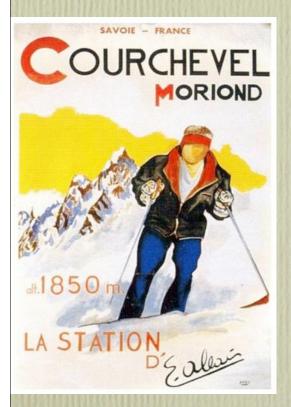
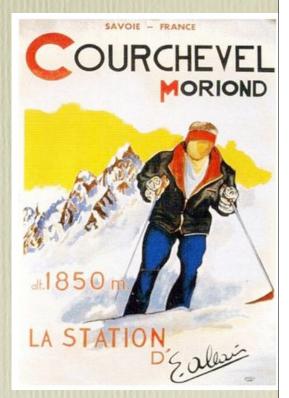
theory summary



Joseph Lykken Fermilab



XLIst Rencontres de Moriond, La Thuile, 11-18 March 2006

innovations of the Rencontres de Moriond:

- first (only?) major international meetings that fully integrate junior physicists
- first major international meetings where theorists and experimentalists actually communicate with each other

the most significant achievement of the Rencontres de Moriond:

"it forces the theorists to change their predictions *twice* per year"

-M. Danilov

thanks and congratulations to all the organizers for yet another successful meeting

and special thanks and congatulations to Jean Tran Thanh Van for 40 years of Rencontres de Moriond!

outline

- are theorists necessary?
- a game of small discrepancies
- SciFi Channel framework for BSM
- the big picture
- neutrino origins
- cosmology 2006
- the future

are theorists necessary?



theorists engage in two types of activity:



- playing around with new/old/stolen ideas for going beyond the standard paradigm (easy, fun, richly rewarded, but potentially useless)
- calculating things within the standard paradigm (useful, but difficult, tedious, and poorly rewarded)

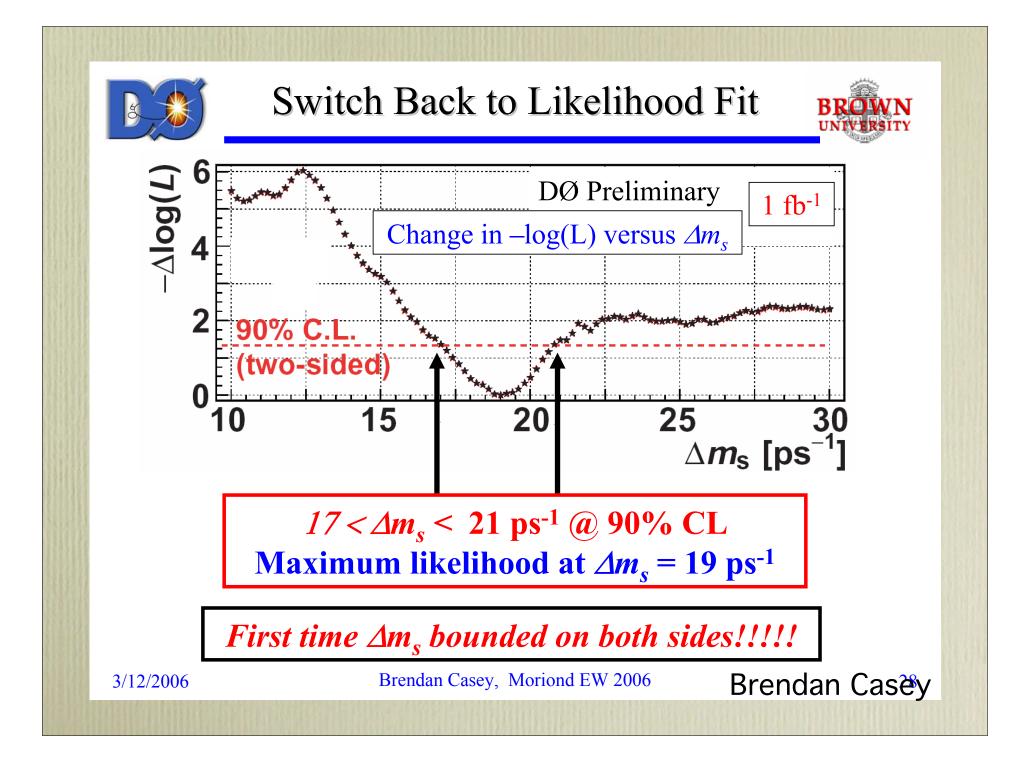
the importance of Standard Model calculations

- the SM still rules (almost) all
- below the energy frontier, new physics means (mostly) rare processes, small discrepancies, small inconsistencies
- at the energy frontier, SM backgrounds are about to get 100-500 times worse (Steve Mrenna)

case in point: B physics

- lots and lots and lots of data
- need precise SM predictions for dozens of observables
- the opportunities for big obvious signals of new physics are dwindling...

Deviations still possible in: Ø fact 2-3 Kt-DATVV few orders of magn. K° -> tT° VV Bs - Bs 30-40% lorder of wagn. BS-DXS UV Bd, s-Dl+lfew order of magn. fact 2 B-DEV My provic B-D \$Ks Francesca Borzumati B => troks N'K By-By WITHOUT touching moderately touching B-DXsl+e-, B-DXsl+e-, B-DXsr



case in point: B physics

- the opportunities for big obvious signals of new physics are dwindling...
- ...so now the game is looking for small discrepancies and small inconsistencies

how do theorists compute B decays?

- combination of electroweak, perturbative QCD, and nonperturbative QCD, further complicated by multiple scales!
- computation of exclusive decays reduces to hadronic form factors, which can be computed from unquenched lattice QCD
- computation of inclusive decays is done using effective Hamiltonians and the Wilsonian operator product expansion

Concezio Bozzi

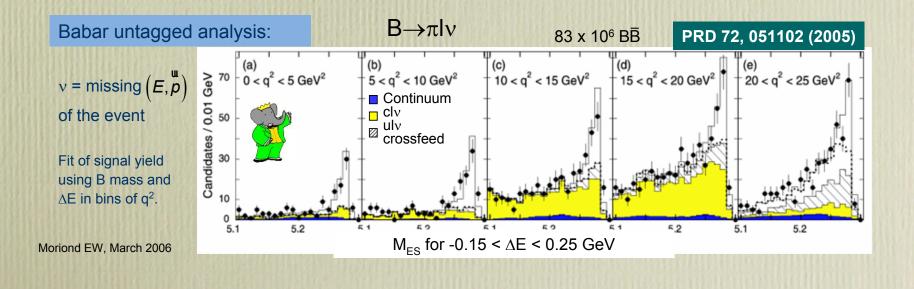
Exclusive |V_{ub}|

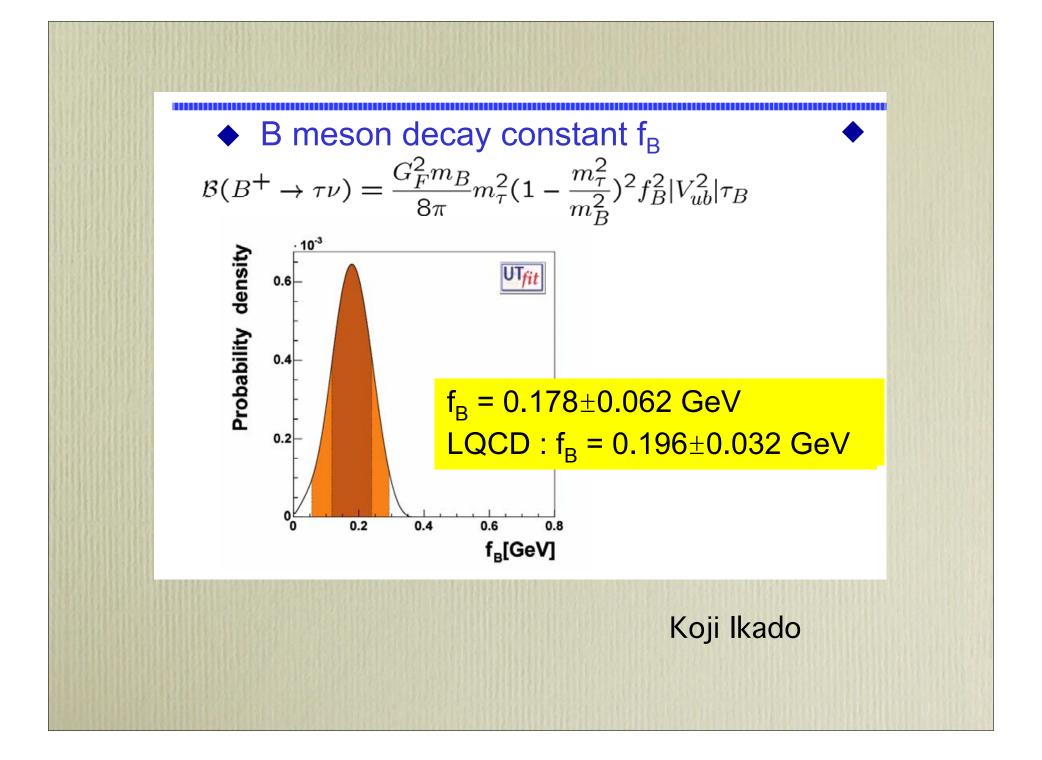
$$\frac{d\Gamma(B \to \pi 1 \nu)}{dq^2} = \frac{G_F^2}{24\pi^3} |V_{ub}|^2 p_\pi^3 |f_+(q^2)|^2$$

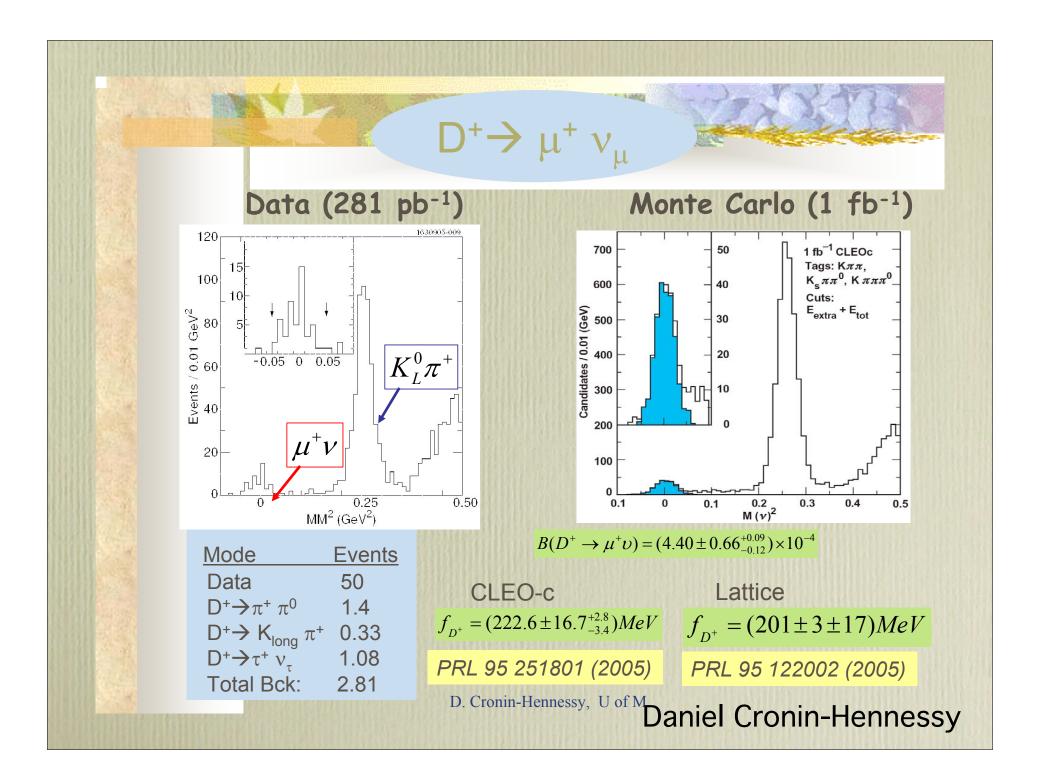
- Theoretically:
 - Uncertainties complementary to inclusive approach
 - FF normalization dominates the error on |V_{ub}| (~10% for all models)

Calculations of $f_+(q^2)$:

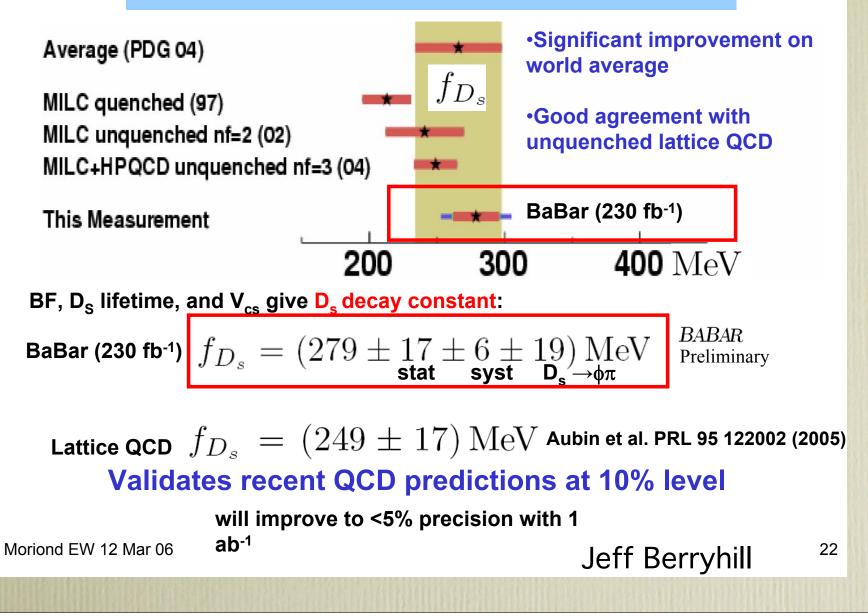
- LQCD (q²>16)
- Light-cone sum rules (q²<14)
- Quark models (ISGW2)
- Experimentally:
 - Good S/B ratio, untagged analyses
 - Small branching fractions
 - Measure q² dependence, compare to theory







$\textbf{D}_{S} \rightarrow \! \mu \nu \text{ Decay Constant } \textbf{f}_{\text{Ds}}$



Summary of the main theoretical limitations.			P. Gambino hep-ph/0510085	
process	quantity	Th error	needs	goal
$B \to D^* l \nu$	$ V_{cb} $	$\sim 4\%$	unquenching, analytic work	1%
$B \to X_c l\nu$	$ V_{cb} $	$\sim 1.5\%$	new pert calculations	<1%
$B o \pi(ho) l u$	$ V_{ub} $	10-15%	2-loop lattice matching etc.	6%
$B \to X_u l\nu$	$ V_{ub} $	$\sim 6-7\%$	more data/synergy with th	< 5%
$B \to X_s \gamma$	BR	$\sim \! 10\%$	NNLO	$<\!5\%$
$B \to \rho^0 \gamma / B \to K^* \gamma$	$\left V_{td}/V_{ts}\right $	10-20%	lattice $SU(3)$ breaking etc	?

- Note: in the K sector, unquenched lattice is already at percent level of accuracy
- e.g. 1% level MILC computation of f_K/f_{π} is input for the KLOE test of CKM unitarity (Matteo Palutan)

inclusive B decays

effective Hamiltonian approach describes inclusive non-leptonic B decays:

$$\langle f|H_{\text{eff}}|i\rangle = \frac{G_{\text{F}}}{\sqrt{2}}\lambda_{\text{CKM}}\sum_{k}C_{k}(\mu)\langle f|Q_{k}(\mu)|i\rangle$$

the Wilson coefs $C_{\bf k}(\mu,\alpha_s)$ are just the scale-dependent couplings of the interactions induced by the operators Q_k. Higher order operators are suppressed by powers of $\Lambda_{\rm QCD}/m_b$

$b \rightarrow s\gamma$ inclusive: BF results						
	fully inclusive	semi-inclusive				
CLE0	BF(b → sy) = $321 \pm 43 \pm 27$ • E (y)>2.0 GeV PRL 87 (2001)					
	$BF(b \rightarrow sy) = 355 \pm 32 \begin{array}{c} +30 \\ -31 \end{array}$ • E(y)>1.8 GeV PRL 93 (2004) 0	61803 • E(y)>2.24 GeV PLB 511 (2001) 151 • 16 modes				
BABAR	$BF(b \rightarrow s\gamma) = 367 \pm 29 \pm 34$ • Lepton-tagged hep-ex/ • E(γ)>1.9 GeV; BF not extrapolated	(0507001 • E(y)>1.9 GeV PRD 72 (2005) 052004 errors are:				
New HFAG average of these measurements using a common shape function for the extrapolation to low photon energies and taking into account the correlated						
Calculatio	h b \rightarrow dy contamination: n of extrapolation factors by following the Electory bar above 225	BF(b \rightarrow sy) = 355 ± 24 ⁺⁹ ₋₁₀ ± 3				
	üller and H. Flächer, hep-ph/050725 10del prediction (NLO):	• $E(y) > 1.6 \text{ GeV}$ stat./syst. combined shape $b \rightarrow dy$ function contamination				
$BF(b \rightarrow sy) = 357 \pm 30$ (see http://www.slac.stanford.edu/xorg/hfag/rare for details)						

• E(y)>1.6 GeV Gambino & Misiak, NPB 611 (2001) 338 Buras, Czarnecki, Misiak, Urban, NPB 631 (2002) 219

⇒ Depressing agreement between theory and experiment!

$\mathbf{B} \rightarrow \mathbf{X_s} \gamma$ predictions

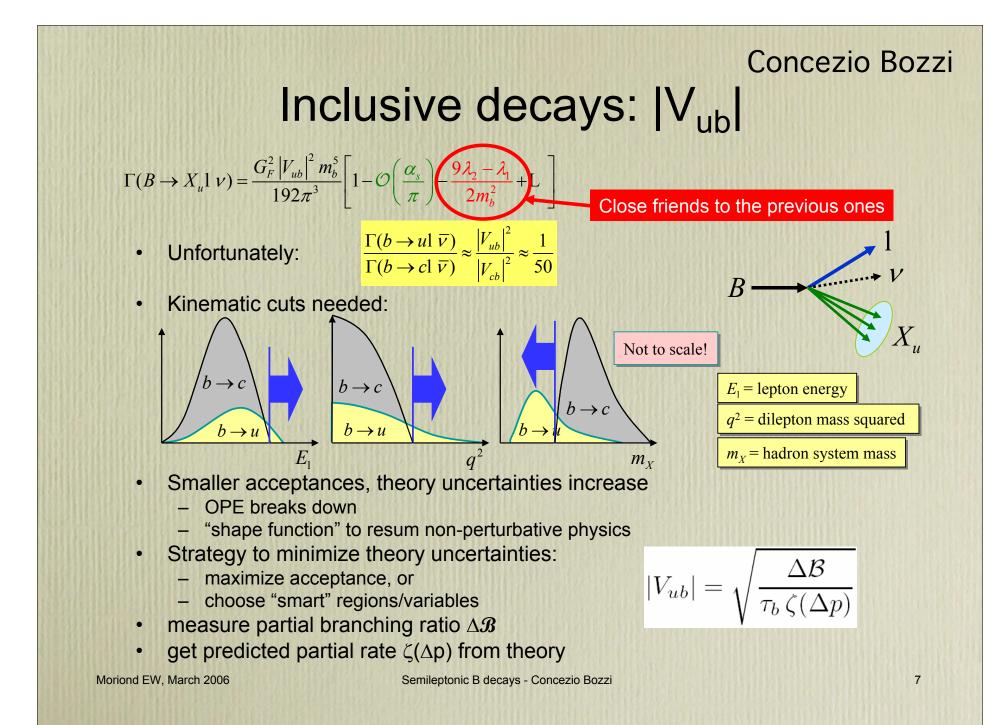
- impressive agreement between NLO theory and data
- but we need a NNLO calculation!
- NLO has too much renormalization scheme-dependence on the charm quark mass (Francesca Borzumati)
- + important effects from the scale

 $\Delta = m_{\mathbf{b}} - 2 E_{\mathbf{min}}^{\gamma} \simeq 1 \,\, \mathrm{GeV}$

T. Becher and M. Neubert hep-ph/0512208

$b \rightarrow s\ell^+\ell^-$ inclusive: sign of C_7

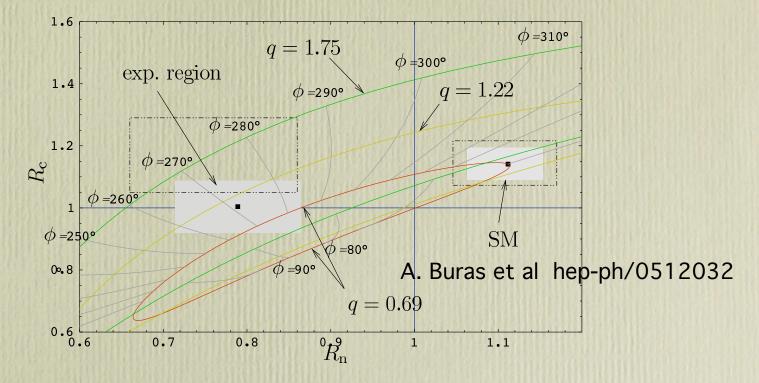
Gambino, Haisch, Misiak, PRL 94 (2005) 061803 $BF(b \rightarrow s\ell^+\ell^-)$ BABAR. weighted SM $C_7 \rightarrow -C_7$ q² range (89M BB) (152M BB) average 8.8 ± 1.0 $q^2 > (2m_u)^2$ 4.1 ± 1.1 5.6 ± 2.0 4.5 ± 1.0 4.4 ± 0.7 $1 < q^2 < 6 \text{ GeV}^2$ 1.57 ± 0.16 1.5 ± 0.6 1.8 ± 0.9 1.6 ± 0.5 3.30 ± 0.25 PRD 72 (2005) PRL 93 (2004) 081802 092005 3.0 2.5 = -1. SUGRA 10.2 2.0 SUGRA example with $C_7 \rightarrow -C_7$ Experiments clearly favor SM-sign for C₇! SM case 1.0 0.5 0.0 Thomas Schietinger 0.0 0.8 0.2 1.0 0.4 0.6 Goto et al., PRD 55 (1997) 4273



the $\mathbf{K} \pi$ puzzle

- difficult theory versus difficult experiment, with a possible inconsistency pointing to new physics
- step 1: isospin analysis of $\mathbf{B} \to \pi \pi$ data to extract hadronic parameters
- step 2: use SU(3) flavor, with known factorizable SU(3) breaking corrections, to apply this to $\mathbf{B} \to \mathbf{K} \pi$
- step 3: predict some ratios, check data:

theory vs data doesn't agree for ratios which are sensitive to EW penguins



this analysis is being improved, but so far the problem is still there (Julie Malcles)

(an analogy based on the SciFi Channel)

- The BSM models were created by man
- They evolved
- They rebelled
- There are many copies
- And they have a plan

- The BSM models were created by man
- They evolved
- They rebelled
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Moriond circa 1983

- BSM theory was supersymmetry, grand unification, technicolor
- the models were primitive
- there was also a small strange community of "neutrino" people
- and a small strange community of "particle-astro" people

- The BSM models were created by man
- They evolved
- They rebelled
- There are many copies
- And they have a plan

Moriond circa 2006

- string theory took the BSM high ground
- supersymmetry models are much more sophisticated, detailed, and ambitious
- supersymmetry has become a framework to describe everything from Higgs to B physics, from inflation to baryogenesis, from unification to LFV, from dark matter to HyperCP
- technicolor mutated into AdS/CFT branes (Francesco Sannino)

- The BSM models were created by man
- They evolved
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they rebelled

- after 30 years, SUSY is still not discovered, despite golden opportunities with LEP, Tevatron, B physics, EDMs, etc (Carlos Munoz)
- mysteries of flavor and of vacuum energy, which SUSY already had trouble with, have gotten worse
- theorists got worried (and bored) and decided to try radically new things...

extra dimensions

- extra dimensions are the other generic prediction of string theory (Mariano Quiros) and anyway are generic new degs of freedom
- they could be infinite but hidden, very large (.1 mm to 10 fm), large (Tev-1), or tiny but warped.
- they could: break SUSY (Yael Shadmi), explain dark matter (Thomas Flacke), explain fermion masses (Gregory Moreau), explain a light Higgs (Mariano Quiros).

Higgs Shmiggs

- theorists are even questioning some of the holy assumptions:
 - models with no Higgs (Sekhar Chivukula)
 - landscape-inspired SUSY (Adam Falkowski), including split-SUSY
- and combining ideas, e.g. Little Higgs and SUSY (Piotr Chankowski)

- The BSM models were created by man
- They evolved
- They rebelled
- There are many copies
- And they have a plan

there are many copies

- despite different theoretical inputs, many BSM models end up looking the same phenomenologically
- this is because they are trying to do the same things
- while simultaneously getting around the bounds from existing data (Guido Mirandella)

there are many copies

- most BSM models have a WIMP dark matter, and thus missing energy signatures at colliders
- the EW precision data imply that the new heavy particles associated with EWSB are:
 - multi-TeV
 - conspiratorial
 - pair-produced (->DM) and minimal flavorviolating

there are many copies

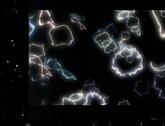
- so some new BSM models look like SUSY (Little Higgs with T-parity, UED,...)
- others resemble each other with new TeVish gauge bosons, top-partners, etc (Little Higgs, Randall-Sundrum, TeV extra dims, GUT-inspired,...)
- and it was already difficult to tell SUSY models apart (Martin White)

status of theories beyond the standard model

- The BSM models were created by man
- They evolved
- They rebelled
- There are many copies
- And they have a plan

replace the standard paradigm by ~2015

the big picture 2006



string unification

supersymmetry

Diol Cl

extra dimensions

- CLIPEE

neutrino origins?

flavor origins?

new TeV scale physics 100 GeV? 1 TeV? 10 TeV?

new long distance physics?

neutrino origins

 $|U_{\alpha i}| = \begin{pmatrix} |U_{e1}| & |U_{e2}| & |U_{e3}| \\ |U_{\mu 1}| & |U_{\mu 2}| & |U_{\mu 3}| \\ |U_{\tau 1}| & |U_{\tau 2}| & |U_{\tau 3}| \end{pmatrix} = \begin{pmatrix} 0.76 - 0.87 & 0.49 - 0.64 & 0.00 - 0.20 \\ 0.20 - 0.53 & 0.42 - 0.72 & 0.58 - 0.82 \\ 0.20 - 0.54 & 0.43 - 0.73 & 0.56 - 0.81 \end{pmatrix}$

Gonzalez-Garcia et al (2005)

Different pattern for leptons and quarks

 $|U_{\alpha i}|_{CKM} = \begin{pmatrix} 0.9739 - 0.9751 & 0.221 - 0.227 & 0.0029 - 0.0045 \\ 0.221 - 0.227 & 0.9730 - 0.9744 & 0.039 - 0.044 \\ 0.0048 - 0.014 & 0.037 - 0.043 & 0.9990 - 0.9992 \end{pmatrix}$

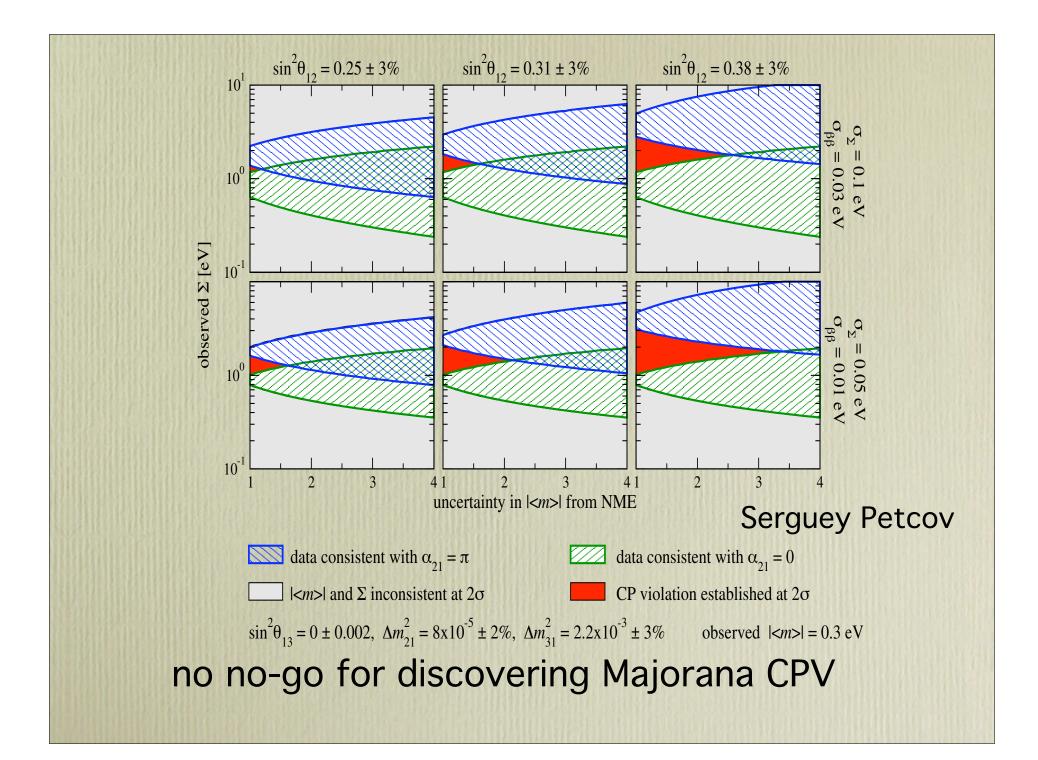
Carlos Pena Garay

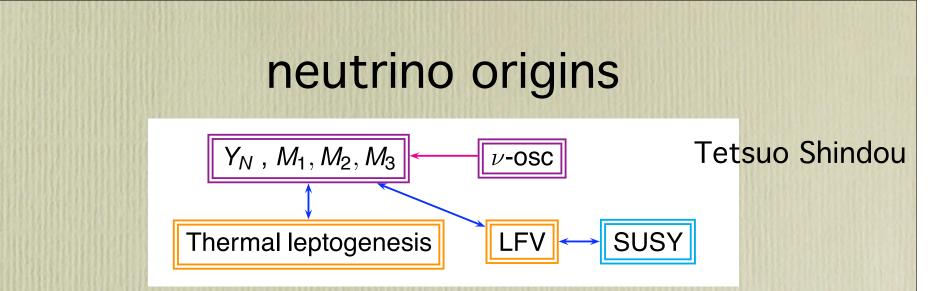
- where did all this come from?
- what are the energy scales where this gets generated?
- is it related to our own genesis?

what we need to know about neutrinos

Serguey Petcov

- Dirac or Majorana masses?
- Mass hierarchy: normal, inverted, quasi-degenerate?
- Absolute scale of masses?
- light steriles? eV, keV?
- relation to dark matter?
- theta_13?
- CP violation? Dirac or Majorana phases?
- lepton flavor violation apart from Majorana masses? related to TeV scale SUSY?
- relation to leptogenesis?
- origin of PMNS masses and mixings: what energy scales and symmetries are involved? relation to CKM?





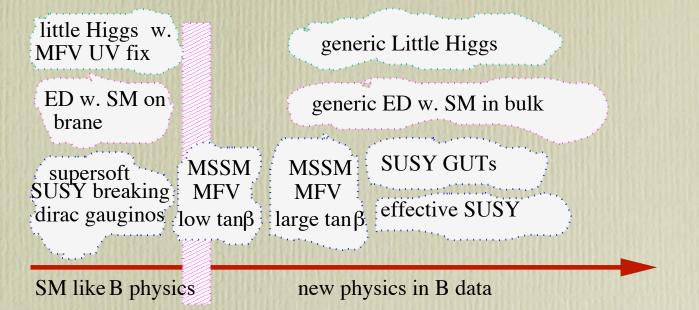
- we heard a story connecting SUSY@LHC to LFV data to discovery of IH and $0\nu\beta\beta$ to leptogenesis
- many such stories may be possible (Thomas Hambye, Fedor Bezrukov, Ernesto Arganda Carreras, Reinhold Rueckl, Pierre Hosteins, David Maybury, Lofti Boubekeur, Michel Tytgat, Thomas Underwood)
- huge long-term challenge to experiments, fertile ground for theorists

classify possible new flavor physics in the quark sector (from Buras and Fleischer)

- Minimal Flavor Violation (MFV): new diagrams, but no new operators. Only source of flavor-changing effects (including CP violation) is the CKM matrix. Examples: THDM-II and CMSSM moderate tan beta
- New operators, but no new CP violation. Example: MSSM with large tan beta
- New CP violation, but no significant contributions from new operators. Example: MSSM with moderate tan beta and nondiagonal squark mass matrices
- General new flavor violation. Examples: generic SUSY, multi-Higgs, Little Higgs, extra dims

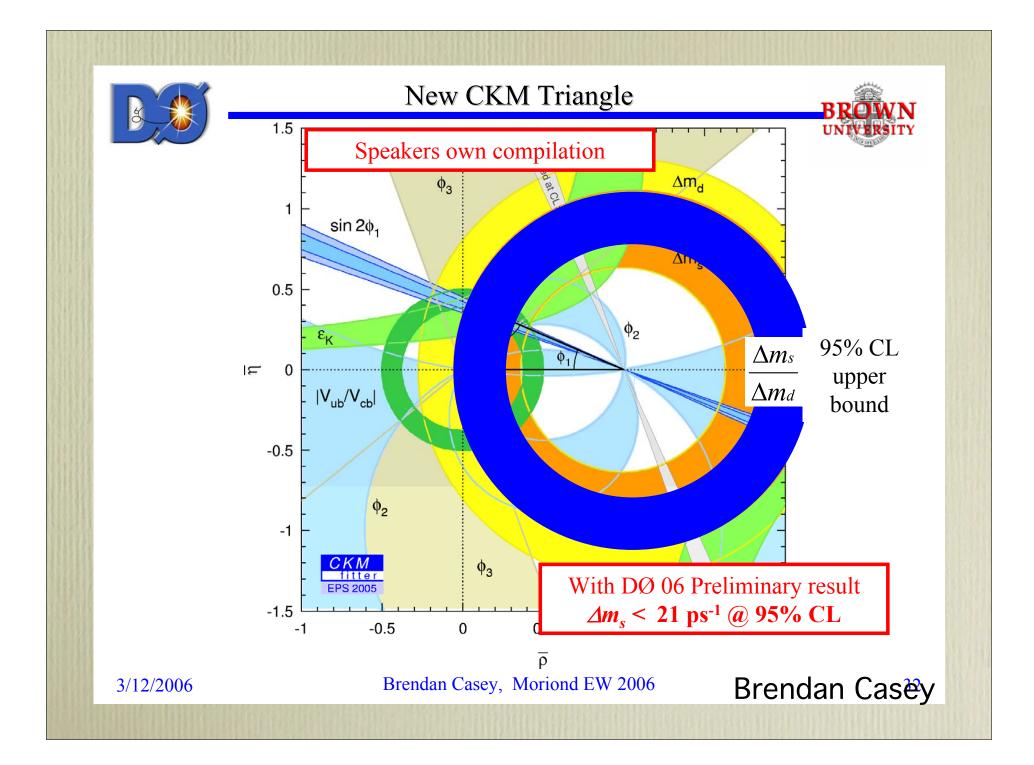
models of EWKSB with NP @ TeV

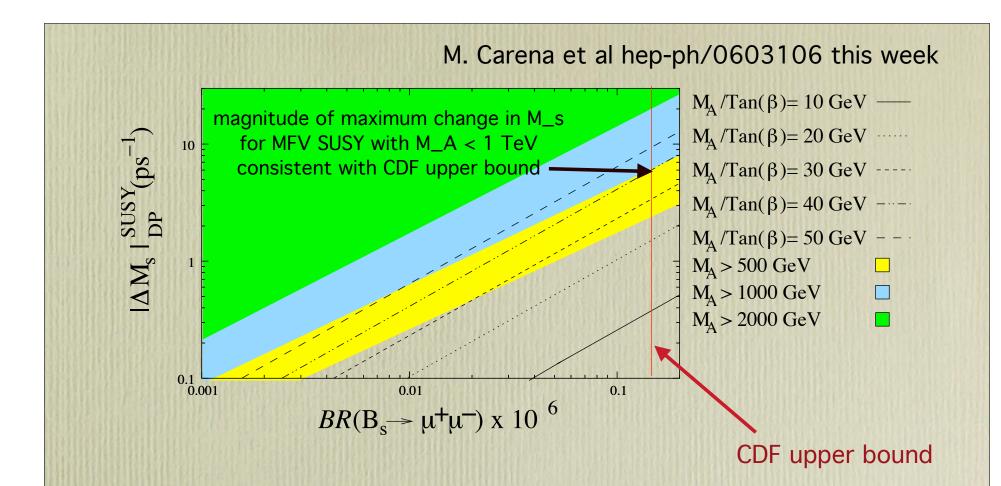
Fig from hep-ph/0207121



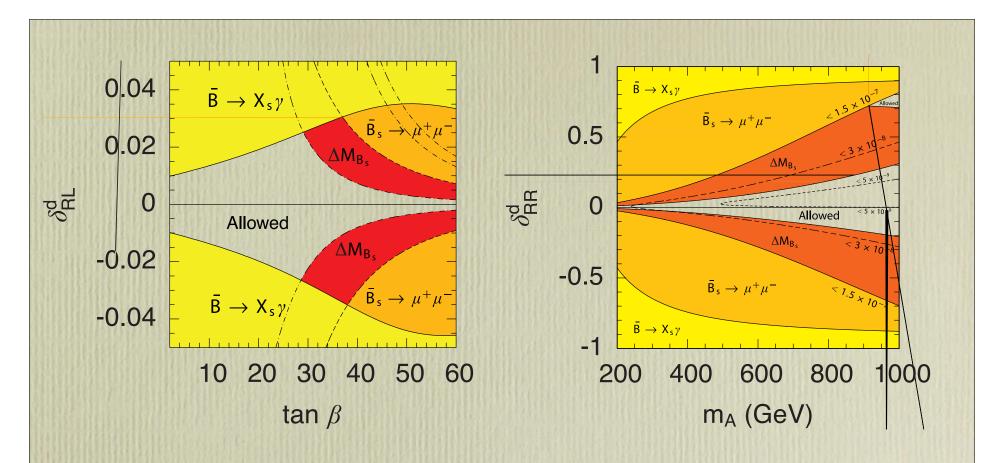
Gudrun Hiller, talk this week at CMS SUSY/BSM

so what does the Dzero result on B_s mixing tell us?





not much change for MFV SUSY



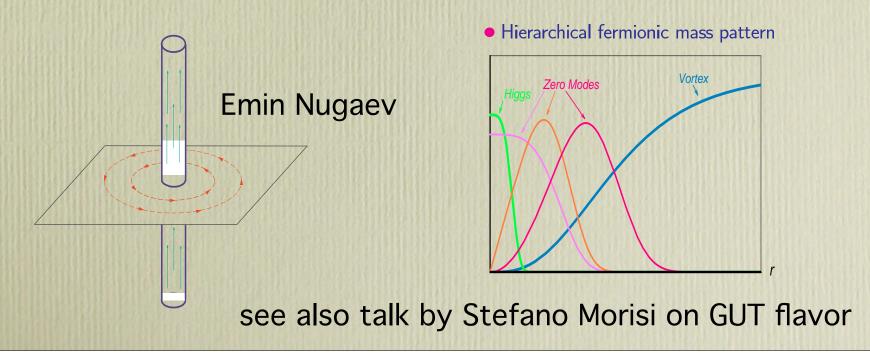
big constraints for generic SUSY

thanks to John Foster, Ken-ichi Okumura, and Leszek Roszkowski for these new plots!! "I was hoping to see some enhancement in B_s mixing as a signal of large atmospheric neutrino mixing and SUSY-GUT, and I am definitely sad"

-Hitoshi Murayama

flavor 2006

- neutrino flavor is a mystery
- the origins of CKM + fermion mass hierarchies are a mystery
- increasingly appears that the new TeV scale physics is MFV, but we don't know why or how
- flavor is a big challenge need new ideas!



cosmology 2006

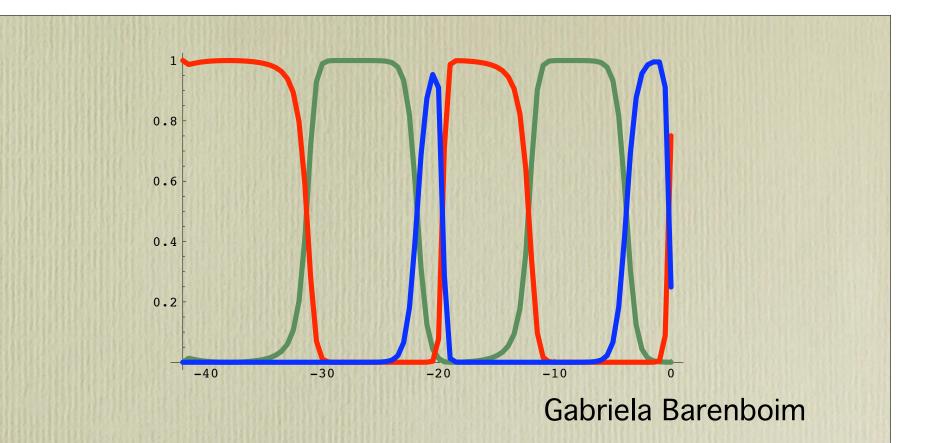
- the exciting WMAP 3-year results arrived during the conference and were reviewed by G. Barenboim and C. Pena Garay
- these results add yet more independent evidence for dark matter, while MOND is under attack from both ends (dwarf galaxies and clusters)
- WIMPs and axions are both well-motivated DM candidates, getting quite constrained by searches
- my guess is that DM will turn out to have several different components (like visible matter)

what we don't know about WIMPS

- the neutralino relic density estimates are strongly dependent on the SUSY model (Martin White).
 Scanning just mSUGRA is not good enough (David Cerdeno).
- Kaluza-Klein DM estimates are based on baby models, could change.
- sneutrino DM was ruled out prematurely (Stephen West)
- don't yet trust models for how WIMPS collect at the centers of galaxies (Malcolm Fairbairn)

the TeV frontier in cosmology

- WIMP relic density estimates assume a standard expansion rate and thermal history between BBN and T ~ 1 TeV
- but we have no independent knowledge of this!
- same is true of the EW phase transition at T~100 GeV, which in turn affects the prospects for EW baryogenesis (Stephan Huber)
- one of the great challenges for particle physicists is to help advance the cosmological frontier from T ~ 1 MeV to T ~1 TeV



- Slinky has a thermal history which satisfies the usual requirements but looks very different before BBN
- we can get such nonstandard cosmologies from a single scalar inflaton, but it has to have a rather strange form:

$$\int d^4 x \, \frac{1}{2} \mathbf{F}(\phi) \partial^{\mu} \phi \partial_{\mu} \phi - \mathbf{V}(\phi)$$
$$\mathbf{F}(\phi) = \frac{3}{8\pi G_N} \frac{\mathbf{c_1} e^{\mathbf{3}\phi} + \frac{4}{3} \mathbf{c_2} e^{\mathbf{4}\phi}}{\mathbf{c_0} + \mathbf{c_1} e^{\mathbf{3}\phi} + \mathbf{c_2} e^{\mathbf{4}\phi}}$$
$$\mathbf{V}(\phi) = \mathbf{c_0} + \frac{1}{2} \mathbf{c_1} e^{\mathbf{3}\phi} + \frac{1}{3} \mathbf{c_2} e^{\mathbf{4}\phi}$$

actually this inflaton doesn't give the Slinky cosmology, in fact it mocks up standard $\Lambda {\rm CDM}$!

inflatons are phenomenological devices!

dark energy?

- all observational evidence for dark energy is indirect
- the current accelerating expansion could be explained in many ways:
 - the FRW approximation is breaking down (Rocky Kolb)
 - the Friedmann eqn is wrong due to modified gravity or extra dimensions (Ignacio Navarro)
 - a tiny cosmological constant
 - quintessence (whatever that is)

the future

"Never trust a theorist"

- S. Ting



The Compact Muon Solenoid Experiment





6 December 2008

Evidence for squark and gluino production in pp collisions at $\sqrt{s} = 14$ TeV

CMS collaboration

Abstract

Experimental evidence for squark and gluino production in pp collisions $\sqrt{s} = 14$ TeV with an integrated luminosity of 97 pb⁻¹ at the Large Hadron Collider at CERN is reported. The CMS experiment has collected 320 events of events with several high E_T jets and large missing E_T , and the measured effective mass, i.e. the scalar sum of the four highest P_T jets and the event \vec{E}_T , is consistent with squark and gluino masses of the order of 650 GeV/ c^2 . The probability that the measured yield is consistent with the background is 0.26%.

Submitted to European Journal of Physics

preview of Moriond 2009

