

# *Cross-fertilization of QCD and statistical physics*

*Stéphane Munier*

*CPHT, École Polytechnique*



Lisbon, June 30, 2006

# Two years of reaction-diffusion

## New insights in high energy QCD from general tools of statistical physics

*Stéphane Munier*

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We show that high energy scattering is a statistical process similar to reaction-diffusion in a system made of a finite number of particles. The Balitsky-JIMWLK equations correspond to the time evolution law for the particle density. The squared strong coupling constant plays the rôle of the minimum particle density. Discreteness is related to the finite number of partons one may observe in a given event and has a sizeable effect on physical observables. Using general tools developed recently in statistical physics, we derive the universal terms in the rapidity dependence of the saturation scale and the scaling form of the amplitude.

Brookhaven, July 16 and July 23, 2004

SM, talk at Brookhaven (2004);  
Iancu, Mueller, SM (2004)  
Enberg, Golec-Biernat, SM (2005)

statistical physics, we are able  
energy behavior of QCD  
(parton model)

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universality class of F-KPP equation

$$1 - u^2(x, t)$$

universality class of sF-KPP equation

$$1 + \sqrt{u(x, t)} \eta(x, t)$$

Fluctuations are related to the finite number of partons in a given event

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## New insights in high energy QCD from general tools of statistical physics

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CPHT,

We show that high energy diffusion in a system made of equations correspond to the squared strong coupling constant. Discreteness is related to the event and has a sizeable dependence developed recently in statistical rapidity dependence of the s

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### Summary

Using general results obtained recently in statistical physics, we are able to derive non-trivial properties on the high-energy behavior of QCD scattering amplitudes **from first principles** (parton model)

Balitsky-Kovchegov equation is in the universality class of F-KPP equation

$$\partial_t u(x,t) = \partial_x^2 u(x,t) - u(x,t) - u^2(x,t)$$

Balitsky-JIMWLK equation is in the universality class of sF-KPP equation

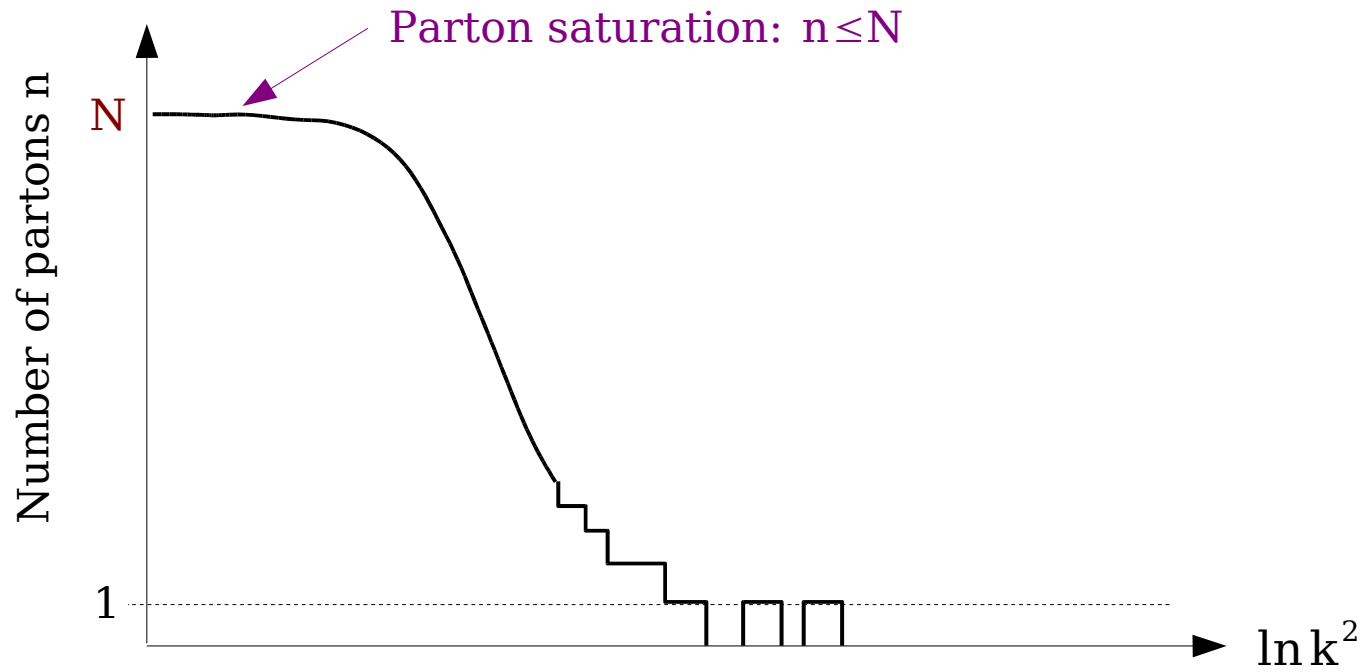
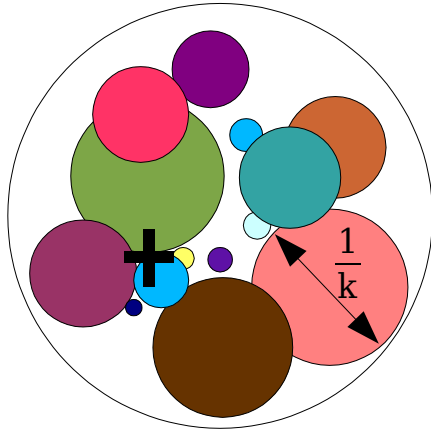
$$\partial_t u(x,t) = \partial_x^2 u(x,t) + u(x,t) - u^3(x,t) + \sqrt{u(x,t)} \eta(x,t)$$

Fluctuations are related to the finite number of partons in a given event

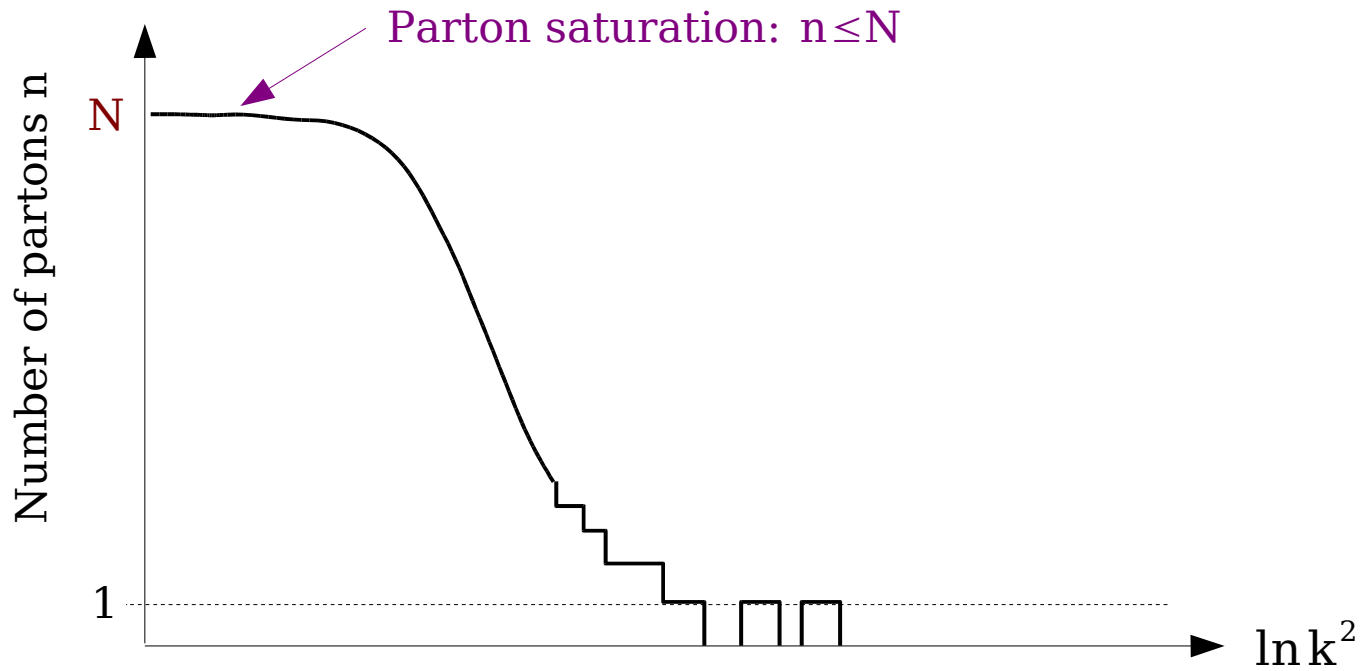
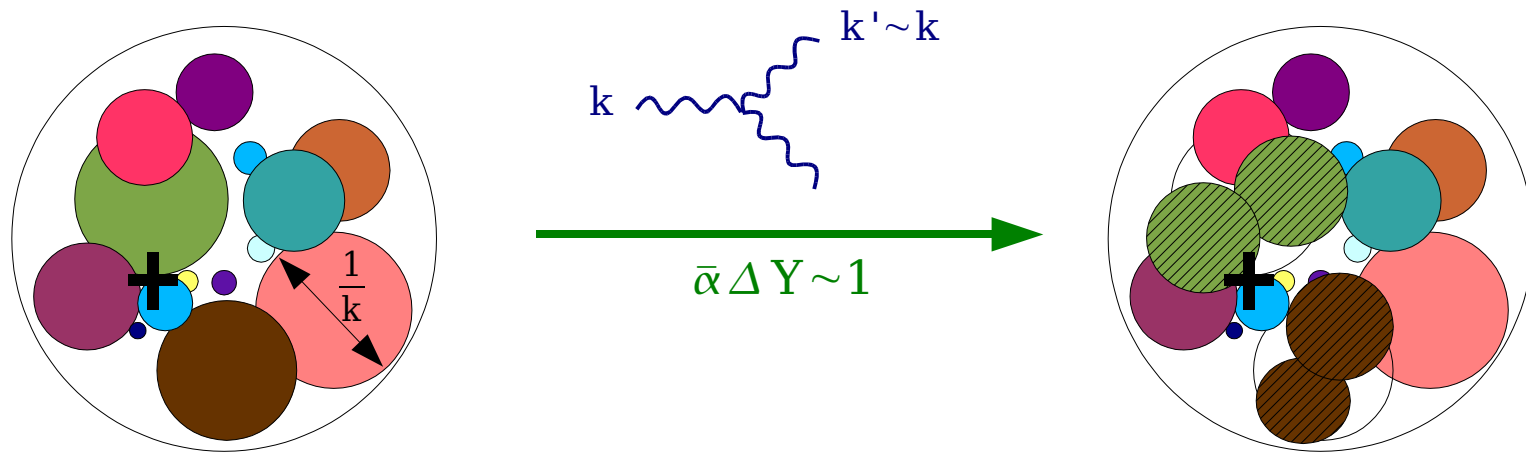
# Outline

- Why the stochastic FKPP equation is relevant to high energy QCD
- A new perspective on its solutions:
  - selective population evolution and statistics of genealogies

# *How a high rapidity hadron looks*



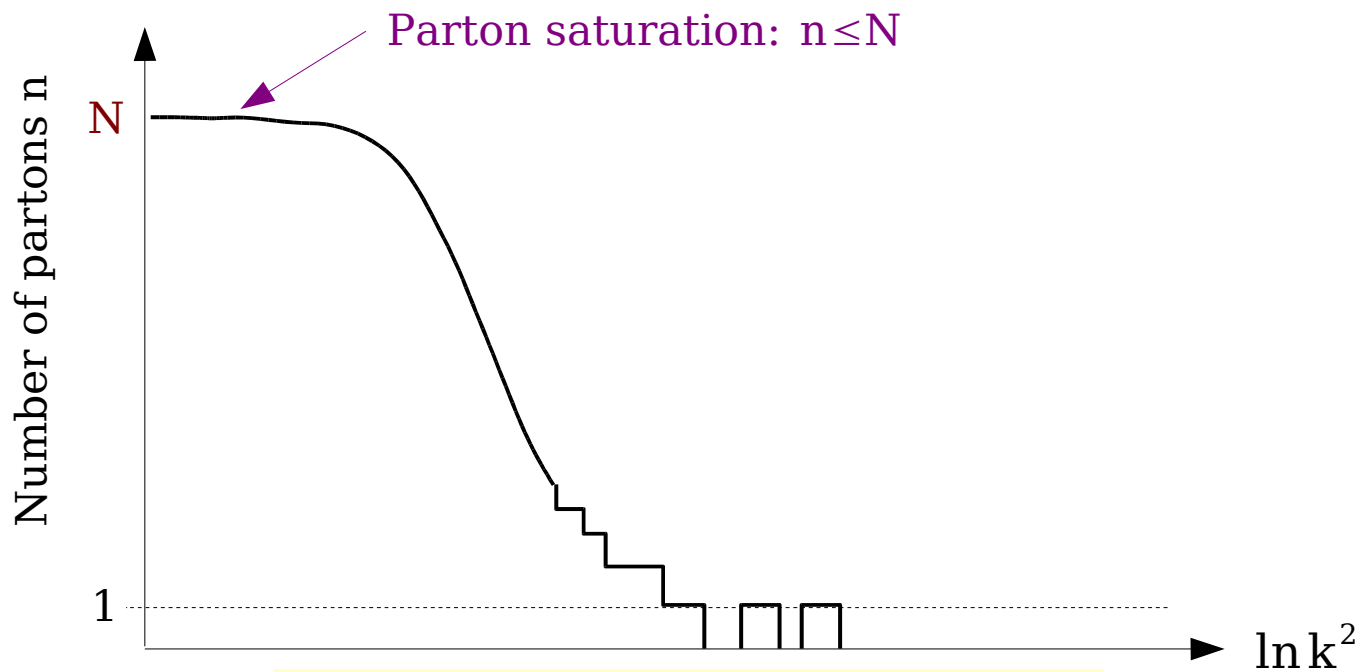
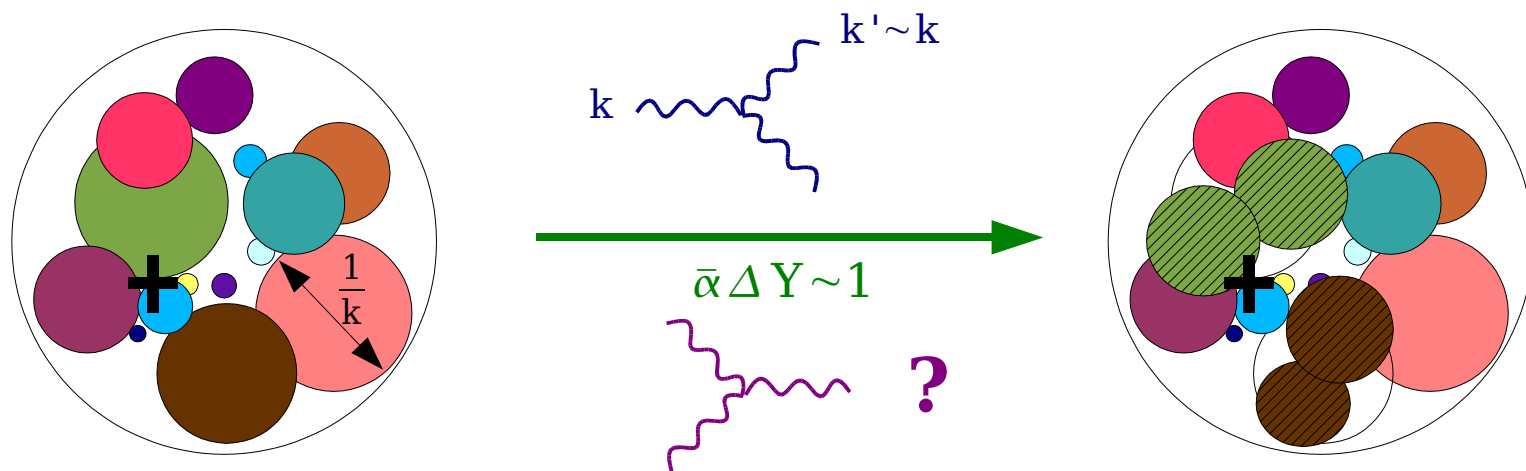
# How a high rapidity hadron looks



$$\partial_{\bar{\alpha} Y} \mathbf{n} = \chi(-\partial_{\ln k^2}) \mathbf{n} + \sqrt{\mathbf{n}} \nu$$

BFKL  $\sim \partial_{\ln k^2}^2 \mathbf{n} + \mathbf{n}$       Noise term due to discreteness

# How a high rapidity hadron looks

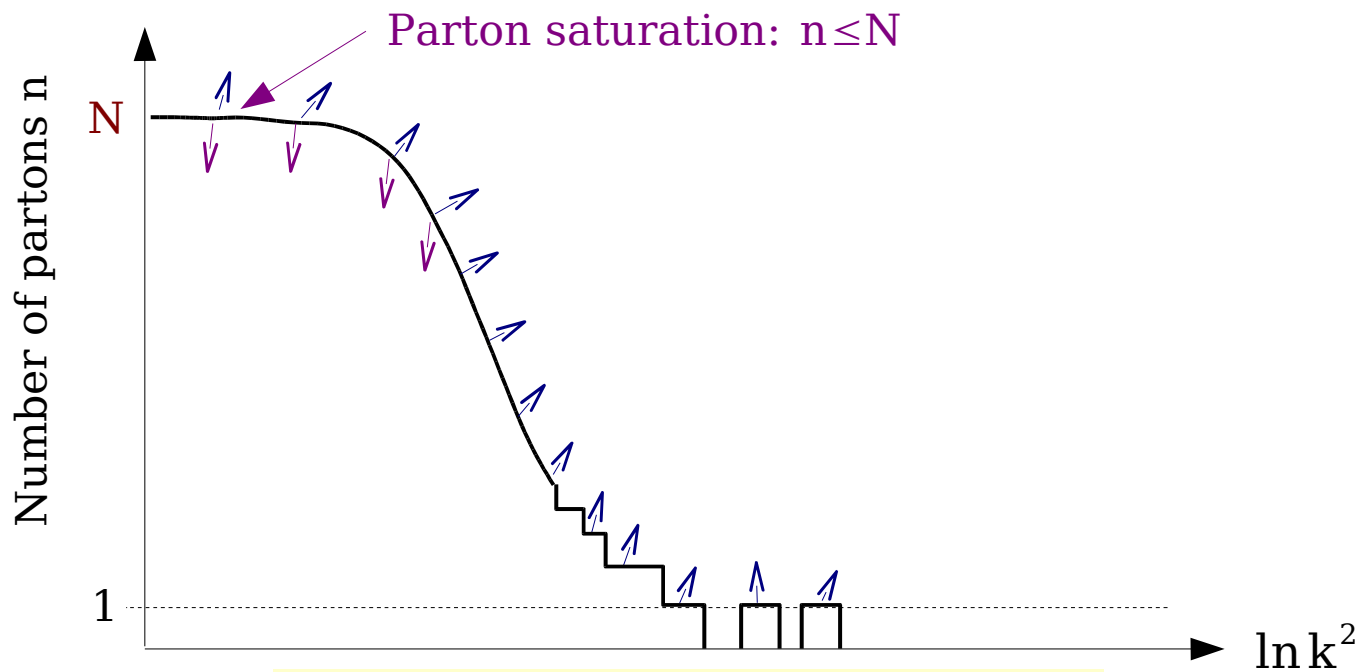
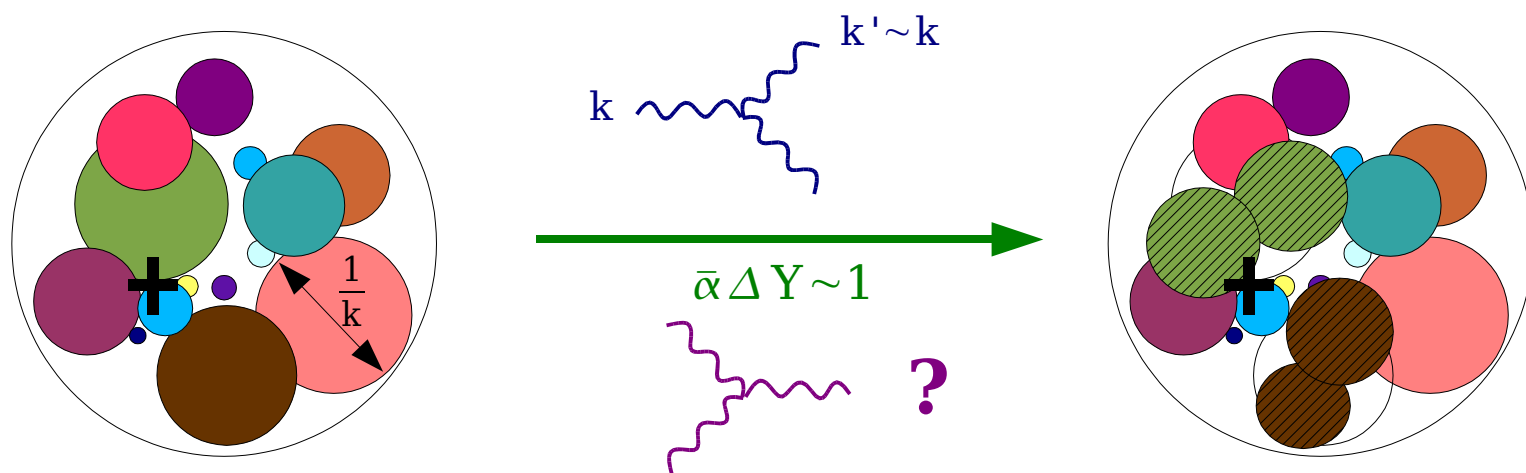


$$\partial_{\bar{\alpha} Y} \mathbf{n} = \chi \left( -\partial_{\ln k^2} \right) \mathbf{n} - \frac{\mathbf{n}^2}{N} + \sqrt{\mathbf{n}} \mathbf{v}$$

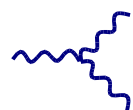
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Noise term due to discreteness

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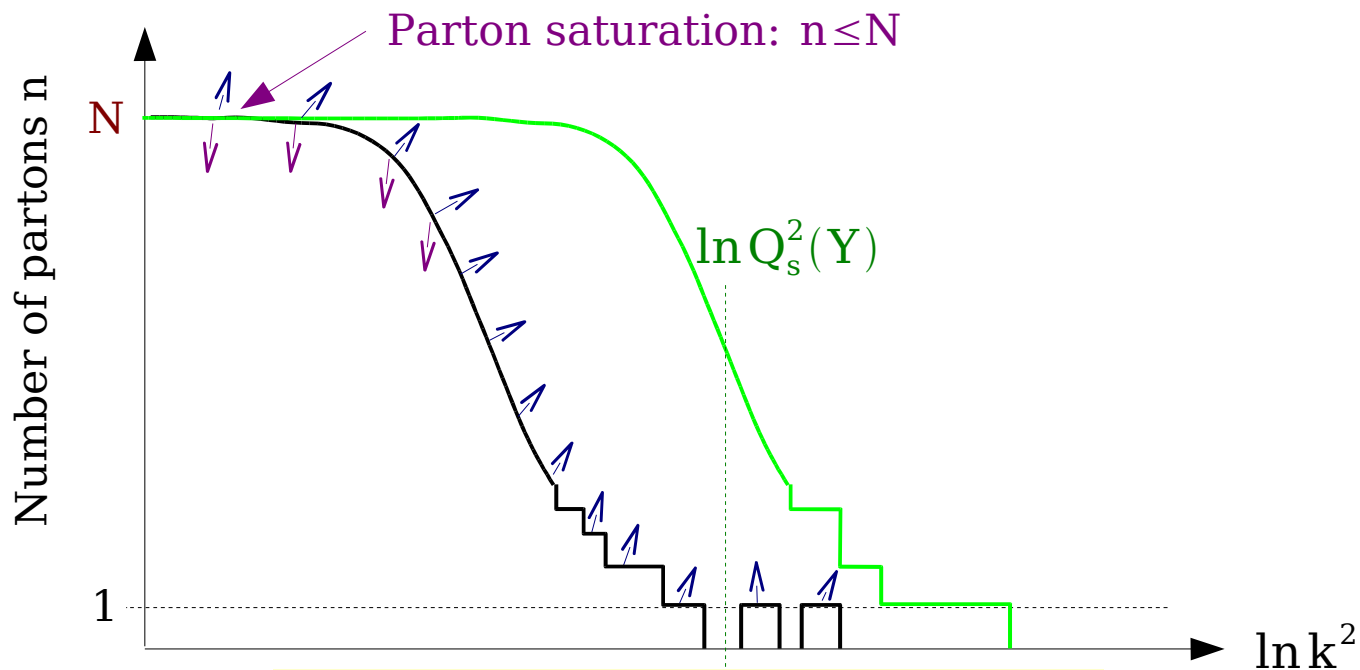
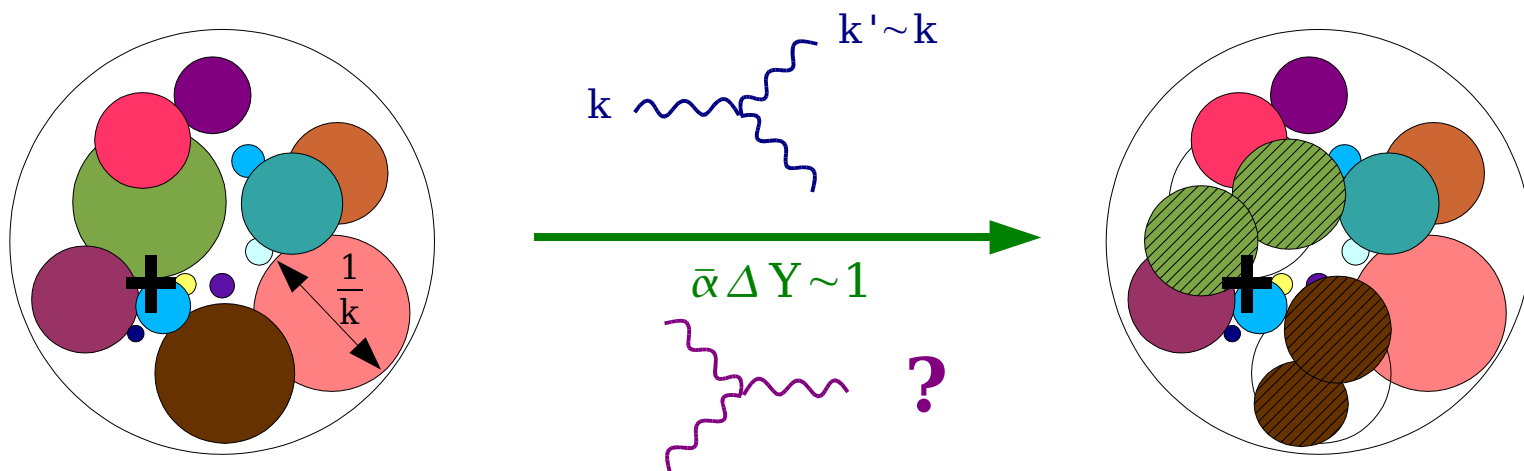
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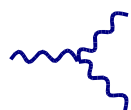
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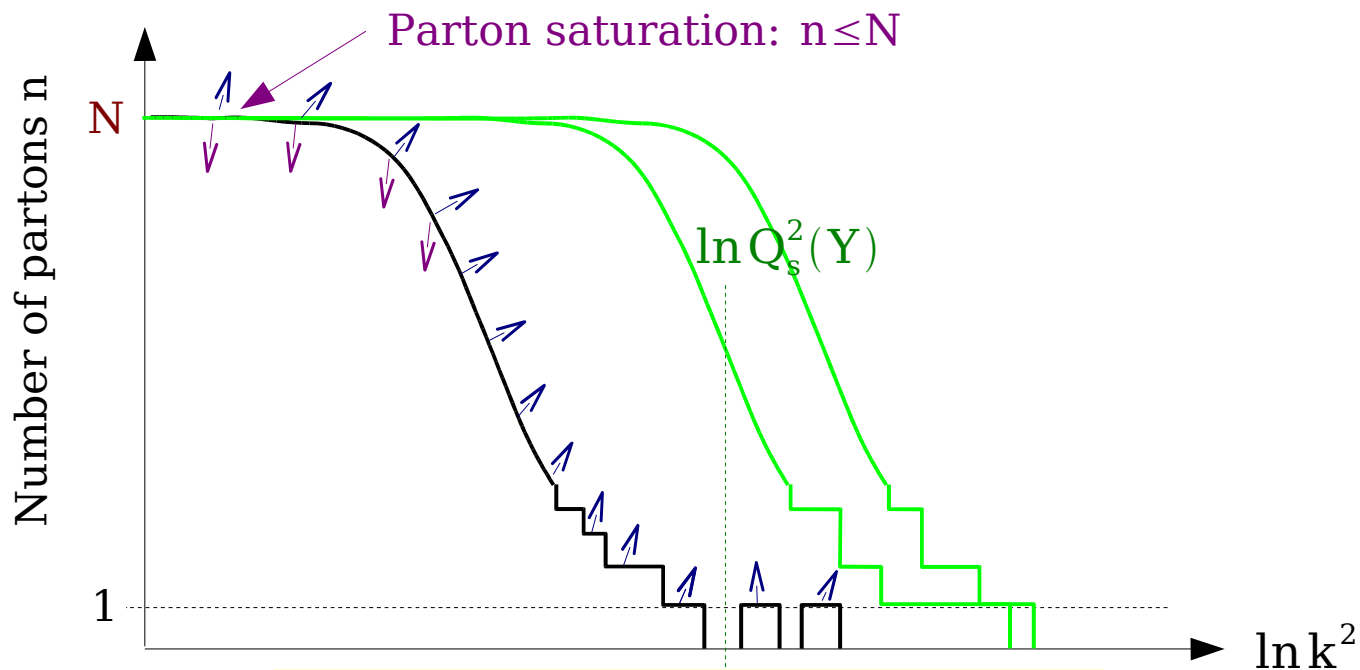
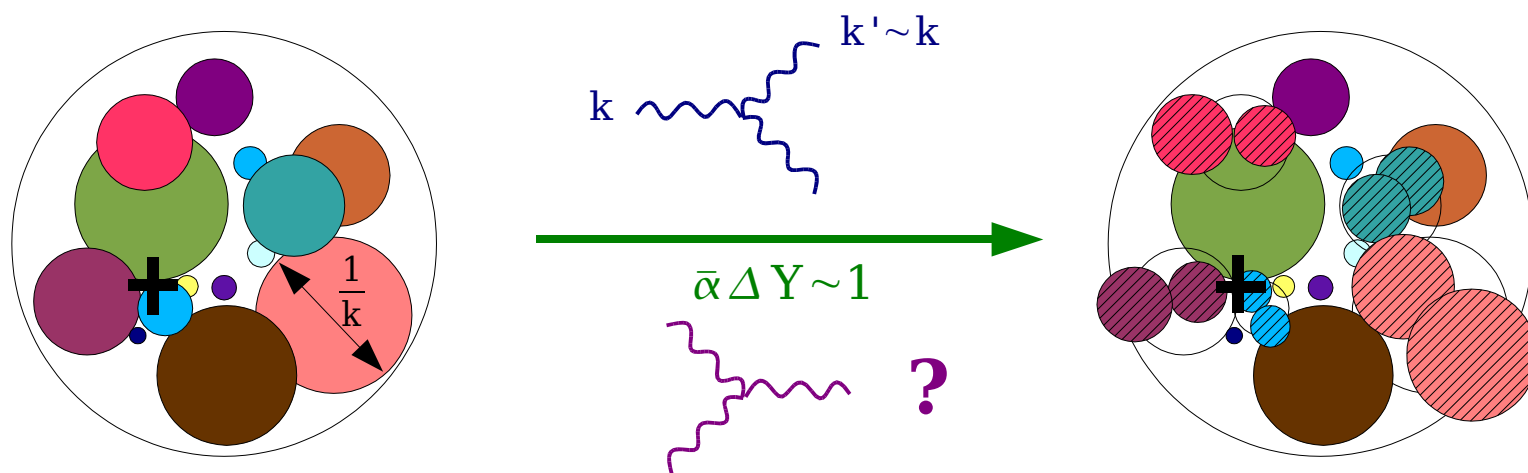


$$\text{BFKL} \sim \partial_{\ln k^2}^2 \mathbf{n} + \mathbf{n}$$

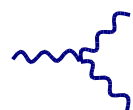


Noise term due to discreteness

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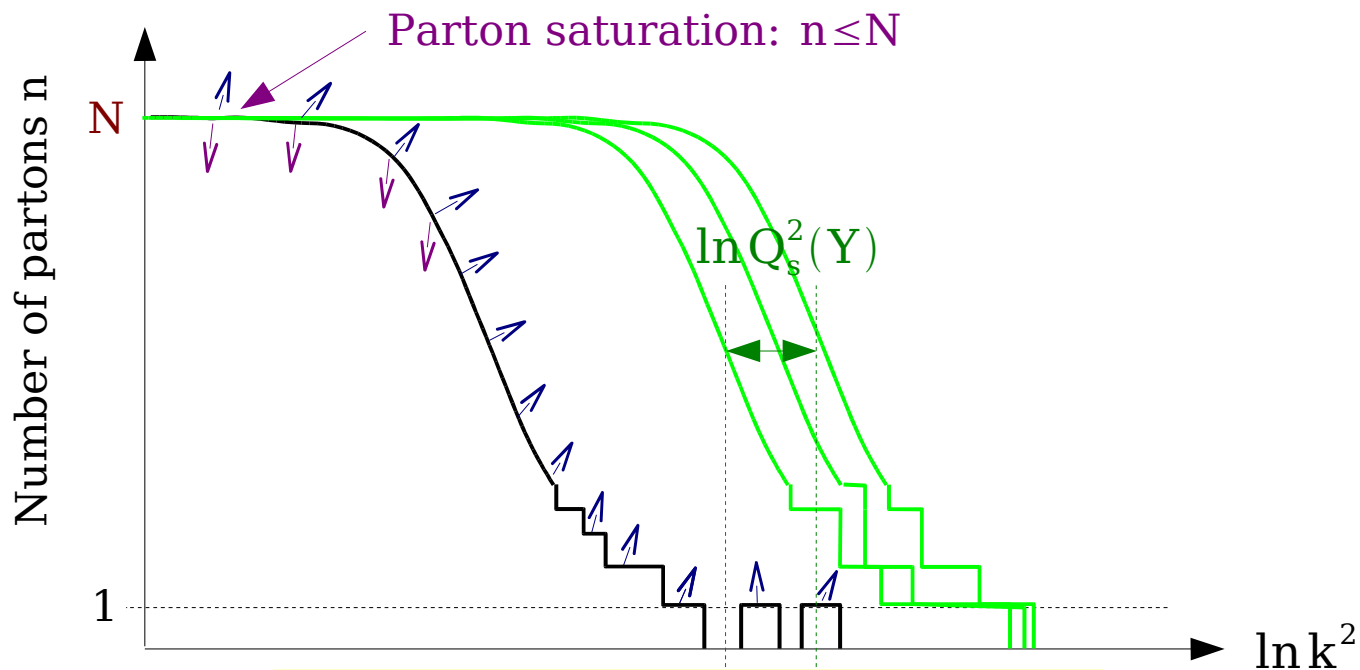
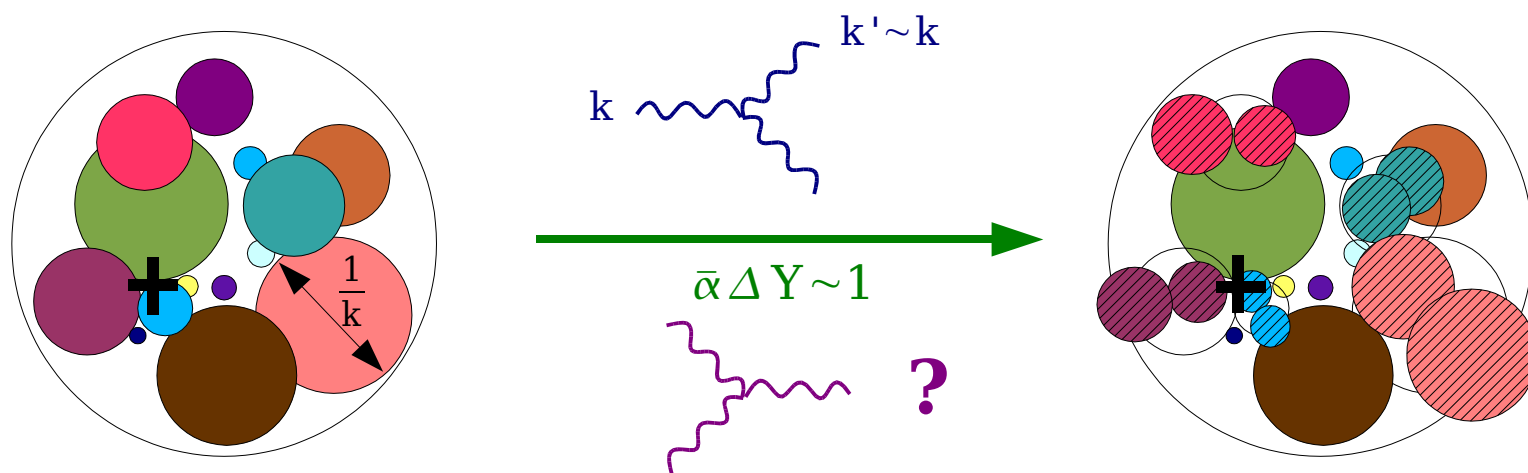


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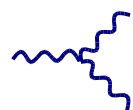


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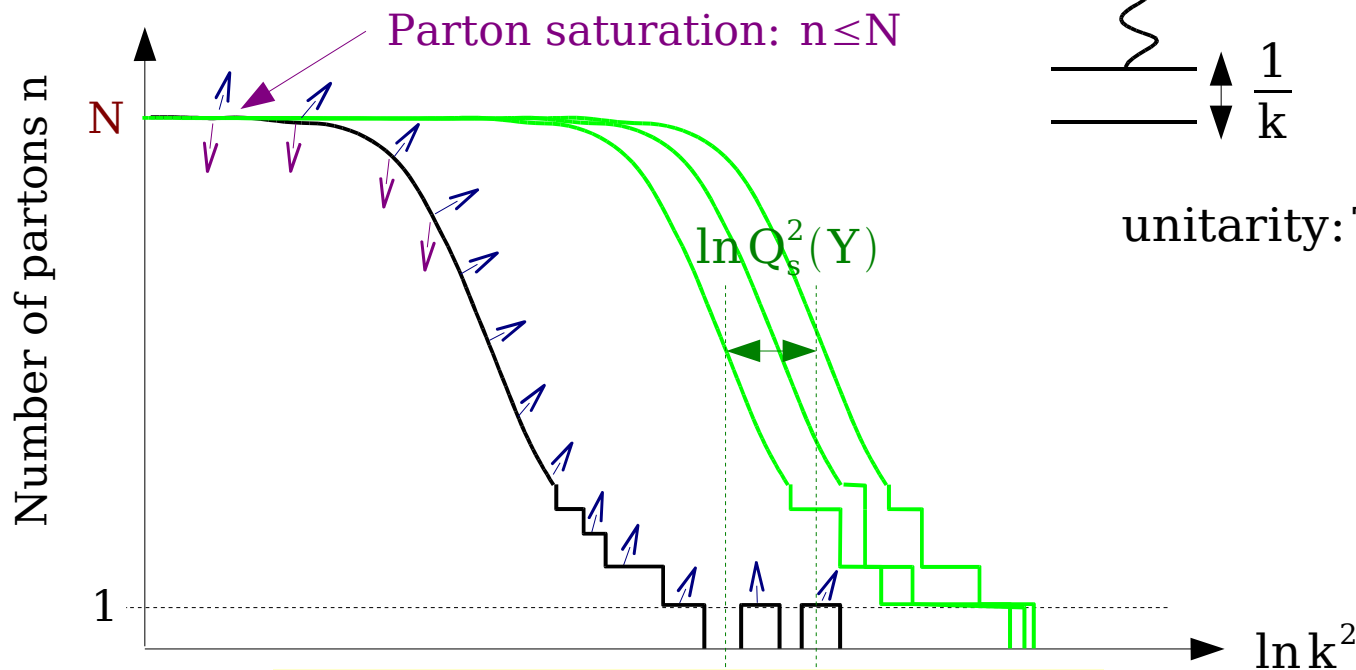
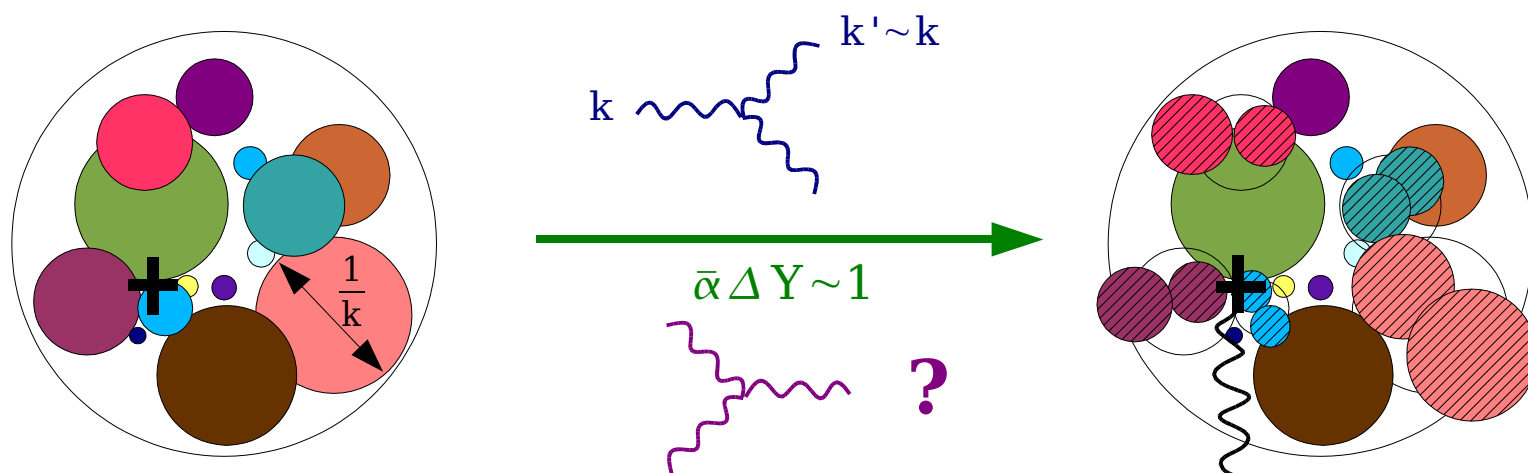


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Noise term due to discreteness

# How a high rapidity hadron looks



$\frac{1}{k}$       $T(k) \sim \alpha_s^2 n(k)$

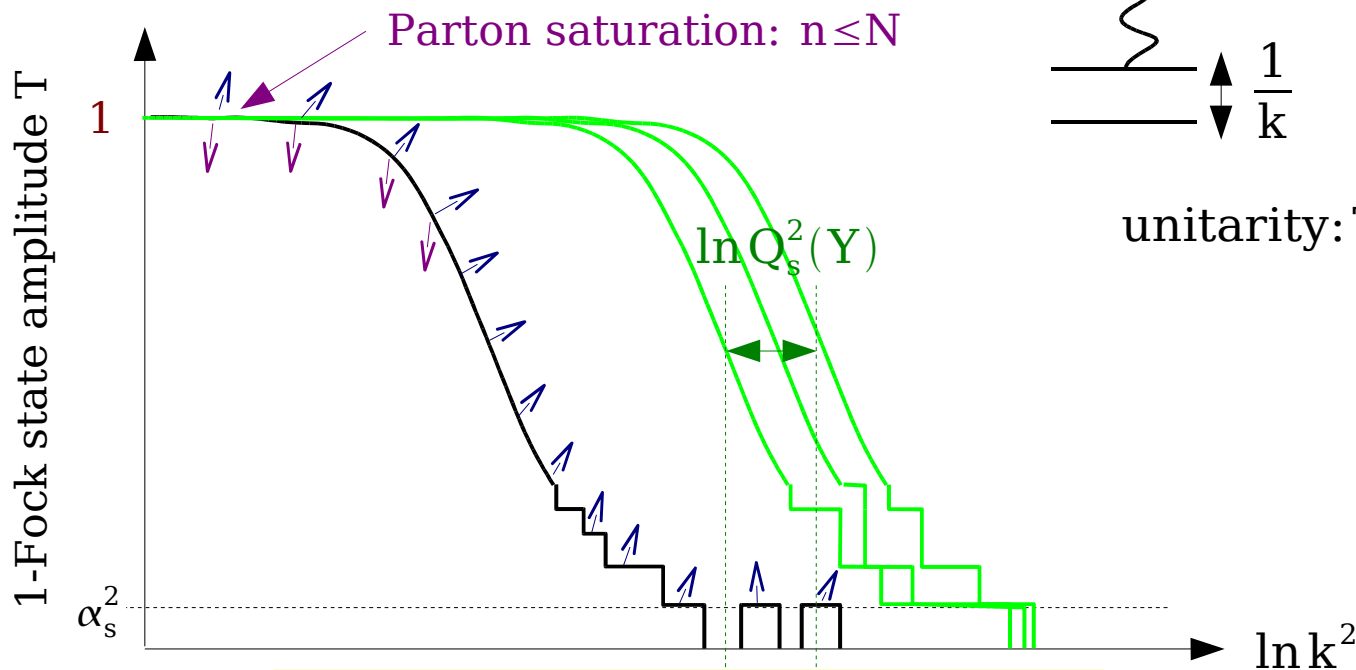
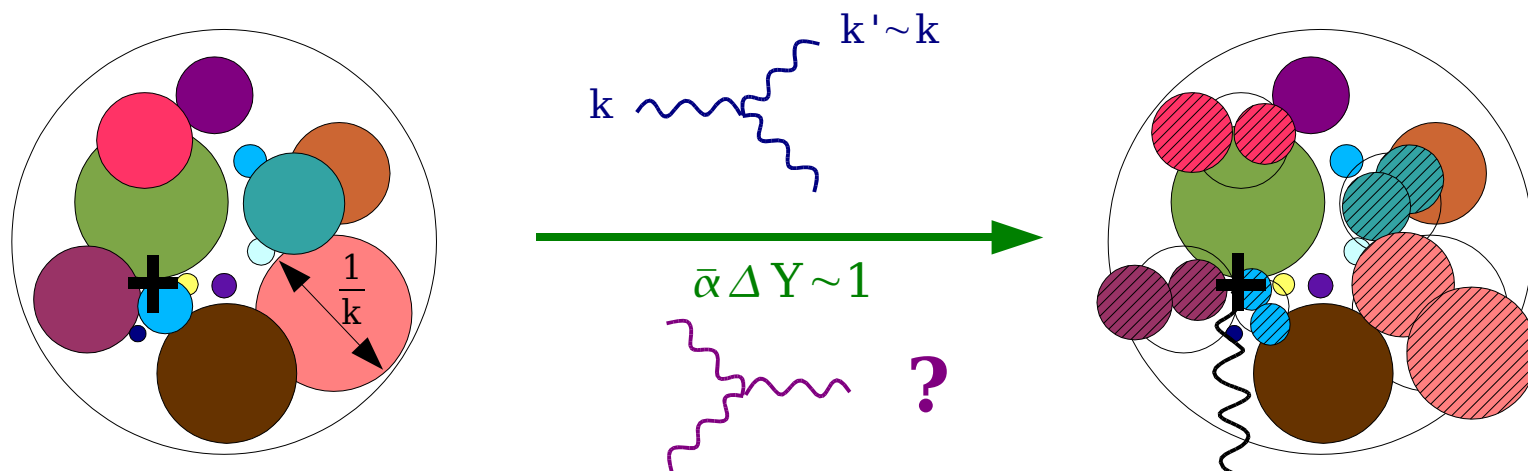
unitarity:  $T \leq 1 \Rightarrow N = \frac{1}{\alpha_s^2}$

$$\partial_{\bar{\alpha} Y} \mathbf{n} = \chi \left( -\partial_{\ln k^2} \right) \mathbf{n} - \frac{\mathbf{n}^2}{N} + \sqrt{\mathbf{n}} \nu$$

$\text{BFKL} \sim \partial_{\ln k^2}^2 \mathbf{n} + \mathbf{n}$

Noise term due to discreteness

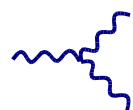
# How a high rapidity hadron looks



$$T(k) \sim \alpha_s^2 n(k)$$

$$\text{unitarity: } T \leq 1 \Rightarrow N = \frac{1}{\alpha_s^2}$$

$$\partial_{\bar{\alpha} Y} T = \chi(-\partial_{\ln k^2}) T - T^2 + \alpha_s \sqrt{T} \nu$$

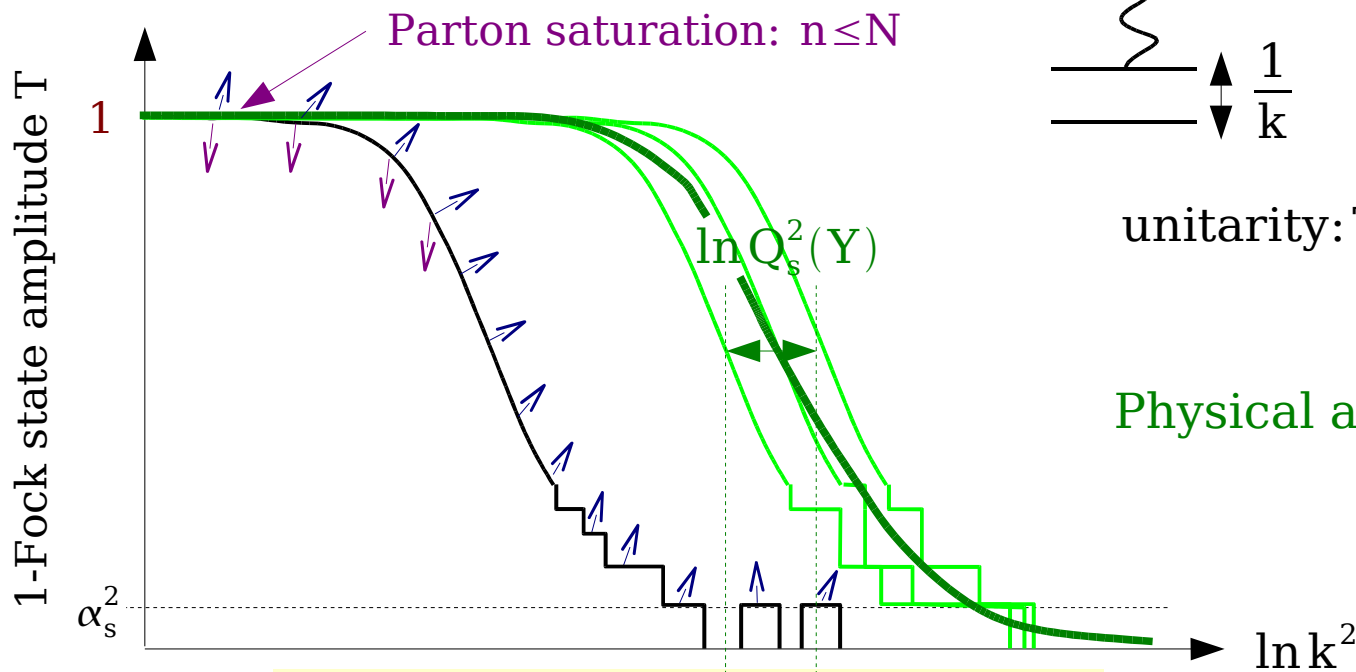
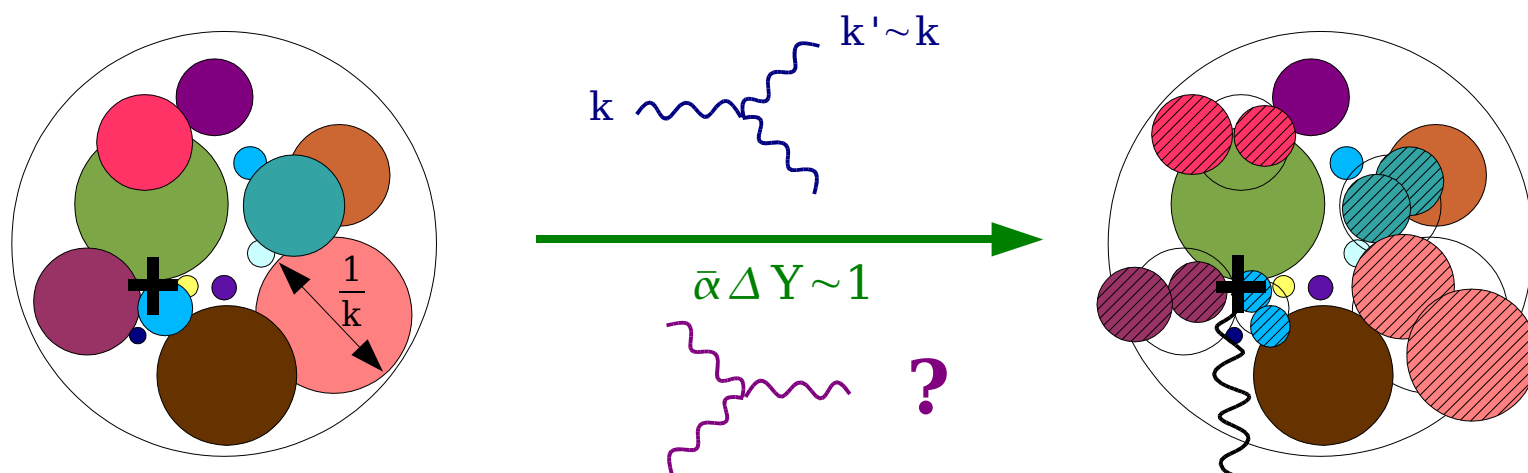


$$\text{BFKL} \sim \partial_{\ln k^2}^2 T + T$$



Noise term due to discreteness

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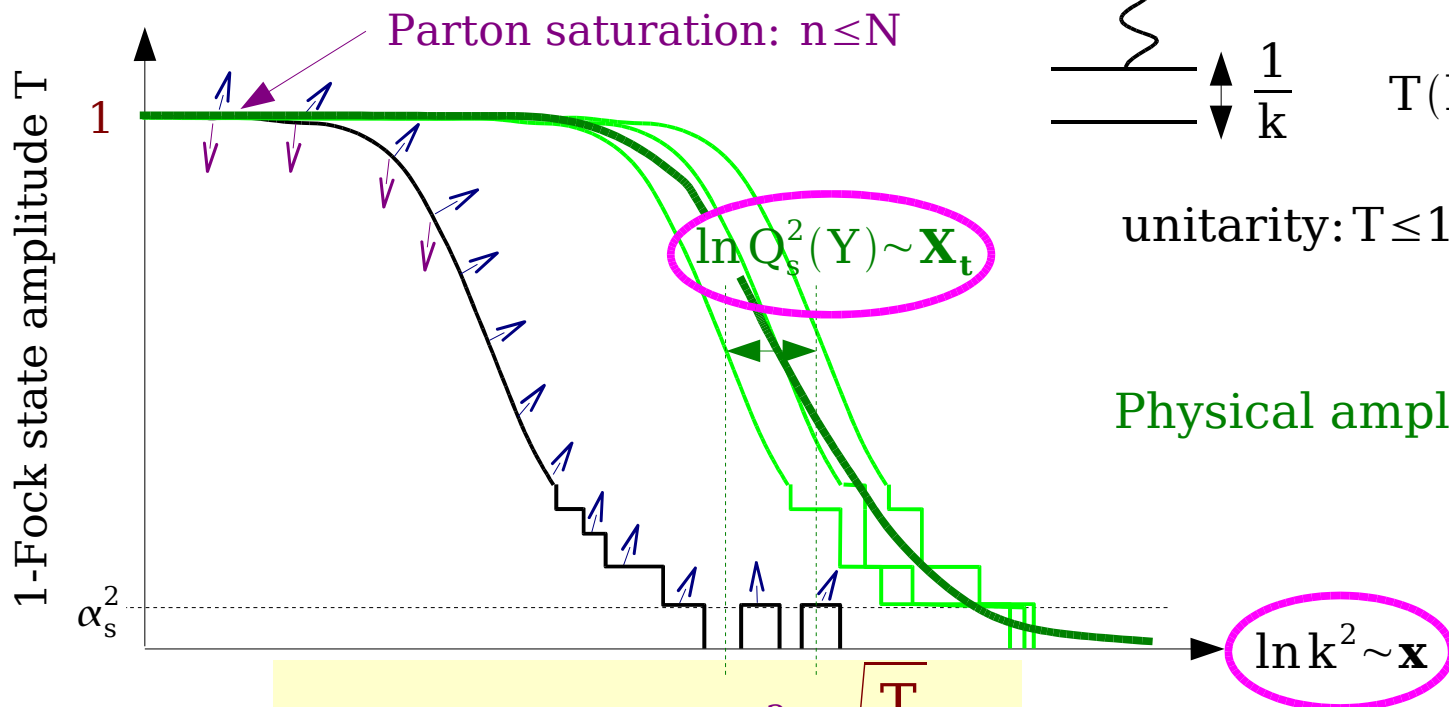
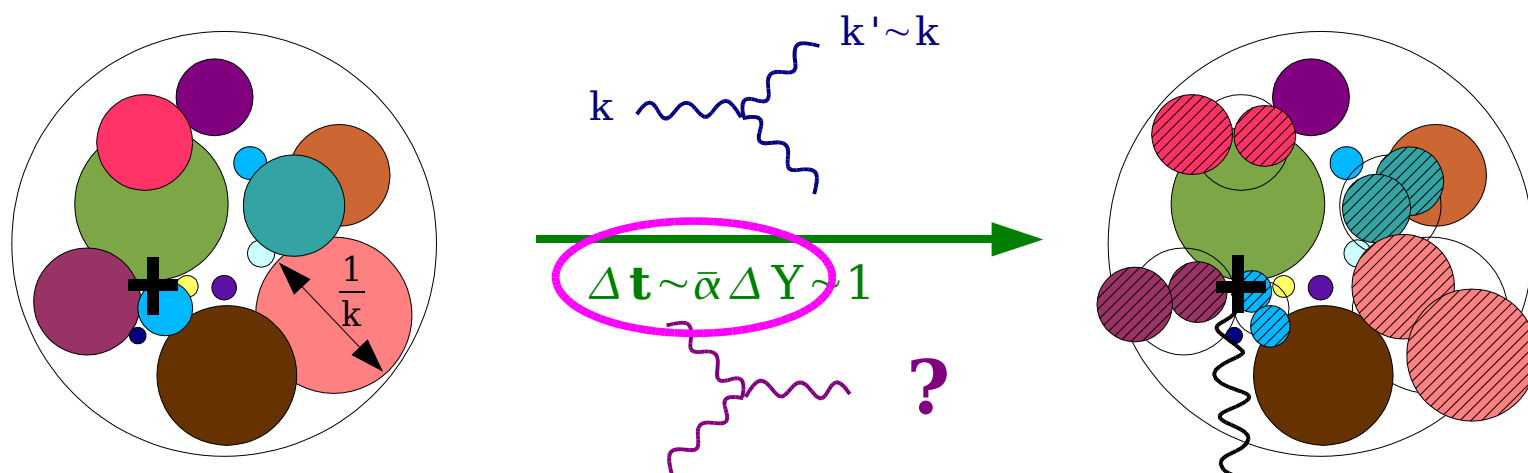
Physical amplitude:  $\langle T \rangle$

$$\partial_{\bar{\alpha} Y} T = \chi(-\partial_{\ln k^2}) T - T^2 + \alpha_s \sqrt{T} \nu$$

$$\text{BFKL} \sim \partial_{\ln k^2}^2 T + T$$

Noise term due to discreteness

# How a high rapidity hadron looks



$$\partial_t T = \chi(-\partial_x) T - T^2 + \sqrt{\frac{T}{N}} \nu$$

diffusive growth  $\sim \partial_x^2 T + T$

Noise term due to discreteness

# Statistics of the position of the front

$$\partial_t \mathbf{T} = \chi(-\partial_x) \mathbf{T} - \mathbf{T}^2 + \sqrt{\frac{\mathbf{T}}{\mathbf{N}}} \nu$$

... Bramson (1986)  
Gribov, Levin, Ryskin (1980)

Brunet, Derrida (1997)  
Mueller, Shoshi (2004)

Brunet, Derrida, Mueller, S.M. (cond-mat/2005, Phys.Rev.E)

$$\langle \mathbf{X}_t \rangle = \left( \frac{\chi(\gamma_0)}{\gamma_0} - \frac{\pi^2 \gamma_0 \chi''(\gamma_0)}{2 \ln^2 \mathbf{N}} \right) t + \left( \pi^2 \gamma_0^2 \chi''(\gamma_0) \frac{3 \ln \ln \mathbf{N}}{\gamma_0 \ln^3 \mathbf{N}} \right) t$$

$$\langle \mathbf{X}_t^n \rangle_{\text{cumulant}} = \pi^2 \gamma_0^2 \chi''(\gamma_0) \frac{\mathbf{n}! \zeta(\mathbf{n})}{\gamma_0^n} \left[ \frac{t}{\ln^3 \mathbf{N}} \right]$$

$(\ln \mathbf{N} \gg 1)$

with  $\frac{\chi(\gamma_0)}{\gamma_0} = \chi'(\gamma_0)$

## Dictionary to QCD

$$t = \bar{\alpha} Y \quad x = \ln k^2$$

$$\mathbf{X}_t = \ln Q_s^2(Y)$$

$\chi = \text{BFKL kernel}$

$$\mathbf{N} = \frac{1}{\alpha_s^2}$$

**Universal results for sFKPP-like equations:** independent of the details, in particular of the saturation mechanism



# Statistics of the position of the front

$$\partial_t T = \chi(-\partial_x) T - T^2 + \sqrt{\frac{T}{N}} v$$

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$$t = \bar{\alpha} Y \quad x = \ln k^2$$

$$X_t = \ln Q_s^2(Y)$$

$\chi$  = BFKL kernel

$$N = \frac{1}{\alpha_s^2}$$

**Universal results for sFKPP-like equations:** independent of the details, in particular of the saturation mechanism

The cumulants are all

- of the same order: long tail distribution!
- proportional to  $\kappa = \frac{t}{\ln^3 N}$

$X_t$  is the sum of  $\kappa$  independent random variables

# Outline

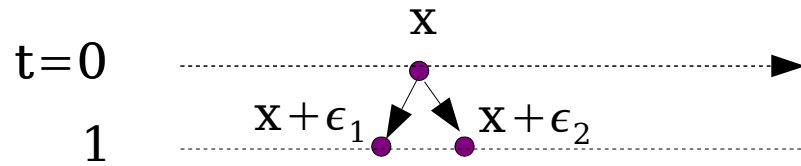
- Why the stochastic FKPP equation is relevant to high energy QCD

- A new perspective on its solutions:

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Brunet, Derrida, Mueller, S.M., Letter on cond-mat/2006  
Extended version to appear

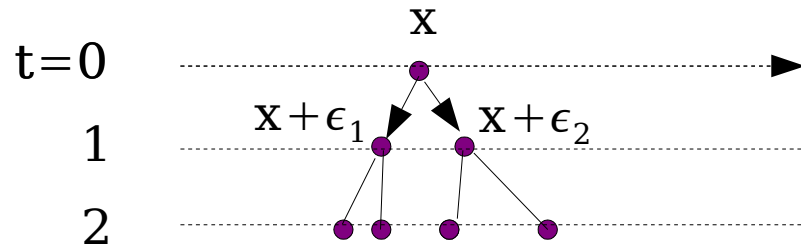
# Selective population evolution



The individual at position  $x$  has  
2 descendants at positions  $x + \epsilon_1$  and  $x + \epsilon_2$

distribution  $\rho(\epsilon)$

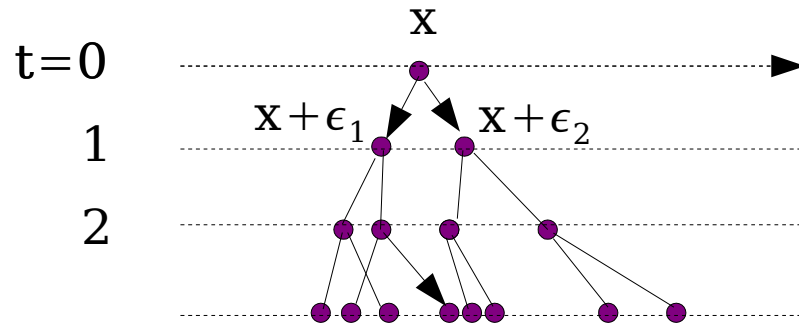
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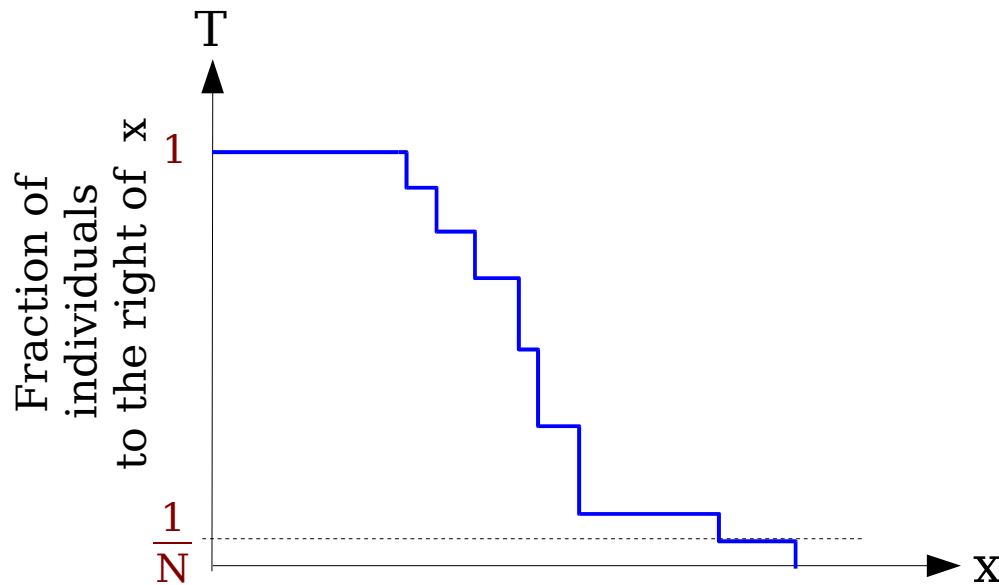
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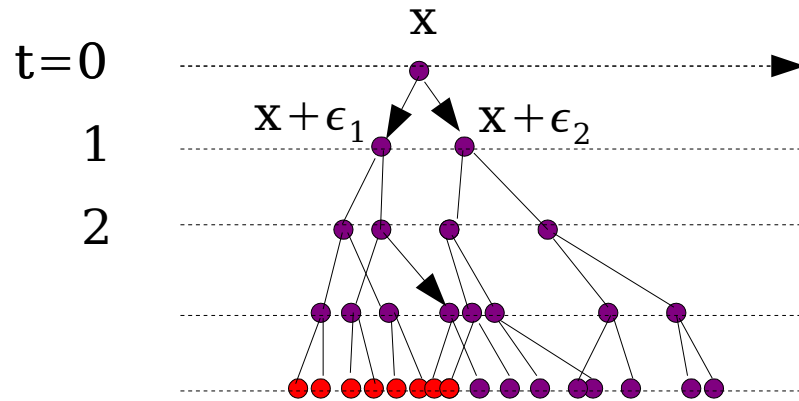
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Keep only the **N=8 rightmost** individuals



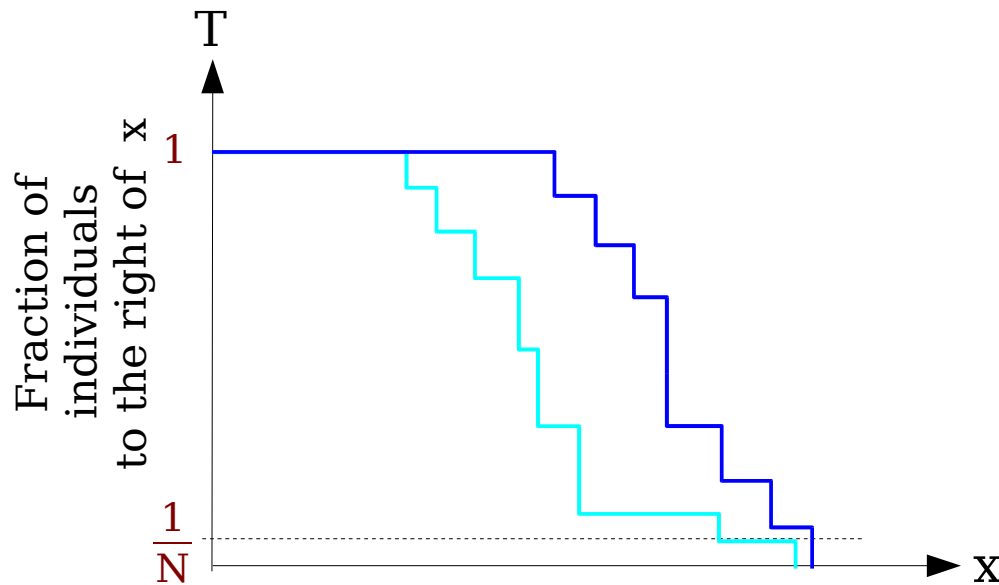
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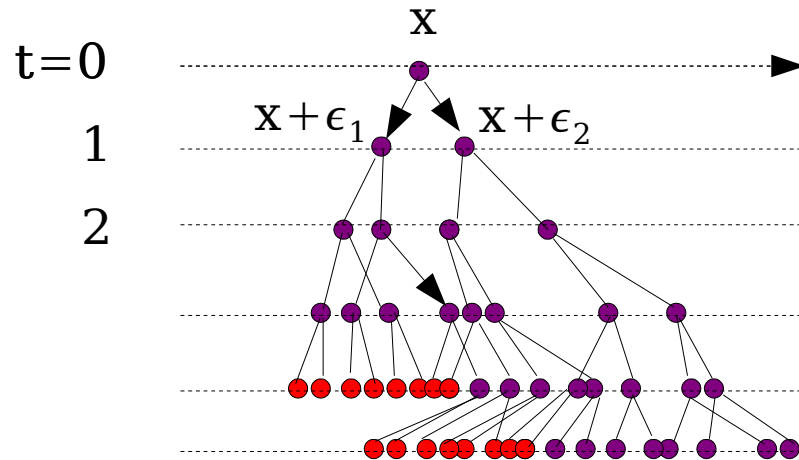
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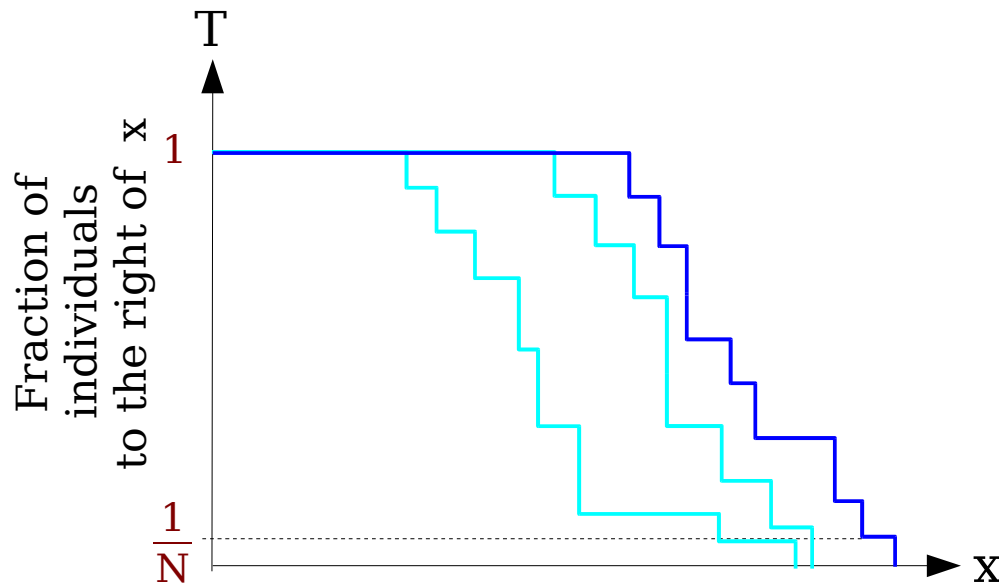
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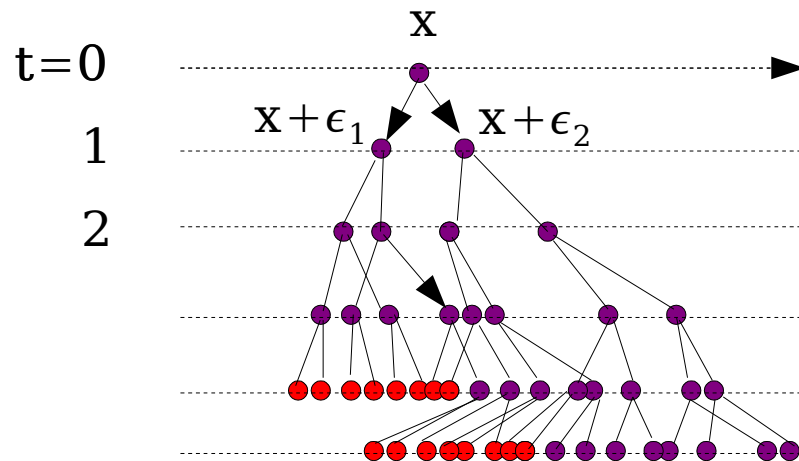
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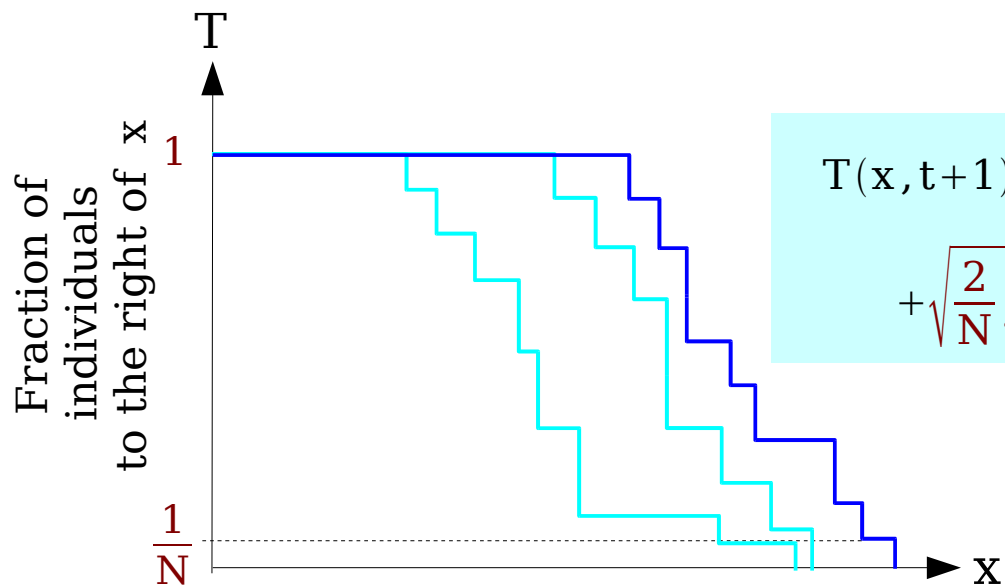
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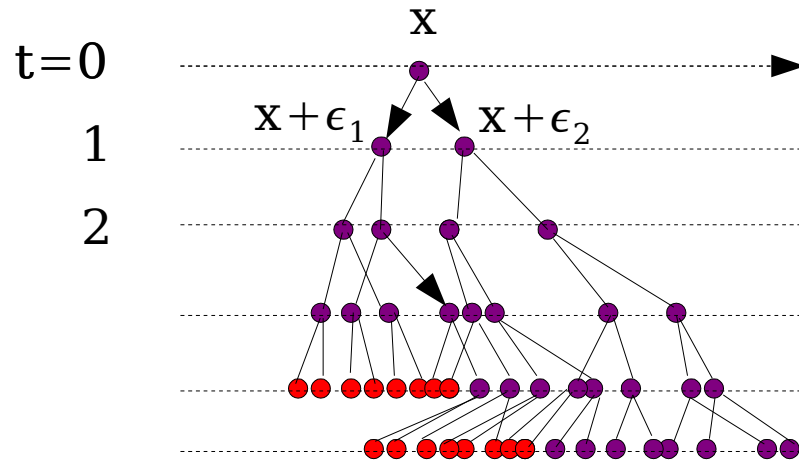
Keep only the **N=8 rightmost** individuals



$$T(x, t+1) = \text{Min} \left( \frac{1}{2 \int d\epsilon \rho(\epsilon) T(x-\epsilon, t)} \right) + \sqrt{\frac{2}{N} \int d\epsilon \rho(\epsilon) T(x-\epsilon, t) (1 - 2 \int d\epsilon \rho(\epsilon) T(x-\epsilon, t))} v(x, t+1)$$



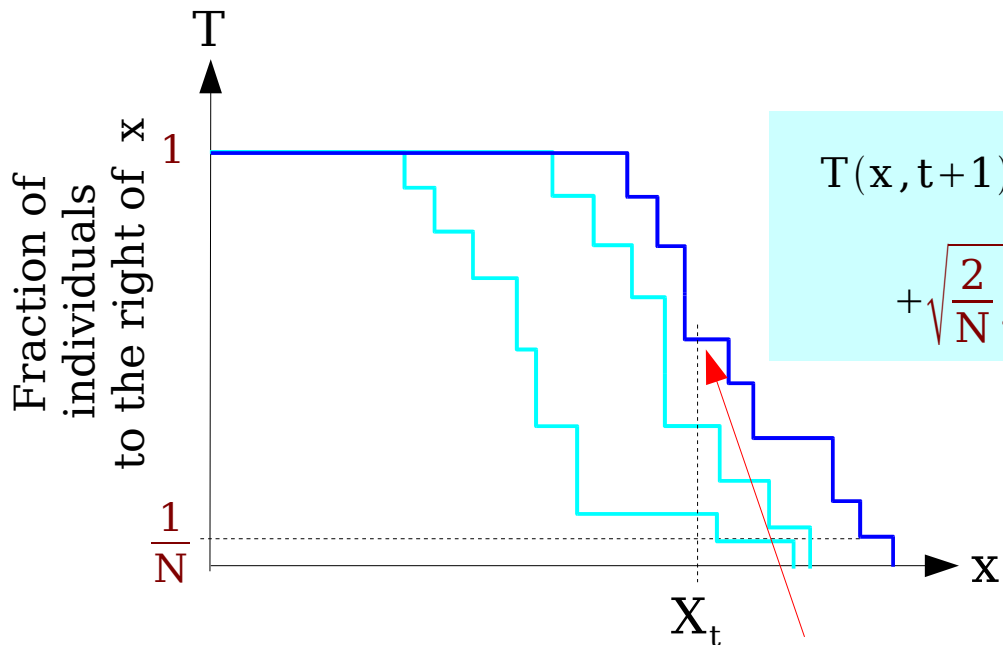
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distribution  $\rho(\epsilon)$

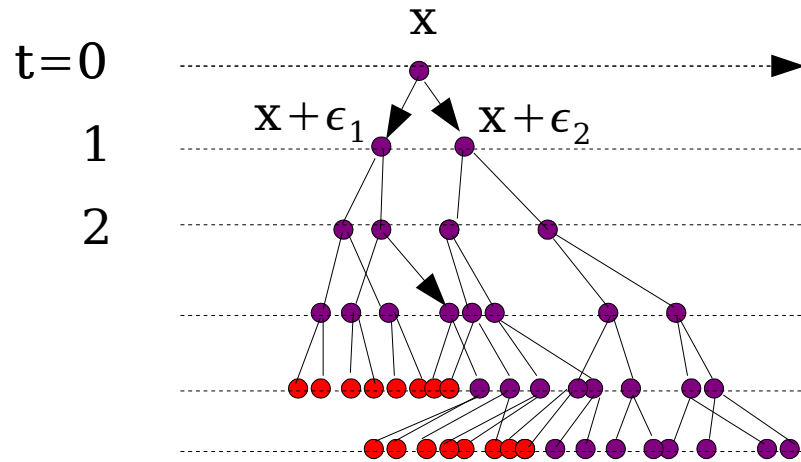
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$$T(x, t+1) = \text{Min} \left( \frac{1}{2 \int d\epsilon \rho(\epsilon) T(x-\epsilon, t)} + \sqrt{\frac{2}{N} \int d\epsilon \rho(\epsilon) T(x-\epsilon, t) (1 - 2 \int d\epsilon \rho(\epsilon) T(x-\epsilon, t))} v(x, t+1) \right)$$

**Traveling wave:** all universal results on the front are applicable with  $\chi(\gamma) = \ln \left( 2 \int d\epsilon \rho(\epsilon) e^{\gamma \epsilon} \right)$

# Selective population evolution

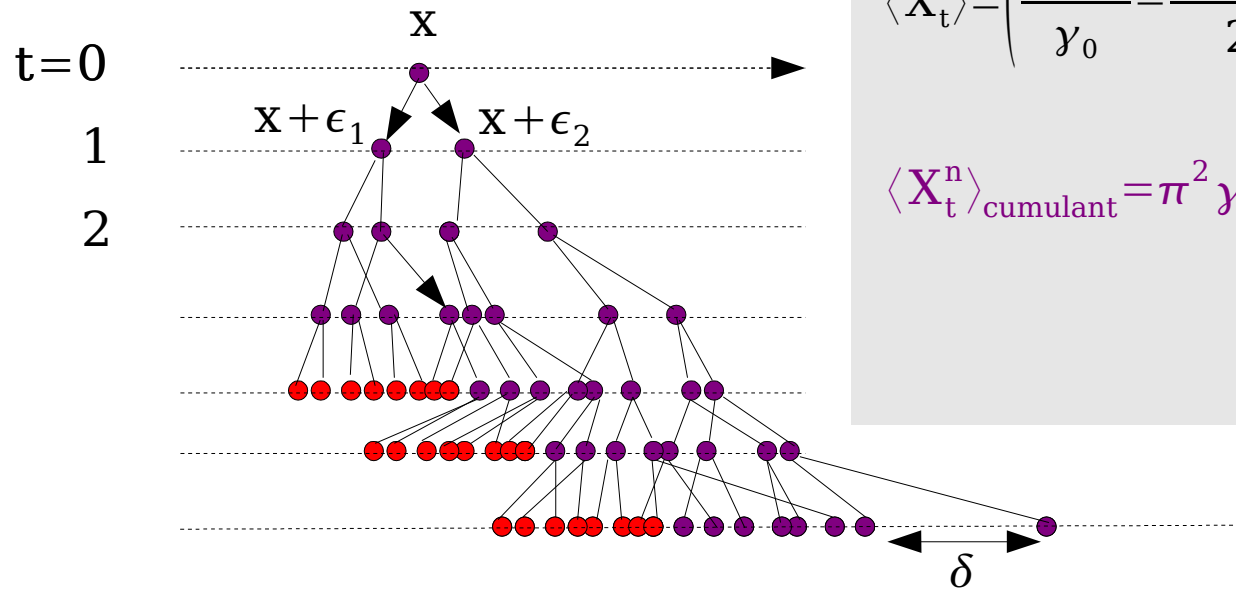


$$\langle X_t \rangle = \left( \frac{\chi(\gamma_0)}{\gamma_0} - \frac{\pi^2 \gamma_0 \chi''(\gamma_0)}{2 \ln^2 N} \right) t + \left( \pi^2 \gamma_0^2 \chi''(\gamma_0) \frac{3 \ln \ln N}{\gamma_0 \ln^3 N} \right) t$$

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$$\text{with } \frac{\chi(\gamma_0)}{\gamma_0} = \chi'(\gamma_0)$$

# Selective population evolution

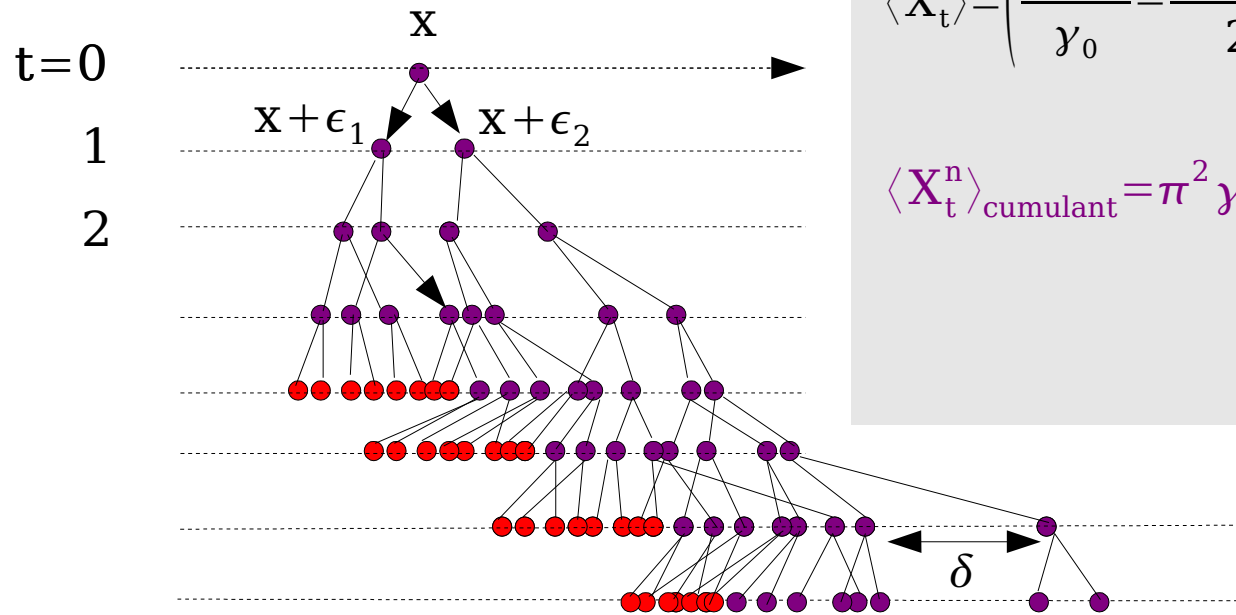


$$\langle X_t \rangle = \left( \frac{x(\gamma_0)}{\gamma_0} - \frac{\pi^2 \gamma_0 x''(\gamma_0)}{2 \ln^2 N} \right) t + \left( \pi^2 \gamma_0^2 x''(\gamma_0) \frac{3 \ln \ln N}{\gamma_0 \ln^3 N} \right) t$$

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# Selective population evolution

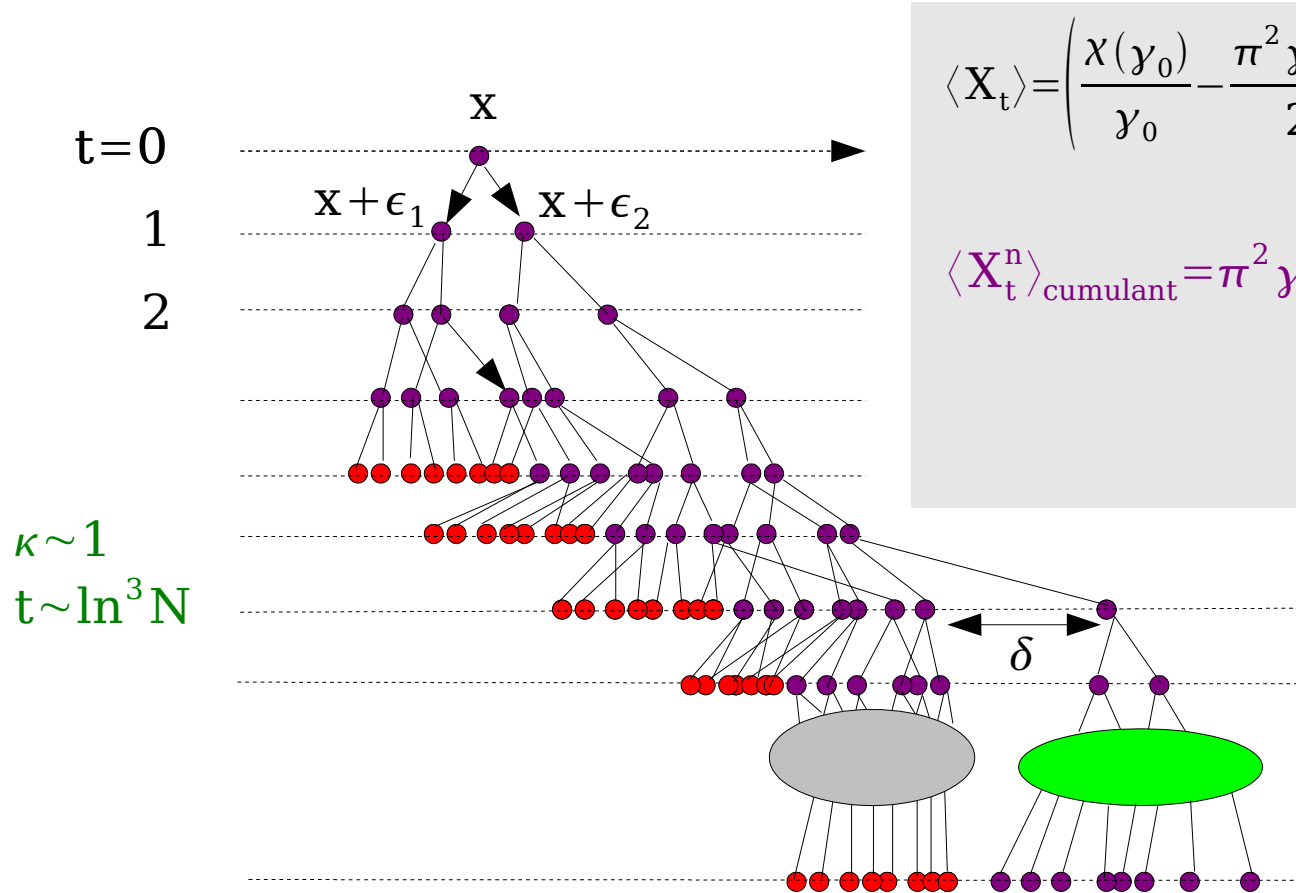


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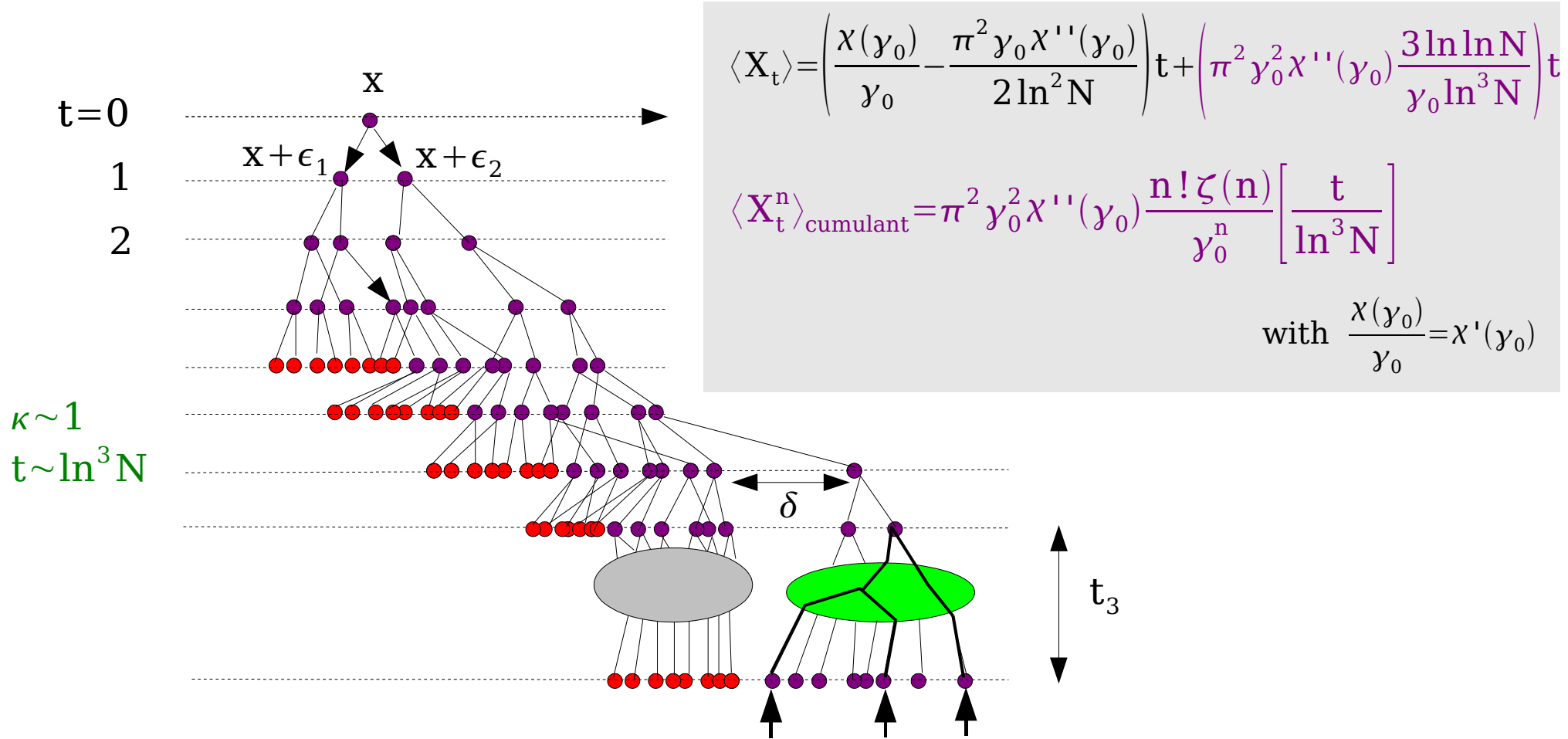


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$$\langle X_t^n \rangle_{\text{cumulant}} = \pi^2 \gamma_0^2 \chi''(\gamma_0) \frac{n! \zeta(n)}{\gamma_0^n} \left[ \frac{t}{\ln^3 N} \right]$$

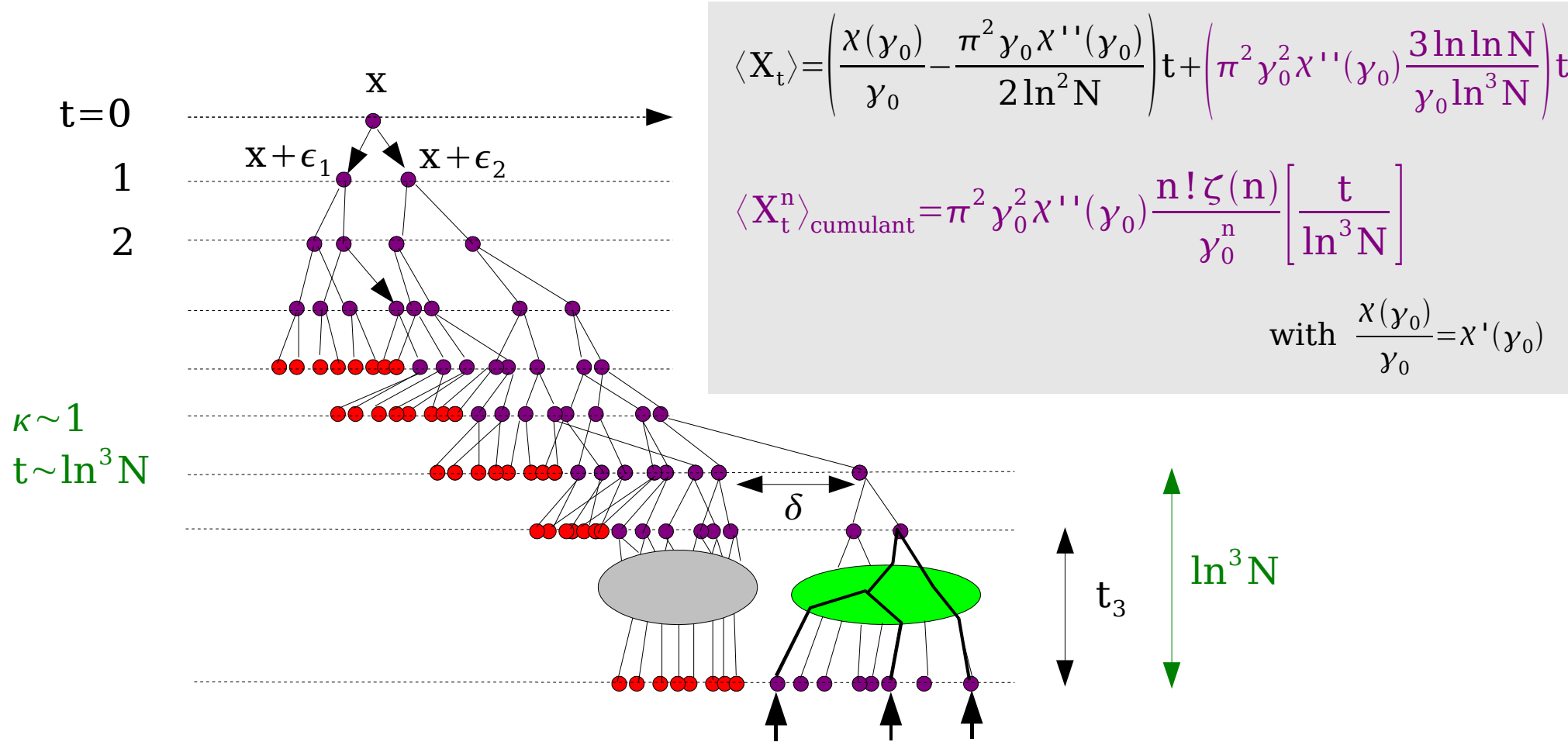
$$\text{with } \frac{\chi(\gamma_0)}{\gamma_0} = \chi'(\gamma_0)$$

# Selective population evolution



**Average number of generations to the first common ancestor of  $k=2, 3, \dots$  randomly chosen individuals?**

# Selective population evolution



**Average number of generations to the first common ancestor of  $k=2, 3, \dots$  randomly chosen individuals?**

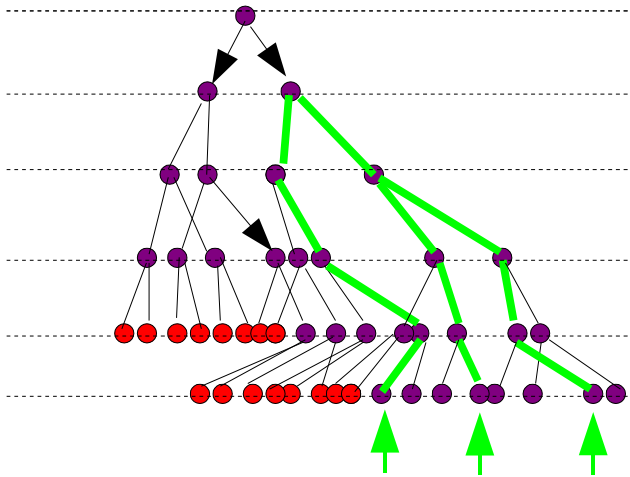
$$\langle t_k \rangle \sim \ln^3 N$$

$$\frac{\langle t_3 \rangle}{\langle t_2 \rangle} = \frac{5}{4}$$

$$\frac{\langle t_4 \rangle}{\langle t_2 \rangle} = \frac{25}{18}$$

# Selective population evolution

With selection



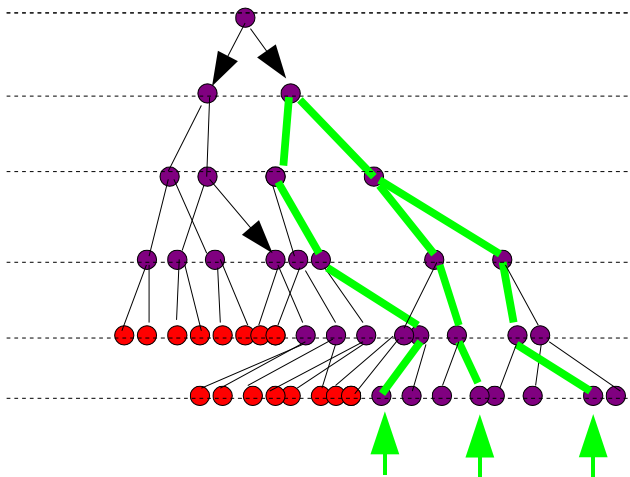
$$\langle t_k \rangle \sim \ln^3 N$$

$$\frac{\langle t_3 \rangle}{\langle t_2 \rangle} = \frac{5}{4} \quad \frac{\langle t_4 \rangle}{\langle t_2 \rangle} = \frac{25}{18}$$



# Selective population evolution

## With selection

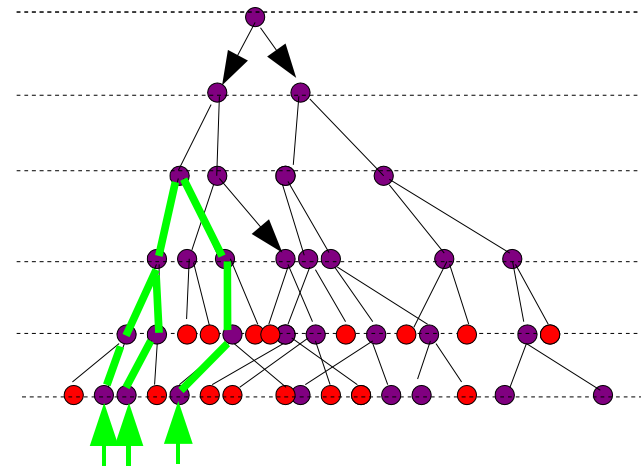


$$\langle t_k \rangle \sim \ln^3 N$$

$$\frac{\langle t_3 \rangle}{\langle t_2 \rangle} = \frac{5}{4}$$

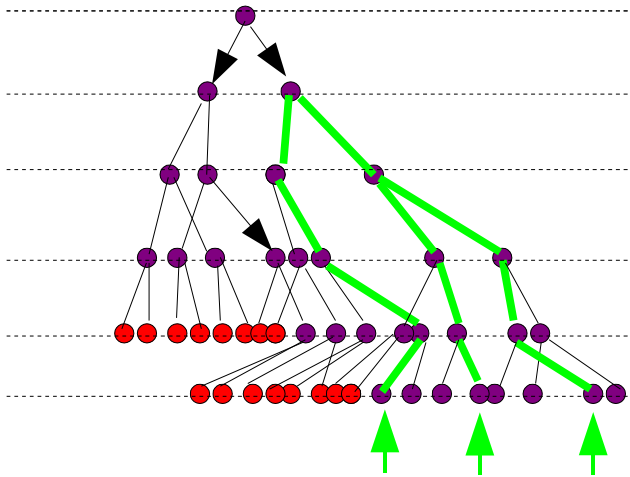
$$\frac{\langle t_4 \rangle}{\langle t_2 \rangle} = \frac{25}{18}$$

## Without selection



# Selective population evolution

With selection

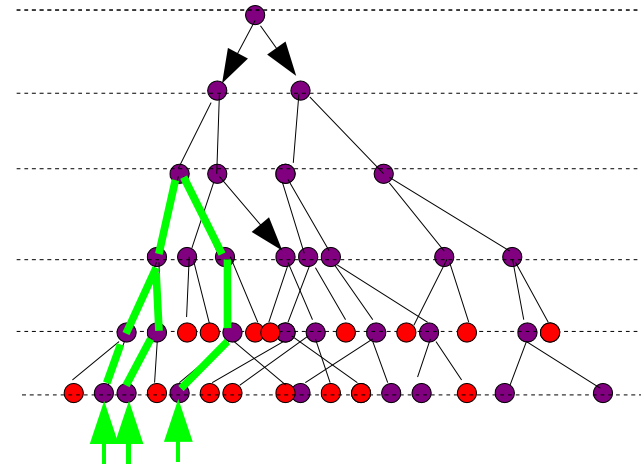


$$\langle t_k \rangle \sim \ln^3 N$$

$$\frac{\langle t_3 \rangle}{\langle t_2 \rangle} = \frac{5}{4}$$

$$\frac{\langle t_4 \rangle}{\langle t_2 \rangle} = \frac{25}{18}$$

Without selection



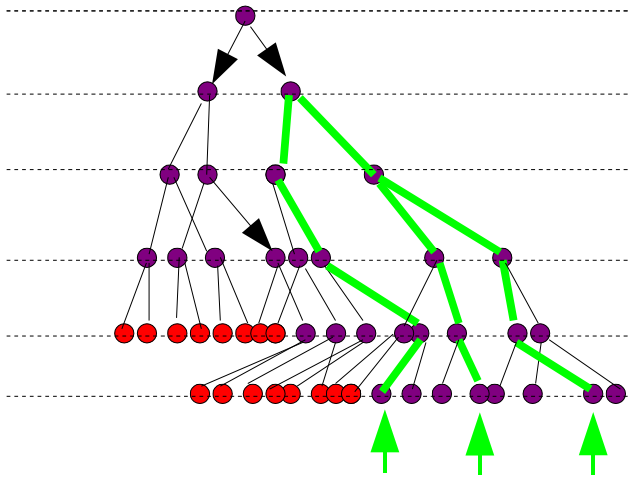
$$\langle t_k \rangle \sim N$$

$$\frac{\langle t_3 \rangle}{\langle t_2 \rangle} = \frac{4}{3}$$

$$\frac{\langle t_4 \rangle}{\langle t_2 \rangle} = \frac{3}{2}$$

# Selective population evolution

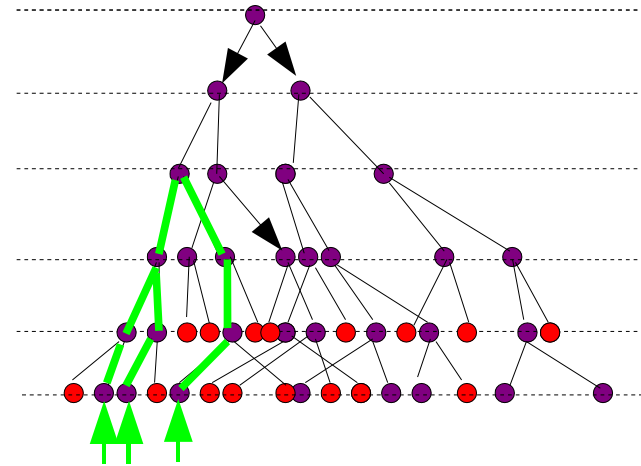
With selection



$$\langle t_k \rangle \sim \ln^3 N$$

$$\frac{\langle t_3 \rangle}{\langle t_2 \rangle} = \frac{5}{4} \quad \frac{\langle t_4 \rangle}{\langle t_2 \rangle} = \frac{25}{18}$$

Without selection



$$\langle t_k \rangle \sim N$$

$$\frac{\langle t_3 \rangle}{\langle t_2 \rangle} = \frac{4}{3} \quad \frac{\langle t_4 \rangle}{\langle t_2 \rangle} = \frac{3}{2}$$

Same statistics as Parisi's trees  
that appear in the **theory of spin glasses!**

# Summary

Reaction-diffusion  
Selective evolution

New results for sFKPP, i.e.:  
- stochastic front propagation  
- properties of population growth

Belongs to the universality class of

Small coupling:  $\alpha_s < 0.1$

High energy QCD

New insights in high energy evolution;  
simple understanding of the fluctuations

New results for QCD amplitudes  $\ln(1/\alpha_s^2) \gg 1$

Trees appear in both contexts  
and look exactly the same!

Spin glasses