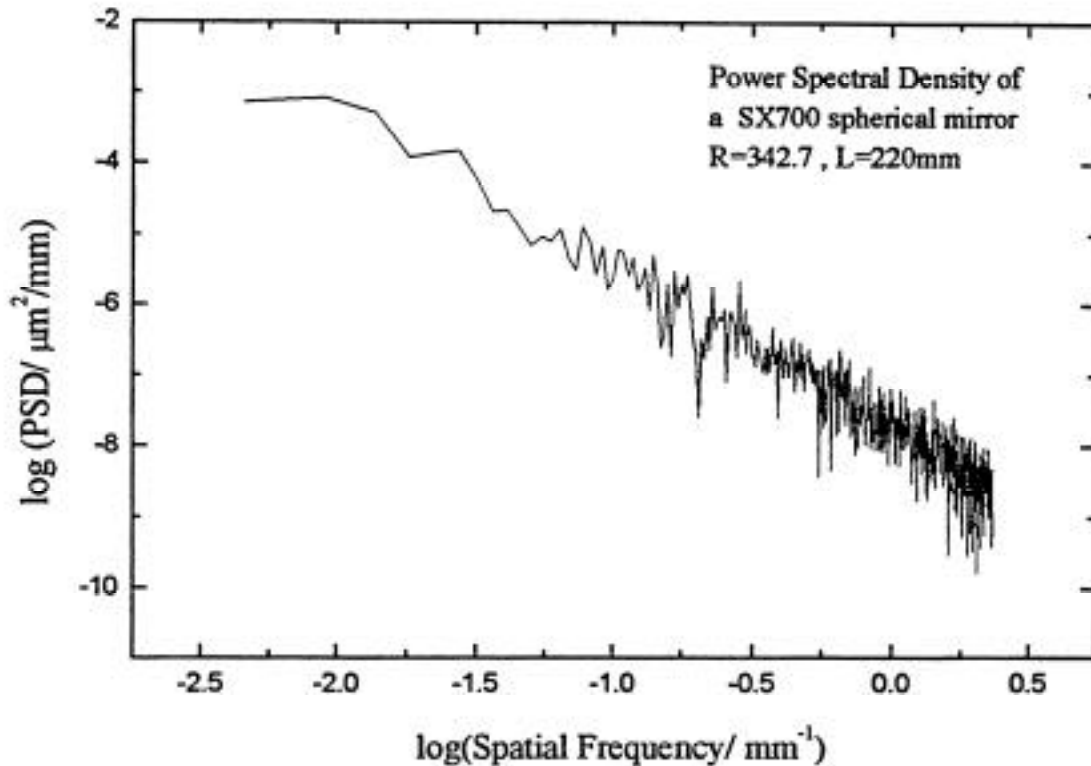
$$S_1(f_x) = \frac{K_n}{f_x^n}$$

$$\mu_1^2 = \frac{(2\pi)^2 K_n}{3-n} \frac{1}{W}^{3-n} - \frac{1}{L}^{3-n}$$

$$\sigma_1^2 = \frac{K_n}{1-n} \frac{1}{\lambda}^{1-n} - \frac{1}{W}^{1-n}$$

$$\sigma_{iw}^2 = \frac{K_n}{2} \Gamma \frac{3-n}{2} W^{n-1}$$

REAL MIRROR EXAMPLE



- Zeiss-made BESSY mirror measured by Zeiss
- Fractal power spectrum
- $S_1 = 2.4 \times 10^{-12} / f^{2.3}$
- Finish PSD not available
- Zeiss bench mark for mirror quality - this is their best routine production - anything better requires new tools to be made

SPECIFICATION AND PERFORMANCE SPREADSHEET



Scattering statistics

Mirror-----> Church paper Zeiss
 Opt Eng Proc SPIE
 Feb 95 3152
 Application--> Church paper SX700
 (test case)

Input data	Unit	Name		
Mirror length	m	L	1	0.22
Wavelength	nm	lam	0.15	2
Grazing angle	mr	alph	3	50
<i>Fractal PSD:</i>				
Kn	$\mu\text{m}^{(3-n)}$	kn	6.64E-09	2.40E-12
n	none	n	1.61	2.3
Acceptable Strehl decr	none		0.1	0.1
Acceptable image growth	none		0.2	0.2
Source size	μm	s	60	100
Source distance	m	sd	400	30
System magnification	none	m	0.25	0.2
Detector width	μm	dw	100	1
Calculated data				
Focal distance	m	F	100	6
System image width	μr	theta	1.012	3.342
Church's P	nm^{-2}	P	3.16E-02	4.93E-02
System coherence width (Constrained<0.95L)	mm	W	69.843	16.931
Strehl:				
Strehl factor	none		0.115	0.854
Roughness decrement	none		0.620	0.057
Figure decrement	none		0.265	0.089
Strehl % decrement	none		88.503	14.582
mu1 Strehl want	μr		0.113	0.374
mu1 Strehl get	μr		0.184	0.352
Sig1 Strehl want	nm		1.258	1.007
Sig1 Strehl get	nm		3.132	0.762
Image:				
Image growth factor	none		1.247	1.095
Image % increment	none		24.688	9.469
sigeff image want	nm		1.779	1.424
sigeff image get	nm		1.977	0.980





