

Abstract – The quest to understand the fundamental nature of matter requires ever higher energy particle collisions, which in turn leads to ever larger and more expensive accelerator facilities. Advanced concepts for electron and positron acceleration are required to reduce the cost and increase the performance of next-generation accelerators. Plasma-based accelerators can sustain electron plasma waves with phase velocities close to the speed of light c and longitudinal electric fields on the order of the nonrelativistic wave breaking field, $E_0 = c m_e \omega_p / e$, where $\omega_p = (4\pi n_e e^2 / m_e)^{1/2}$ is the plasma frequency at an electron density n_e [1]. For $n_e \sim 10^{18} \text{ cm}^{-3}$, $E_0 \sim 100 \text{ GV/m}$. Massively parallel particle-in-cell (PIC) simulations are required to simulate both laser-driven (LWFA) [2] and beam-driven (PWFA) [3] concepts, in order to support on-going experiments and to explore new ideas. We summarize recent successes in the use of parallel PIC codes VORPAL [4], OSIRIS [5] and QuickPIC [6] to validate computations with experimental data, to benchmark codes with independent implementations and to benchmark reduced PIC algorithms. Code performance and representative algorithms are discussed in the context of past work and future challenges.

Code Validation & Benchmarking

• VORPAL simulations of LWFA are validated against experimental data from LBNL [7]:

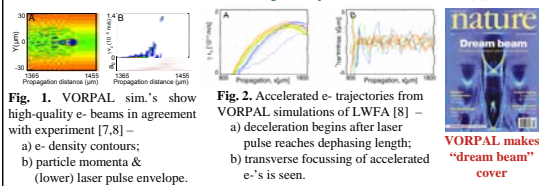


Fig. 1. VORPAL sim. show high-quality e- beams in agreement with experiment [7,8] – a) e- density contours; b) particle momenta & (lower) laser pulse envelope.

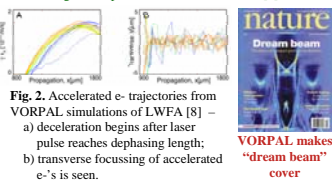


Fig. 2. Accelerated e- trajectories from VORPAL simulations of LWFA [8] – a) deceleration begins after laser pulse reaches dephasing length; b) transverse focussing of accelerated e-'s is seen.

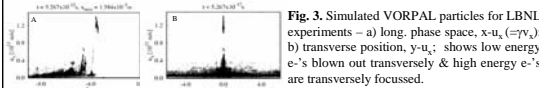


Fig. 3. Simulated VORPAL particles for LBNL experiments – a) long phase space, $x-u$, ($-p_x$); b) transverse position, $y-u$; shows low energy e-'s blown out transversely & high energy e-'s are transversely focussed.

• 3D OSIRIS simulations of LWFA are validated against experimental data from RAL [9]:

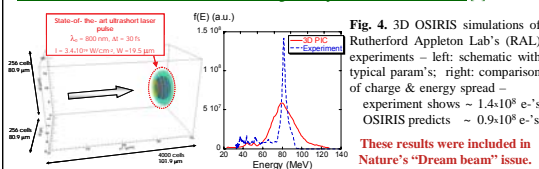


Fig. 4. 3D OSIRIS simulations of Rutherford Appleton Lab's (RAL) experiments – left: schematic with typical param; right: comparison of charge & energy spread – experiment shows $\sim 1.4 \times 10^8$ e-'s OSIRIS predicts $\sim 0.9 \times 10^8$ e-'s

• Simulations of tunneling ionization validated against experimental data from LBNL [10]:



Fig. 5. a) spectrum of 800 nm laser pulse (LBNL data), blue-shifted after tunnel-ionizing He gas over many Rayleigh lengths; b) FFT spectrum from 2D PIC simulation, using ADK model; c) spatially-resolved FFT from simulation shows most blue-shifting occurs at front of pulse.

• PWFA simulations show tunneling ionization can self-generate the necessary plasma [11]:

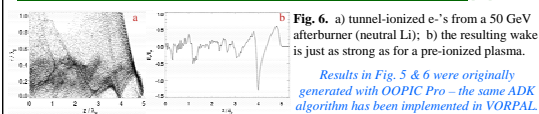


Fig. 6. a) tunnel-ionized e-'s from a 50 GeV afterburner (neutral Li); b) the resulting wake is just as strong as for a pre-ionized plasma.

• 3D QuickPIC w/ tunneling ionization is validated against E-164x PWFA experiments [12,13]:

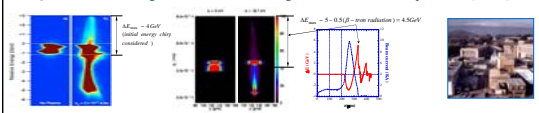
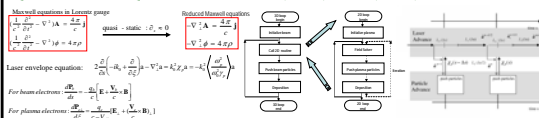


Fig. 7. QuickPIC can accurately model these E-164x experiments at SLAC in 3D, with tunneling ionization physics, and 100x speed-up over explicit PIC.

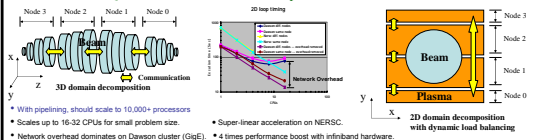
Note: the ADK algorithm first implemented in OOPIC Pro (and now VORPAL) was benchmarked against the alternate implementation in OSIRIS; QuickPIC was then benchmarked against OSIRIS.

Parallel Algorithms & Reduced Models

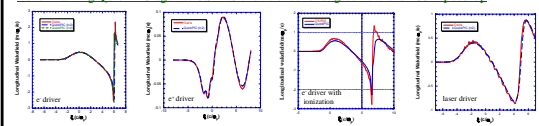
• Quasi-static model in QuickPIC works with beam (PWFA) & laser (LWFA) drivers:



• Parallelization of QuickPIC will enable use of $\sim 10^4$ processors in the future:



• Benchmarking QuickPIC against OSIRIS shows excellent agreement and 100x speed-up:



• Ponderomotive guiding center (i.e. envelope) model [14] of laser pulse is a powerful technique:

$$(\Delta + \partial_z)(A + \bar{A}) = -J - j + \nabla \phi, \quad (\Delta_\perp + i\omega \partial_t + \frac{\partial}{\partial z} \partial_z) a = \chi a \quad (\partial_z + \partial_t) a = \frac{c}{2i\omega} (1 + \frac{\partial}{\partial z} \partial_z) (\chi + \Delta_\perp) a$$

$$j = -\bar{a} \sum \frac{q_i \rho_i}{m_i}, \quad \bar{a} = \frac{1}{2} a(x, y, t) e^{i\omega t + cc} \quad \xi = x - ct, \tau = t \quad \chi = -ik \sum \frac{q_i \rho_i}{m_i}$$

• VORPAL's envelope model [15] agrees with PIC for $a_0 \sim 1$, showing $\sim 200x$ speed-up:

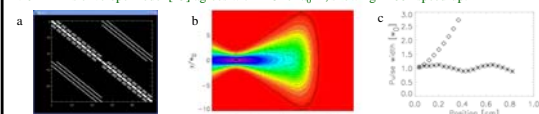


Fig. 8. a) sparsity pattern of the discretized envelope evolution equation; b) contours of laser pulse envelope in vacuum; c) envelope width in vacuum (diamonds) and in a plasma channel (stars). VORPAL can now simulate LBNL's 3 cm capillary discharge channel in a few processor-hours.

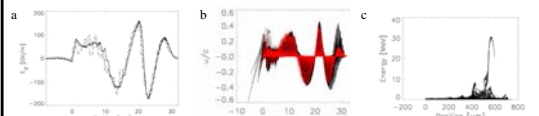


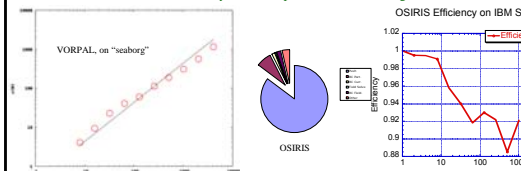
Fig. 9. VORPAL results: a) accelerating wave fields for PIC (solid) & envelope; b) normalized velocities for envelope (red) & PIC; c) background e-'s can be trapped for $a_0 = 2.5$

• Perfectly Matched Layer (PML) absorbing BC's [16] have been implemented in VORPAL:

- 1) these efficiently absorb a wide range of electromagnetic wavelengths, regardless of propagation angle
- 2) a 10-20 cell buffer region is required around the simulation domain (mesh be thicker than the longest relevant wavelengths)
- 3) a modest number of PIC macro-particles entering the PML region does not appear to cause problems
- 4) in some cases, significant speed-up for LWFA simulations is obtained by greatly reducing the transverse size of the domain
- 5) for long LWFA simulations with a moving window, reflected waves are a concern. PML's directly address this issue

Code Performance & Future Concepts

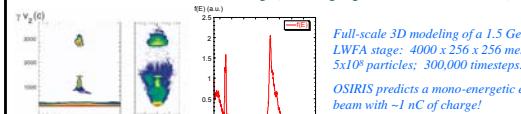
• VORPAL and OSIRIS scale efficiently to $>1,000$ processors on "seaborg" IBM SP at NERSC:



VORPAL and OSIRIS are expected to scale to $>10^4$ processors for sufficiently large problem sizes (e.g. longer Ti:Sapphire laser pulses for lower-density LWFA).

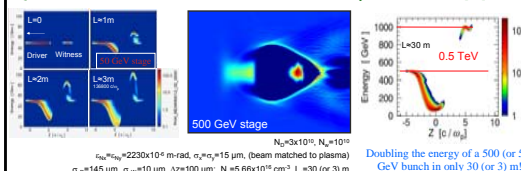
QuickPIC will require some code development to reach this scale of parallel computing – use of pipelining for long, 3D drivers:

• 3D OSIRIS simulation of 1.5 GeV LWFA stage (interesting "light source" candidate as well):

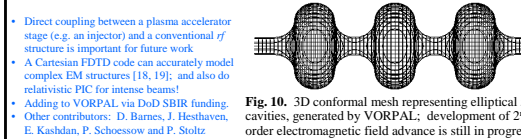


Full-scale 3D modeling of a 1.5 GeV LWFA stage: $4000 \times 256 \times 256$ mesh; 5×10^6 particles; 300,000 timesteps. OSIRIS predicts a mono-energetic e-beam with ~ 1 nC of charge!

• QuickPIC can simulate a TeV PWFA afterburner in 3D with only 5,000 node hours [17]:



• VORPAL will have 2nd-order accurate conformal BC's (work in progress):



• Some agenda items for future work include:

- Simulation tools must continue moving from qualitative physics to quantitative prediction.
- In particular, we must simulate in detail the e-beam emerging from the plasma – including emittance, energy spread, bunch length & total charge
- Investigate mesh refinement, fluid & Vlasov models, high-order field & particle algorithms
- Push our codes to $\sim 10^4$ processors; continue improving serial performance
- Model upcoming 1 GeV LWFA (LBNL, RAL, LOA) and 10 GeV PWFA (SLAC) exp's
- Push our reduced PIC models to the 100 GeV - 1 TeV range; get scaling laws correct
- Model e-cloud physics (LHC, ILC damping ring); add circular/elliptical pipes to QuickPIC

[1] E. Esarey et al., IEEE Trans. on Plasma Sci. 24, 252 (1996).
 [2] T. Tajima & J. Dawson, Phys. Rev. Lett. 43, 267 (1979).
 [3] C. Joshi et al., Nature 311, 525 (1984).
 [4] C. Nieter & J. Cary, J. Comp. Phys. 196, 448 (2004).
 [5] R. Hemker, Ph.D. thesis, UCLA (2000).
 [6] R. Fonseca et al., Lect. Notes Comput. Sci. 2331, 342 (2002).
 [7] C. Huang et al., J. Comp. Phys. (submitted).
 [8] C.G.R. Geddes et al., Nature 431, 538 (2004).
 [9] C.G.R. Geddes et al., Proc. Part. Accel. Conf. (2005), in press.
 [10] S.P.D. Mangels et al., Nature 431, 535 (2004).
 [11] D. Dimitrov et al., Proc. Advanced Accel. Workshop, AIP 647, 192 (2002).
 [12] D. Bruhwiler et al., Phys. Plasmas 10, 2022 (2003). (Invited)
 [13] M. Zhou et al., Proc. Part. Accel. Conf. (2005), in press.
 [14] M. Hogan et al., Phys. Rev. Lett. (2005), accepted.
 [15] D. Gordon et al., IEEE Trans. Plasma Science 28, 1135 (2000).
 [16] P. Messmer et al., Proc. Part. Accel. Conf. (2005), in press.
 [17] S. Gedney, IEEE Trans. Anten. & Prop. 44, 1630 (1996).
 [18] C. Huang et al., Proc. Part. Accel. Conf. (2005), in press.
 [19] S. Dey & R. Mitra, IEEE Micro, Guided Wave Lett. 7, 273 (1997).
 [19] I. Zagorodnov et al., Int. J. Num. Model. 16, 127 (2003).

