



Project Status Overview: Optical Design Optimization

LSST Telescope Final Design October 3, 2005

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This work was performed under the auspices of the U.S. Department of Energy by University of California, Lawrence Livermore National Laboratory under Contract W-7405-Eng-48.

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Summary of accumulated changes to baseline

- The CA ID of the secondary is 1.8 m to facilitate installation of the camera through the hole in the secondary.
- Space of 5.0 cm (down from 8.0 cm) between clear apertures of the primary and tertiary mirrors.
 - The CA ID of the primary, joint line and the CA OD of the tertiary are respectively: 5116, 5066 and 5016 mm.
 - The overall sag of the M1 / M3 monolith is 680 mm.
- Warren Davison suggests reducing the diameter of the hole in the tertiary to 1054.7 mm, from 1160 mm.









- 8.36 m diameter beam
- Full field of view of 3.5 degrees
- Etendue of 318 m² deg²
- Focal ratio f/1.25
 - EFL: 10.45 m
 - Plate scale: 50.7 μm/arc-sec
- Flat focal plane
- Five photometric filters
 - g-r-i-z-y pass bands, 410 nm -1028 nm
 - Images calculated for 5 equally spaced and weighted wavelengths spanning the pass band
- Primary and tertiary mirror :
 - monolithic
 - 5 cm between beams on each mirror
- Camera
 - 3 cm between beams at L1 input surface
 - Camera can be inserted in hole of secondary: 1.8 m diameter CA ID
- Telescope by itself is corrected on-axis
- No ADC

Three fused spherical silica lenses

- L1
 - Maximum diameter, :1.62 m
 - Edge thickness at CA diameter is 3.4 cm
- L2
 - Central thickness of 3.0 cm
 - Minimum space to filter at least 30 cm
 - Accommodate filter change mechanism
 - L3
 - Vacuum barrier
 - >2.5 cm space to focal plane
 - Central thickness sufficient to provide "bulletproof" design for fracture safety
 - Diameter/thickness ratio of 12.17 yields
 6.0 cm thickness for 73 cm lens
- Fused silica filters
 - First surface concentric about chief ray
 - CT and 2nd surface curvature optimized for image quality
 - Minimum center thickness of 1.35 cm





Refine design regarding throughput, camera location, secondary asphericity while maintaining imaging performance

- 1. Add small amounts of asphericity on L_2 and L_3 to:
 - Reduce asphericity on secondary mirror
 - Improve null testing for L₂ and L₃
 - L_2 is now fixed; no imaging advantage for moving L_2
- 2. Study small changes in focal ratio: f/1.25 to f/1.20
 - Further reduce asphericity on secondary
 - L₂ cantilever and/or secondary diameter increase slightly





Bonus: asphericity on secondary can be reduced from 132 microns to 17 – 100 microns



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- Decreasing the L₁ cantilever can:
 - increase the diameter of the secondary
 - decrease the vignetting
 - increase the asphericity of the secondary
- Decreasing the focal ratio of the telescope can:
 - decrease the asphericity of the secondary
 - Either the L₁ cantilever or secondary diameter will increase











- Look at designs with secondary asphericity 17- 34 microns
 - There are relatively small increases of the L1-FP vertex distance
 - increase from 2.63 to 2.75 2.81 m should be acceptable
 - The potential advantages for fabricating and testing the secondary mirror are more significant
- Ranges
 - f/1.25 to f/1.2236 (plate scale 50.66 to 50.00 microns/arc-sec)
 - Primary radius 19.6 to 19.9 m
 - L1 cantilever 2.7 to 2.8 m
 - M2 CA diameter 3.37 to 3.44 m
 - Telescope length 6.32 to 6.45 m





- Comments on next slide [slide 10]
 - Designs A-H increase the L1 cantilever from 2.6 m to 2.7 m
 - Designs A-C with biconcave L2 have 21 to 25 microns asphericity
 - Long radius on 1st surface of L2 leads to testing difficulties since radius must be within + /- 50 mm or so
 - Designs D-H have a plano-concave L2
 - Ideal for fabrication and testing
 - Asphericity increases to 30 to 34 microns
 - Decreasing the focal ratio
 - Decreases the vignetting as it increases the diameter of M2
 - Slightly improves the image sizes
- Designs D-H are reasonable but look at what is required to further reduce asphericity on secondary



L1 cantilever increased to 2.7 meters; f/1.25, f/1.24 and f/1.2236



A-C, L2 is biconcave; D-H, L2 is plano-concave

LSST : Sept 30,2005		Units	А	В	С	D	Е	F	G	Н
FNBR			1.2500	1.2400	1.2336	1.2500	1.2500	1.2400	1.2336	1.2336
	Plate scale	μ m/sec	50.66	50.26	50.00	50.66	50.66	50.26	50.00	50.00
Casting	Required depth: M1M3	mm	676	676	676	676	671	672	676	674
M1	Radius	m	-19.681	-19.695	-19.688	-19.681	-19.832	-19.811	-19.685	-19.795
M1	Max. Dep.: BF parabola	μm	109	113	116	114	117	119	120	123
M2	Diameter (Optical CA)	m	3.40	3.41	3.42	3.40	3.41	3.43	3.42	3.44
M2	Max. Dep.: BSF	μm	25	21	21	32	34	31	30	31
M3	Max. Dep.: BSF	μm	411	409	406	407	394	391	394	388
L1	Diameter (Optical CA)	m	1.58	158	1.58	1.58	1.58	1.58	1.58	1.58
L1	Center Thickness	mm	-72.13	-70.18	-69.99	-72.67	-71.95	-70.70	-69.76	-72.63
L2	Diameter (Optical CA)	m	1.056	1.042	1.046	1.058	1.048	1.048	1.044	1.064
L2	Surface1: Radius	m	29.93	39.78	54.53	plano	plano	plano	plano	plano
L2	Surface2: Radius	m	-2.914	-2.954	-2.95	-2.651	-2.727	-2.796	-2.864	-2.789
	Max. Dep.: BSF	μm	260	253	268	213	230	245	233	244
L2-Inter Filtor	LZ IO FILLEI SPACE	mm	-300	-300	-300	-300	-300	-300	-300	-310.7
Comoro	Juliace Raz.RD Filler	m	-0.403	-0.223	-0.122	-0.3	-0.300	-0.174	-0.104	-0.000
Calliera			2.710	2.709	2.700	2.711	2.090	2.091	2.700	2.094
	Overall System Length ¹	m	6.352	6.331	6.317	6.35	6.37	6.35	6.32	6.33
	Integrated Etendue(A Ω)	m ² deg ²	318.9	319.2	319.7	318.9	319.5	319.9	319.7	320.3
	on-axis	%	62.55	62.55	62.55	62.55	62.55	62.55	62.55	62.55
	full-field	%	55.47	55.67	55.72	55.58	56.01	56.19	55.80	56.37
	Integrated	%	60.39	60.45	60.54	60.39	60.49	60.57	60.54	60.65
	Vignetting (% from center)	%	11.32	11.01	10.93	11.14	10.46	10.17	10.79	9.89
Image Size	e (Worst case 80% Diffraction	n Encircled	l energy Dia	a.)						
	g: 410 - 552 nm	sec	0.305	0.300	0.299	0.328	0.307	0.298	0.303	0.280
	r : 550 -694 nm	sec	0.273	0.259	0.250	0.280	0.248	0.234	0.257	0.236
	i: 694 - 847 nm	sec	0.265	0.252	0.242	0.267	0.236	0.222	0.248	0.215
	y: 847 - 930 nm	sec	0.278	0.261	0.247	0.274	0.239	0.225	0.252	0.219
	z: 960 - 1028 nm	sec	0.290	0.274	0.260	0.282	0.245	0.234	0.259	0.231





- Designs I M: Minimum asphericity of 17 microns achieved for L1 cantilever of ~2.8 m
- All designs I M should be acceptable
- Designs K-M look at reducing L1 CA size from 1.58 m to 1.55 m
 - Powers of L1 and L2 increase slightly (and weight, thickness)
 - Reducing the focal ratio increases the space between L2 and the filter; could be helpful for filter interchange
 - Size of L2 increases slightly
- Recommendation: Designs K-M should be passed by small group to further discuss



Increasing L1 cantilever to 2.8 m reduces secondary asphericity to minimum possible values; K-M trades M2 diameter and vignetting vs. L1 size



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LSST : Sep	LSST : Sept 30,2005		I	J	К	L	М
FNBR			1.2500	1.2400	1.2500	1.2400	1.2336
	Plate scale	μ m/sec	50.66	50.26	50.66	50.26	50.00
Casting	Required depth: M1M3	mm	675	677	677	677	677
M1	Radius	m	-19.840	-19.839	-19.842	-19.838	-19.835
M1	Max. Dep.: BF parabola	μm	102	104	100	105	111
M2	Diameter (Optical CA)	m	3.39	3.40	3.37	3.40	3.42
M2	Max. Dep.: BSF	μ m	17	17	17	17	17
M3	Max. Dep.: BSF	μm	400	393	400	401	403
L1	Diameter (Optical CA)	m	1.56	1.58	1.550	1.550	1.550
L1	Center Thickness	mm	-74.410	-74.220	-78.400	-79.14	-81.60
L2	Diameter (Optical CA)	m	1.058	1.060	1.078	1.084	1.098
L2	Surface1: Radius	m	plano	plano	planc	plano	plano
L2	Surface2: Radius	m	-2.609	-2.660	-2.508	-2.545	-2.541
L2	Max. Dep.: BSF	μ m	230	253	224	235	233
L2-filter	L2 to Filter space	mm	-305.3	-312.2	-300	-323.9	-355.8
Filter	Surface1&2:RD i-filter	m	-5.899	-5.731	-5.744	-5.69	-5.628
Camera	L1 1st vertex to M2	m	2.777	2.811	2.804	2.780	2.762
	Overall System Length ¹	m	6.43	6.42	6.45	6.42	6.39
	Integrated Etendue(A Ω)	m ² deg ²	318.3	318.2	317.6	318.8	319.5
	on-axis	%	62.55	62.55	62.55	62.55	62.55
	full-field	%	55.72	55.53	55.59	55.80	56.12
	Integrated	%	60.27	60.26	60.13	60.37	60.50
	Vignetting (% from center)	%	10.92	11.22	11.13	10.79	10.28
	a: 410 - 552 nm	sec	0.276	0.264	0.274	0.267	0.264
	r : 550 -694 nm	sec	0.228	0.209	0.232	0.216	0.209
	i: 694 - 847 nm	Sec	0.226	0.202	0.225	0,208	0.201
S	v: 847 - 930 nm	sec	0.233	0.207	0.232	0.215	0.209
ST Telescope t 3.2005	z: 960 - 1028 nm	sec	0.247	0.215	0.237	0.218	0.209



Current baseline Detector/WFS Layout

20 Curvature Sensors

