

# An Update on Reducing the Uncertainty in Solar Radiometric Measurements

Daryl Myers, Ibrahim Reda, Stephen Wilcox  
National Renewable Energy Laboratory  
Golden Co 80401

Alisha Lester  
Smith College, Northampton, MA 01603

[daryl\\_myers@nrel.gov](mailto:daryl_myers@nrel.gov)

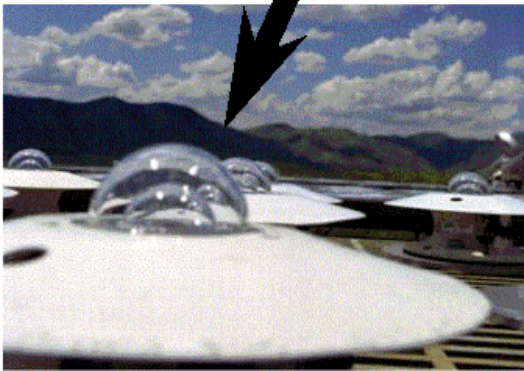
May 26-27, 2005  
Athens, Greece



# Solar Radiation Component Equation

Total (global)  $G$ , Direct beam,  $B$ , Diffuse Sky (scattered),  $D$ ,

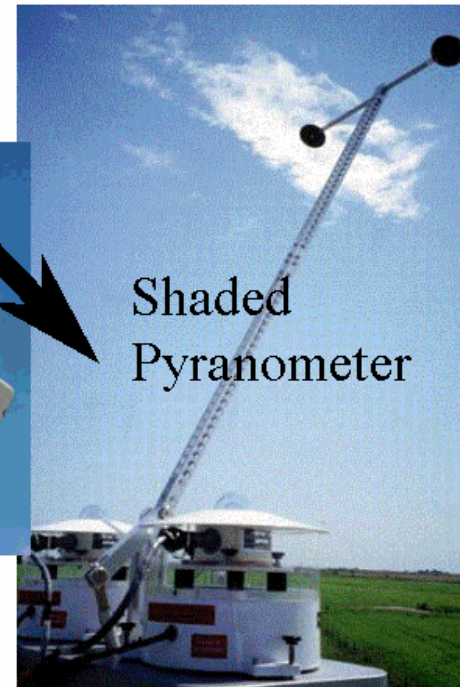
$$G = B \cos(i) + D, \quad i = \text{incidence angle}$$



Pyranometer



Pyrhelioscope



Shaded  
Pyranometer

# Solar Radiometer Responsivity Issues

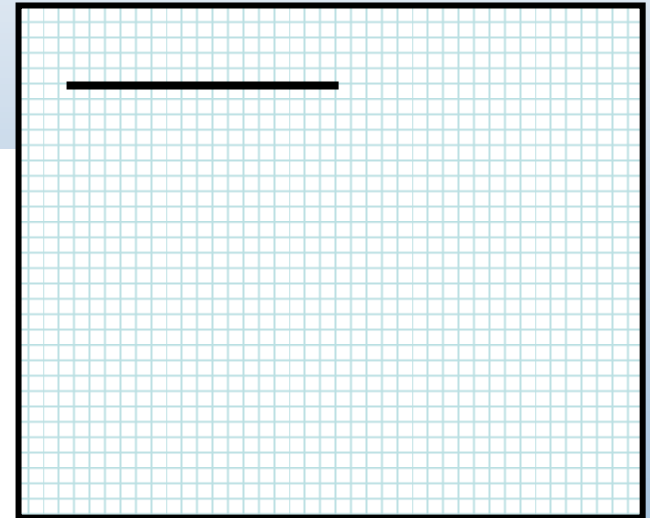
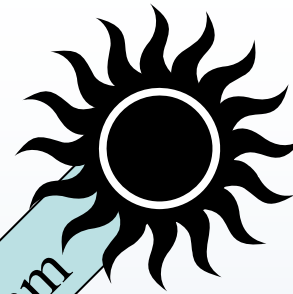
- Pyranometer Thermopile Offset IR voltage
  - Corrections at calibration time
  - Post-hoc correction schemes based on
    - "cosine response" though DAY and Year
    - IR radiation exchange error voltage
- Pyrheliometer Environmental Influences
  - Correction at calibration time  $\alpha$ 
    - Wind speed,  $d\text{Temperature}/dt$ , Irradiance
  - Post-hoc correction based on  $Ws, dT, I$

# Shade/Unshade Response

Beam x Cos(z)

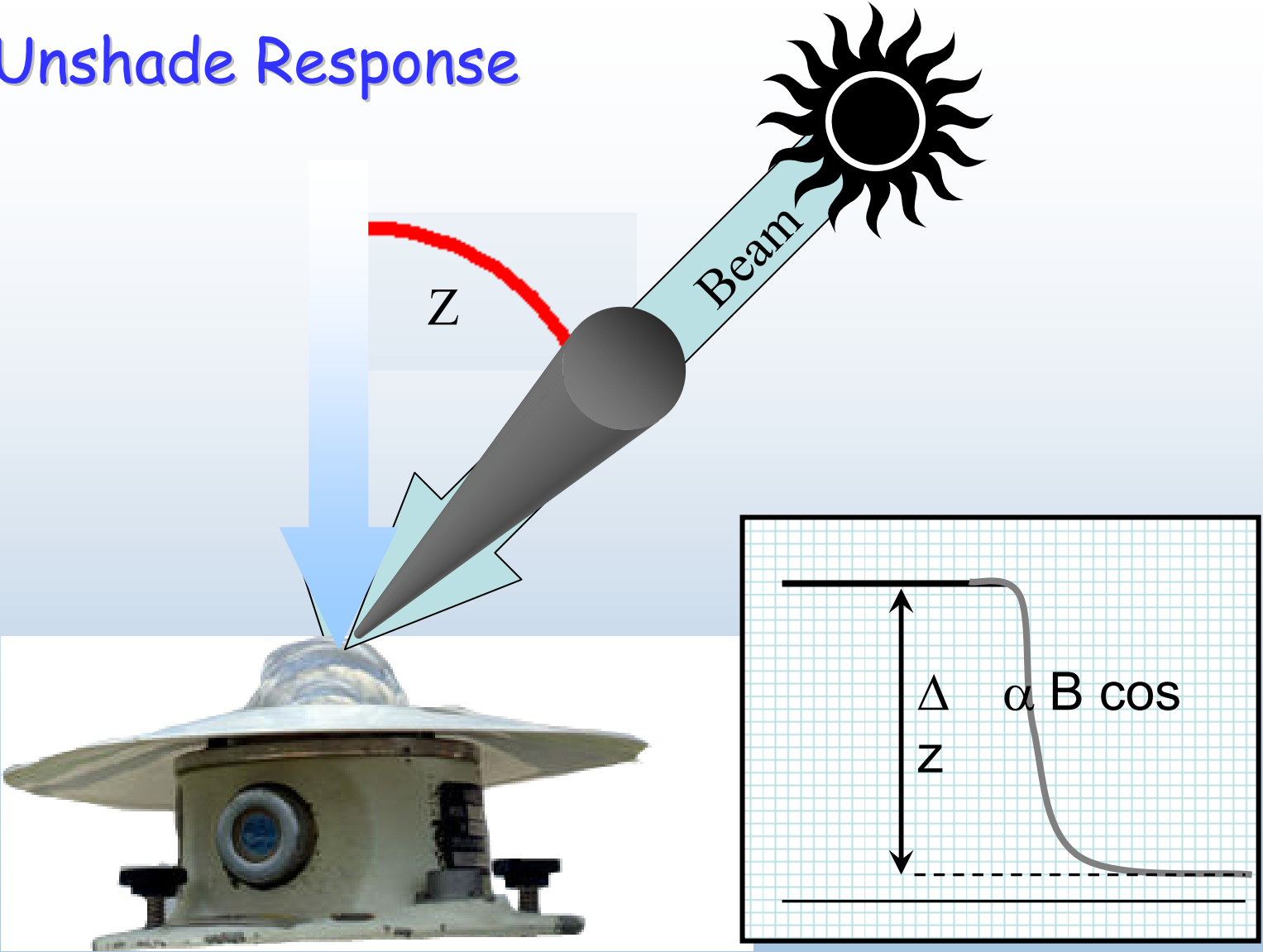
Z

Beam



Reda, I., T. Stoffel, D. Myers, *A Method to Calibrate a Solar Pyranometer for Measuring Reference Diffuse Irradiance*. *Solar Energy*, 2003. 74: p. p. 103-112.

# Shade/Unshade Response



Reda, I., T. Stoffel, D. Myers, *A Method to Calibrate a Solar Pyranometer for Measuring Reference Diffuse Irradiance*. *Solar Energy*, 2003. 74: p. p. 103-112.

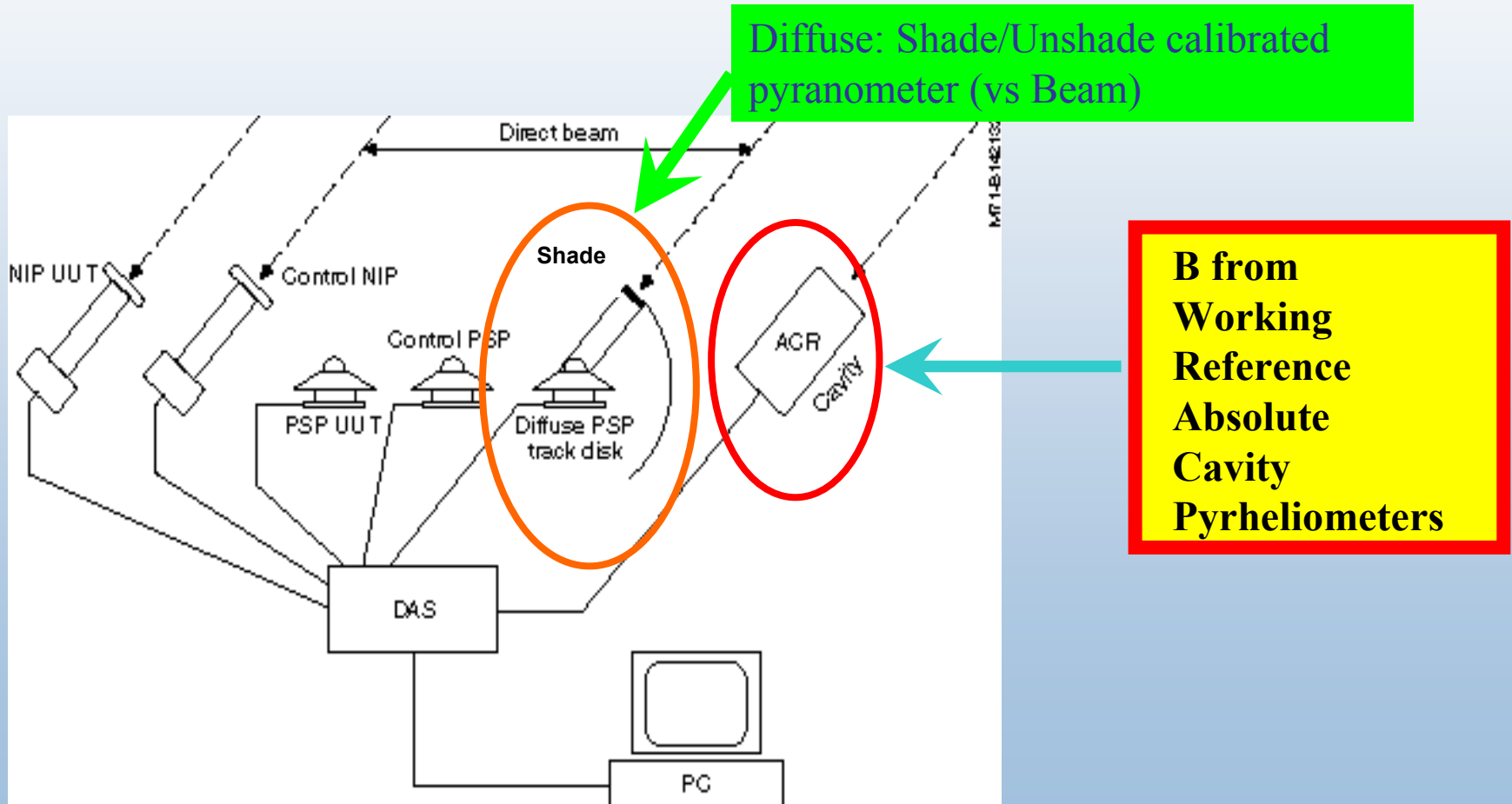
# Component Sum Method

$$R_s = \frac{U}{[B \cdot \cos(z) + D]}$$

U = Test Pyranometer signal volts;

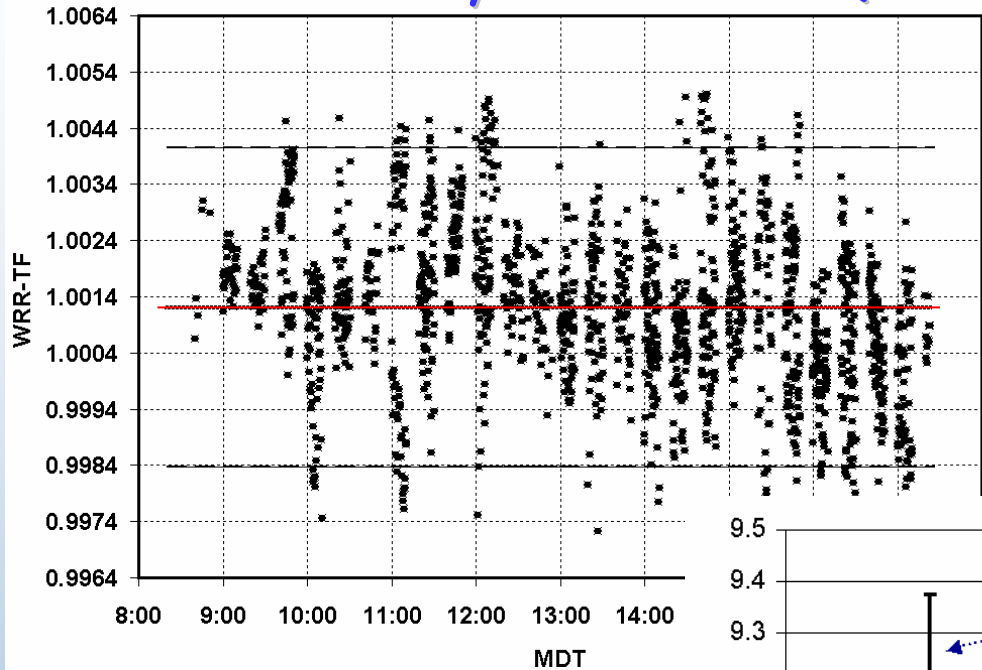
D= Diffuse

B = Beam radiation; Z = Zenith Angle



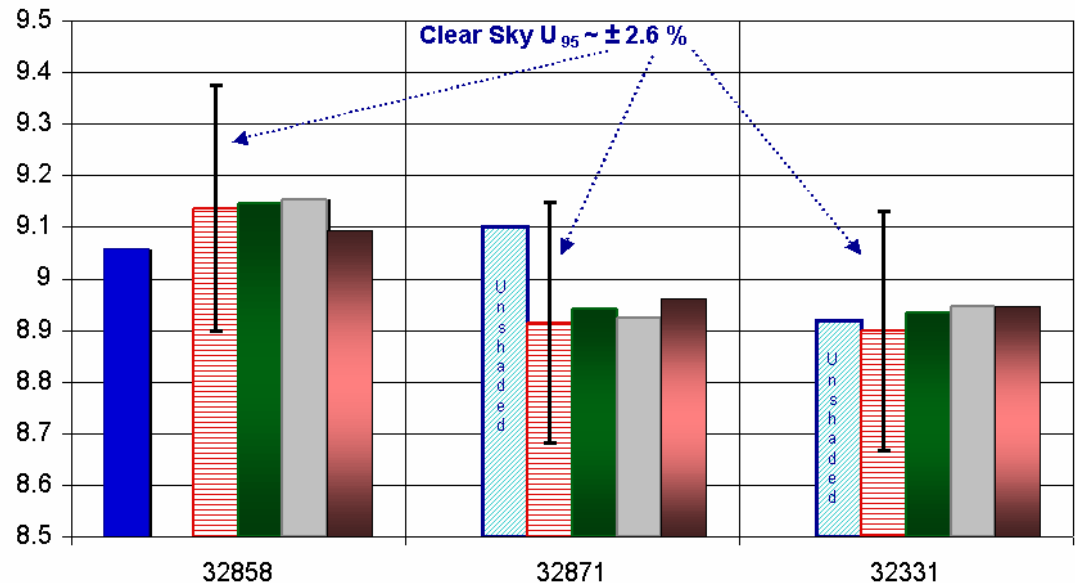
# Empirical Reference Irradiance Uncertainties

## Cavity Reference (World Radiometric Reference) Transfer Uncertainty

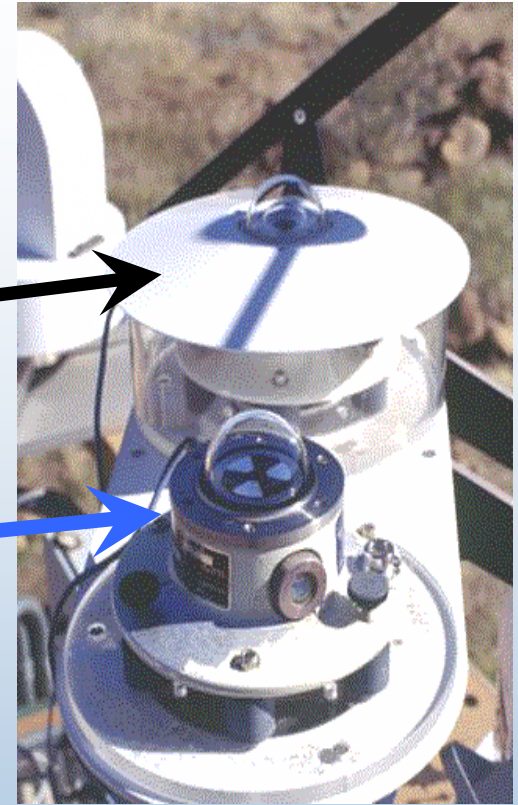
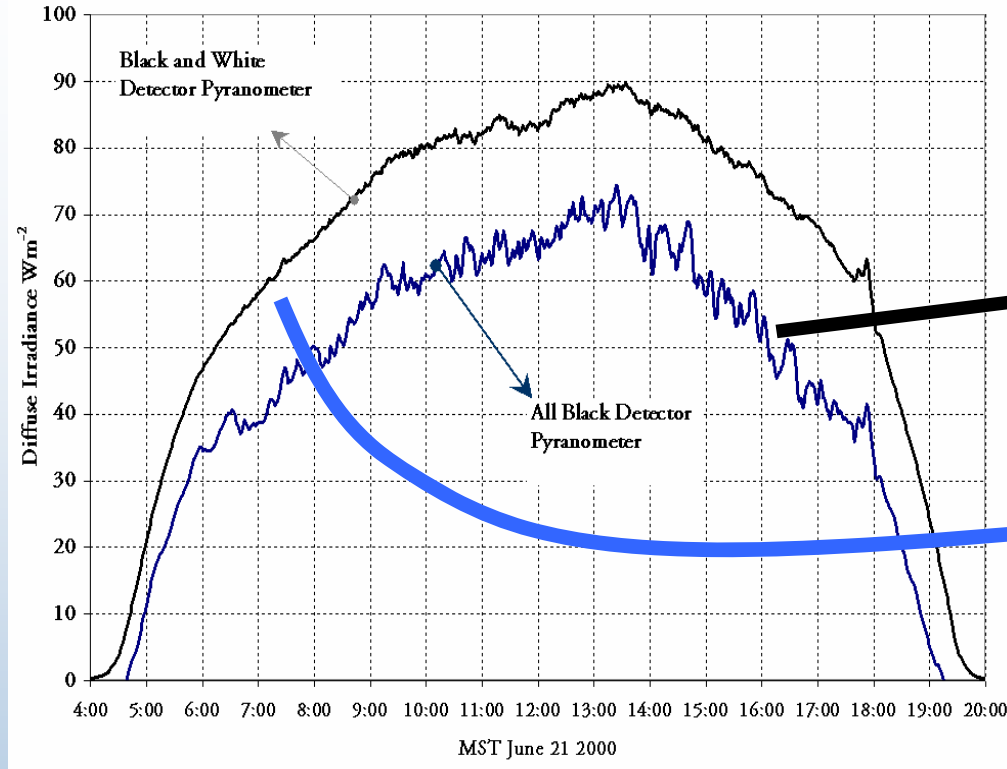


$\pm 0.35\%$

Diffuse  
Pyranometer  
Calibration  
Uncertainty



# Pyranometer Offset Error Signal

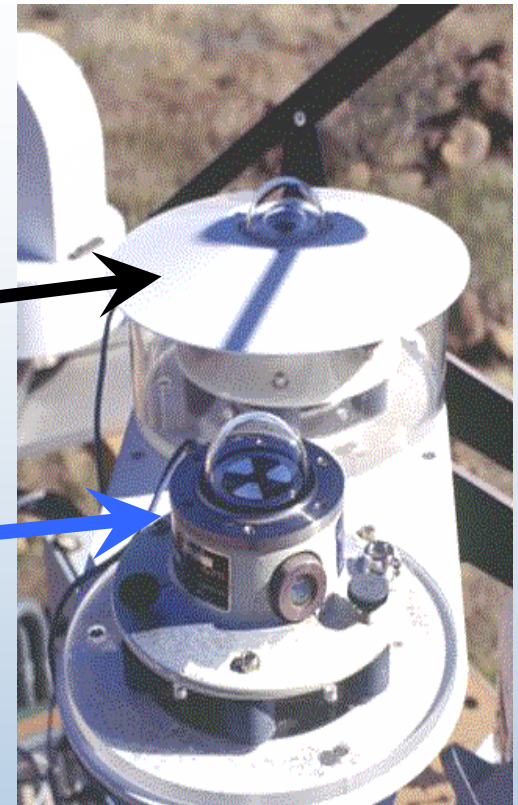
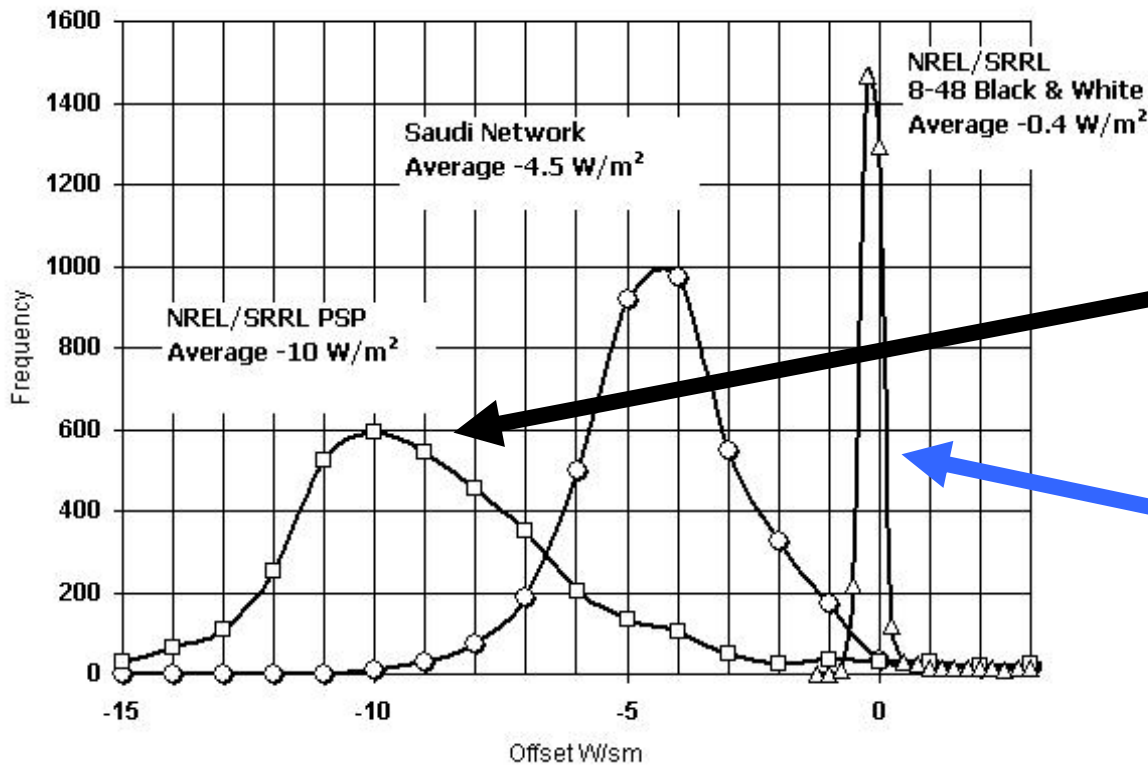


**All Black thermopile detectors with reference junctions in instrument body are never in thermal equilibrium; suffer  $-5 \text{ W/m}^2$  to  $-20 \text{ W/m}^2$  thermal offset. Offset produced by INFRARED exchange between detector & Sky/domes. Black & White reference and hot junctions in same thermal conditions, low thermal offsets. All-black unshaded units possess offset!**

Dutton, E. G., J. J. Michalsky, T. Stoffel, B. W. Forgan, J. Hickey, T. L. Alberta, I. Reda, *Measurement of Broadband Diffuse Solar Irradiance Using Current Commercial Instrumentation with a Correction for Thermal Offset Errors*. Journal of Atmospheric and Oceanic Technology, 2001. 18(3): p. 297-314



# Pyranometer Offset Error Signal



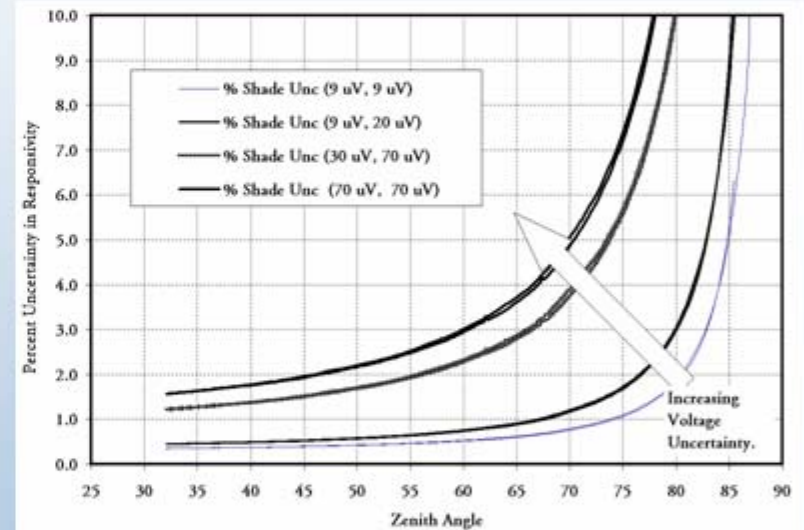
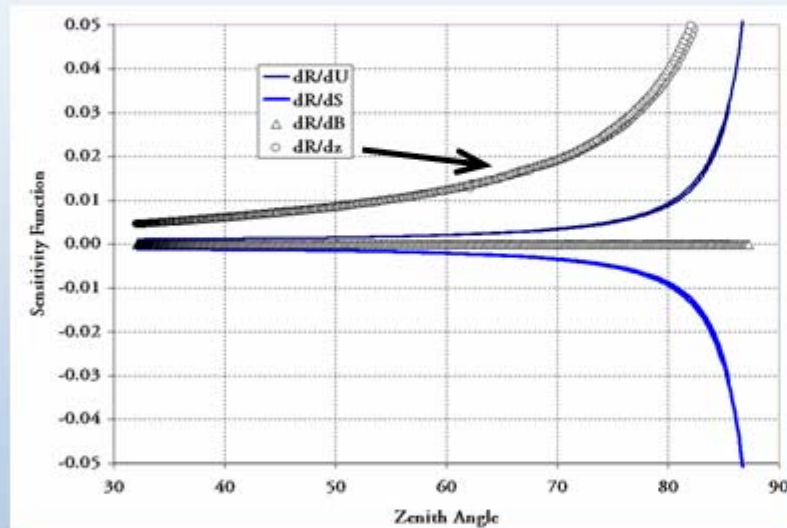
**All Black thermopile detectors with reference junctions in instrument body are never in thermal equilibrium; suffer  $-5 \text{ W/m}^2$  to  $-20 \text{ W/m}^2$  thermal offset. Offset produced by INFRARED exchange between detector & Sky/domes. Black & White reference and hot junctions in same thermal conditions, low thermal offsets. All-black unshaded units posses offset !**

Dutton, E. G., J. J. Michalsky, T. Stoffel, B. W. Forgan, J. Hickey, T. L. Alberta, I. Reda, *Measurement of Broadband Diffuse Solar Irradiance Using Current Commercial Instrumentation with a Correction for Thermal Offset Errors*. Journal of Atmospheric and Oceanic Technology, 2001. 18(3): p. 297-314

# Shade-Unshade Calibration

$$R_{\text{shade}} = \frac{(U + e_U) - (S + e_S)}{[B \cdot \text{Cos}(z)]}$$

$$U_{\text{shade}}^2 = (\partial_U R_s * e_U)^2 + (\partial_S R_s * e_S)^2 + (\partial_B R_s * e_B)^2 + (\partial_z R_s * e_z)^2$$



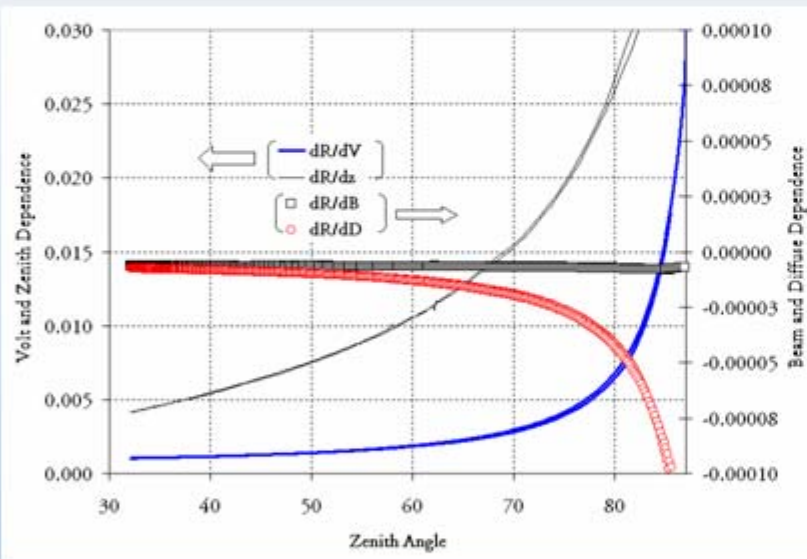
Sensitivity functions for shade-unshade calibration. Sensitivity to shade (negative line) and unshade (positive line) voltages are mirror image of each other. Greatest sensitivity is to zenith angle (circles). Negligible sensitivity to beam uncertainty.

Total Uncertainty in shade-unshade calibrations versus zenith angle for various uncertainties in voltage measurement with fixed beam ( $4 \text{ Wm}^{-2}$ ) and z angle ( $0.06^\circ$ ) uncertainty. Arguments in parenthesis are uncertainty in shade unshade voltages, respectively

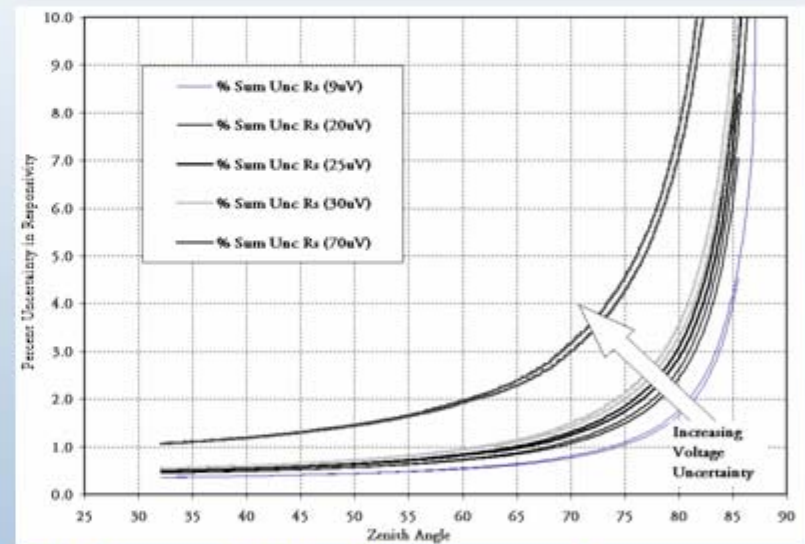
# Component Sum Calibration

$$R_{\text{sum}} = \frac{B \cdot \cos(z) + D}{U}$$

$$U_{\text{sum}}^2 = (\partial_U R_s * e_U)^2 + (\partial_D R_s * e_D)^2 + (\partial_B R_s * e_B)^2 + (\partial_z R_s * e_z)^2$$

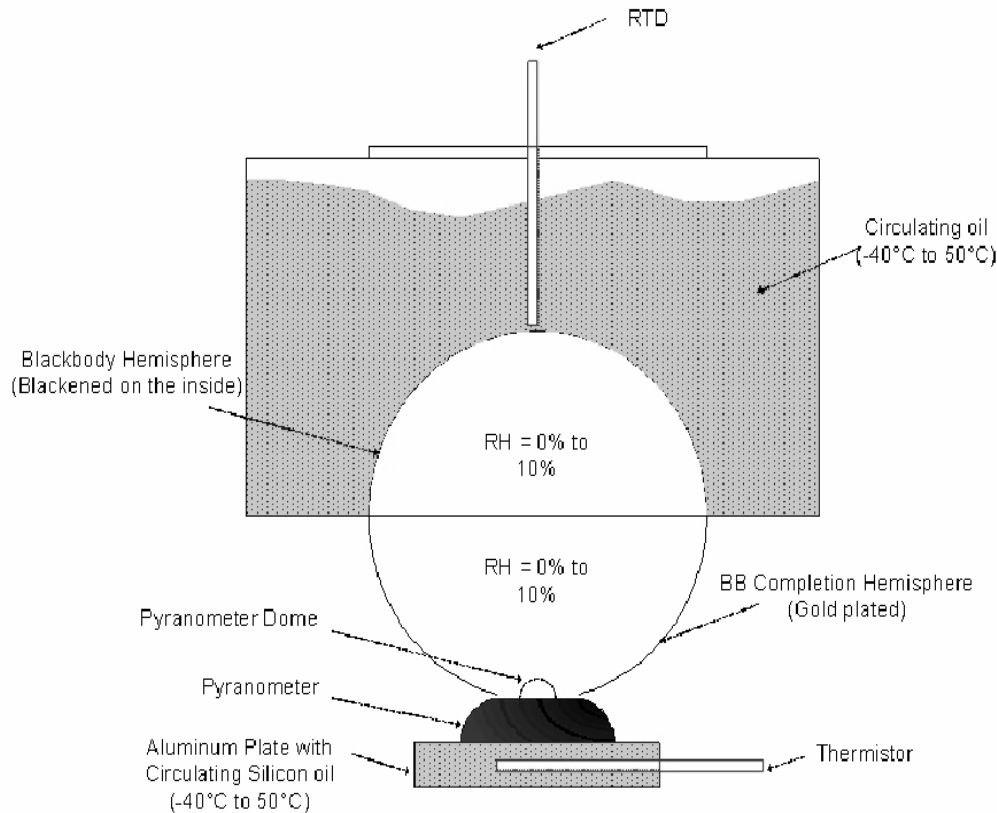


Sensitivity functions for component summation calibration. Sensitivity to beam (square) and diffuse (circle) irradiances are much less (right scale) than to voltage (heavy line) and zenith angle (light line) (left scale).



Total uncertainty in component sum calibrations as a function of zenith angle for various uncertainties in voltage measurement (in parenthesis), and fixed beam ( $4 \text{ Wm}^{-2}$ ), zenith angle ( $0.06^\circ$ ), and diffuse ( $2 \text{ Wm}^{-2}$ ) uncertainty.

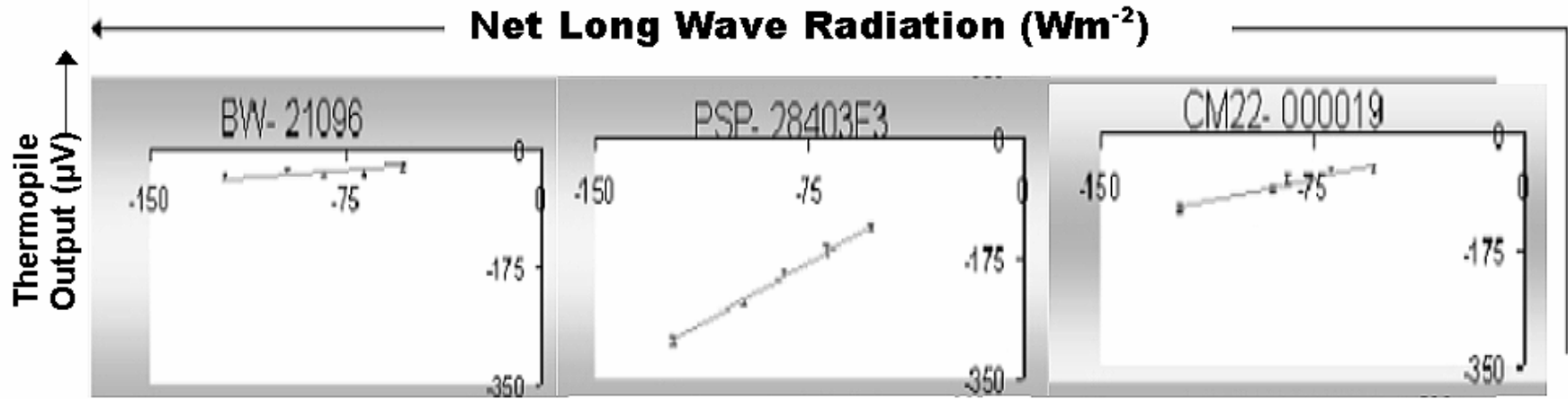
# Characterize shortwave pyranometer net-IR response using Blackbody IR system



$W_{NET}$ $W_{bb} - W_c$ ( $Wm^{-2}$ )	Pyran Signal ( $\mu V$ )	NET IR Rs(bb) $\mu V W^{-1} m^{-2}$
-110.8	-243.7	<b>2.20</b>
-60.3	-130.8	<b>2.17</b>
-131.6	-289.5	<b>2.20</b>
-71.3	-154.7	<b>2.17</b>
-83.6	-183.9	<b>2.20</b>

Reda, I., J. Hickey, C. Long, D. Myers, T. Stoffel, S. Wilcox, J.J. Michalsky, E.G. Dutton, D. Nelson, *Using a Blackbody to Calculate Net-Longwave Responsivity of Shortwave Solar Pyranometers to Correct for Their Thermal Offset Error During Outdoor Calibration Using the Component Sum Method*. Journal of Atmospheric and Oceanic Technology, 2005. In Press.

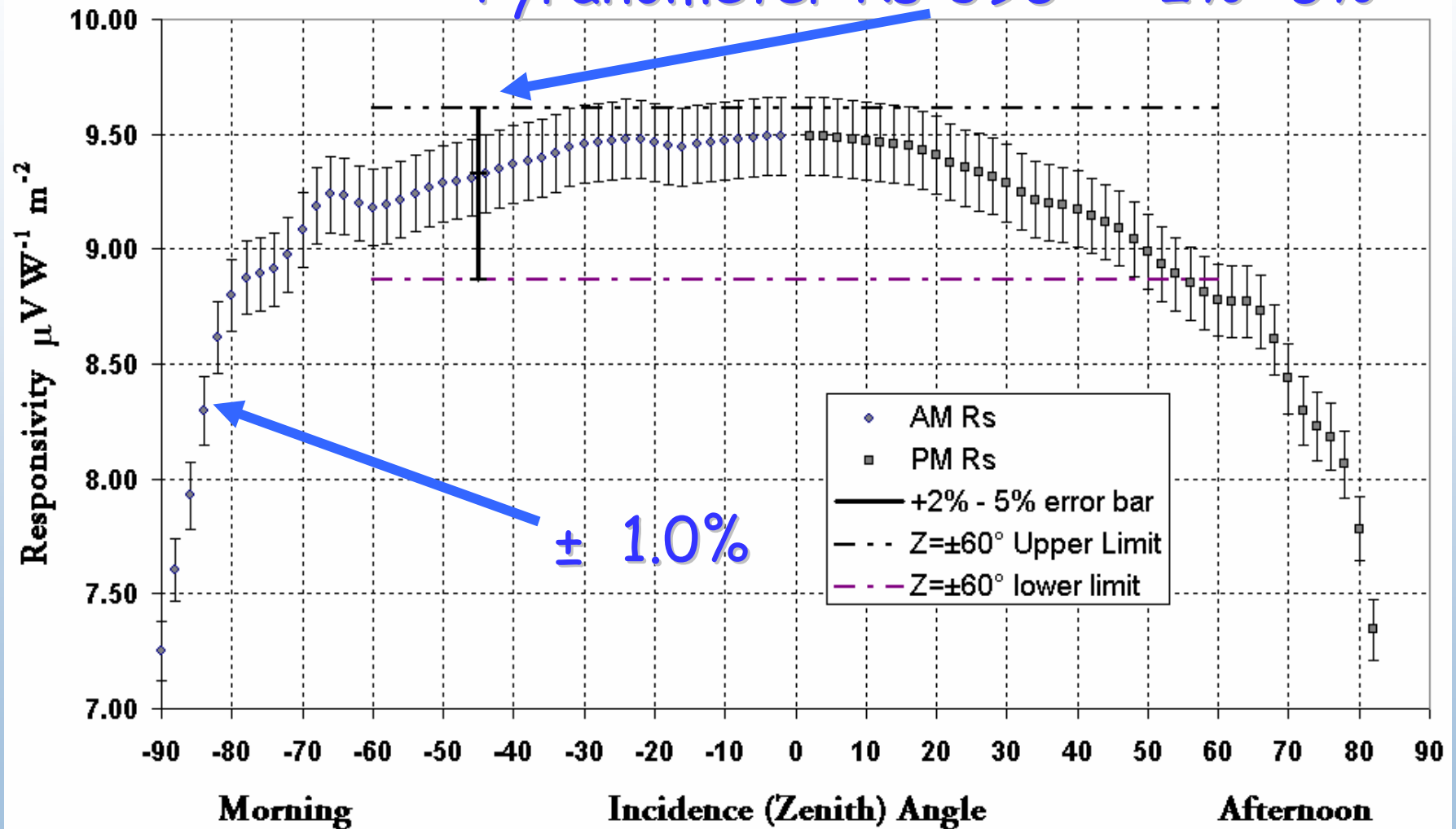
# Shortwave pyranometer signals in response to net infrared (longwave) radiation



Pyranometer Model	# Tested	RS <sub>bb</sub> (µV/Wm <sup>-2</sup> )	RS <sub>MFR</sub> (µV/Wm <sup>-2</sup> )
EPLAB 8-48	2	0.8314	9.465
K&Z CM-22	1	0.8872	9.300
<b>EPLAB PSP</b>	<b>12</b>	<b>2.1757</b>	<b>8.46</b>
Spectrosun SR-75	1	1.1851	8.69

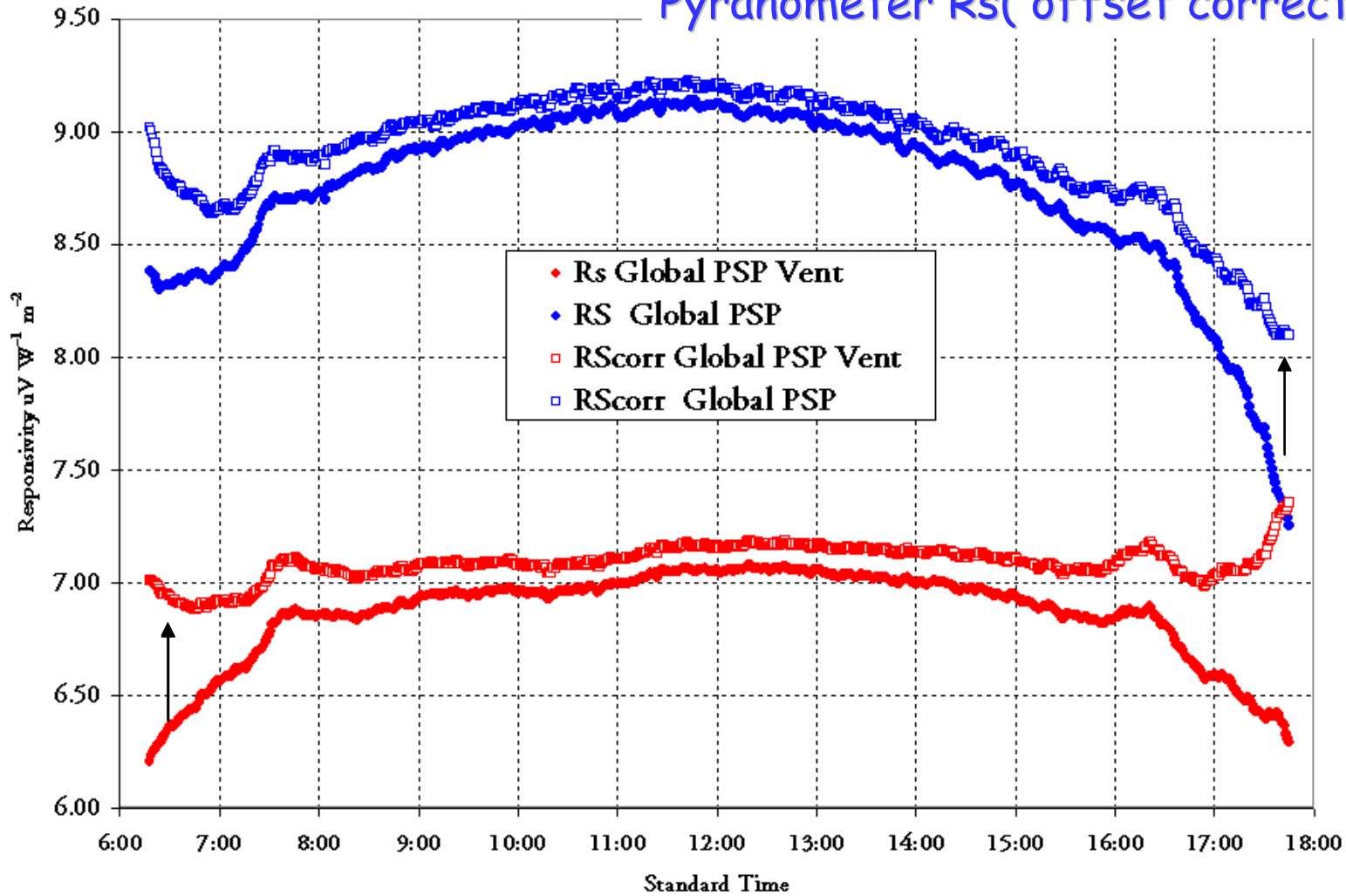
# Pyranometer Responsivity Calibration Results

Pyranometer Rs U95 = +2% -5%



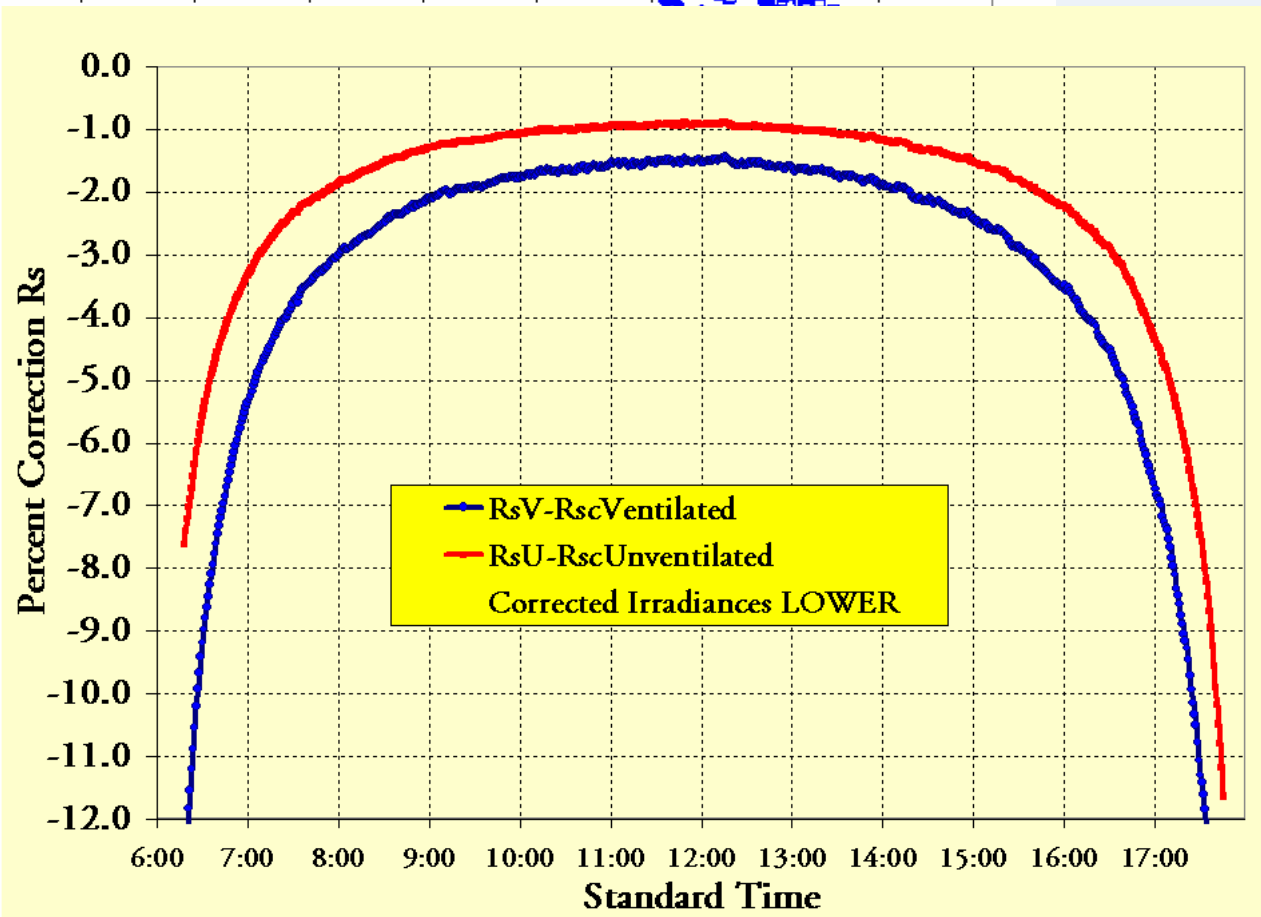
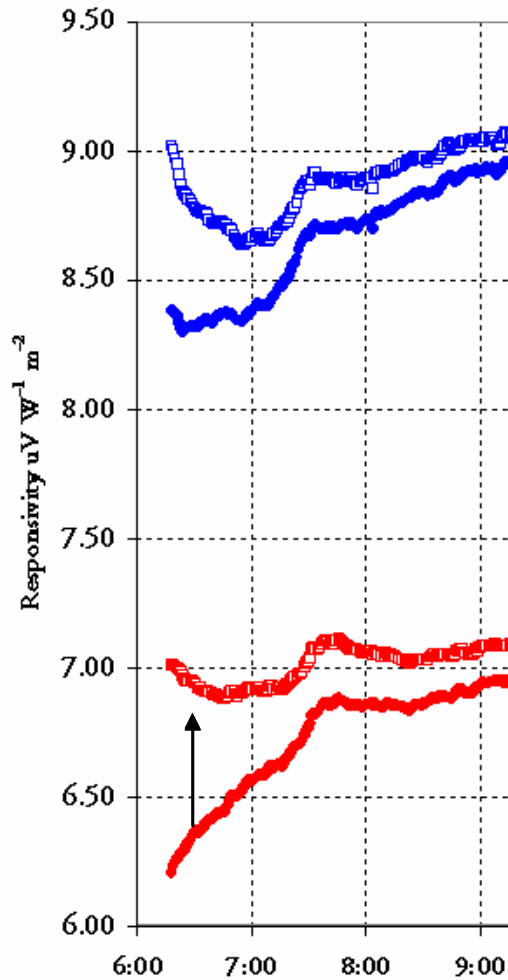
# Pyranometer Calibration IR Corrections

Pyranometer Rs (offset corrected)



# Pyranometer Calibration IR Corrections

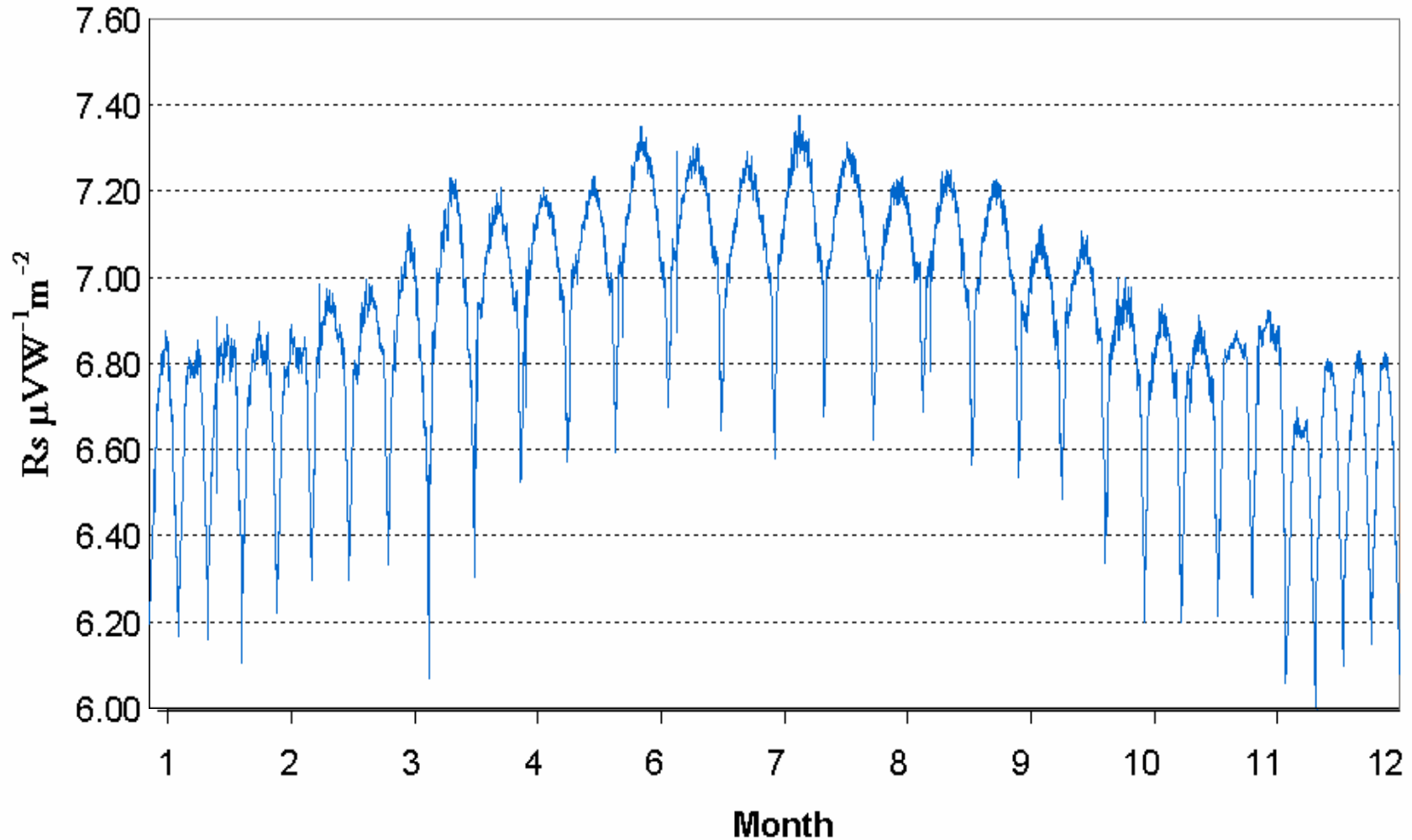
Pyranometer Rs (offset corrected)





# "Daily" Calibration & Characterization

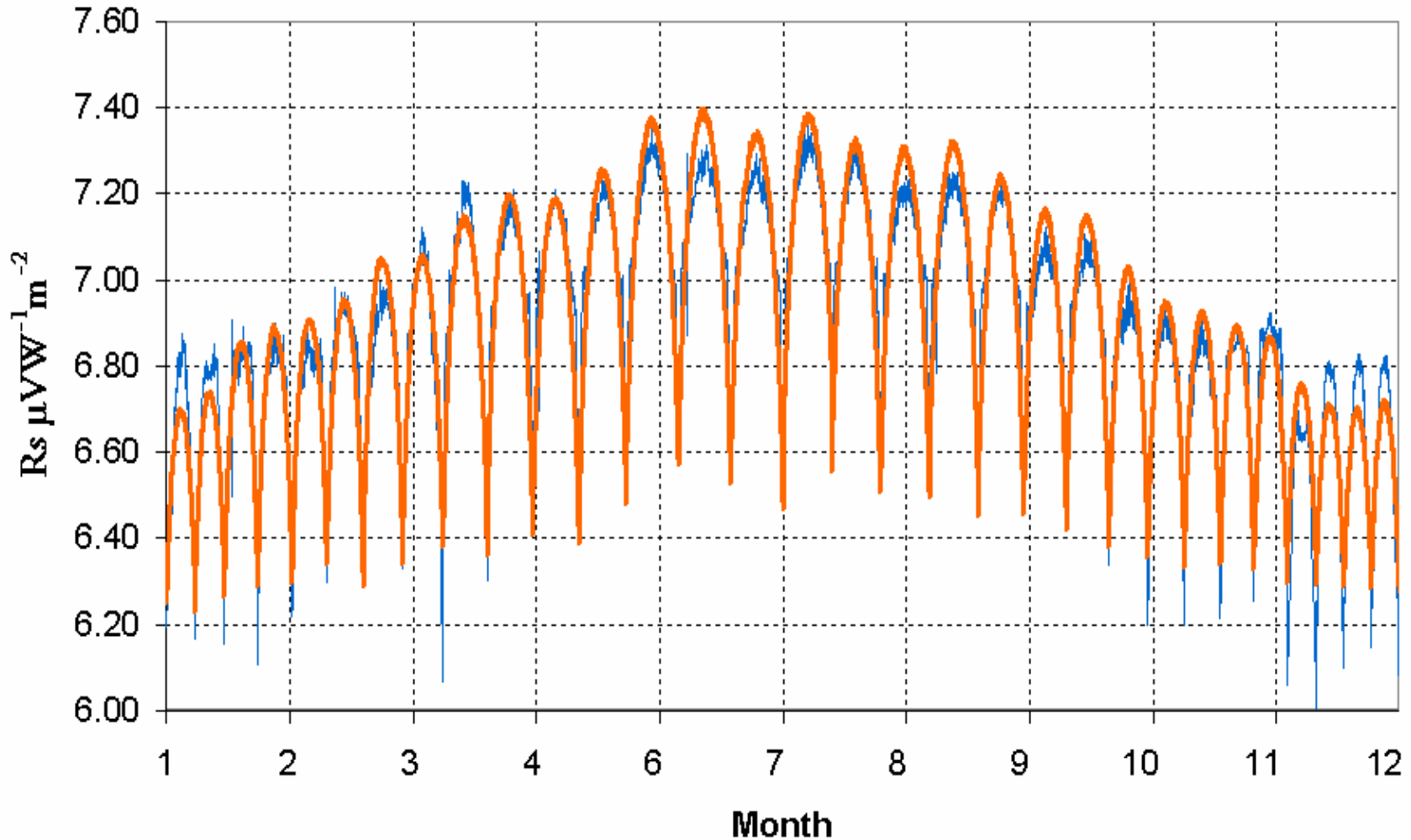
## Pyranometer $R_s$ through the Year



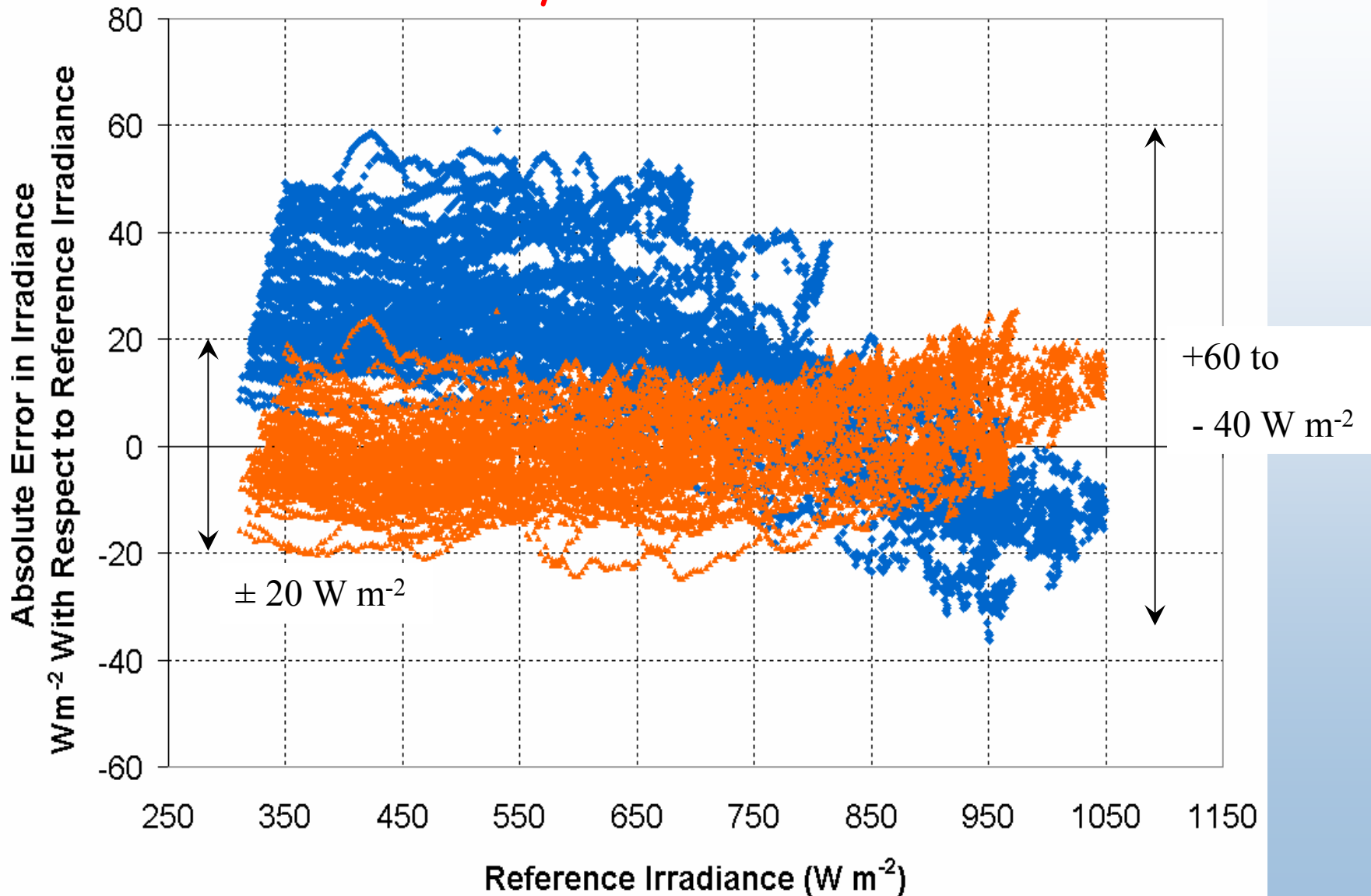
# "Daily" Calibration & Characterization

Pyranometer  $R_s$   
through the Year

$$R_{S_{u8-48}}(Z, D, I) = 0.6392 \ln[\cos(Z)] - 0.0936 \cos[(D - 2\pi \cdot 360)/365] + 0.0008 I + 7.545$$

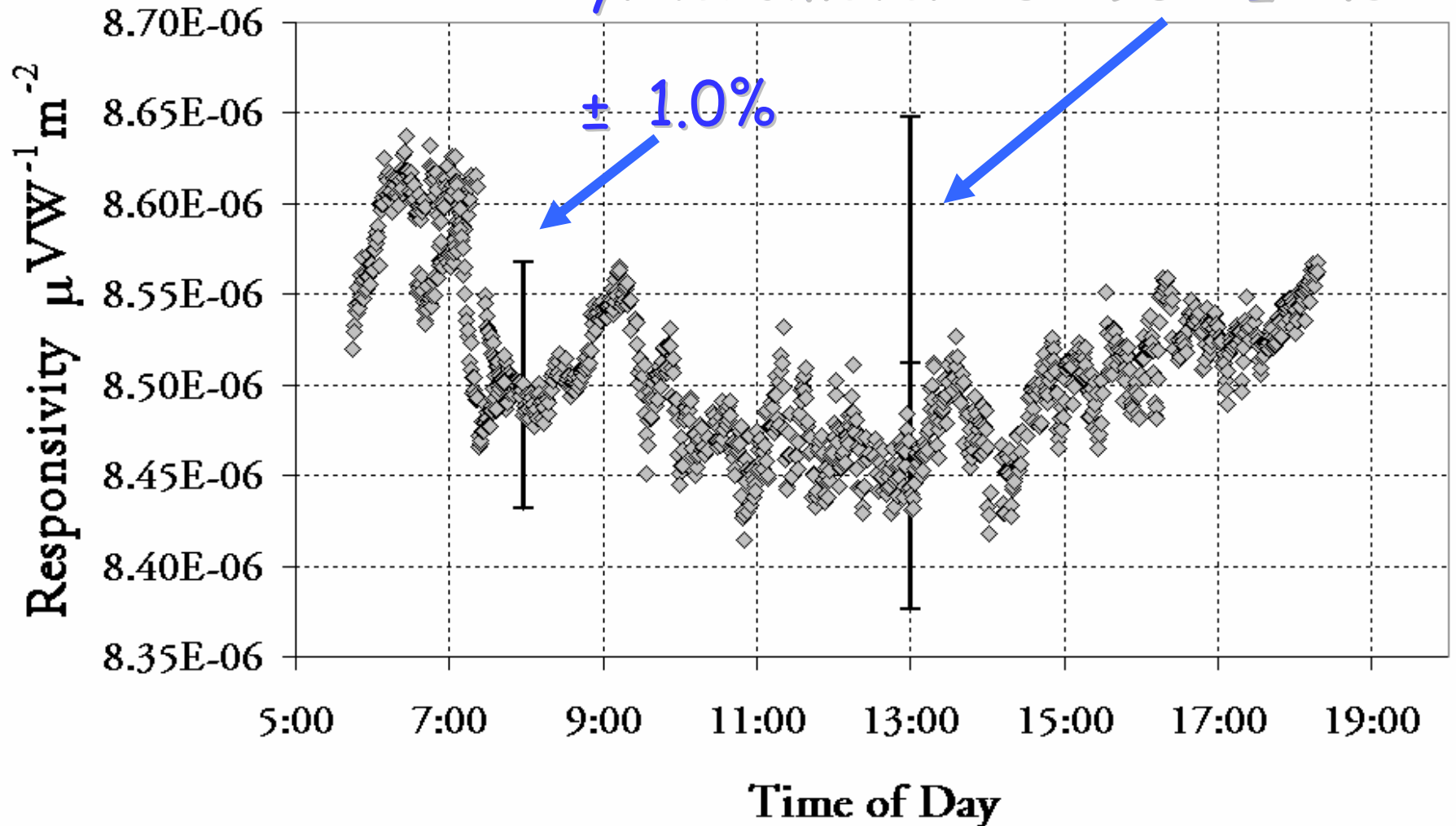


# RS<sub>8-48U</sub> (Z, DA, IR) Applied to Pyranometer mV One year of data



# Pyrheliometer Rs Calibration Results

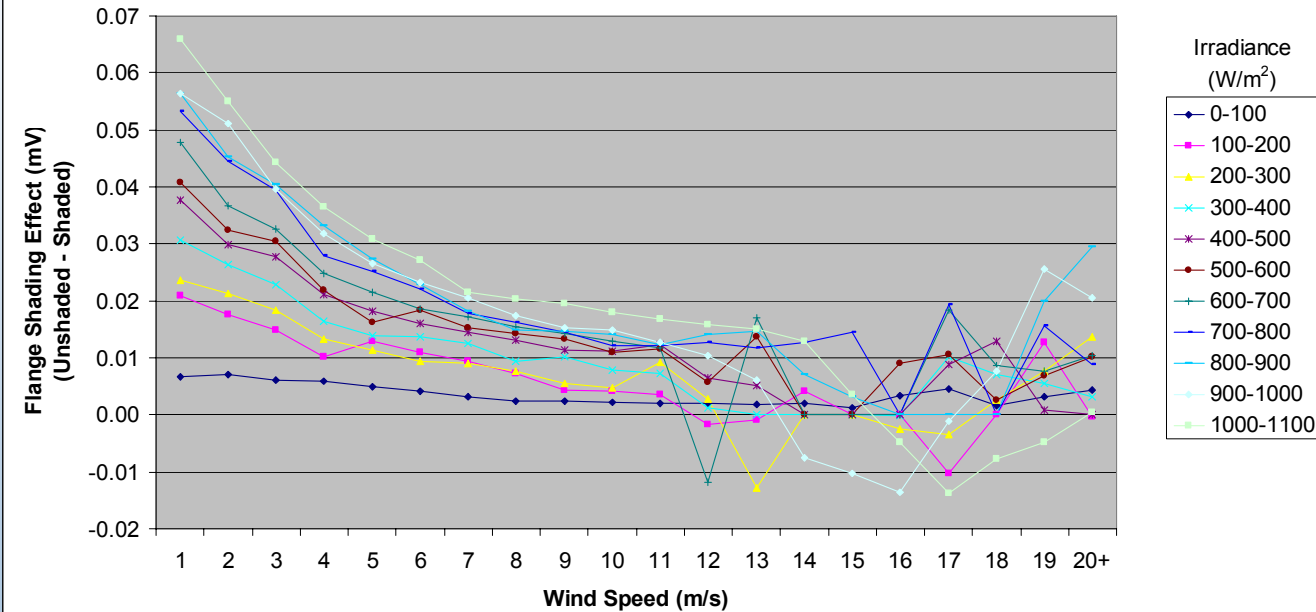
Pyrheliometer Rs U95 =  $\pm 1.8\%$



# Pyrheliometer Rs environmental influences (flange, window, instrument)



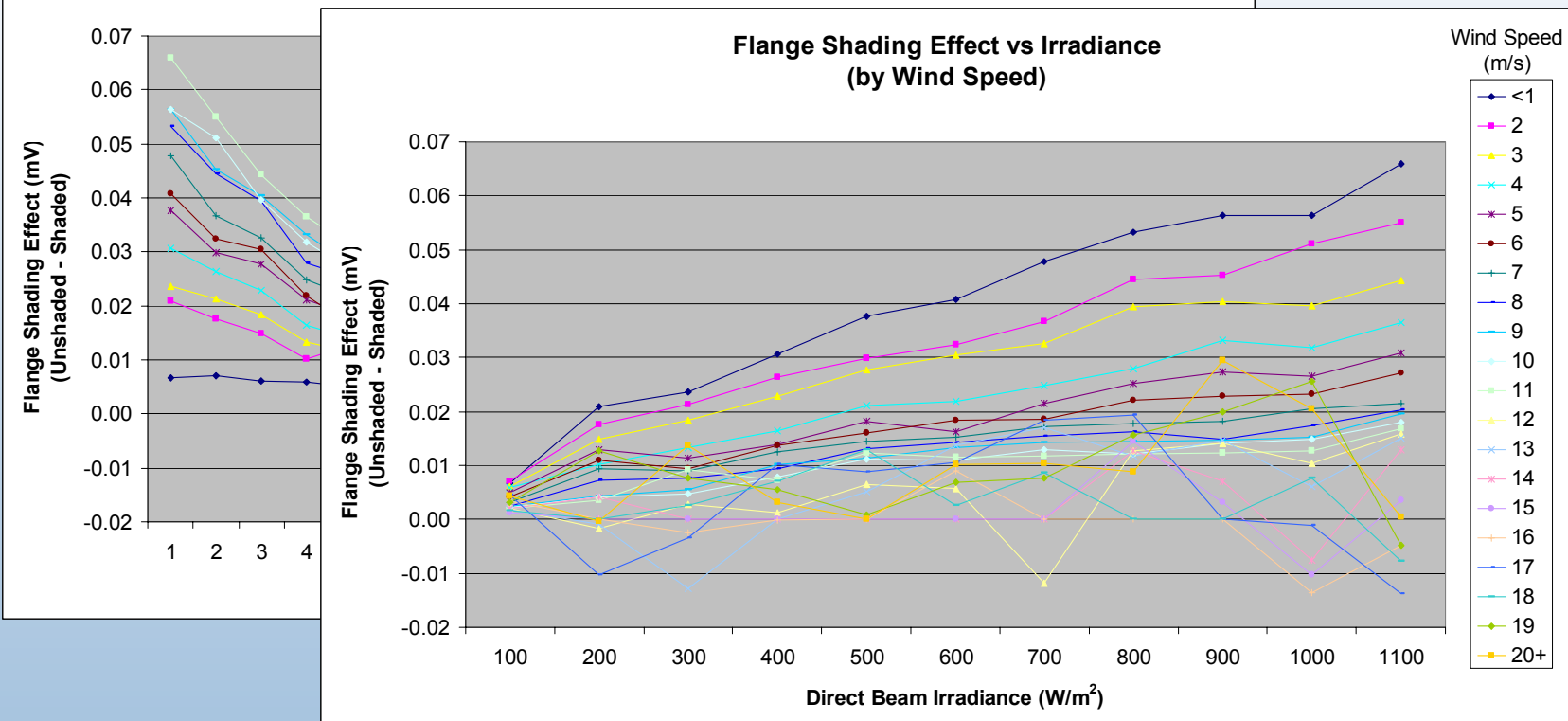
Flange Shading Effect vs Wind Speed  
(by Irradiance)



# Pyrheliometer Rs environmental influences (flange, window, instrument)



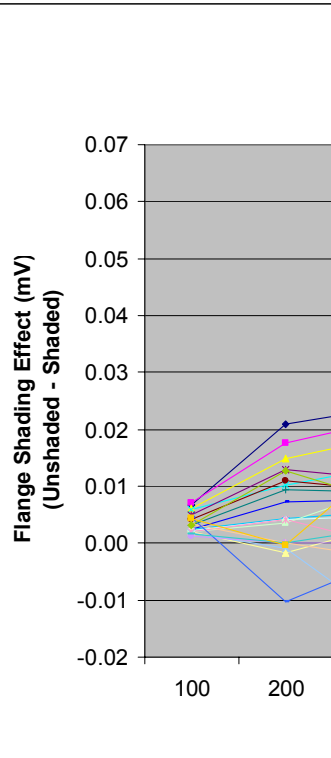
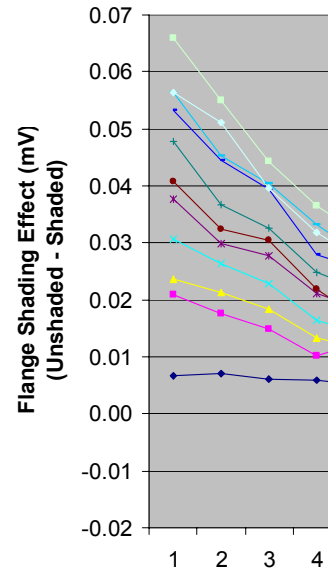
Flange Shading Effect vs Wind Speed  
(by Irradiance)



# Pyrheliometer Rs environmental influences (flange, window, instrument)

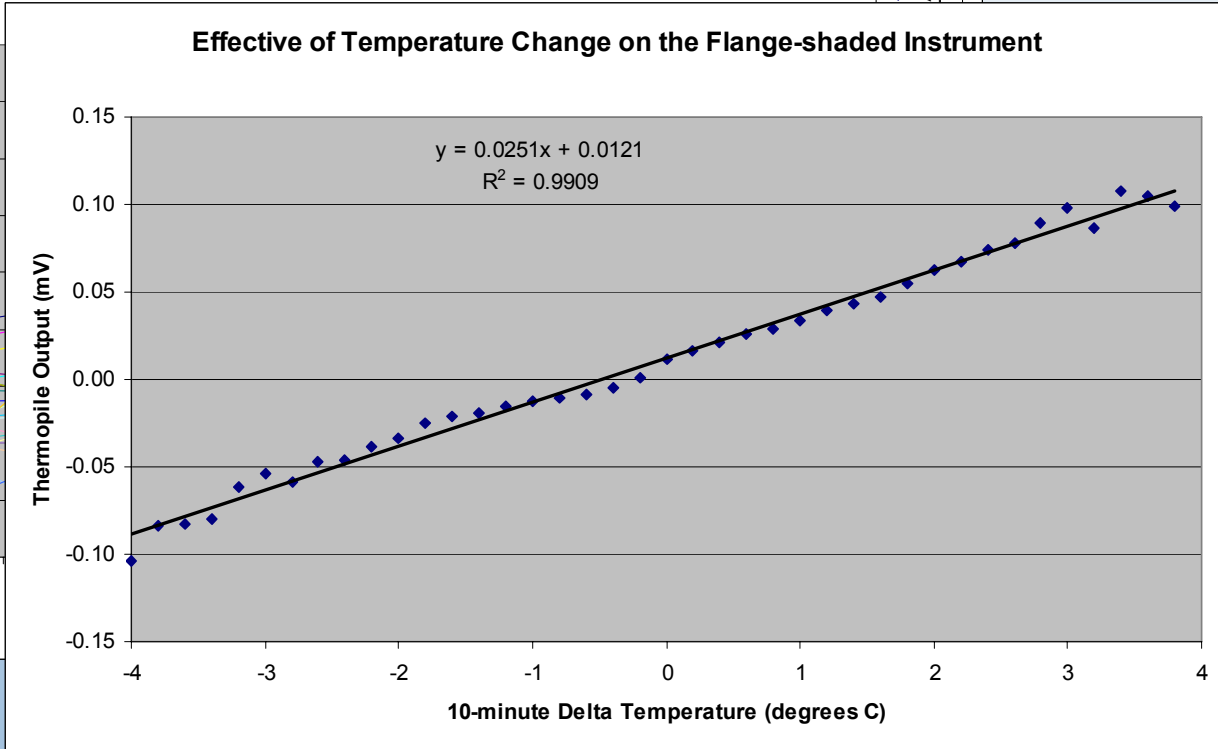


**Flange Shading Effect vs Wind Speed (by Irradiance)**



**Flange Shading Effect vs Irradiance (by Wind Speed)**

Wind Speed (m/s)  
◆ <1



# Instrument Calibration & Characterization

Shade window



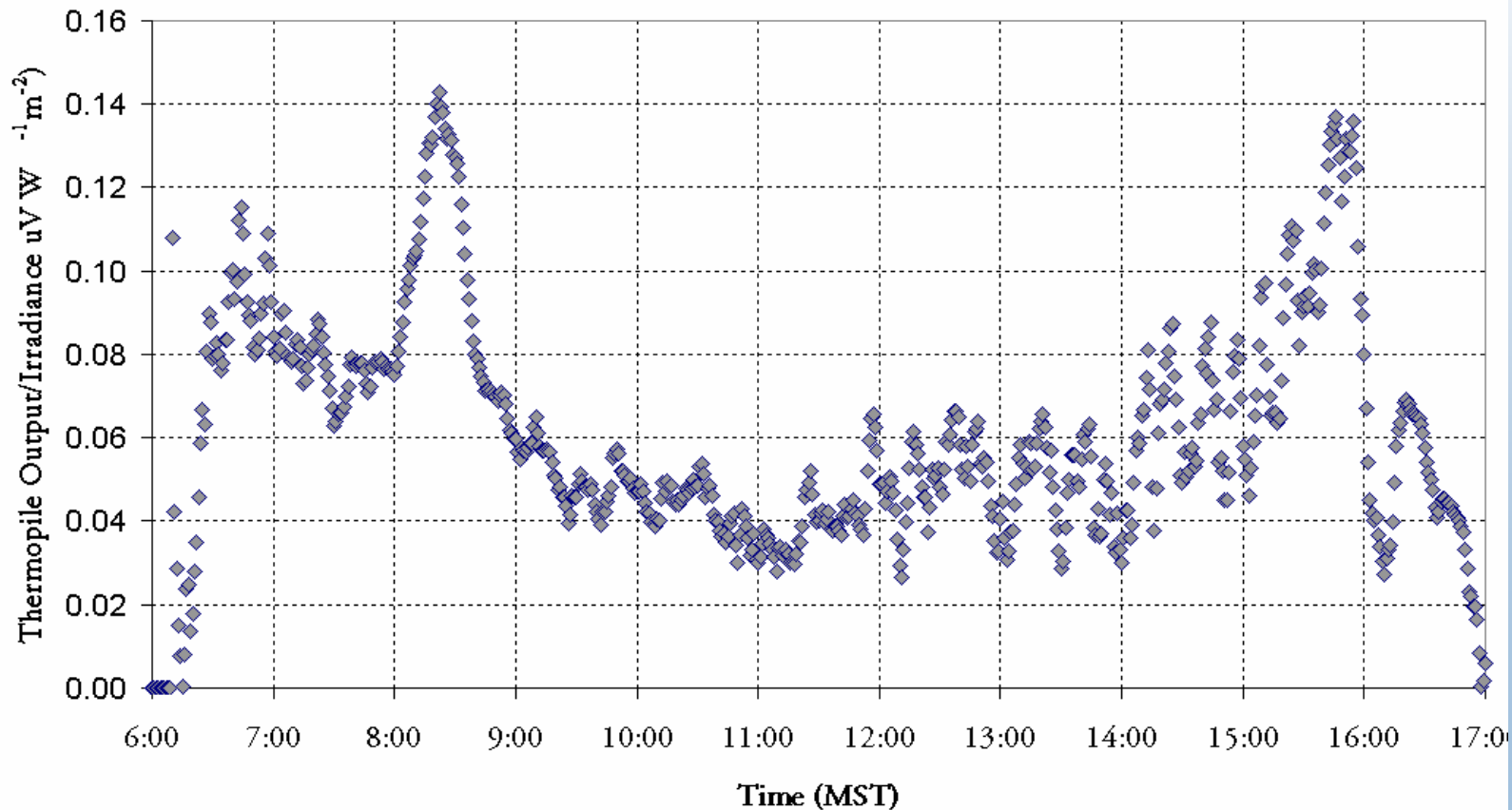
Minus

Shade Flange



DNI

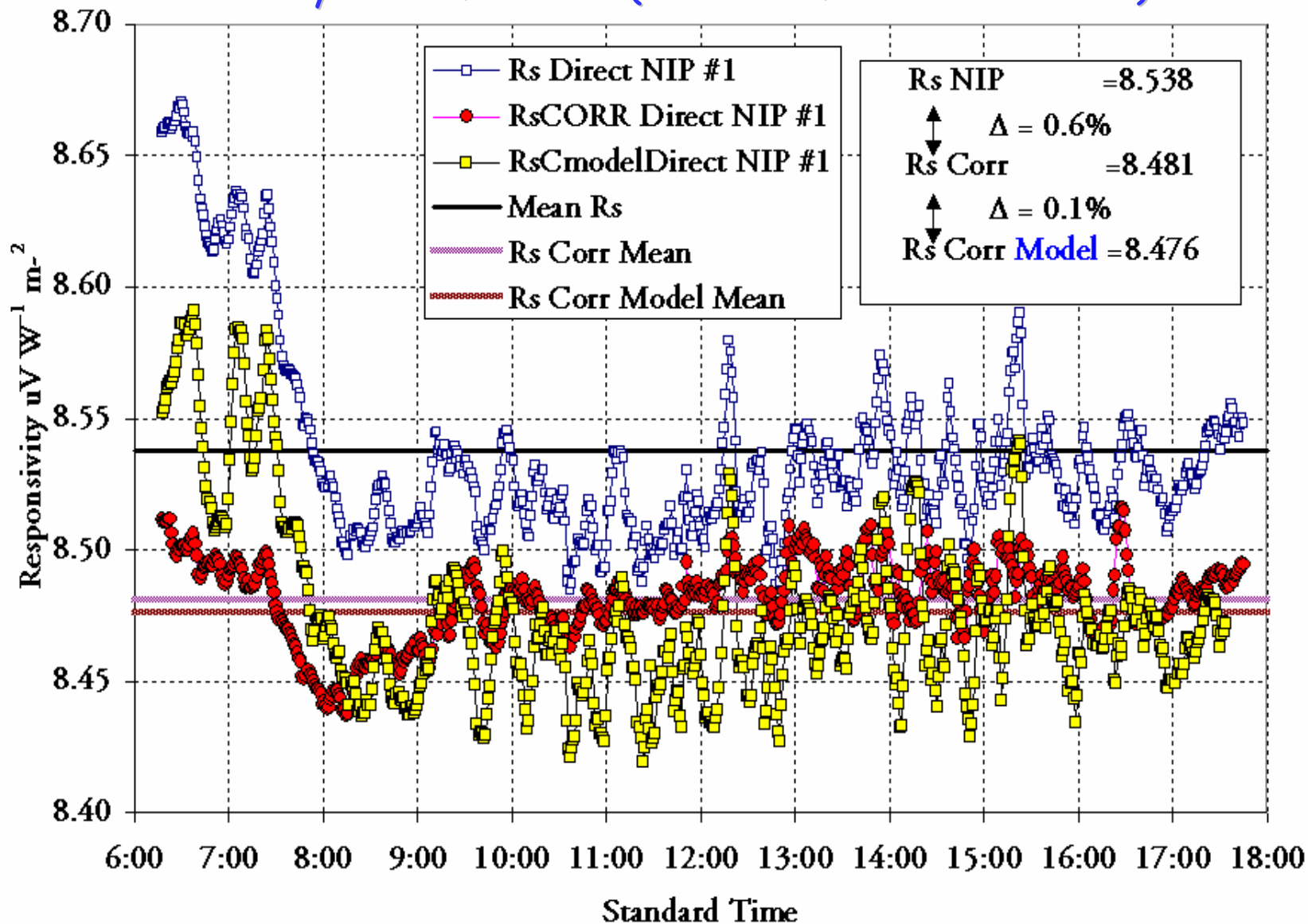
Pyrheliometer Responsivity variations from environmental influences (on flange, instrument)





# Pyrheliometer Calibration Corrections

## Pyrheliometer Rs(-Environment corrected-)



# Solar Radiometer Responsivity Issues

- Pyranometer Thermopile Offset IR voltage
  - Corrections at calibration time -
    - Monitor IR; reduce U95 by  $\frac{1}{2}$  to 1%  $\forall Z$
  - Post-hoc correction schemes based on
    - "cosine response" though DAY and Year
      - Global data uncertainty 60  $Wm^{-2}$   $\rightarrow$  20  $Wm^{-2}$
- Pyrhelimeter Environmental Influences
  - Correction at calibration time  $\propto Ws, dT/dt, DNI$ 
    - Reduce U95 offset in  $R_s \sim 1/3$ ; 0.6% to 0.2%
  - Post-hoc correction based on  $Ws, dT/dt, I$   
Research continues!