Pressure of the standard model

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Strong and Electroweak Matter '06 May 11, 2006

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JHEP 0601:060,2006, hep-ph/0510375 & JHEP 0603:011,2006, hep-ph/0512177

Why do we want to know the pressure?

- To understand the hydrodynamics of heavy-ion collisions (only the QCD degrees of freedom are thermalized)
- Cosmology: the relic densities of particles left over from the Big Bang (WIMPs)
 - $-\sim 10\%$ deviation from the ideal gas pressure leads to $\sim 1\%$ change in the relic density of WIMPs (Hindmarsh & Philipsen, Phys. Rev. D71, 087302 (2005))
- Theoretically important

$$\mathcal{Z} = \int \mathcal{D}\varphi \, \mathrm{e}^{-S_E[\varphi]}, \quad p = \frac{T}{V} \ln \mathcal{Z}.$$

QCD pressure



Kajantie *et al.*, Phys. Rev. D67, 105008 (2003).

- Pressure of QCD (with massless quarks) is studied extensively, expansion stretched to the last perturbatively computable term
- The expansion does not converge well unless the temperature is asymptotically high.

The electroweak sector

- Studies of electroweak thermodynamics centered on the electroweak phase transition (electroweak baryogenesis).
- Differences to the QCD computation
 - Presence of fundamental scalars driving the phase transition (crossover)
 - Another mass scale at tree level (Higgs mass), $p(T)\sim T^4+\nu^2T^2+\nu^4$
 - Many independent couplings, formal powercounting: $g^2 \sim g'^2 \sim \lambda \ (\sim g_{Y, {
 m top}}^2 \sim g_s^2)$, $\nu^2 \lesssim g^2 T^2$
 - Weakly coupled theory, expect the expansion to converge better.
 - Weakly coupled even close to the crossover transition!

Calculation of the pressure

- Straightforward perturbative expansions in the coupling constants inhibited by infrared divergences, $p(T)/T^4 \sim 1 + g^2 + g^4(1/\epsilon + 1) + \dots$ $(d = 4 2\epsilon)$
- Can be cured by resummation
- Convenient framework: effective field theories
 - separate different mass scales into successive effective theories by integrating out the large scales one by one.

$$\mathcal{L}_{\rm E} = \frac{1}{4} F^a_{ij} F^a_{ij} + \frac{1}{2} (D_i A^a_0)^2 + \frac{1}{2} m_D^2 A^a_0 A^a_0 + |D_i \Phi|^2 + m_3^2 \Phi^{\dagger} \Phi + \mathcal{L}_{\rm E,int} (A^a_0, \Phi) m_D^2 \sim g^2 T^2, \quad m_3^2 \sim -\nu^2 + g^2 T^2, \quad g_3^2 = g^2 T \quad \text{etc.}$$

- Recipe for calculating the pressure:
 - Compute the strict perturbative expansion of pressure in the full 4-dimensional theory, schematically

$$p_{\rm E}(T) \sim T^4 \left[1 + g^2 + g^4 \left(\frac{1}{\epsilon} + 1 + \ln \frac{\Lambda}{T} \right) + \mathcal{O}(g^6) \right].$$

 Construct the 3-dimensional effective theory and compute its pressure, schematically

$$\frac{p_{\rm M}(T)}{T} \sim m^3 + g_3^2 m^2 \left(\frac{1}{\epsilon} + 1 + \ln\frac{\Lambda}{m}\right) + g_3^4 m (1 + \ln\frac{\Lambda}{m}) + \mathcal{O}(g_3^6).$$

 \implies Pressure given by $p(T) = p_{\rm E}(T) + p_{\rm M}(T) + \mathcal{O}(g^6)$

$$\frac{p(T)}{T^4} \sim 1 + g^2 + g^3 + g^4(1 + \ln g) + g^5(1 + \ln g) + \mathcal{O}(g^6).$$

Numerical results

SU(2) + Higgs theory:



No large deviation from the ideal gas result, perturbative expansion appears to be converging well.

Scale dependence:



Only weak scale dependence, no significant reduction in the scale dependence as the expansion is stretched from $\sim g^2$ to $\sim g^5$.

Full standard model:



QCD contributions dominate (# of QCD degrees of freedom = 79 of the total of 106.75)

Full standard model – corrections due to massless QCD:



Large QCD contributions remain due to heavy top!

Number of bosonic degrees of freedom:



EoS variable $w(T) = p(T)/\epsilon(T)$ and the speed of sound $c_s^2(T) = p'(T)/\epsilon'(T)$:



What happens close to the electroweak crossover?



• Theory remains weakly coupled at the transition, can hope to use similar methods to calculate the pressure near the transition.

• \implies construct another effective theory by integrating out the adjoint scalars A_0^a

$$\mathcal{L}_{E2} = \frac{1}{4} F^a_{ij} F^a_{ij} + \left| D_i \Phi \right|^2 + \widetilde{m}_3^2 \Phi^{\dagger} \Phi + \widetilde{\lambda}_3 \left(\Phi^{\dagger} \Phi \right)^2,$$

$$\widetilde{m}_3^2 \sim \underbrace{m_3^2}_{\sim g^3 T^2} - g_3^2 m_D \sim g^3 T^2.$$

Then the contribution from the effective theories divides into two parts:

$$p_{M}(T) = p_{M1}(T) + p_{M2}(T),$$

$$\frac{p_{M1}(T)}{T} \sim m_{D}^{3} + g_{3}^{2}m_{D}^{2}(1/\epsilon + 1 + \ln\frac{\Lambda}{m_{D}})$$

$$+ g_{3}^{4}m_{D}(1/\epsilon + 1 + \ln\frac{\Lambda}{m_{D}}) + \dots,$$

$$\frac{p_{M2}(T)}{T} \sim \tilde{m}_{3}^{3} + \tilde{g}_{3}^{2}\tilde{m}_{3}^{2}(1/\epsilon + 1 + \ln\frac{\Lambda}{\tilde{m}_{3}}) + \dots$$



• Singular behavior cured.

• A new class of diagrams resummed to the expansion.

Conclusions

- Pressure of the full standard model is calculated to three loop order.
- Deviation from the ideal gas is large, the expansion of the electroweak sector converges better.
- New type of terms in the expansion:
 - Another mass scale leading to qualitatively different type of terms in the expansion, $\nu^2 T^2, \ \nu^4.$
 - Terms of order $g^5 \ln g$ from the fundamental scalars.
- Potential concequences to WIMP relic density.