

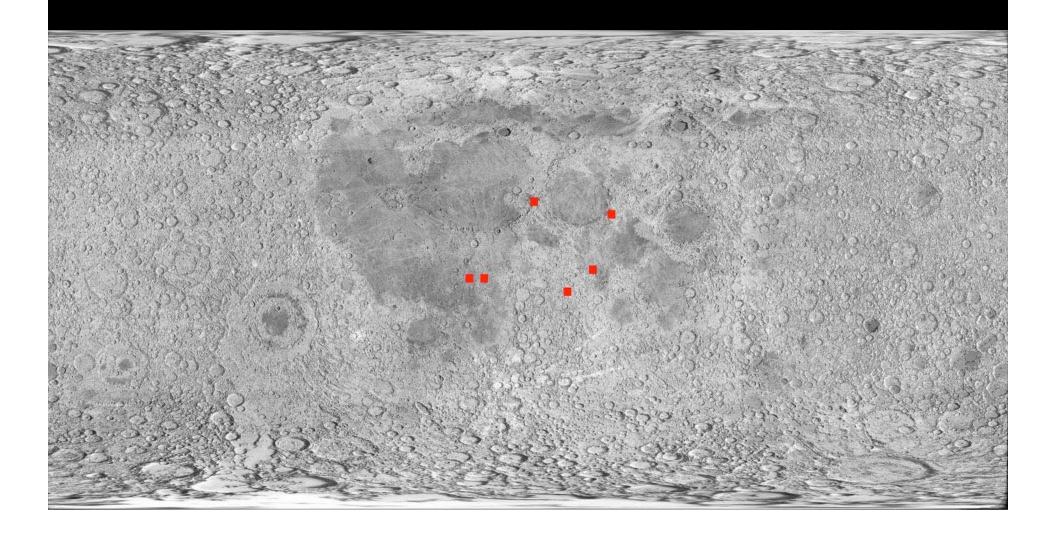
Moon Mineralogy Mapper: Return to the Moon by way of India

[M³ or "M-cube"] Carle Pieters, Principal Investigator Brown University

Apollo Exploration 1969 - 1972



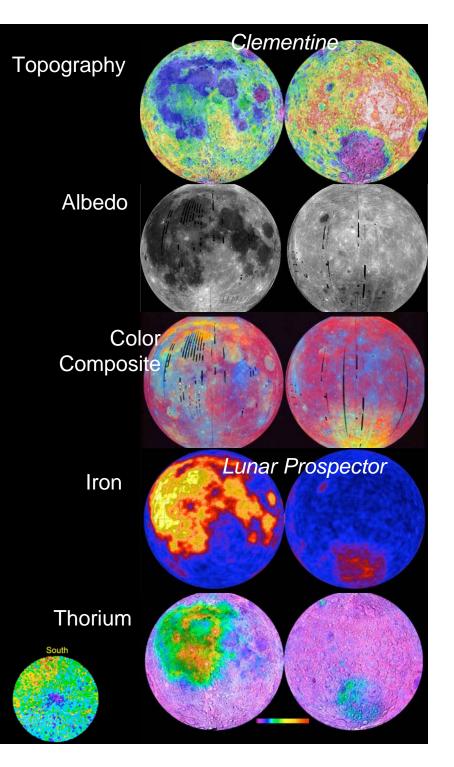
Landing sites are equatorial areas on the lunar near-side.



Paradigm Shifts from two small missions 1994 &1998

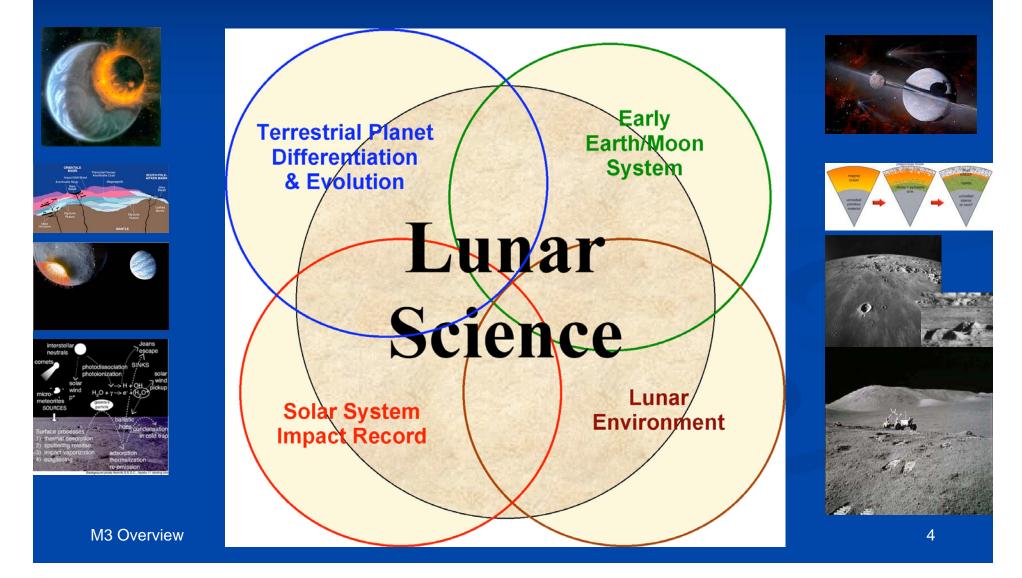
- The enormous South-Pole Aitken basin dominates the feldspathic farside of the Moon.
 - Largest and oldest lunar basin
 - Minor basalt fill
 - Iron-rich interior (lower crust/mantle)
- Heat producing elements were concentrated on the lunar nearside (Apollo sites) early in lunar history.
- The poles are unusual environments and may accumulate volatiles.

Polar H

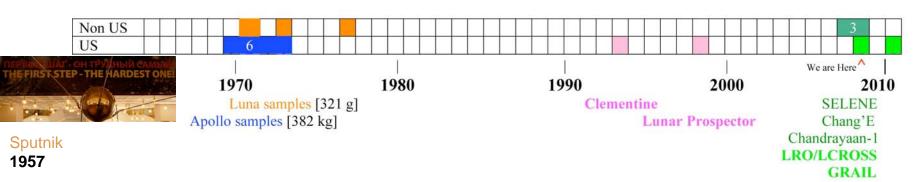


Overarching Science Themes See NRC/NAS 2007 Report:

http://books.nap.edu/catalog.php?record_id=11954

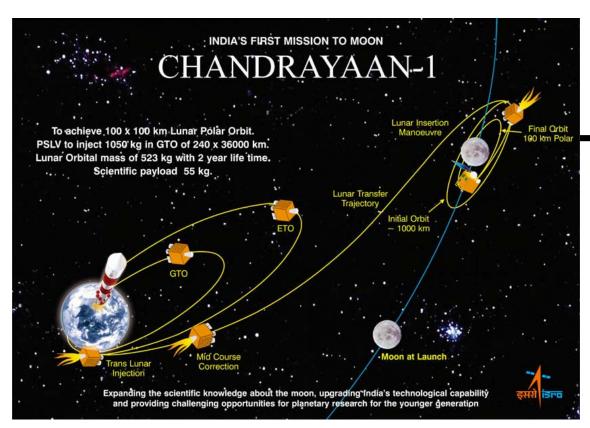


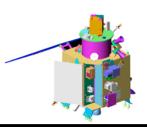
Lunar Exploration Timeline: Return after a long Drought





- The Apollo/Luna samples brought new and fundamental understanding of planetary evolution (and the Earth-Moon system).
- After decades of neglect, two very small missions were sent to the Moon. The small pulse of new data sparked several paradigm shifts.
- A fleet of sophisticated modern sensors are now at last exploring the Moon.



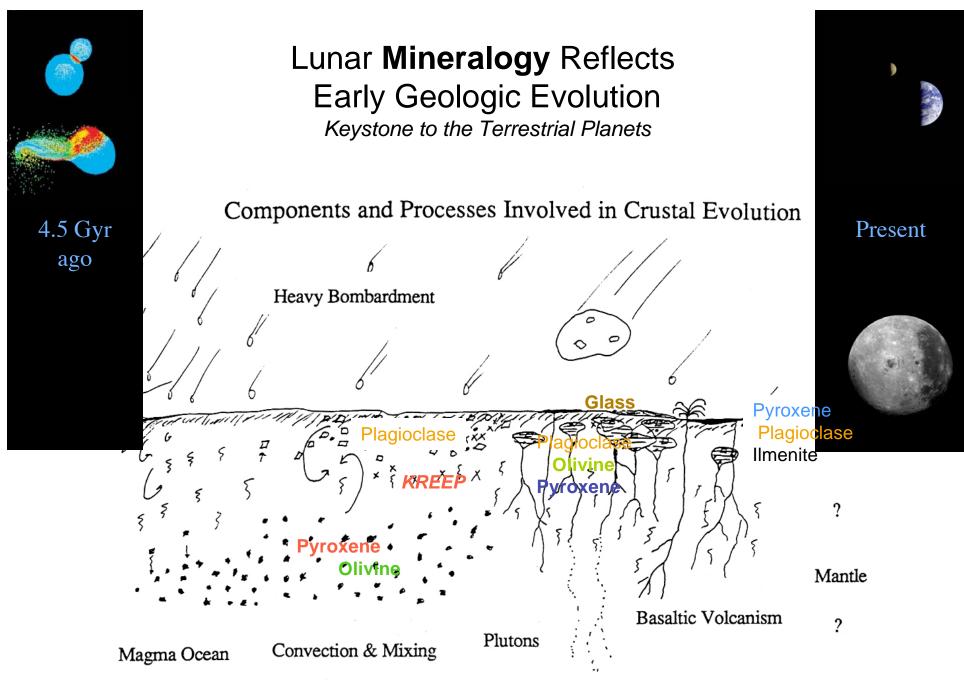


M3: A Guest Instrument on Chandrayaan-1, ISRO's first mission to the Moon

ISRO: Indian Space Research Organization

- What: Near-infrared imaging spectrometer to measure and map the mineral composition of the Moon at high resolution
- Why: Valuable science data to understand the evolution of the Moon.
- Selection: Summer 2004 selected by ISRO for foreign payload "short list"; Feb. 2005 selected for funding by NASA as a Discovery Mission of Opportunity after an extensive peer-review process; Confirmation 2/2006
- When: Deliver 2007, Launch middle 2008. Two-year operations at Moon





Serial Magmatism



M3 Objectives



- Primary **Science** Goal: Characterize and map the lunar surface composition in the context of its geologic evolution. This translates into several science sub-topics to be addressed.
- Primary **Exploration** Goal: Assess and map the Moon mineral resources at high spatial resolution to support planning for future, targeted missions.





Moon Mineralogy Mapper (M3)



Chandrayaan-1 launches in mid 2008 on an Indian spacecraft

- 100 km circular polar Orbit
- Two year mission duration

M3 is a NASA Discovery "Mission of Opportunity"

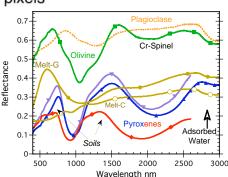
- Team led by PI: C. Pieters
- Designed and built at JPL

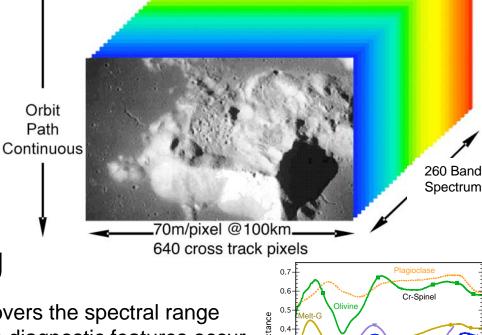
M3 is a pushbroom imaging spectrometer

Two spatial dimensions One spectral dimension M3 covers the spectral range where diagnostic features occur for all common rock-forming minerals and hydrous phases [0.43 to 3.0 µm].

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ISPO





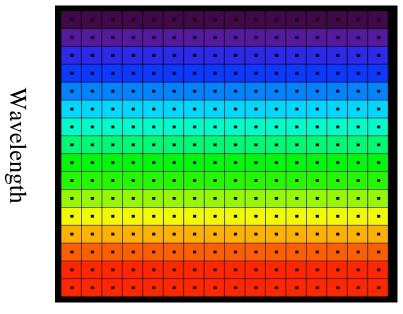
M3 Pushbroom Imaging Spectrometer



Unique Characteristics of M3 Spectrometer



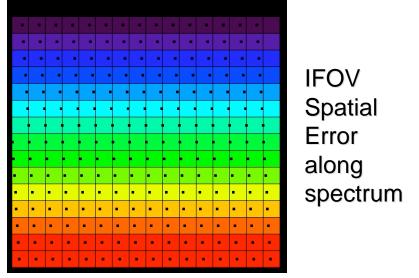
M3 design provides spectral and spatial Uniformity: <10% of a pixel across and along detector FOV.



Spatial Cross Track Sample

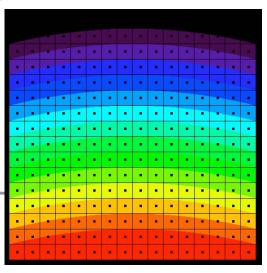






Common Problem Modes of other spectrometers

Spectral Error across FOV

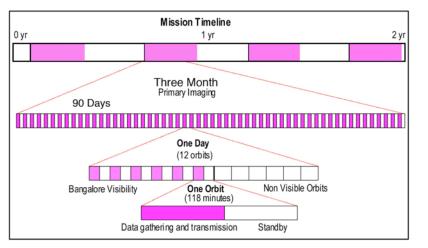




M3 Measurements extend over Two Years



Chandrayaan-1 Two-year mission plan: Four optical periods with high sun lighting



Data volume limited by downlink options.

M3 Measurement plan:

All M3 Reflectance Spectra

- 0.70 to 3.0 μm [0.43 to 3.0 μm achieved]
- 40 km FOV, contiguous orbits
- high SNR
- 1 Gbyte/orbit
- Targeted Mode: Optimum
 - Resolution (100 km orbit):
 - 70 m/pixel spatial
 - 10 nm spectral [260 bands]
 - 3 optical periods [20 50% coverage]
 - 10 to 12 deg latitude/orbit
- Low Resolution Mode: Global Coverage
 - Resolution (100 km orbit):
 - 140 m/pixel spatial
 - 20 & 40 nm selected (86 bands, ~3x spectral averaging)
 - 1 optical period [~100%]

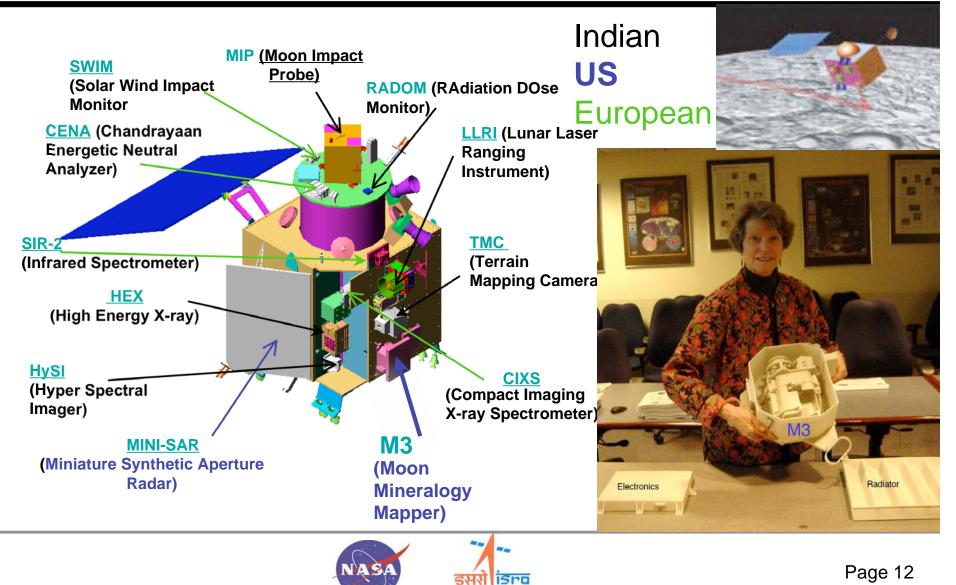






M3 is Part of a Highly International Exploration Mission



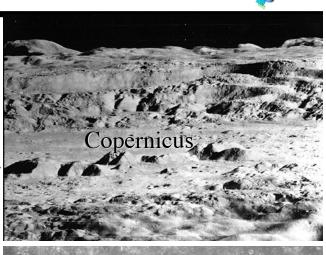


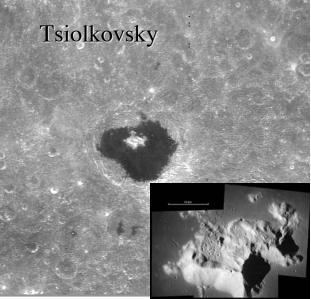


M3 Science Goals Address Planetary Science Issues



- Origin and Evolution of the lunar crust and mantle.
 - Evaluate primary crustal components and their distribution across the highlands.
 - Characterize the diversity and extent of different types of basaltic volcanism.
 - Map fresh craters [probes to the interior; impact record].
- Identify and assess deposits containing volatiles.
- Identify and evaluate concentrations of unusual/unexpected minerals.



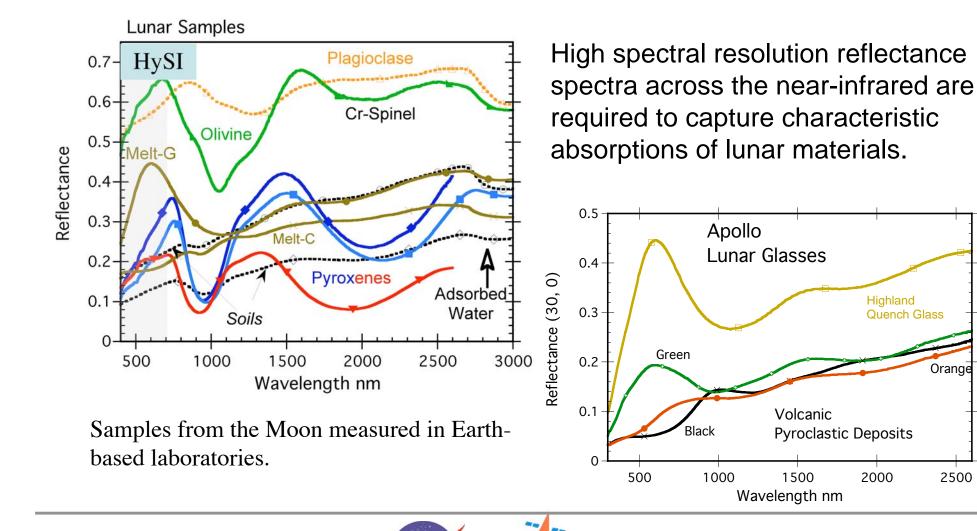






Minerals Exhibit Highly Diagnostic Absorption Bands





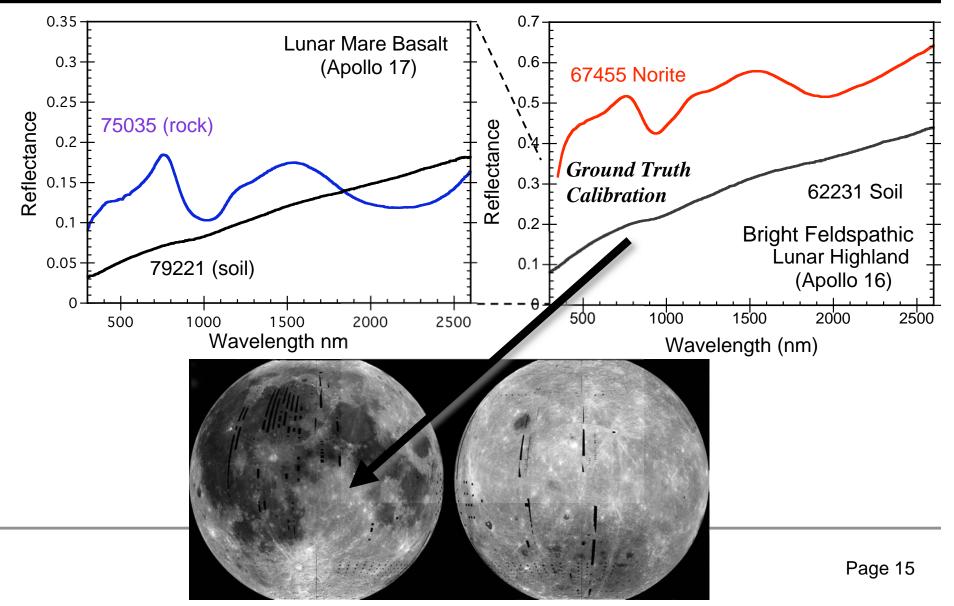
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Lunar Samples Provide Ground Truth

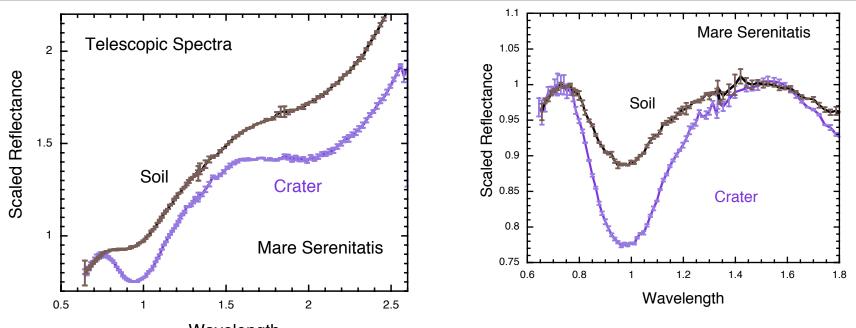


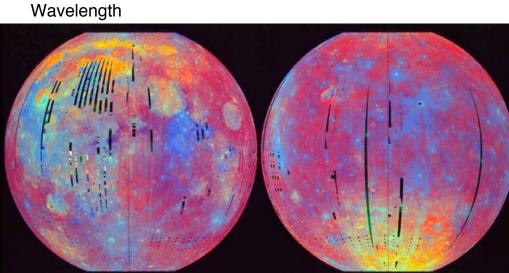




Mineral Characteristics are Weaker in Well-developed (mature) Soils





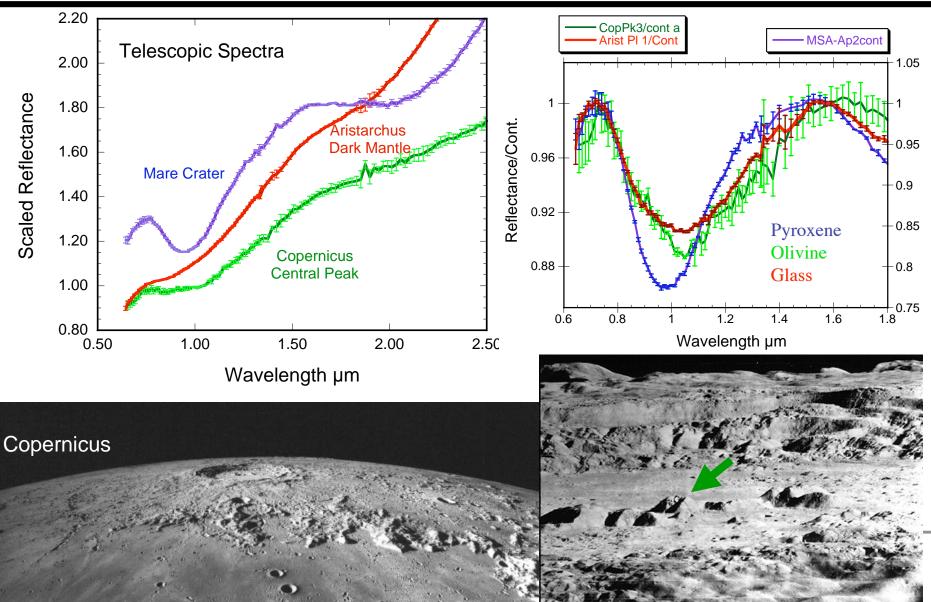


Clementine Color Composite R: 750/415 G: 750/950 B: 415/750



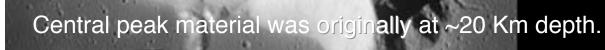
Mineral Characterization Requires High Spectral Resolution



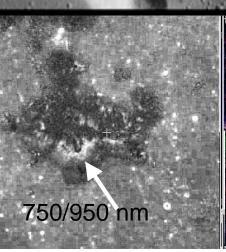


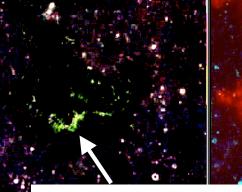


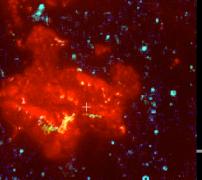
Pieters and Tompkins 1999 JGR



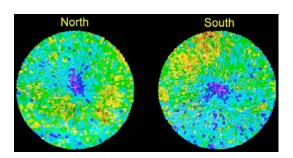
750 nm Albedo





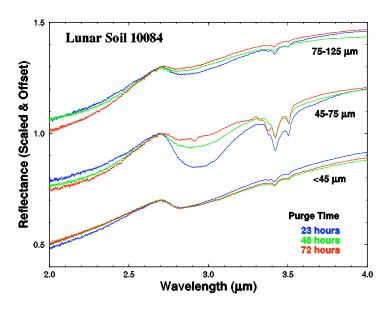


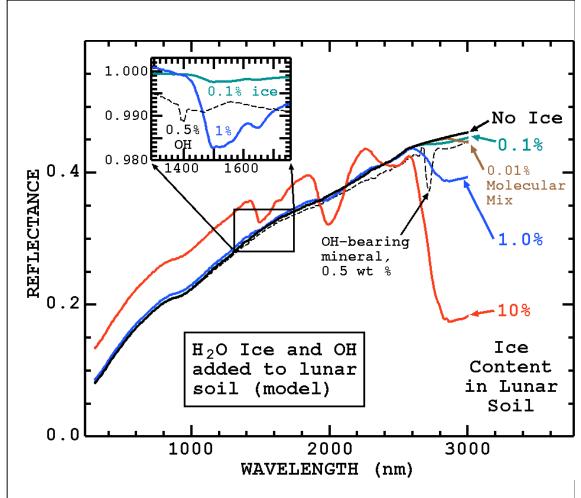
Mafic rich zones in central peaks:



Lunar Prospector Hydrogen at the poles Water ice?

Anhydrous lunar soils do not give up adsorbed water easily.





M3 Near-IR+ spectroscopy can uniquely identify the presence of small amounts of OH or H2O

(surface scattered solar radiation).



M3 January 2007



- Last viewing before radiator and thermal blankets attached
- Preparation for final thermal vacuum testing and calibration









Milestones



- M3 ship to India: August 2007
- Chandrayaan-1 instrument integration and testing: Fall-Winter 2007
 - Chandrayaan-1 launch from India: Middle 2008
 - Chandrayaan-1 operations at the Moon: few weeks later!
 - Initial science report: late 200
- Continue science analyses



