

THE DEVELOPMENT OF X-RAY ASTRONOMY

PROF. RICCARDO GIACCONI

Presented at Symposium

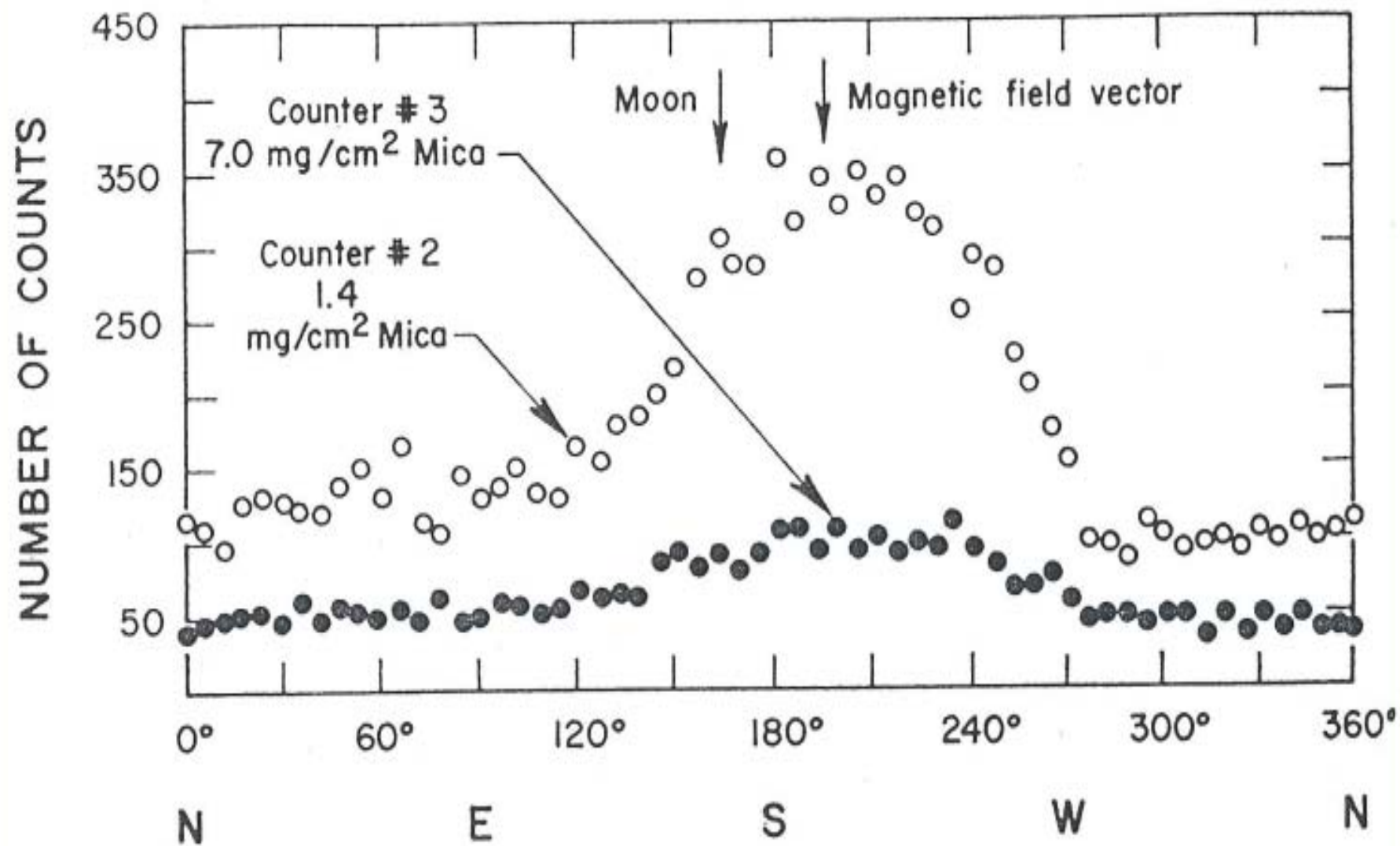
FOUR YEARS OF CHANDRA OBSERVATIONS

September 16, 2003

GIACCONI, CLARK, ROSSI (1960)

SUN	< 20 Å	CORONAL EMISSION	$\sim 10^6 \text{ CM}^2 \text{ S}^{-1}$
SUN AT 8 LIGHT YEARS	< 20 Å	CORONAL EMISSION	$2.5 \times 10^4 \text{ CM}^2 \text{ S}^{-1}$
SIRIUS IF $L_X \sim L_{\text{OPT}}$	< 20 Å	? NO CONVECTIVE ZONE	$0.25 \text{ CM}^2 \text{ S}^{-1}$
FLARE STARS	< 20 Å	SUNLIKE FLARE?	?
PECULIAR A STARS	< 20 Å	$B \sim 10^4$ GAUSS LARGE B PARTICLE ACCELERATION	?
CRAB NEBULA	< 25 Å	SYNCHROTRON $E_e ? 10^{13} \text{ eV}$ IN $B = 10^4$ GAUSS LIFETIMES?	?
MOON	< 23 Å	FLUORESCENCE	$0.4 \text{ CM}^2 \text{ S}^{-1}$
MOON	~ 20 Å	IMPACT FROM SOLAR WIND ELECTRONS $\Phi_e = 0-10^{13} \text{ CM}^2 \text{ S}^{-1}$	$0-1.6 \times 10^3 \text{ CM}^2 \text{ S}^{-1}$
SCO X-1	2-8 Å	?	$28 \pm 1.2 \text{ CM}^2 \text{ S}^{-1}$

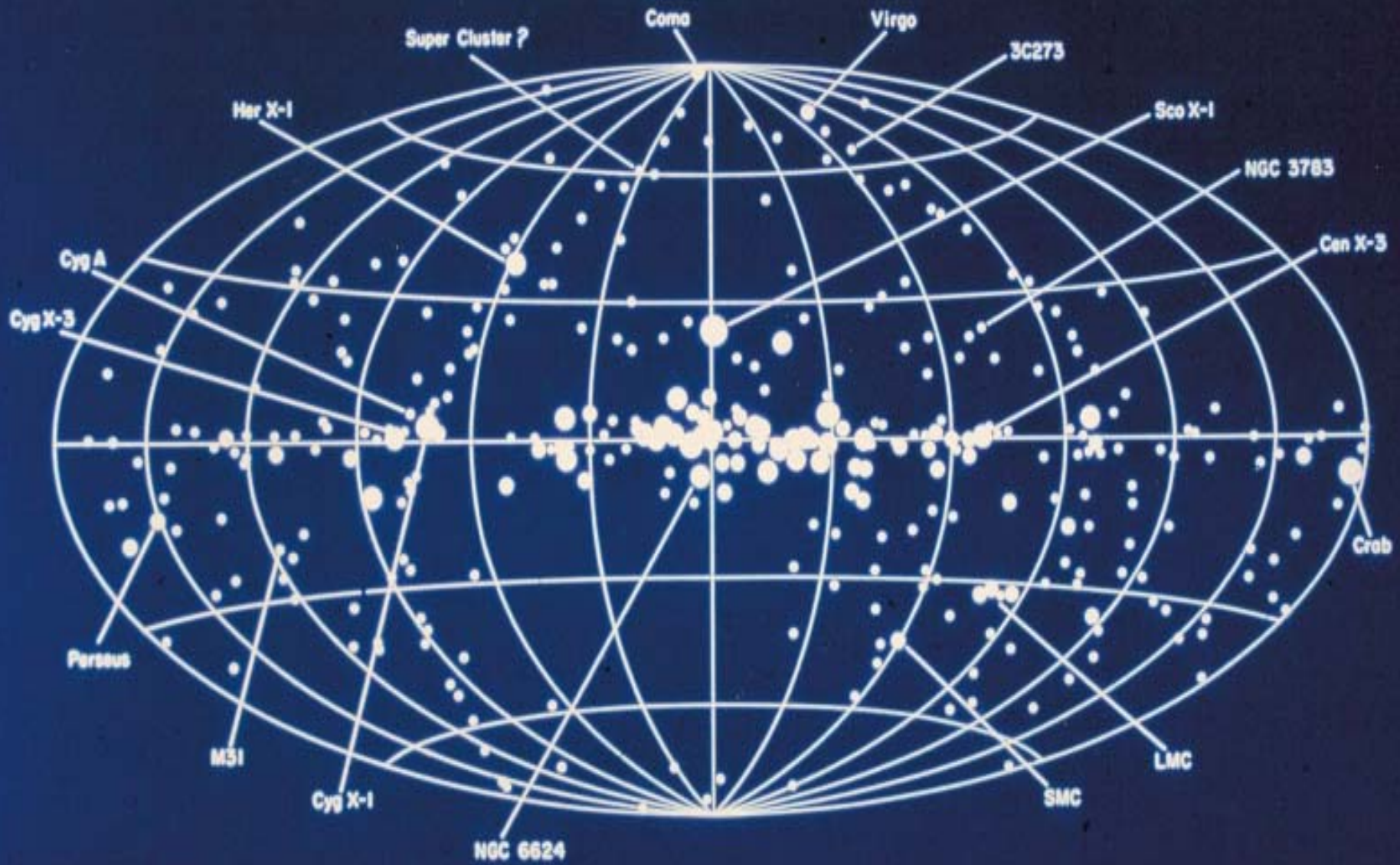
ScoX-1



UHURU satellite (artist's image)



THE FOURTH UHURU CATALOG



Hercules X-1 artist concept





THE DISCOVERY OF X-RAY BINARY SYSTEMS

- \$ EXISTENCE OF BINARY STELLAR SYSTEMS CONTAINING A NEUTRON STAR OR A BLACK HOLE**
- \$ EXISTENCE OF BLACK HOLES OF STELLAR MASS**
- \$ MEASURE OF THE MASS, RADIUS, MOMENT OF INERTIA AND EQUATION OF STATE FOR NEUTRON STARS (DENSITY 10^{15} GR/CM³)**
- \$ A NEW SOURCE OF ENERGY DUE TO GRAVITATIONAL INFALL (100 TIMES MORE EFFICIENT PER NUCLEON THAN FUSION)**
- \$ A MODEL (GENERALLY ACCEPTED) FOR THE NUCLEUS OF ACTIVE GALAXIES AND QUASARS**

Declination (2000.0)



THE DISCOVERY OF X-RAY EMISSION FROM CLUSTERS OF GALAXIES

- **EXISTENCE OF GAS**

$$T = 10^7 - 10^8 \text{ K}$$

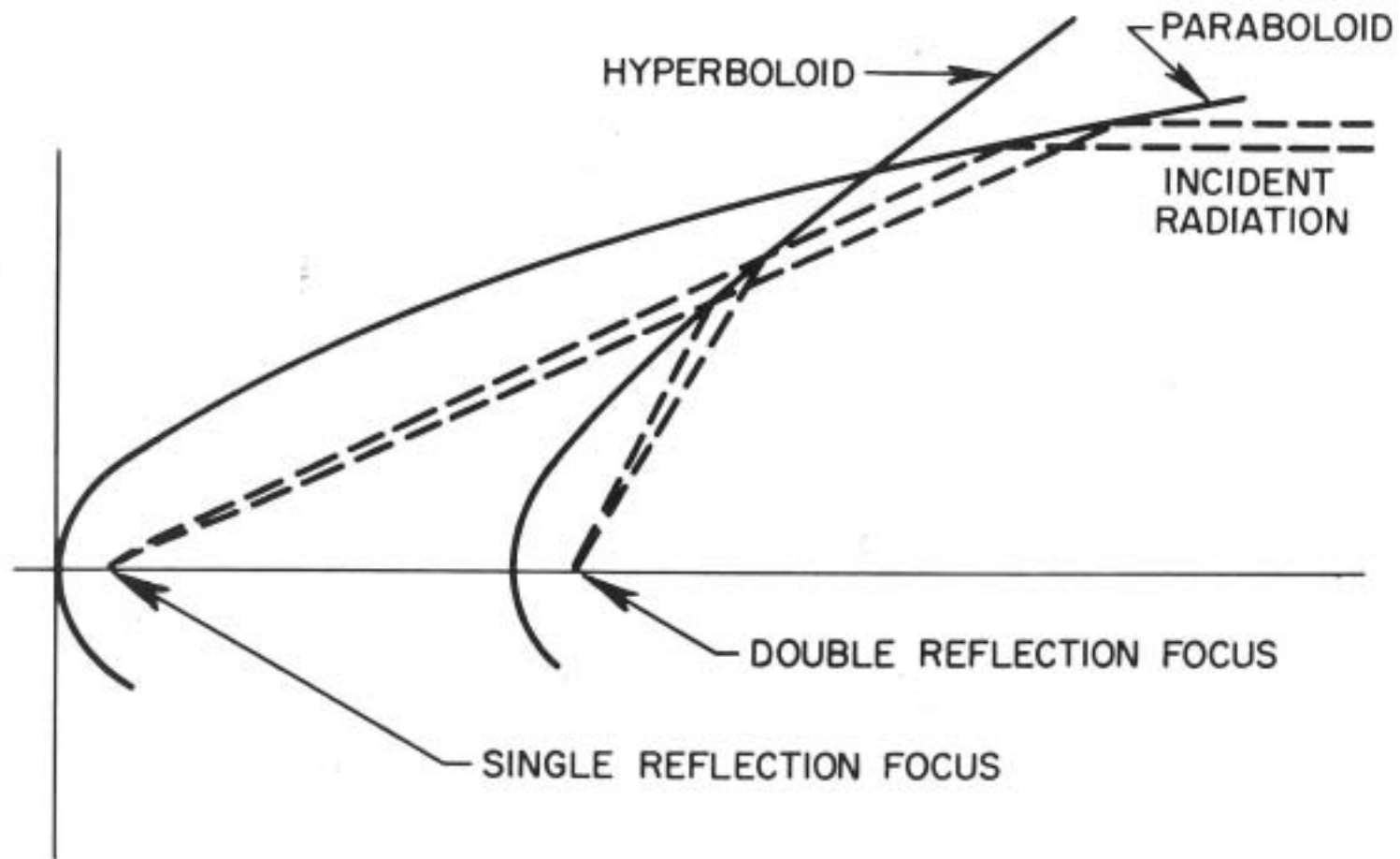
$$L_X = 10^{42} - 10^{45} \text{ ERG S}^{-1}$$

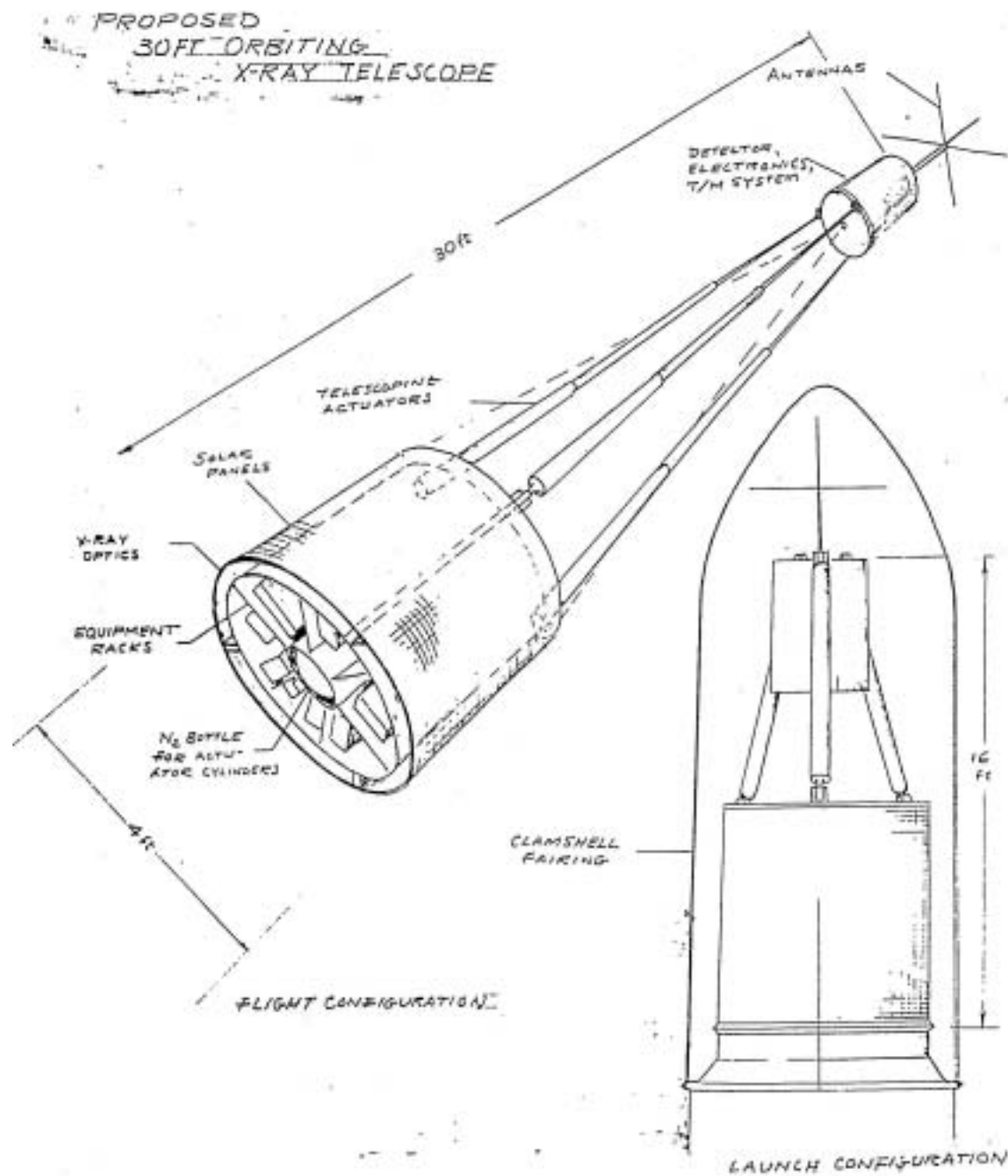
$$\text{DENSITY } 10^{-2} - 10^{-3} \text{ CM}^{-3}$$

$$\text{MASS IN GAS} \sim \text{MASS IN GALAXIES}$$

- **STRUCTURE AND DISTRIBUTION OF MASS IN CLUSTERS**
- **MANY CLUSTERS NOW FORMING – SUBSTRUCTURES = A YOUNG UNIVERSE**
- **TRACERS OF LARGE SCALE STRUCTURE**
- **FORMATION, EVOLUTION AND DYNAMIC DEVELOPMENT OF CLUSTERS;
SPECTRUM OF INITIAL FLUCTUATIONS; CHEMICAL EVOLUTION**
- **H₀ THROUGH SUNAYEV-ZELDOVICH EFFECT
(INVERSE COMPTON ON 3K RADIATION)**

FOCUSSING X-RAY TELESCOPE



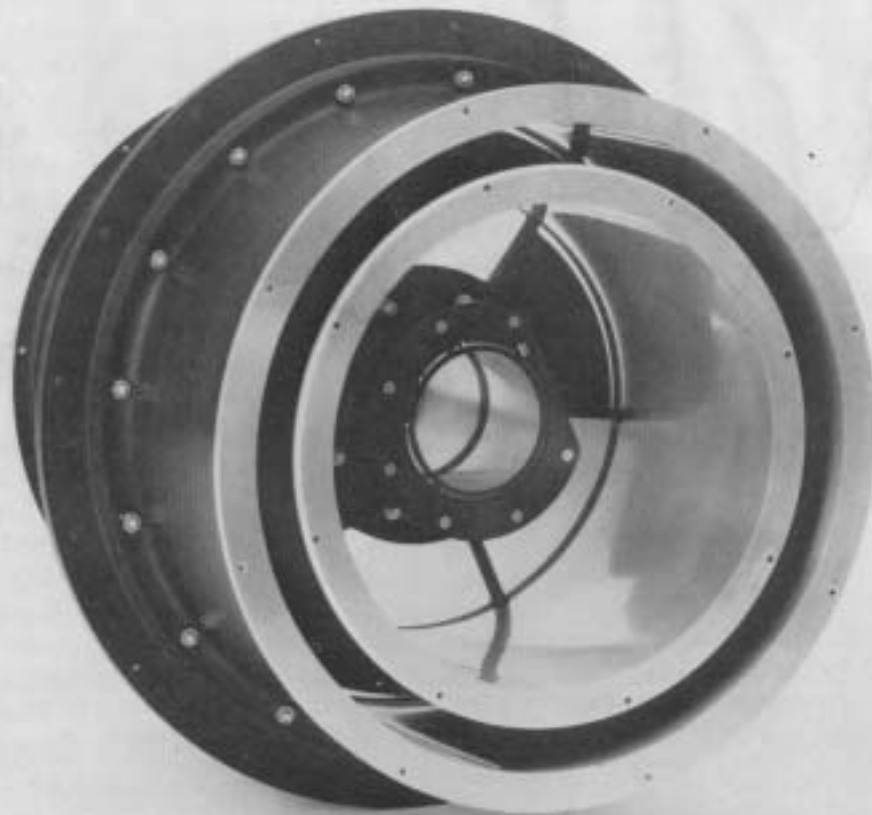


X-RAY TELESCOPE



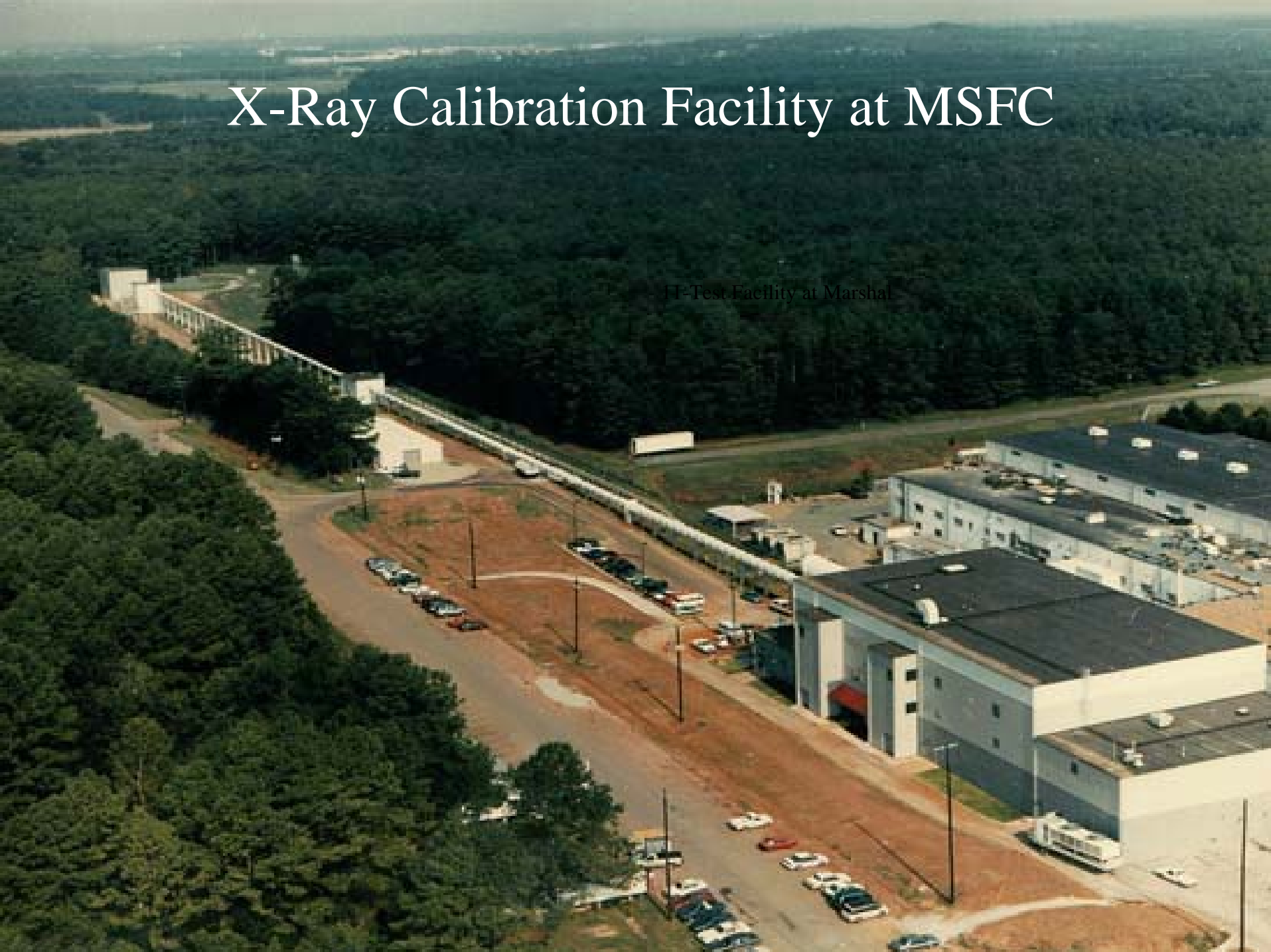


So54 Sky Lab mirror



X-Ray Calibration Facility at MSFC

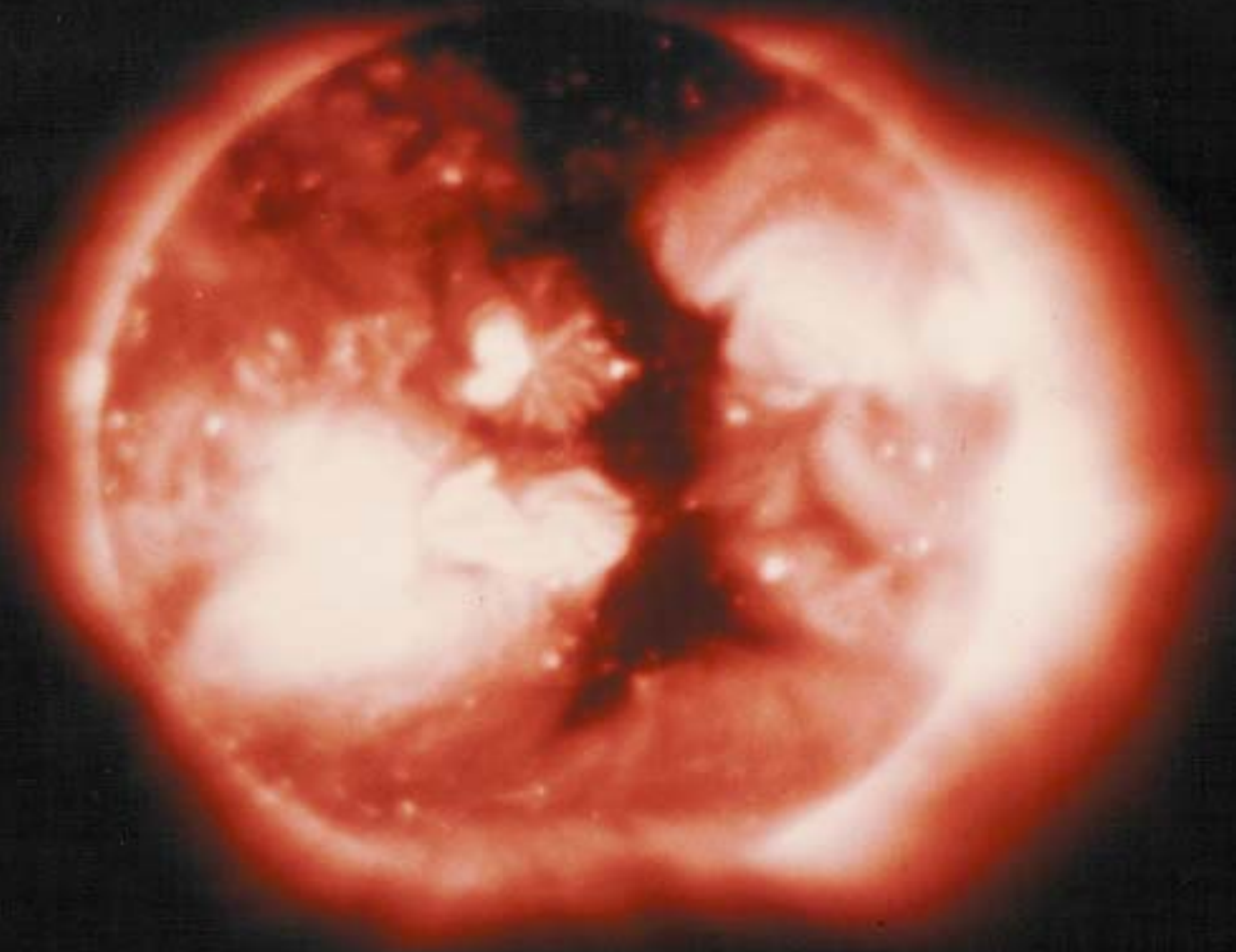
H-Test Facility at Marshall



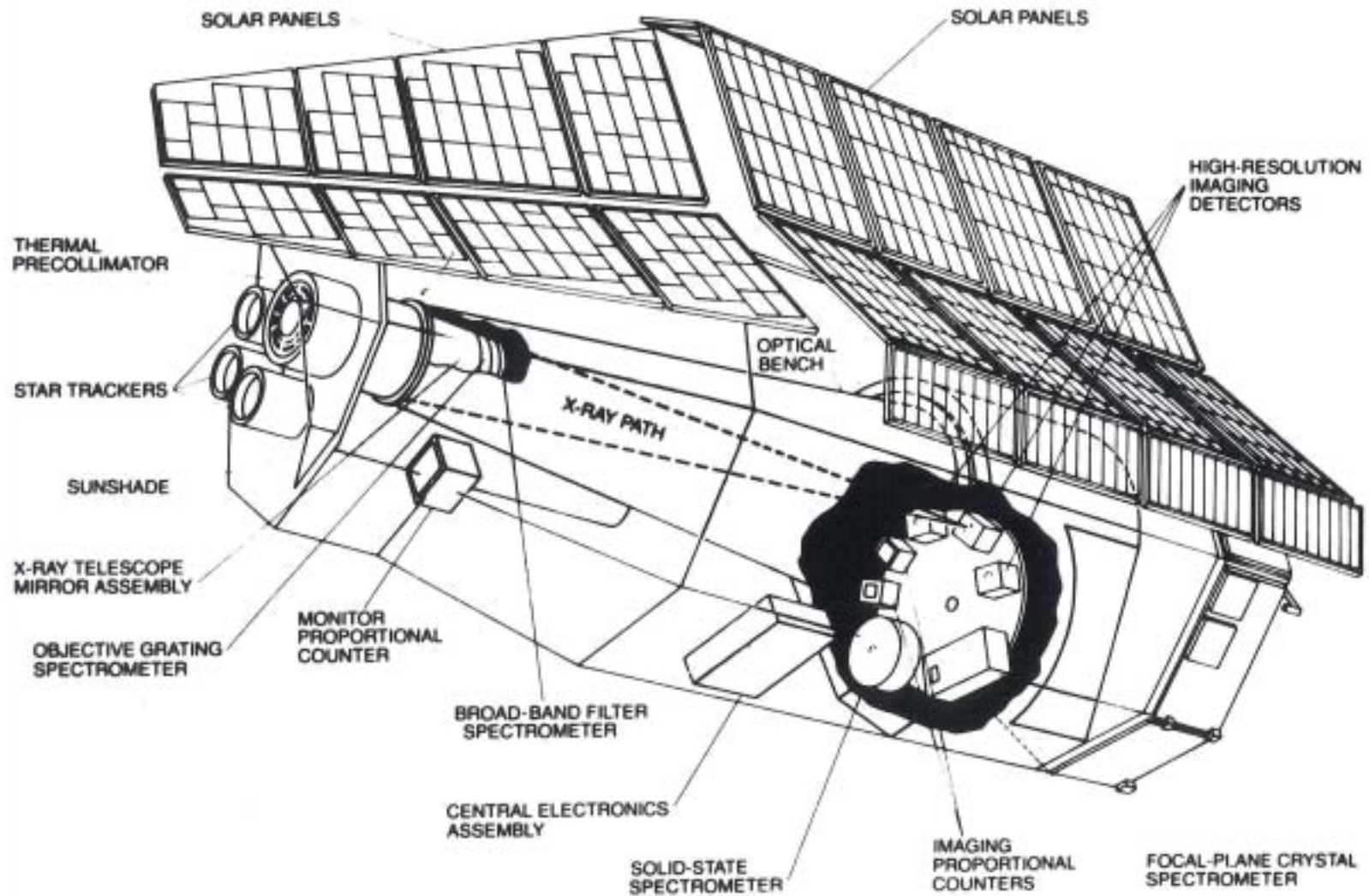
Sky Lab



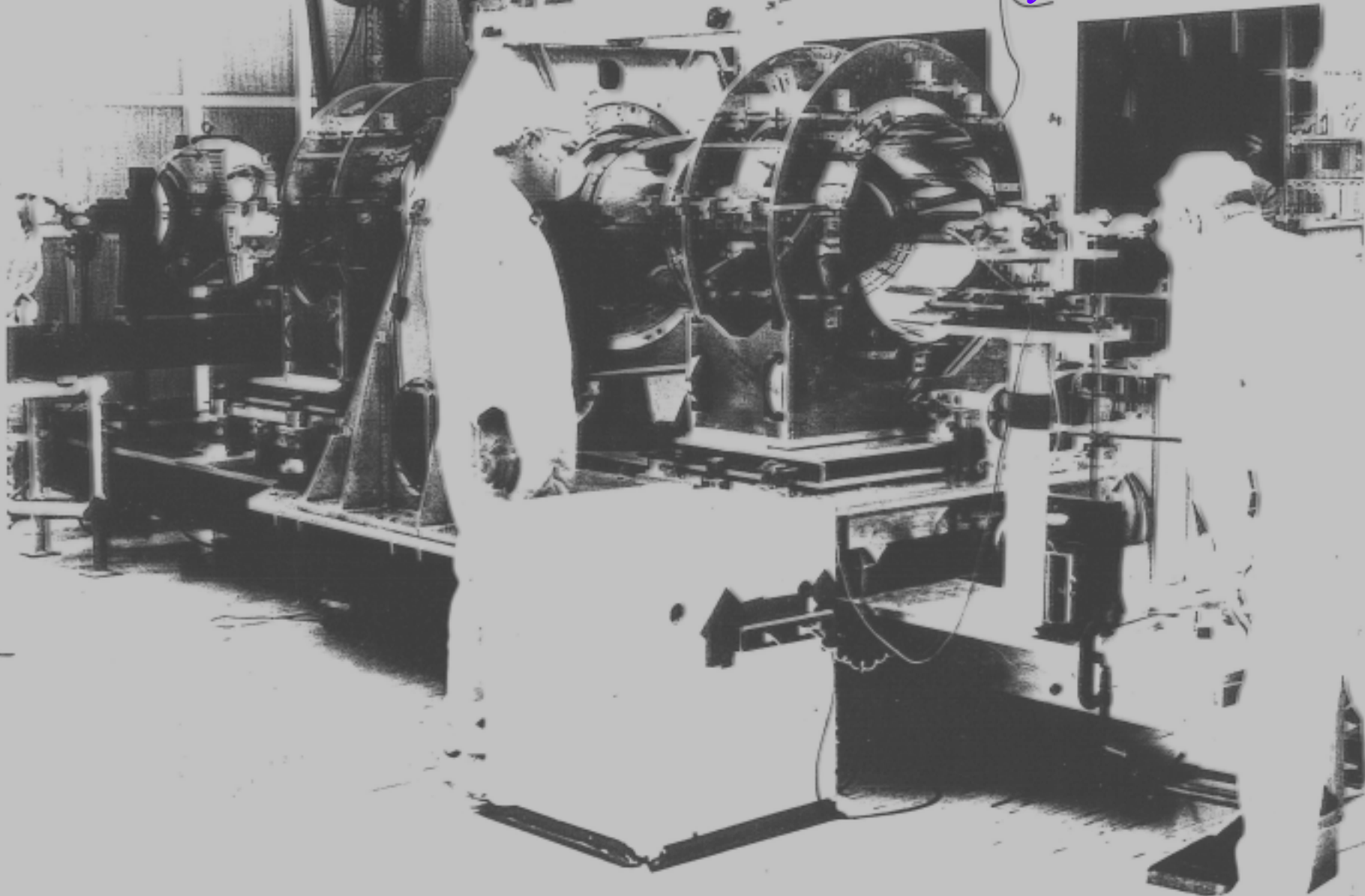
Solar results



Einstein Observatory HEAO-B (schematic)



Einstein Observatory HEAO-B

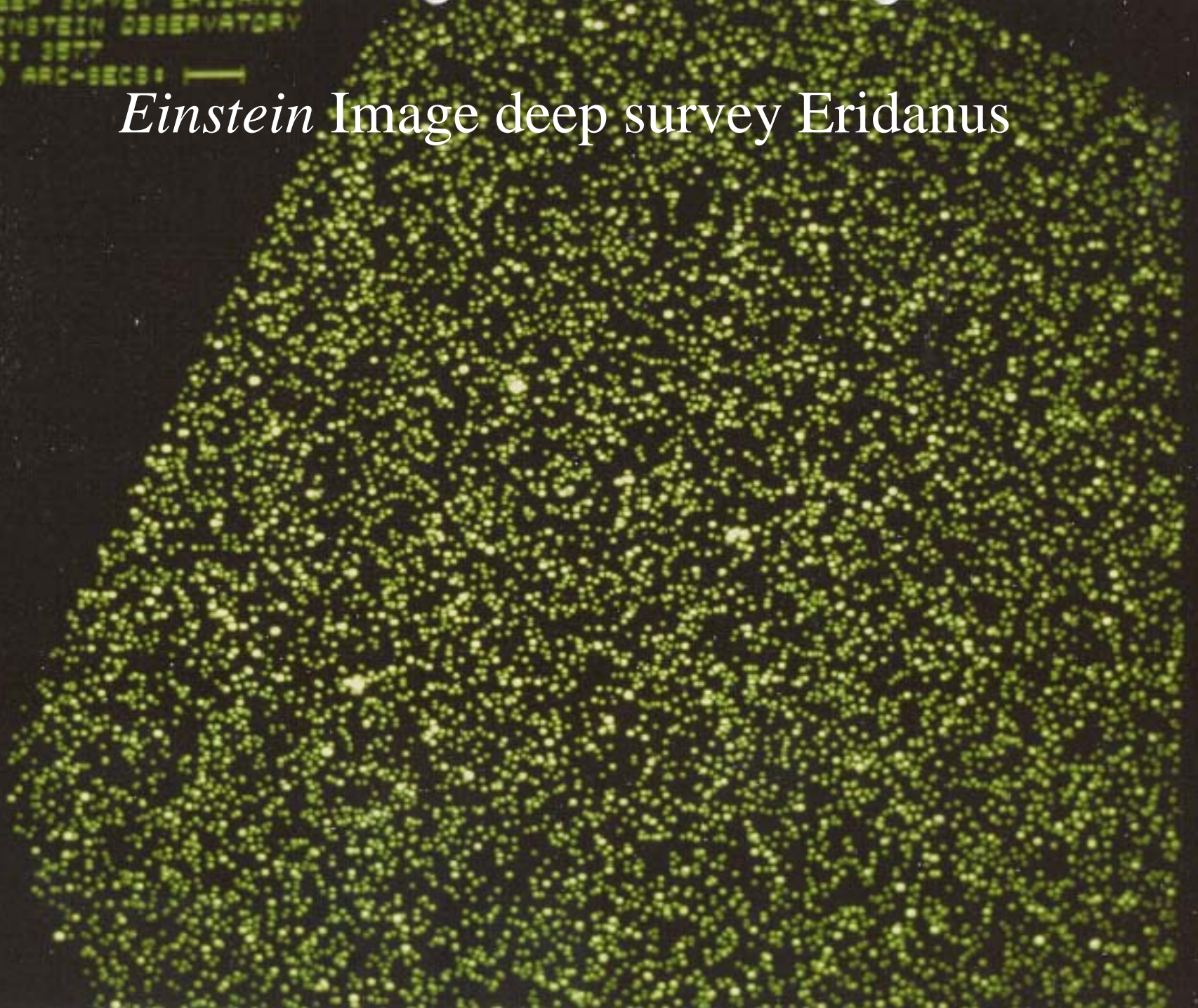


CELESTIAL OBJECTS OBSERVED WITH EINSTEIN

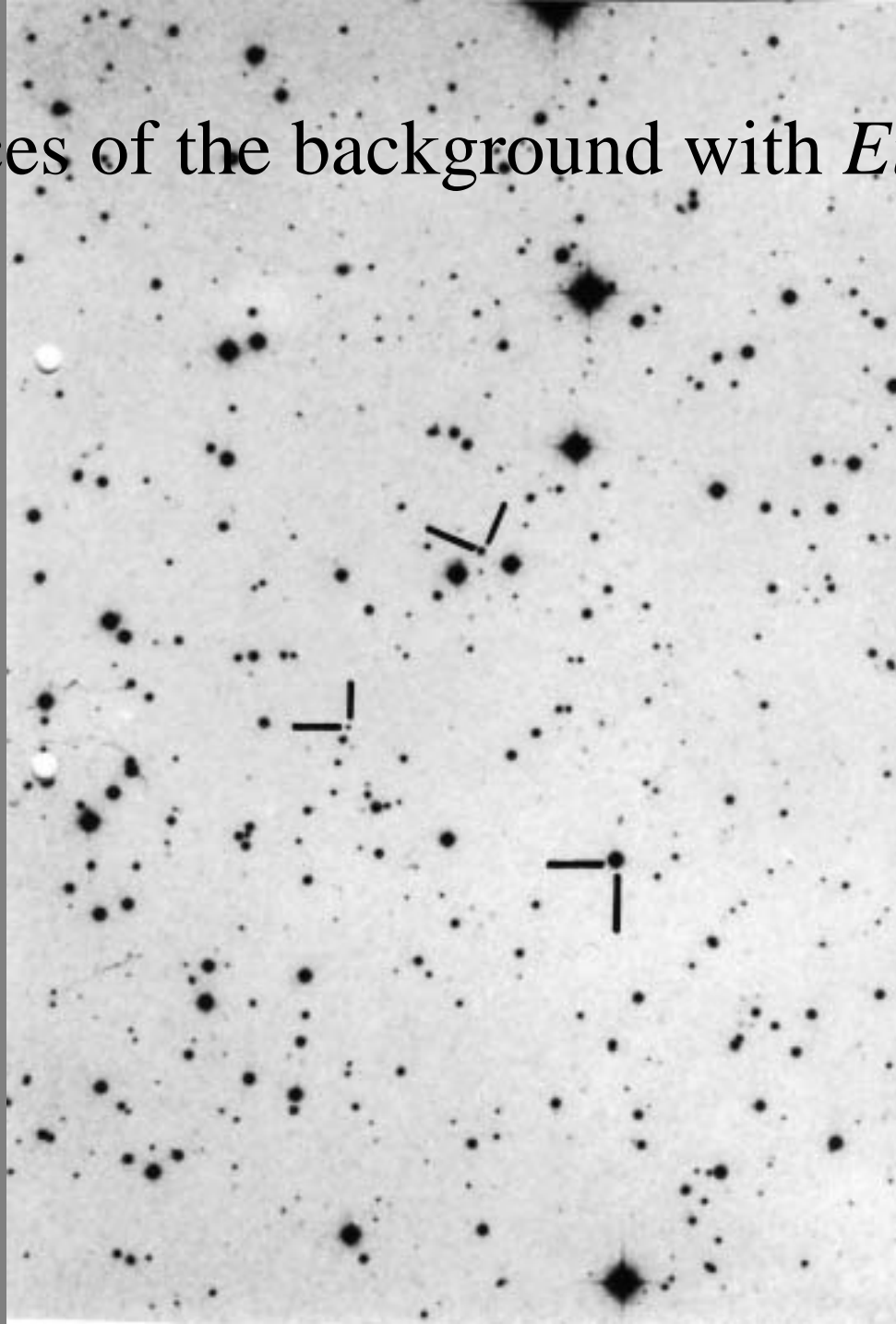
- **AURORA ON JUPITER**
- **X– RAY EMISSION FROM ALL TYPES OF MAIN SEQUENCE STARS**
- **NOVAS AND SUPERNOVAS**
- **PULSARS**
- **BINARY X–RAY SOURCES AND SUPERNOVAS IN EXTERNAL GALAXIES**
- **NORMAL GALAXIES**
- **NUCLEI OF ACTIVE GALAXIES**
- **QUASARS**
- **GROUPS AND CLUSTERS OF GALAXIES**
- **SOURCES OF THE EXTRAGALACTIC X–RAY BACKGROUND**

DEEP SURVEY ERI-100
EINSTEIN OBSERVATORY
HRI 3577
60 ARC-SECS1 —

Einstein Image deep survey Eridanus



Sources of the background with *Einstein*

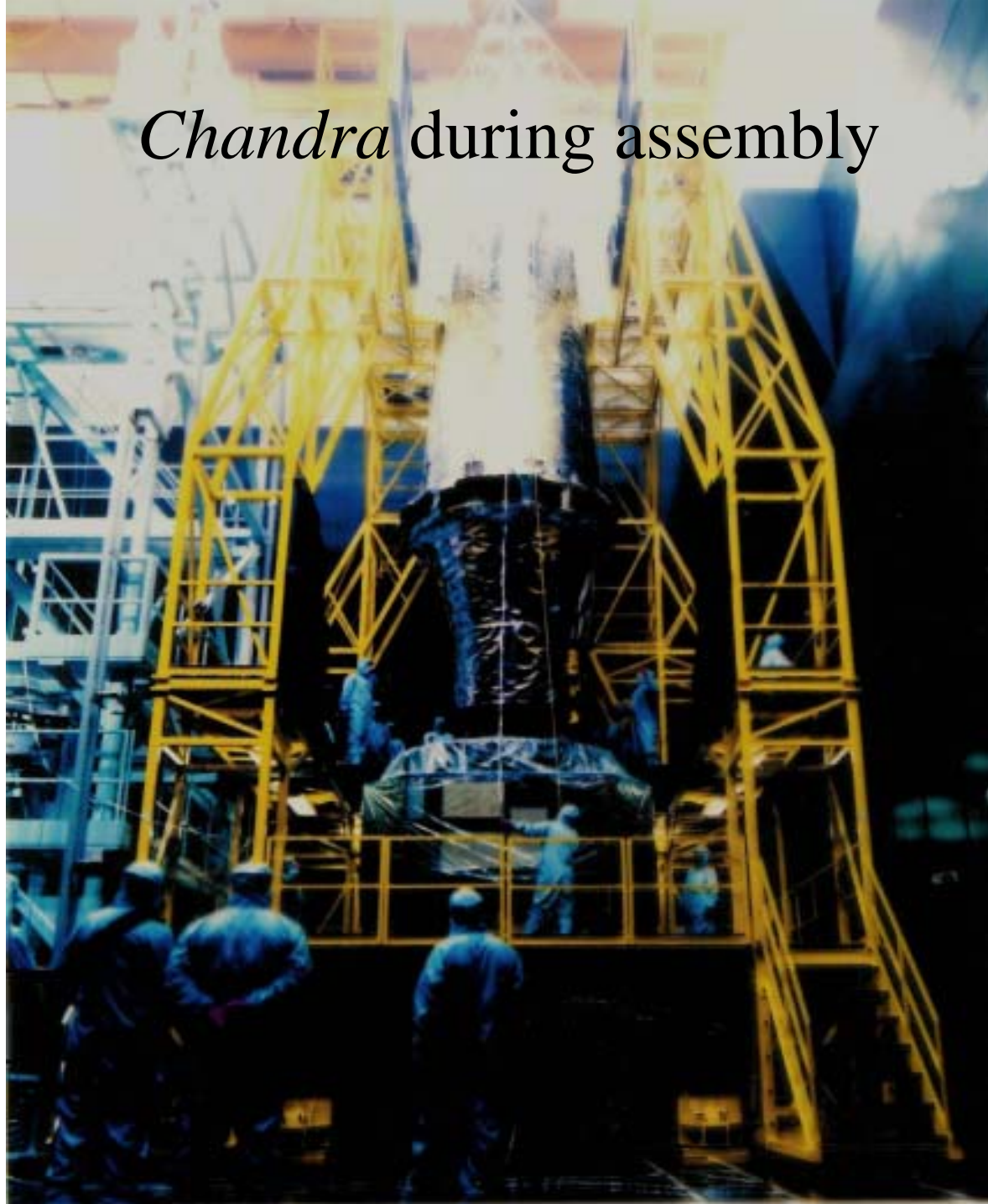


M31 CENTER REGION
30 ARC-SECS: 1—1
EINSTEIN OBSERVATORY

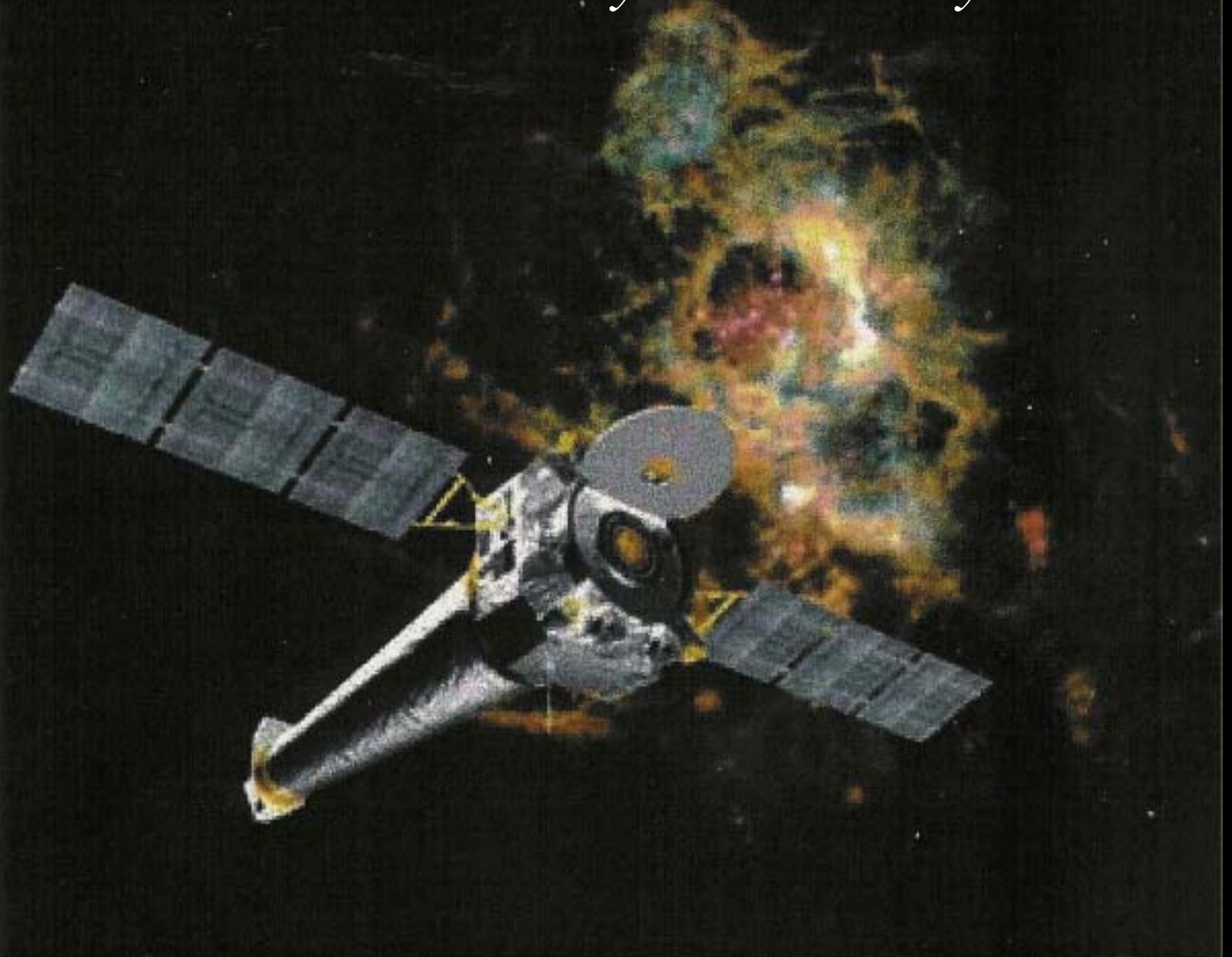
M31 with *Einstein*



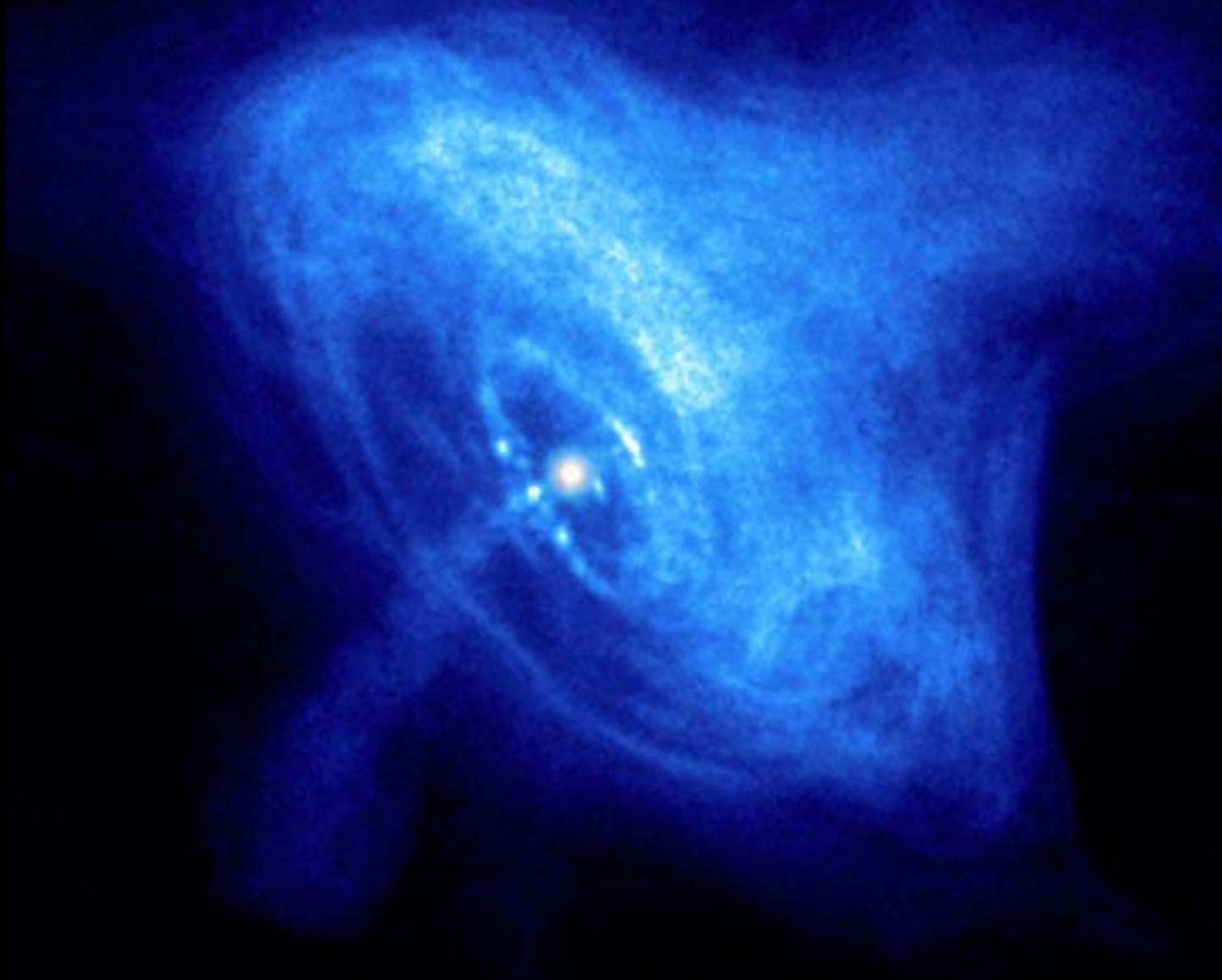
Chandra during assembly



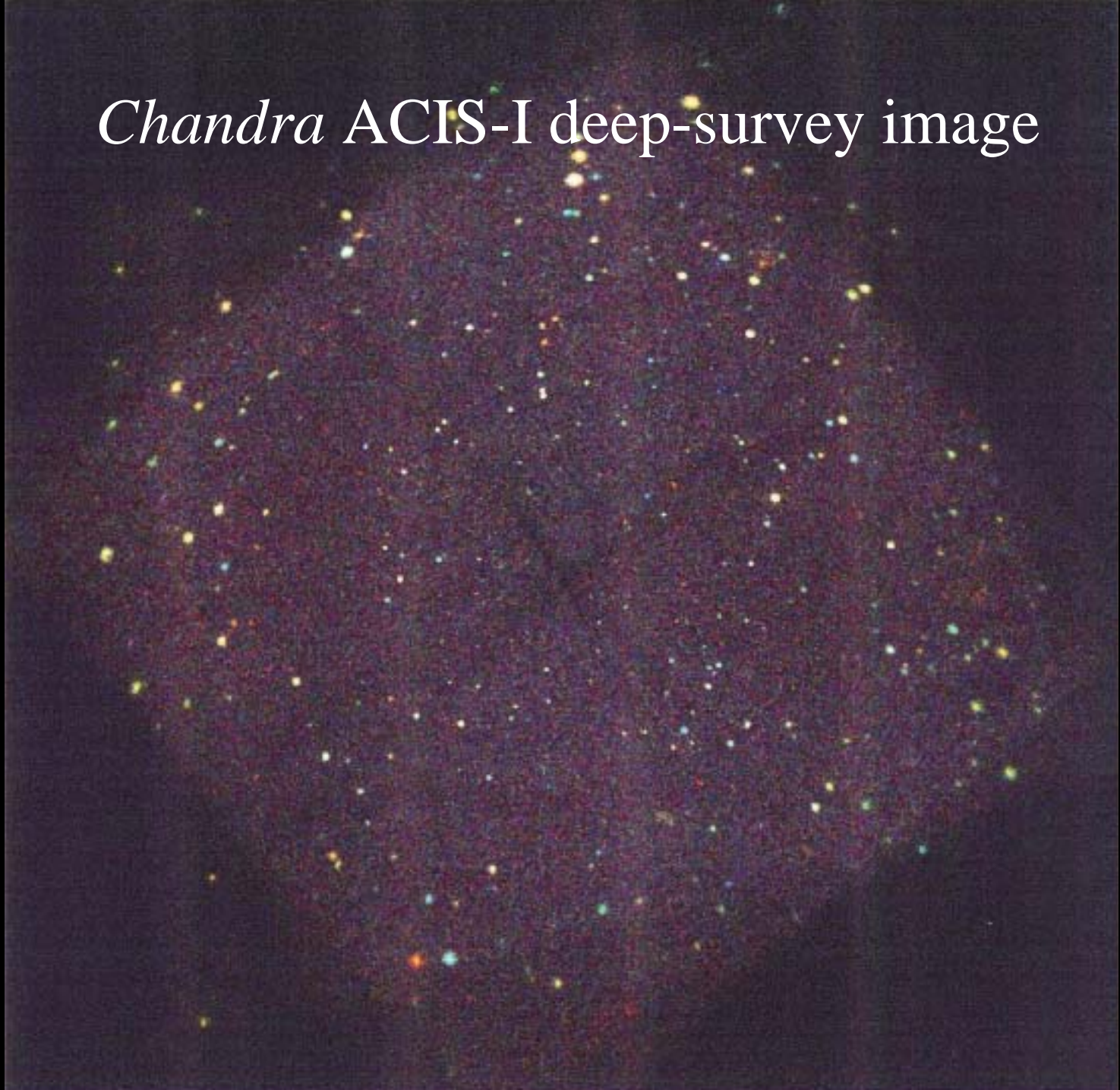
Chandra X-ray Observatory



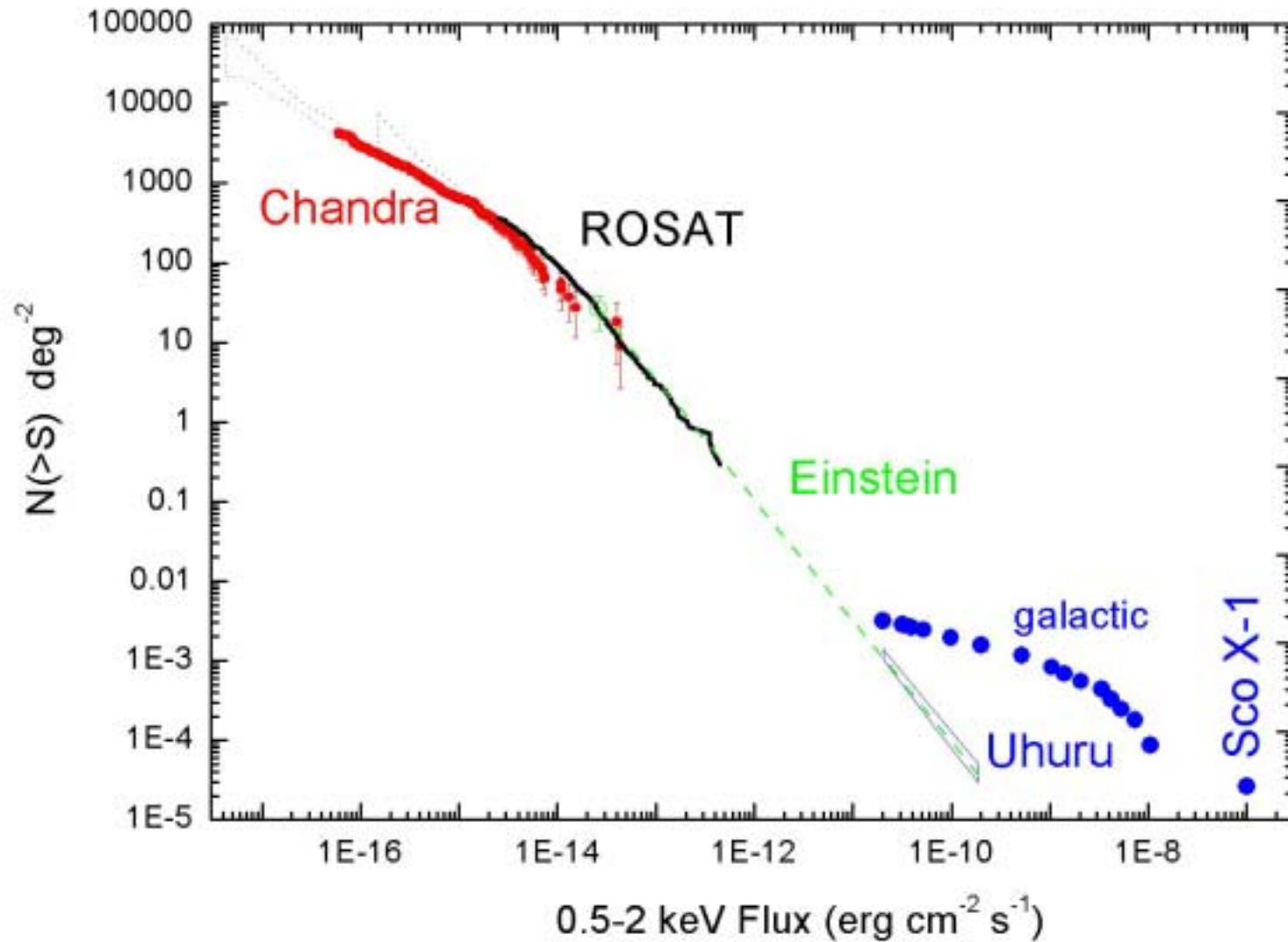
Crab Nebula

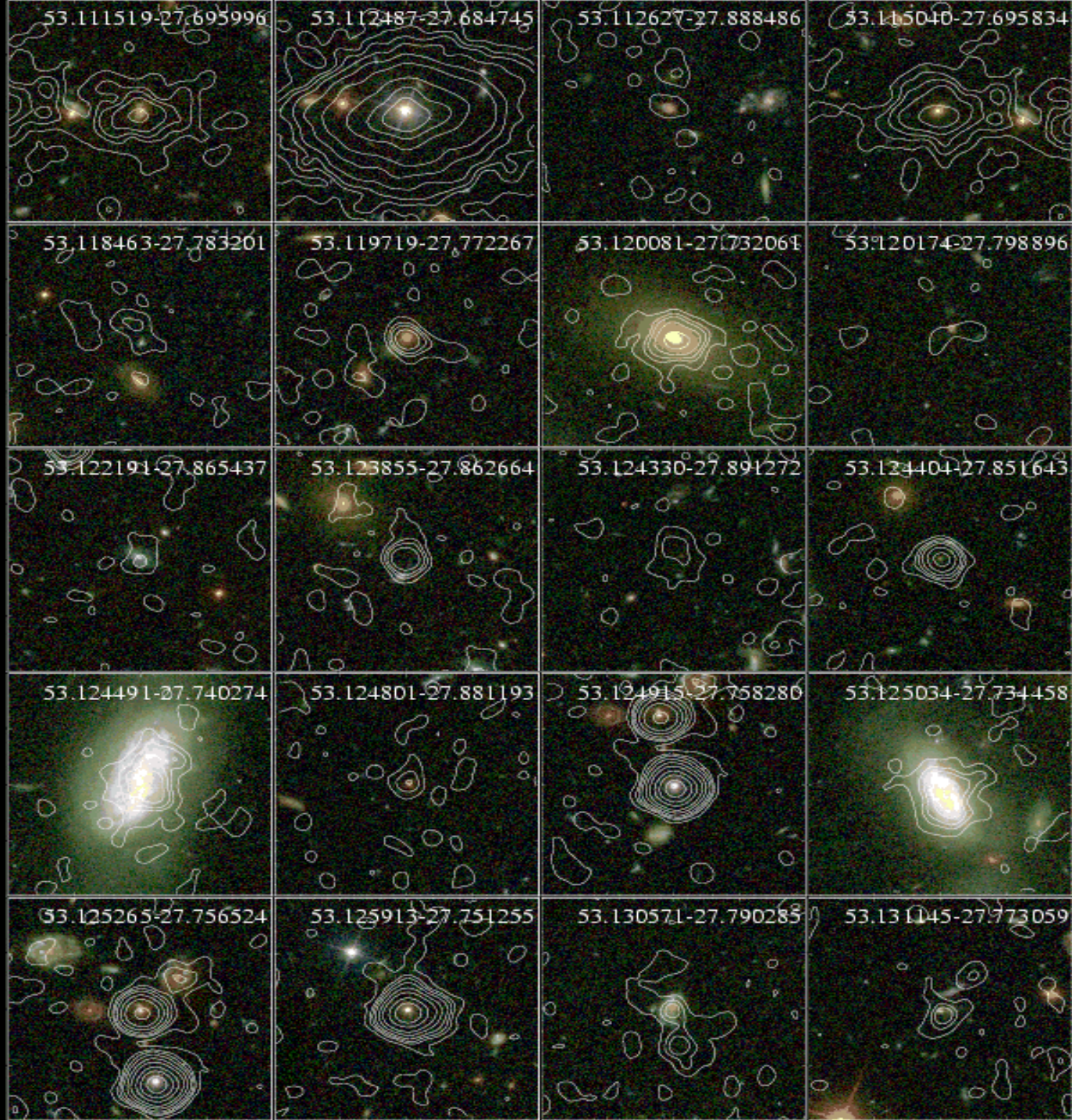


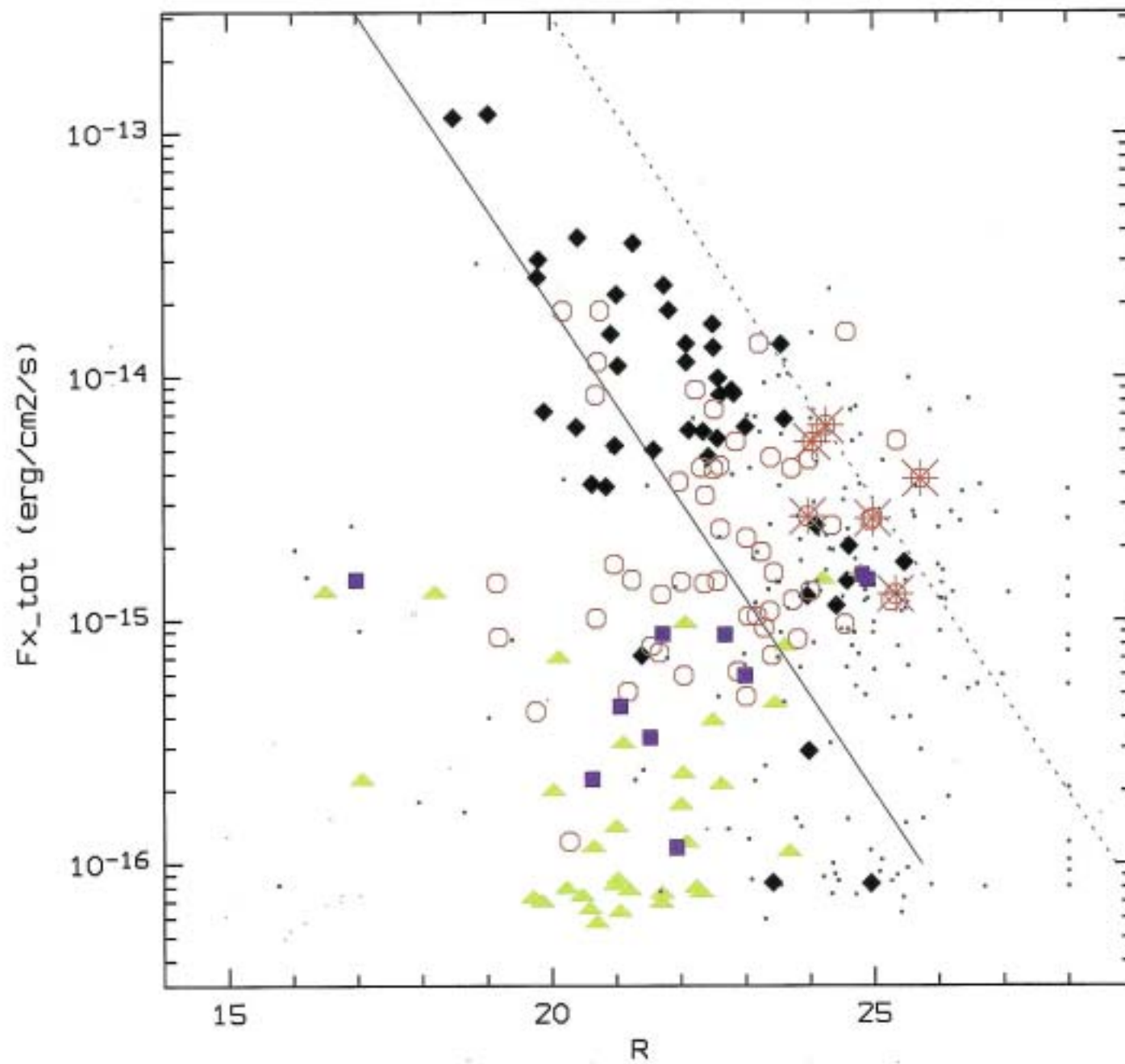
Chandra ACIS-I deep-survey image



Increase in sensitivity







THE X-RAY BACKGROUND (.5 – 10 KEV)

\$ > 95% XRB DUE TO SOURCES

\$ XRB SPECTRUM = ? SPECTRA OF SOURCES

\$ SPECTRUM EVOLUTION

\$ SOURCES: AGN TYPE I ($L = 10^{42} - 10^{45}$ $H \sim 0.5$)
AGN TYPE II ($L = 10^{41} - 10^{44}$ $H > 0.0$)
GALAXIES ($L < 10^{42}$ $H \sim -1$)
CLUSTERS OF GALAXIES
UNIDENTIFIED $\sim 10\%$

\$ SOURCE DENSITY 3000/SQ DEGREE
(10^8 ALL SKY) AT $S = 5 \times 10^{-17}$ ERG CM⁻² SEC⁻¹
HIGHEST DENSITY AGN'S (BH'S)

\$ 1/3 OF SOURCE IN CDFS ARE IN SHEETS $Z = 0.67, 0.73$
(LOW SIGMA) $Z = 1.04, 1.62$ AND 2.57

\$ DEEP SURVEYS STUDY EVOLUTION OF GALAXIES – AGN'S –
CLUSTERS OF GALAXIES & LARGE STRUCTURES

